

INTRODUCTION

Molecular nitrogen and methane are the main precursors of Titan's atmospheric chemistry driven by high energy electrons from Saturn's magnetosphere and solar UV photons.

A great number of laboratory experiments have been carried out to mimic Titan's atmospheric chemistry. But, up to now, N₂/CH₄ mixtures have been submitted to a unique type of energy source: UV light, high energy particles or electric discharges.

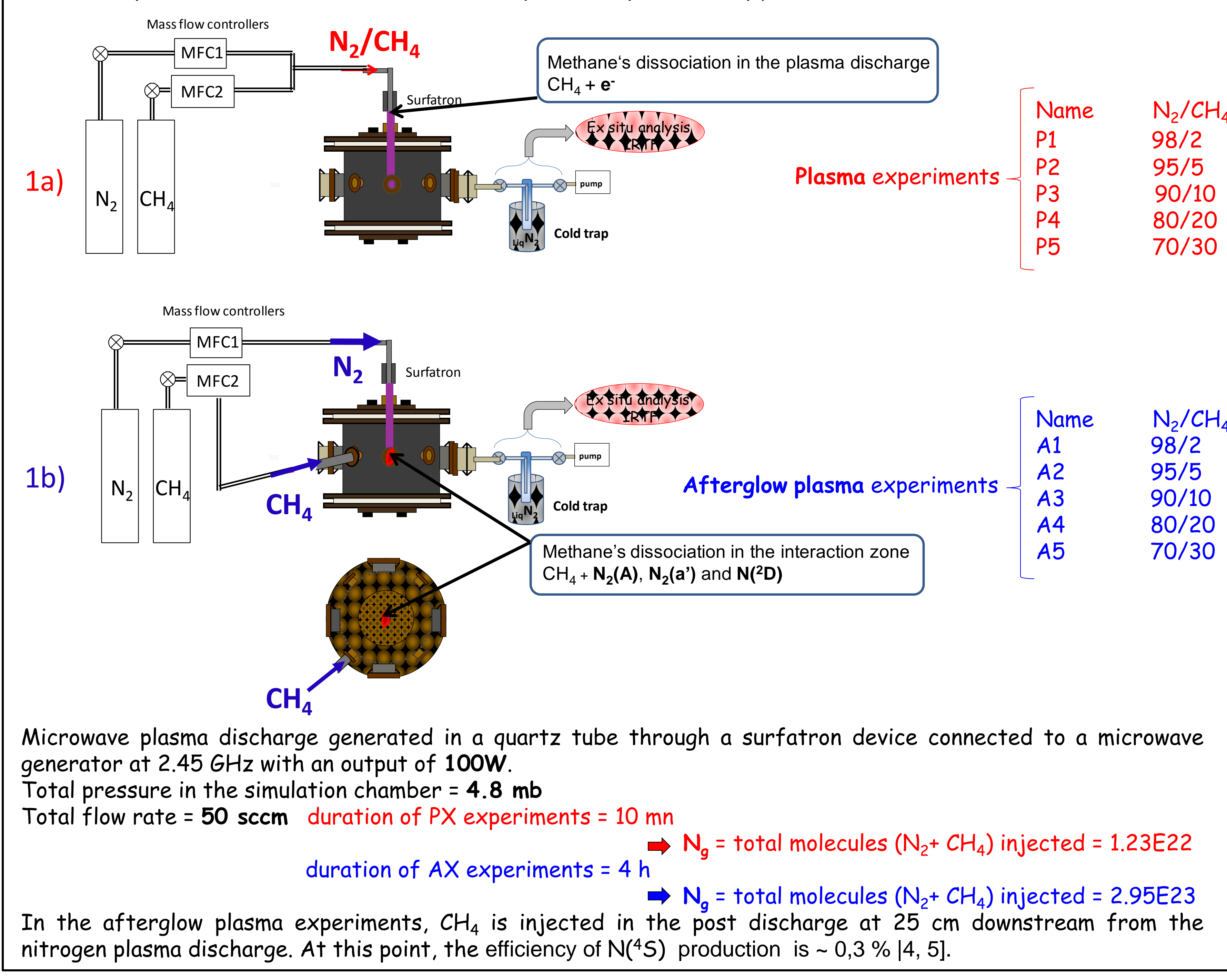
We are developing a new program of simulations [1] aiming to **improve the representativeness towards Titan's condition in term of energy deposition** which consists in submitting a N₂/CH₄ mixture in a flow reactor to both electrons (microwave plasma discharge) and photons (Lyman-alpha delivered by a continuous H₂/He lamp).

In a first step, N₂/CH₄ chemistry has been initiated using microwave **plasma** discharges or **afterglow plasma** discharges.

A detailed kinetic model has been developed to **simulate the results obtained for the post-discharge experiments**. Theoretical predictions are then compared to experimental measurements.

EXPERIMENTAL

Titan's experimental simulations have been undertaken using N₂/CH₄ **plasma** (figure 1a) as well as N₂/CH₄ **afterglow plasma** (figure 1b) with different mixing ratio of N₂ and CH₄. Ex-situ qualitative and quantitative analysis of the resulting gas mixture recovered in a cold trap has been performed, for the first time, by IRTF spectroscopy.

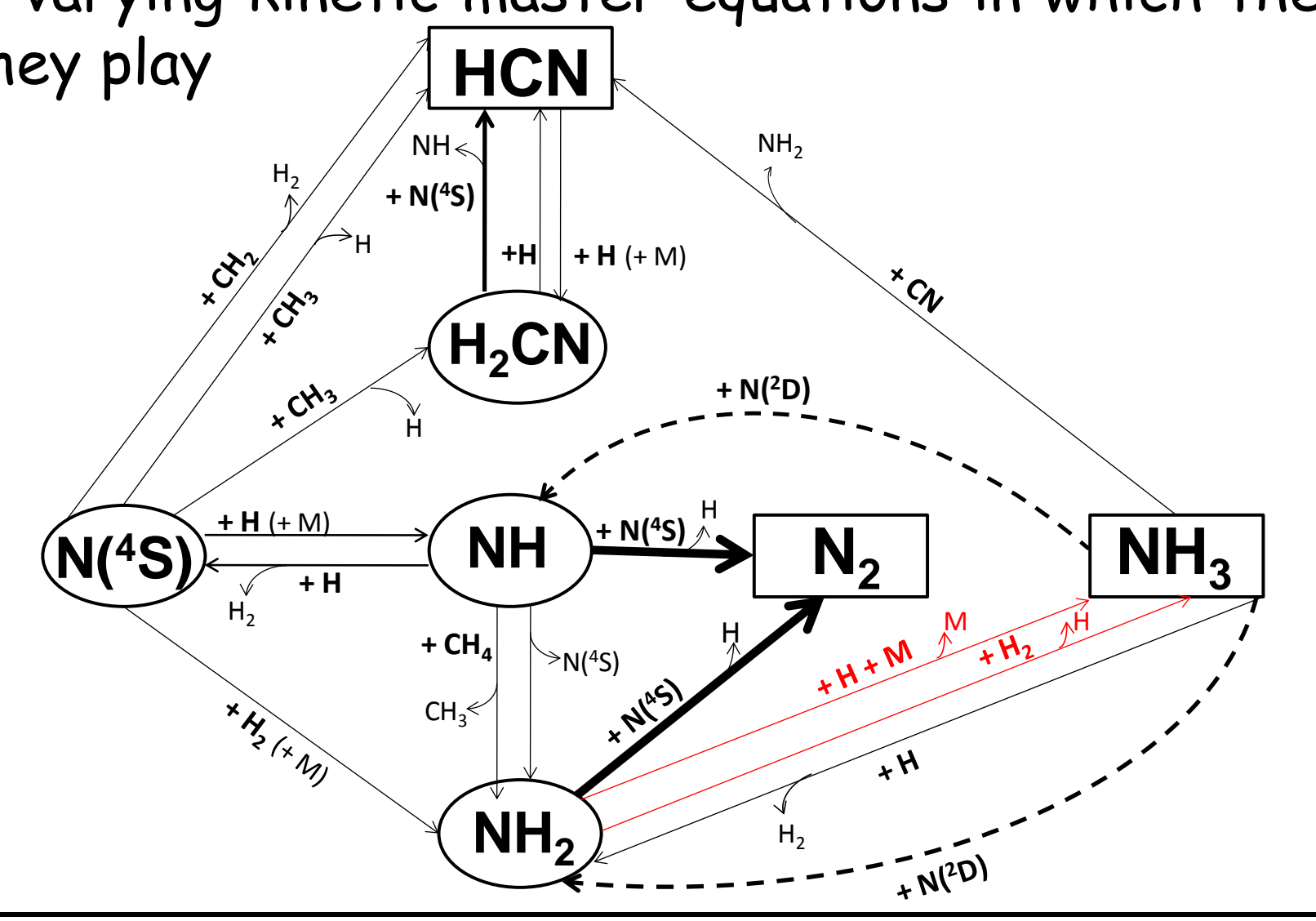


KINETIC MODEL

A self-consistent kinetic model for the afterglow of a flowing microwave discharge in pure N₂, in which various percentages of CH₄ are introduced has been developed by Pintassilgo and co-workers [2,3,4]. This model has **two separate modules**, one for the **discharge** and another for the **post-discharge**, the stationary solutions obtained in the first module being the initial conditions for the second one.

The first module considers the various atomic and molecular species produced in the discharge, i.e. ions, neutral species such as ground state N(⁴S) and metastables N(²D), N(²P) or N₂(A³Σ_u⁺) and N₂(a¹Σ_g⁻). The latter may be either carried out to the post discharge or created in the post-discharge. Their concentrations and time evolutions are studied thanks to the second module of the model by considering a system of coupled time-varying kinetic master equations in which the electron impact processes can be neglected since they play a very minor role.

In the post-discharge, the nitrogen metastables mentioned above are able to dissociate methane producing CH₃, CH₂, CH radicals. These species along with H and H₂ further react with atomic nitrogen N(⁴S) to produce more complex species, to produce more complex species starting with HCN. Additionally, the model takes into account reactions leading to NH₃ via NH₂ radical.



REFERENCES

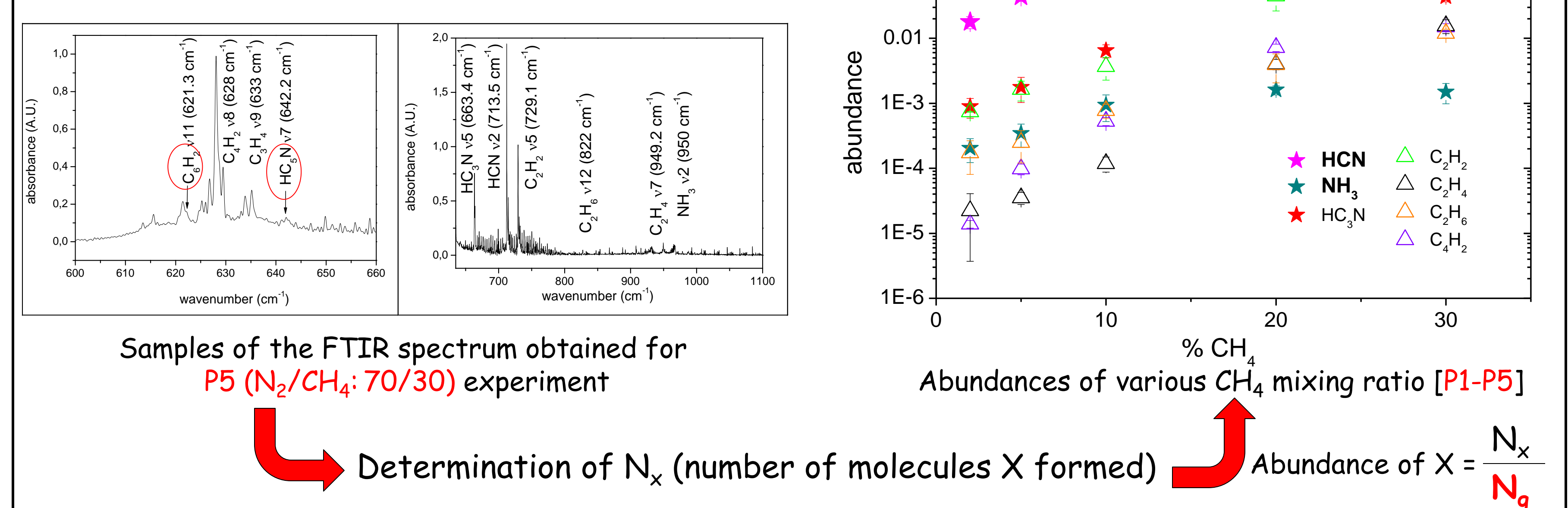
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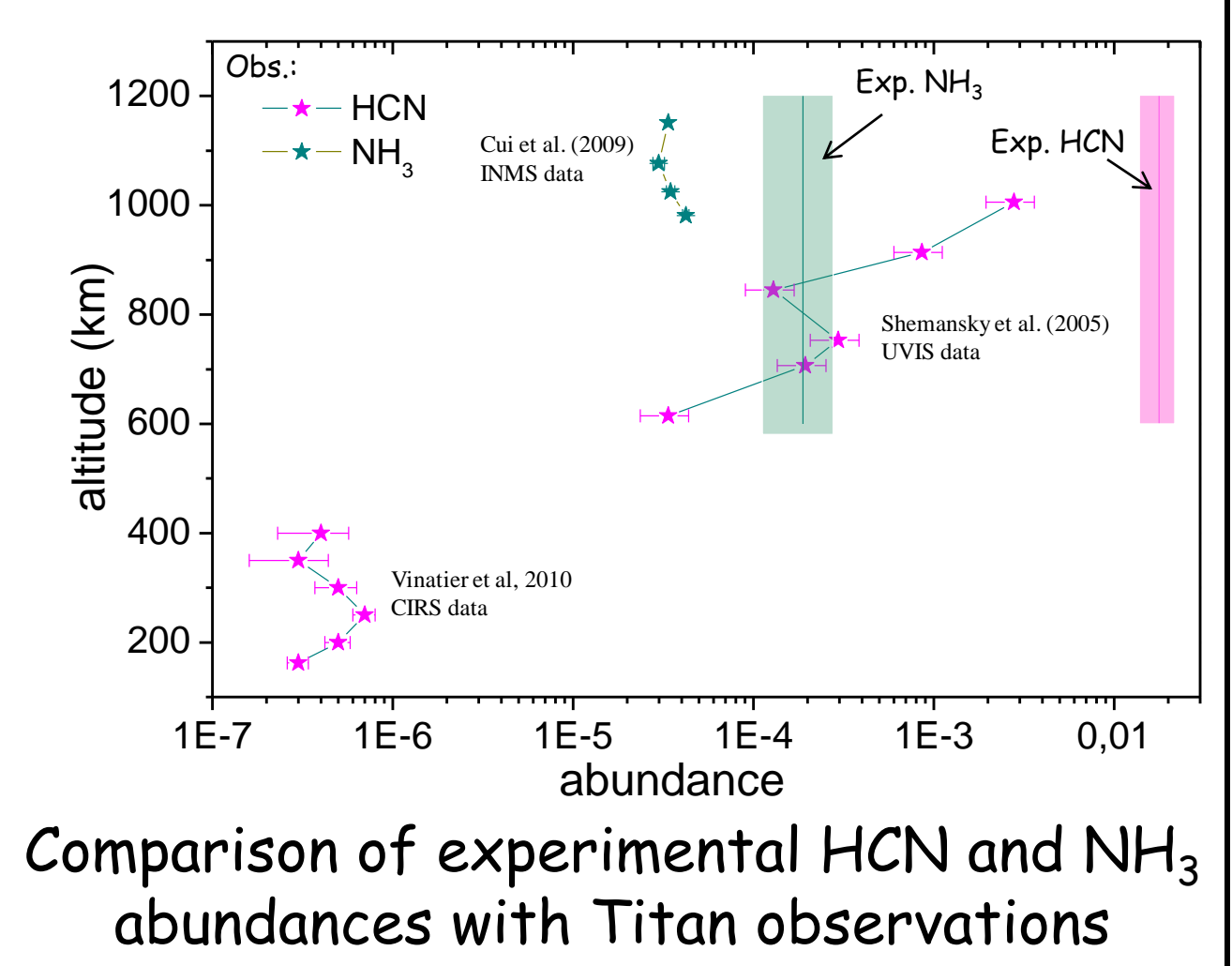
RESULTS

Plasma experiments



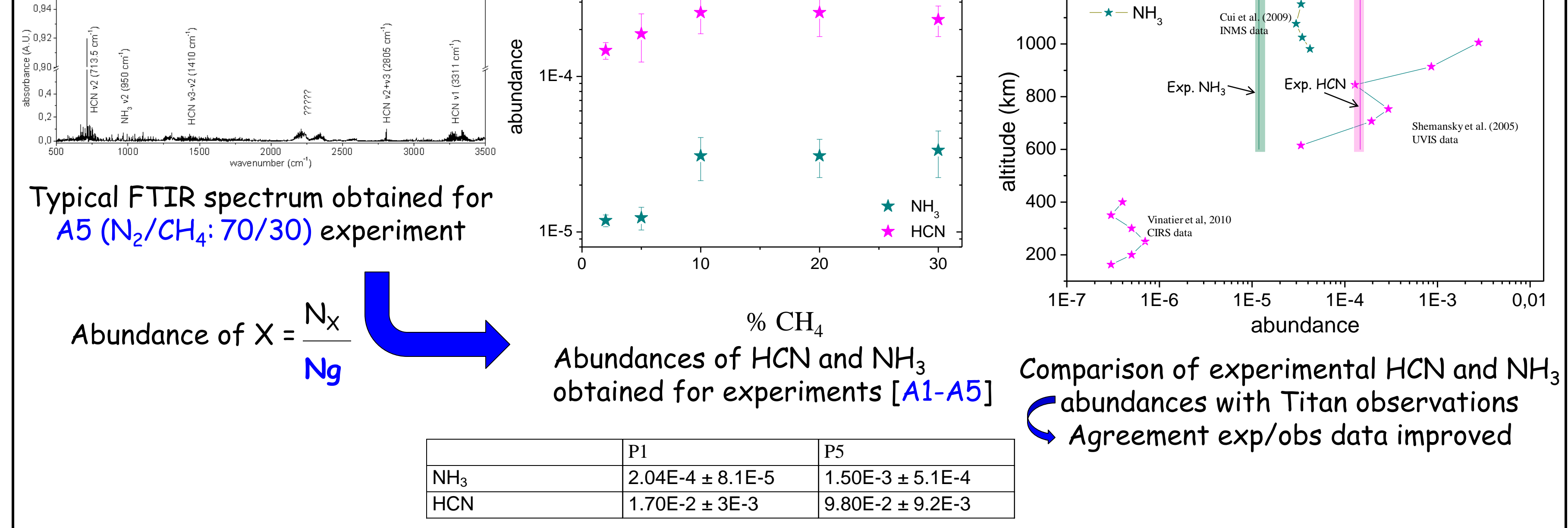
The experimental results are consistent with those obtained for previous plasma simulations (Coll et al, 1995) and are in relative good agreement with the observational data in the upper atmosphere of Titan (900 to 1200 km) but...

Plasma experiments	Titan's upper atmosphere (>500km)
high energetic electrons	high energetic electrons + photons
T gas ~ 100K [5]	T lower than 180 K



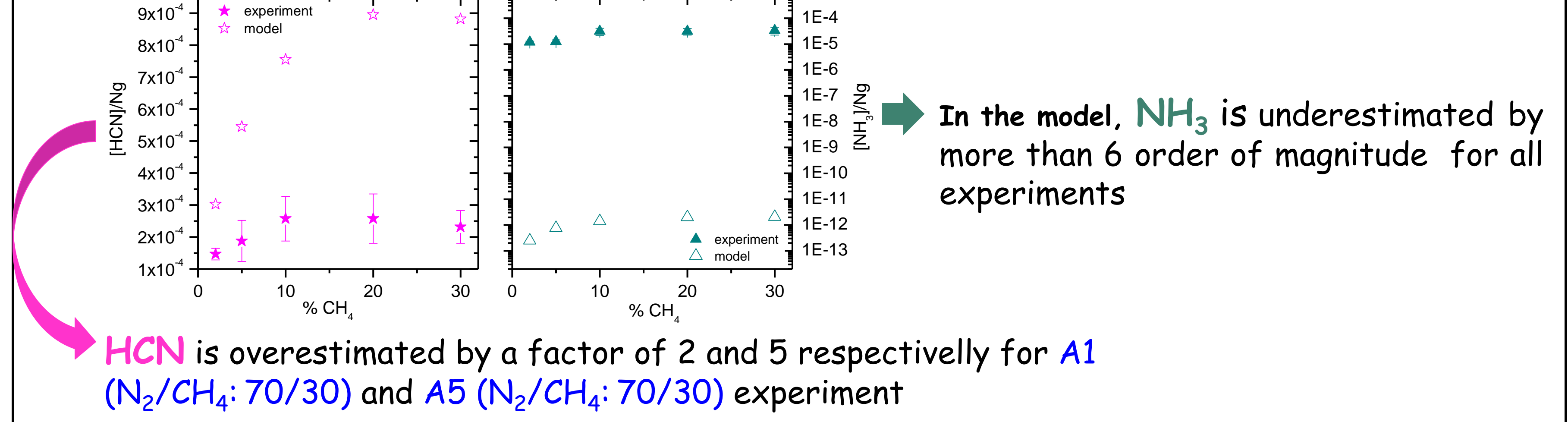
Such microwave plasma experiments can not be considered as representative simulations of Titan's atmosphere.

Afterglow plasma experiments (representativeness towards Titan's temperature improved as T gas ~ 300K)



Unlike plasma experiments, only two main species are observed here: HCN and NH₃

Comparison with theoretical predictions



Conclusion

In the chemical scheme used to describe the evolution of the post discharge, the most efficient reactions lead to the formation of HCN as well as N₂ regeneration.

The reactions involved in the formation of NH₃ (in red in the scheme) are too slow in the gas phase to reproduce the experimental abundance of this compound.

Therefore, surface processes must be taken into account in the model in order to reproduce the production of NH₃.

WORK IN PROGRESS

