

## Implication on the study of methane photolysis at Lyman alpha (121.6 nm)

M.-C. GAZEAU<sup>1</sup>, Y. BENILAN<sup>1</sup>, ET. ES-SEBBAR<sup>1,2</sup>, A. JOLLY<sup>1</sup>, E. ARZOUMANIAN<sup>1</sup>, N. FRAY<sup>1</sup> and H. COTTIN<sup>1</sup>

<sup>1</sup> LISA (Laboratoire Interuniversitaire des Systèmes atmosphériques), UMR CNRS 7583, Universités Paris-Est Créteil Val de Marne (UPEC) and Paris Denis Diderot, CNRS-UMR 7583, IPSL, 61, avenue du Général de Gaulle, 94010 Créteil Cedex, France

<sup>2</sup> King Abdullah University of Science and Technology (KAUST), Clean Combustion Research Center (CCRC), Thuwal 23955-6900, Kingdom of Saudi Arabia

### I. INTRODUCTION

**Context:** VUV Photons from gas discharges are efficient sources of electromagnetic radiation for photochemical studies in laboratory and simulations of planetary atmosphere. In such experiments, extensive investigation of the VUV emission as the function of the plasma parameters (photons flux and wavelength dependency) is required.

**Objective:** we want to increase the photon flux delivered by a H<sub>2</sub>/He lamp at Lyman alpha (121.6 nm) by determining the accurate plasma parameters to be used. The effect of addition of rare gas (argon) on the monochromaticity of the lamp is also studied.

**Approach:** the emission spectrum of the lamp is recorded by a VUV spectrometer. Its photon flux is measured by chemical actinometry using CO<sub>2</sub> gas as an actinometer. The method is based on the temporal measurements of CO density issued from photo-dissociation of CO<sub>2</sub> at Lyman alpha using FTIR spectrometer. The temporal dependence of the CO density is determined by systematic comparison of experimental spectra with synthetic ones using spectroscopic parameters from the GEISA database [1].

### II. EXPERIMENTAL

#### 1. VUV source

The plasma is created in a quartz tube (i.d. 8 mm, e.d. 10 mm, length 20 cm) in which the flow of H<sub>2</sub>/He (/Ar) gas mixture is injected. The microwave excitation is coupled to the plasma through a McCarroll cavity (Optos Instruments, Inc.) powered by a microwave generator (SAIREM-2.45 GHz-300W). Excited H atoms generated in the plasma are responsible for the VUV emission at Lyman alpha

#### 2. Emission spectrum of the H<sub>2</sub>/He(/Ar) microwave discharge lamp:

The lamp is connected to a VUV spectrometer in order to record its VUV emission spectrum (Fig. 1)

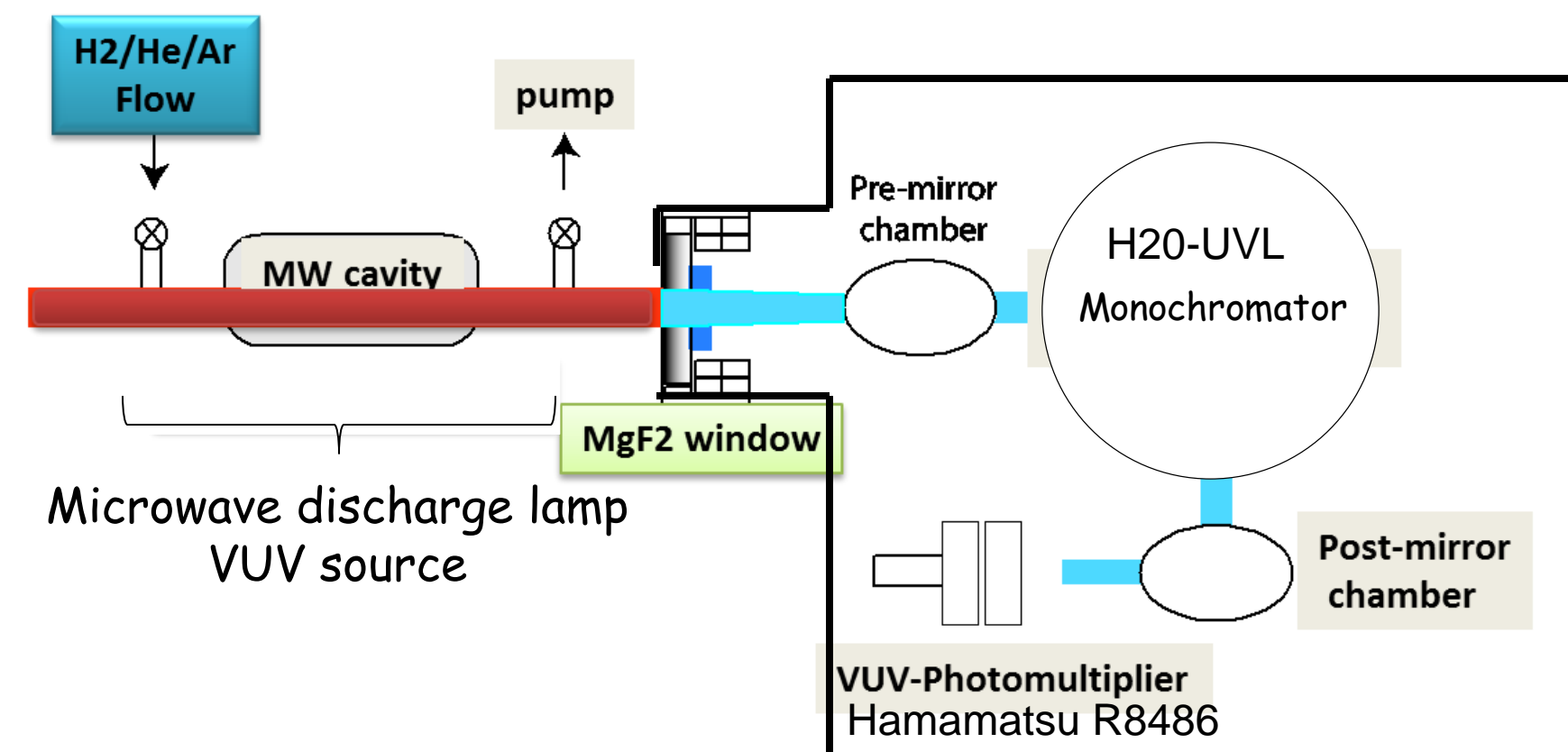


Fig. 1: The experimental arrangement of the microwave plasma source connected to a VUV spectrometer.

#### 3. Photon Flux of the H<sub>2</sub>/He(/Ar) microwave discharge lamp (Chemical Actinometry):

The lamp is connected to a multi-reflexion white cell containing CO<sub>2</sub> (Fig. 2).

The CO issued from CO<sub>2</sub> photolysis is monitored by a Fourier Transform Infrared Spectrometer. A spectroscopic model is used to deduce the CO column density [2] (Fig. 3).

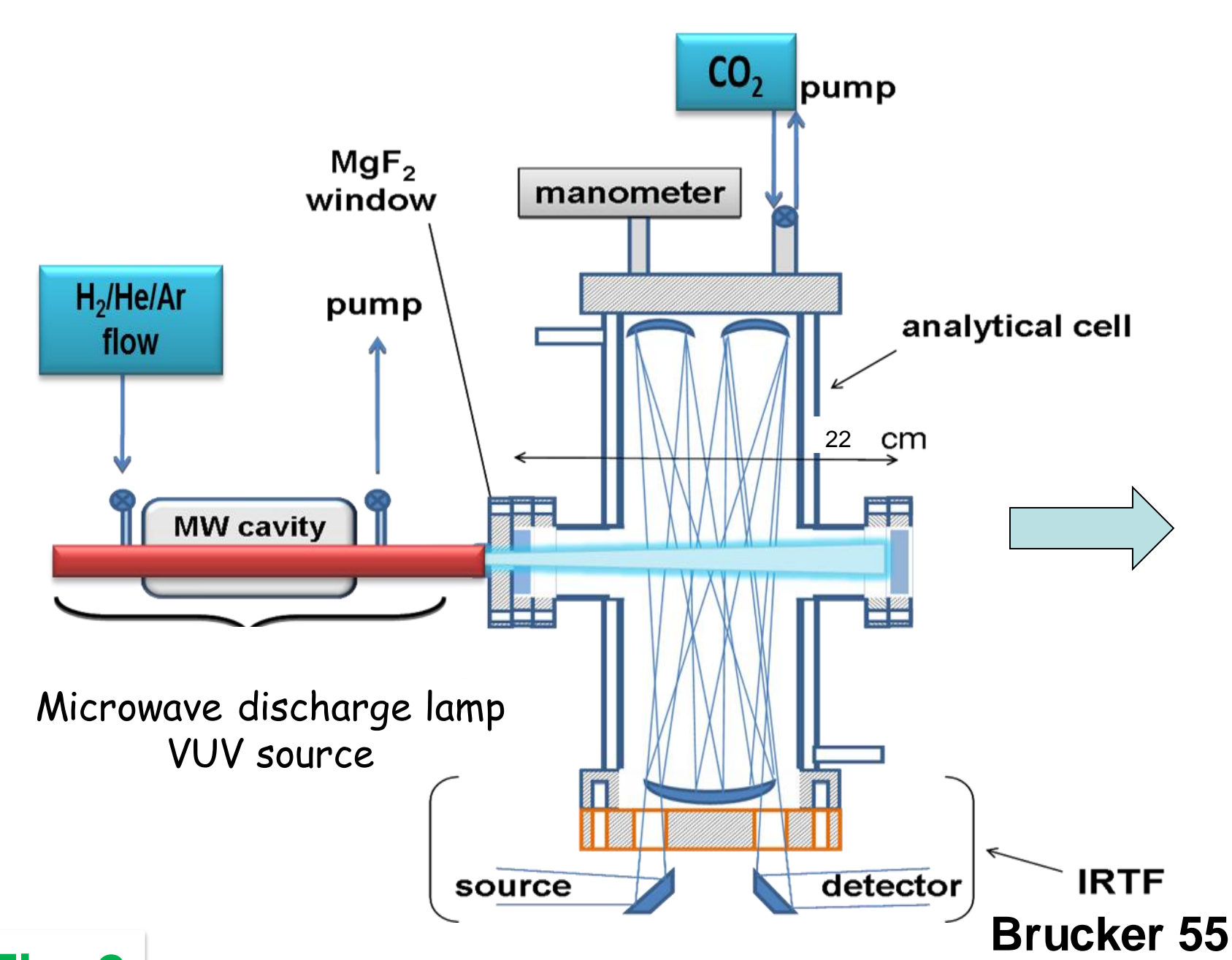


Fig. 2

SETUP to study the photolysis of CO<sub>2</sub> by a microwave discharge lamp.  
Volume of the multireflexion white cell:  $V_{cell} = 2600 \text{ cm}^3$   
Optical pathlength = 500 cm

The total number of Lyman alpha photons entering the cell per second,  $\Phi_{121.6}$ , is calculated from the determination of the CO production rate ( $d[CO]/dt$ ) (Fig. 4) using equation 1:

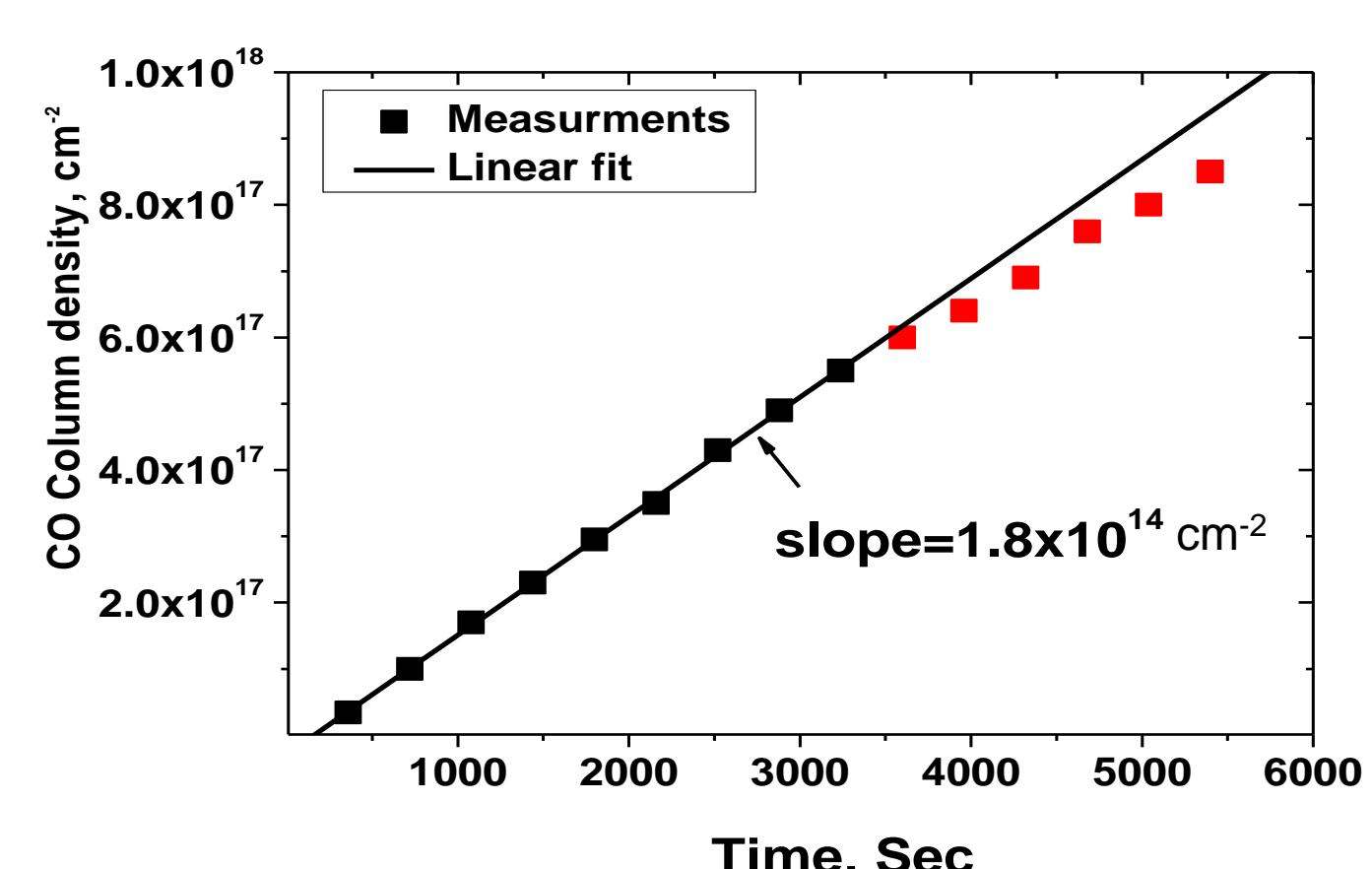


Fig. 4: CO column density as the function of the irradiation time.

$$d[CO]/dt = \text{slope} \times V_{cell} / \text{Optical pathlength} = 9.36 \cdot 10^{14} \text{ s}^{-1}$$

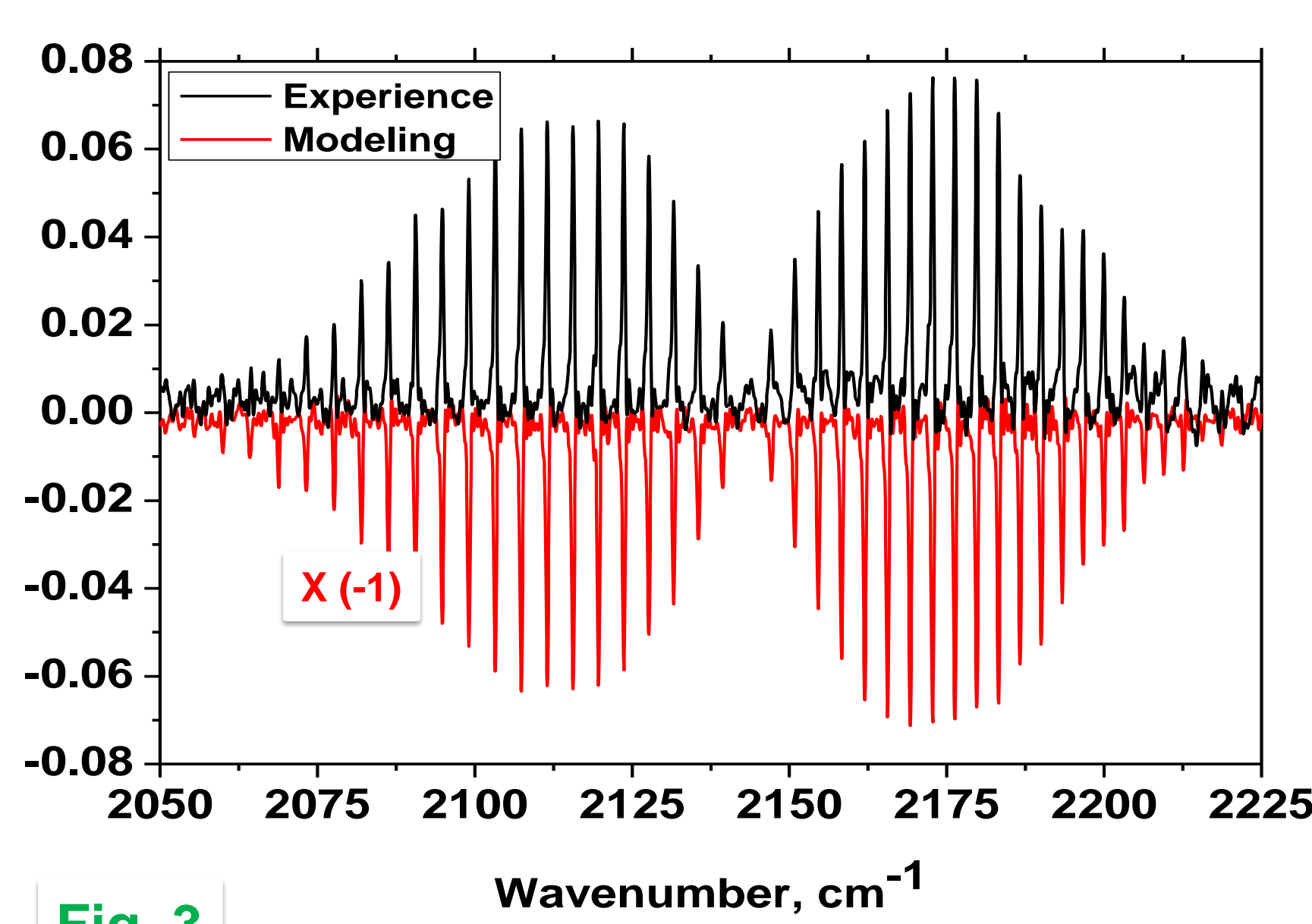


Fig. 3

CO production from CO<sub>2</sub> photolysis monitored in the ~ 2040-2240 cm<sup>-1</sup> range (spectral resolution: 0.5 cm<sup>-1</sup>). Comparison between experiment and theoretical data.

$d[CO]/dt = \Phi_{121.6} \cdot \eta \cdot \beta$  Eq. 1.

$\eta$  is the quantum yield of CO production from CO<sub>2</sub> photolysis at 121.6 nm assumed equal unity.  
 $\beta_{CO_2}$  is the fraction of the absorbed flux by CO<sub>2</sub> which is given by the Beer-Lambert law  $\beta_{CO_2} = 1 - \exp(-\sigma_{CO_2} \cdot I \cdot p)$  where  $\sigma = 1.61 \text{ atm}^{-1} \text{ cm}^{-1}$  is the absorption cross section of CO<sub>2</sub> with value at 121.6 nm taken from [3],  $l=22 \text{ cm}$  is the length of the cell and  $p$  is the gas pressure of CO<sub>2</sub> (4 Torr).

Using Eq. 1., the flux of photons delivered from the H<sub>2</sub>/He discharge lamp is  $5.5 \cdot 10^{15} \text{ s}^{-1}$ .

### III. RESULTS

#### 1. VUV spectrum of H<sub>2</sub>/He microwave discharge lamp, influence of the plasma parameters:

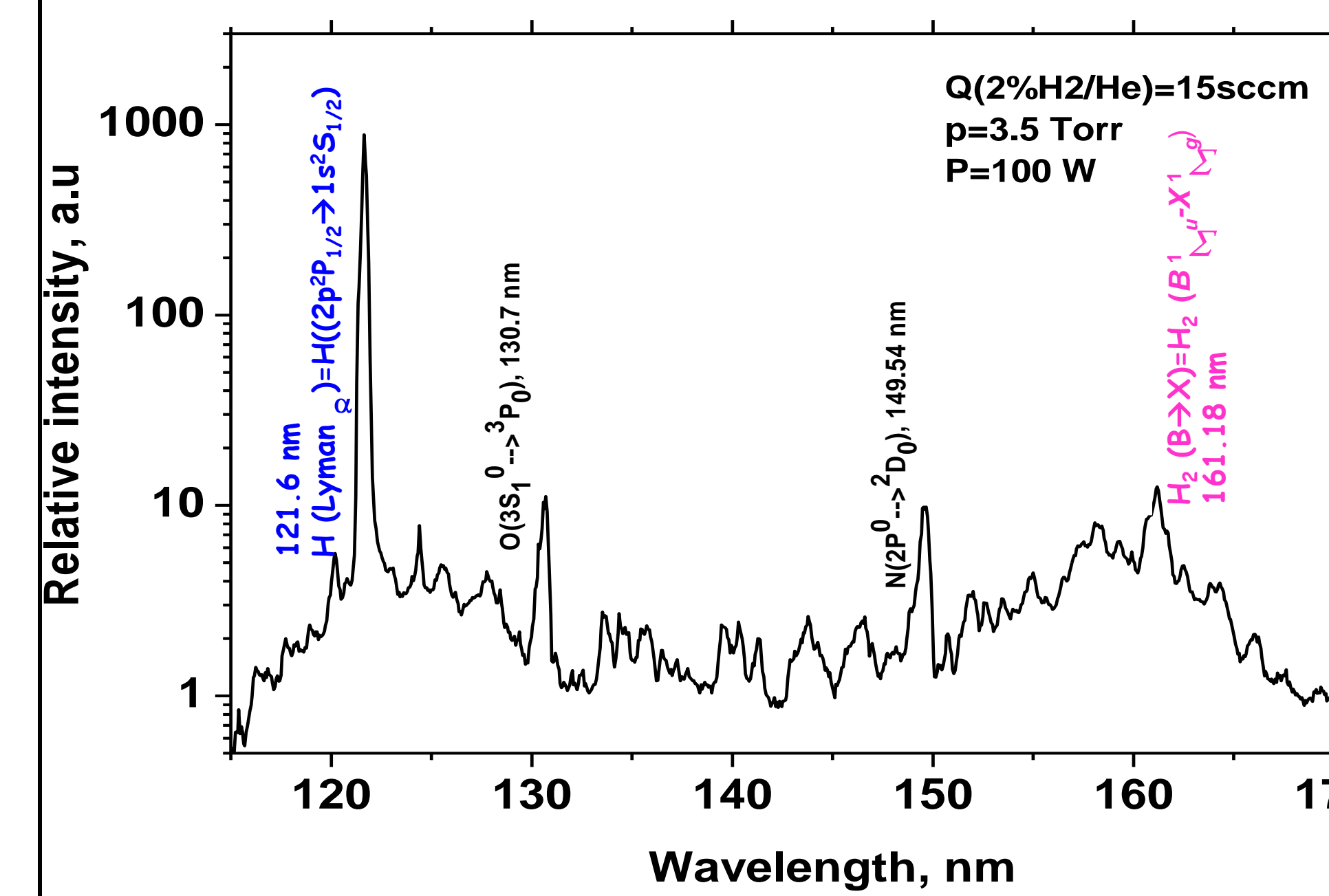


Fig. 5: A typical VUV spectrum between 115 and 170 nm. The main features emission are identified. Experimental conditions: 2% H<sub>2</sub>/He, p=3.5 Torr, flow rate=15 sccm and P=100 W.

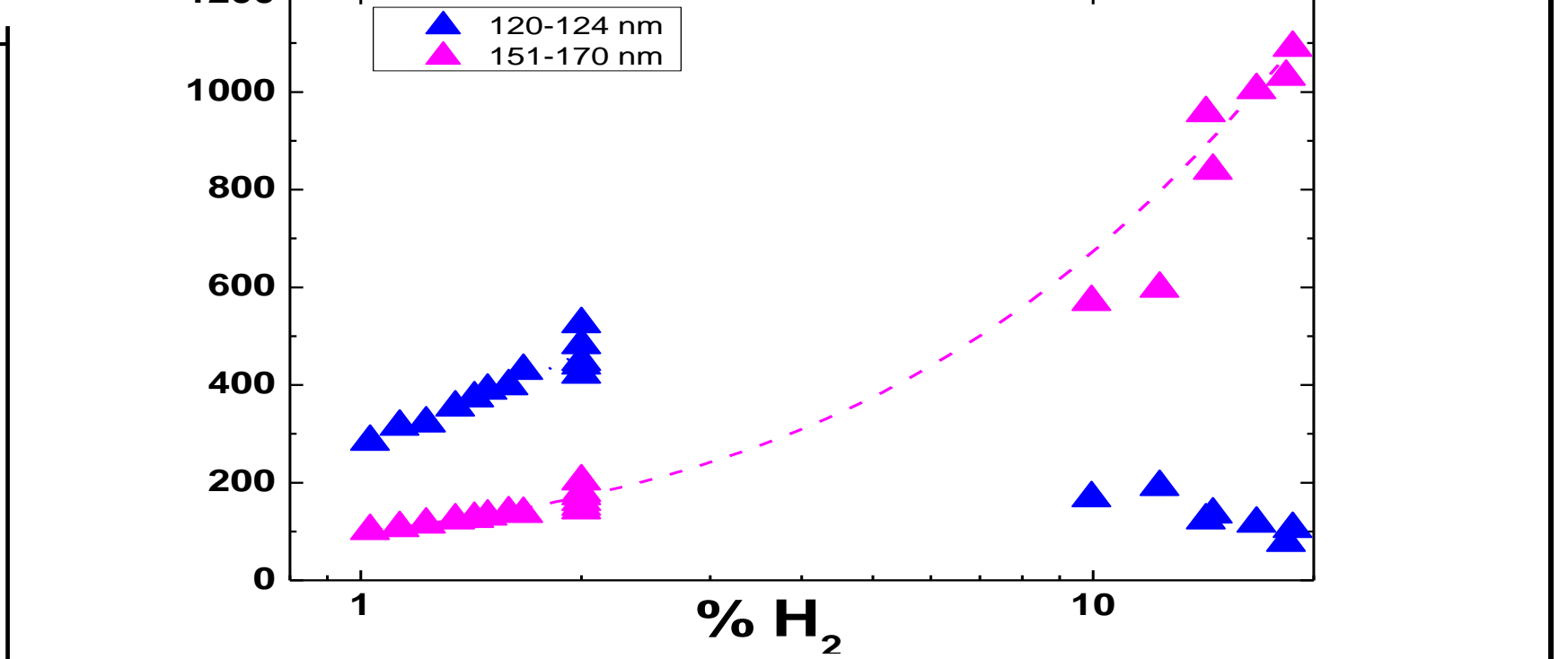


Fig. 6: Dependence of integrated line of H (Lyman alpha) and H<sub>2</sub> (B→X) with the % H<sub>2</sub> in He. p= 3.5 Torr and P=100 W.

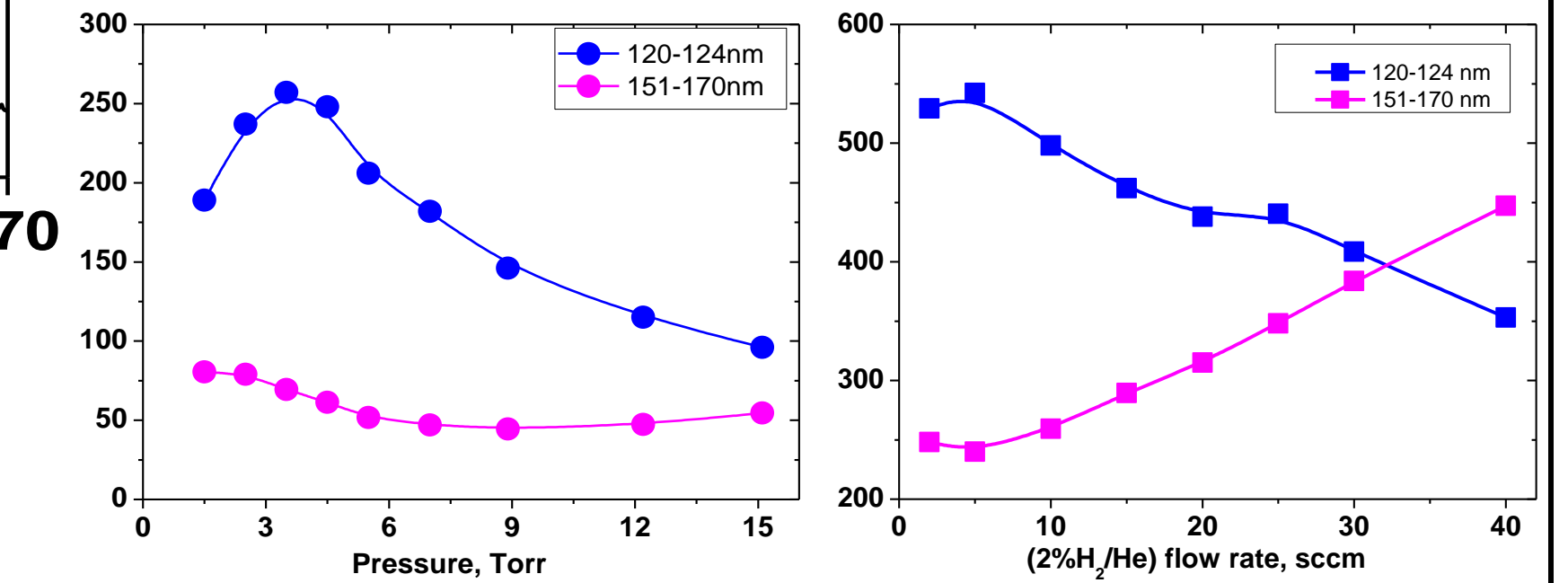


Fig. 7: Dependence of integrated line of H (Lyman alpha) and H<sub>2</sub> (B→X) with the gas pressure (left panel) and flow rate (right panel). 2% H<sub>2</sub>/He and P=100 W.

➔ The intensity of the H (Lyman alpha) emission line is optimized within plasma conditions (P=100W): ~ 2-5% H<sub>2</sub> in He, p = 3.5 Torr, flow rate ~ 5-10 sccm

#### 2. H(Lyman alpha)/H<sub>2</sub>(B→X) emission, effect of the addition of Ar in the H<sub>2</sub>/He mixture:

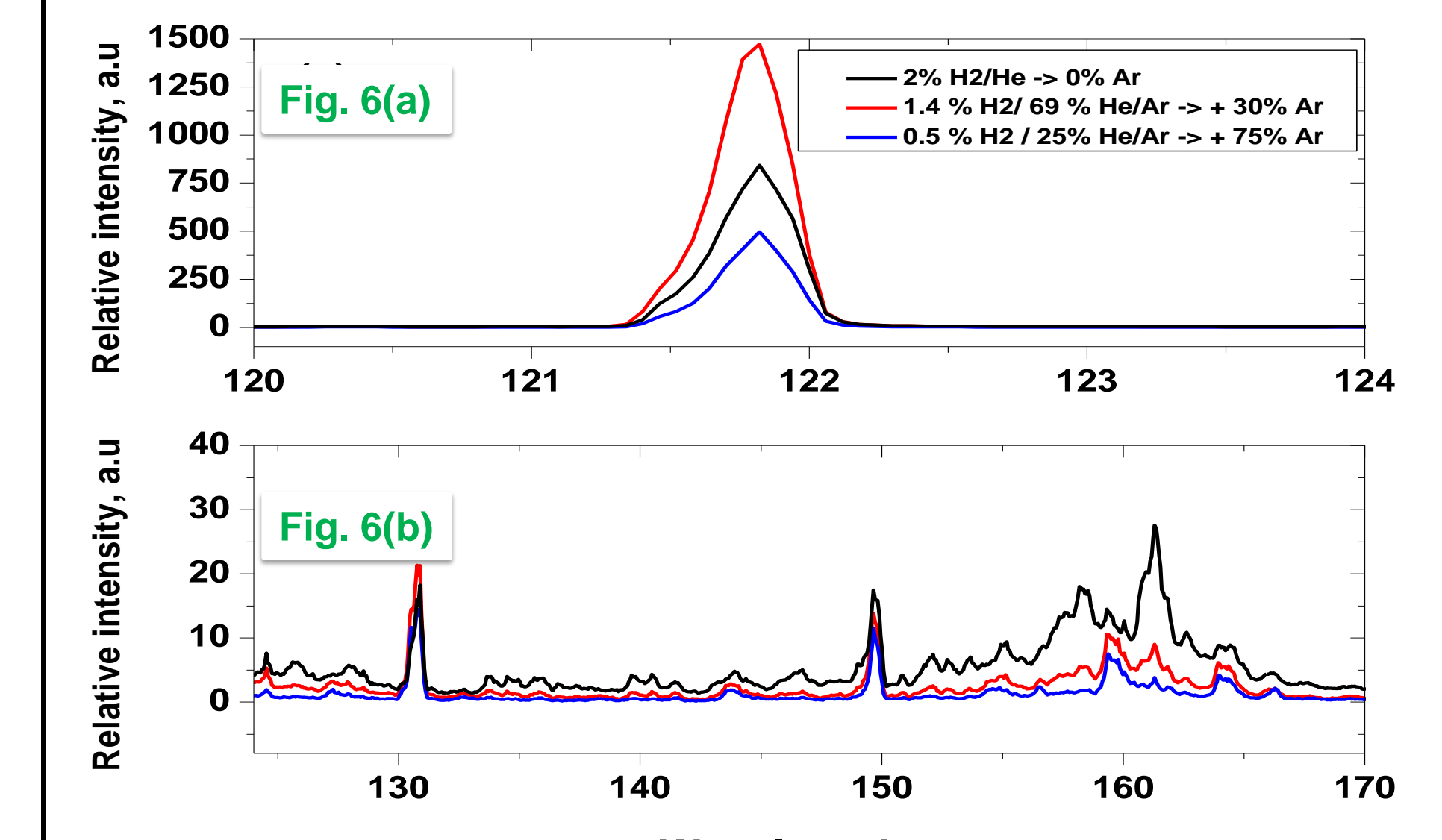


Fig. 6: VUV line emission in the range (a) 120-124 nm and (b) 124-170 nm as a function of the Ar addition in H<sub>2</sub>/He.

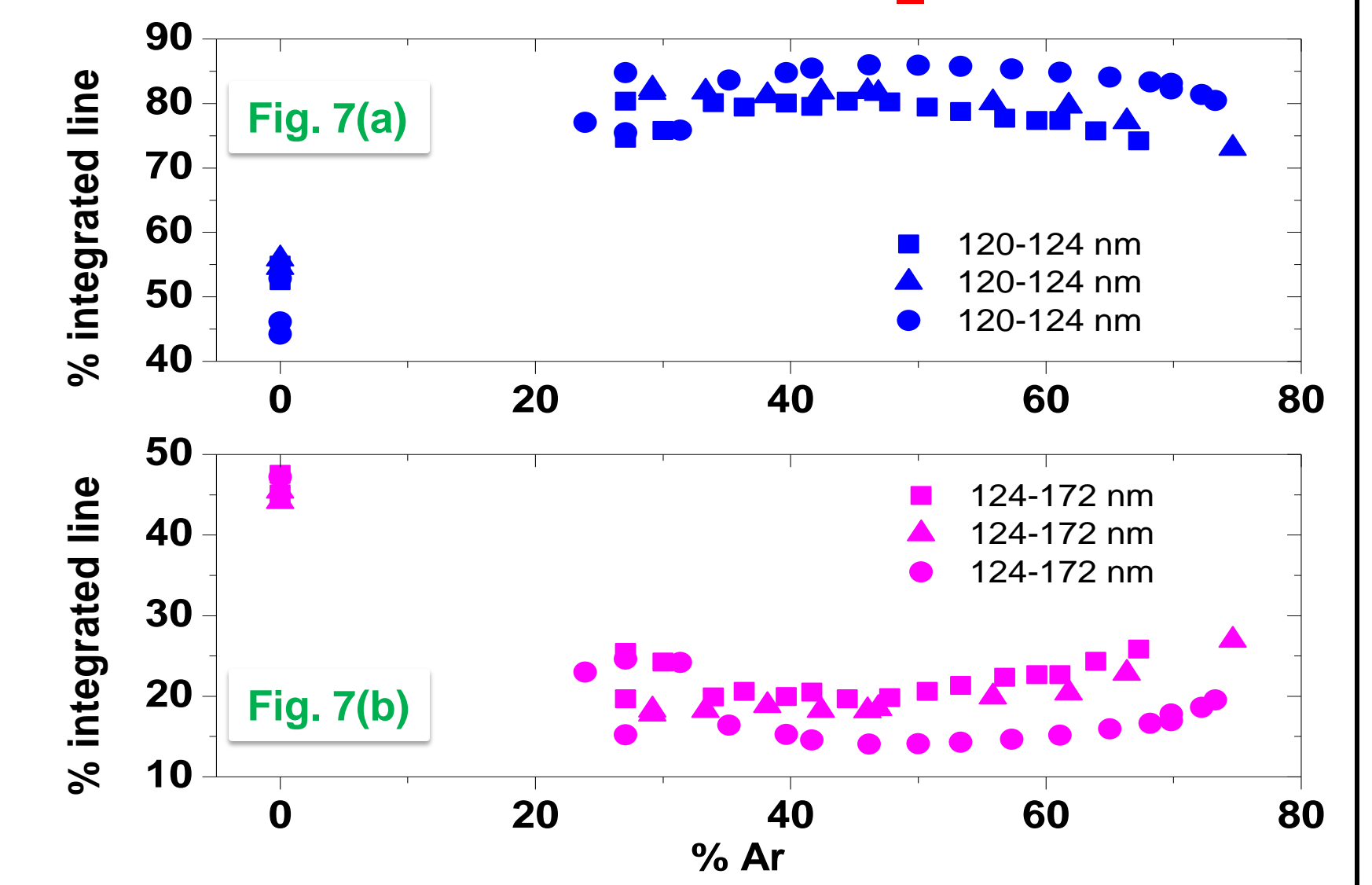


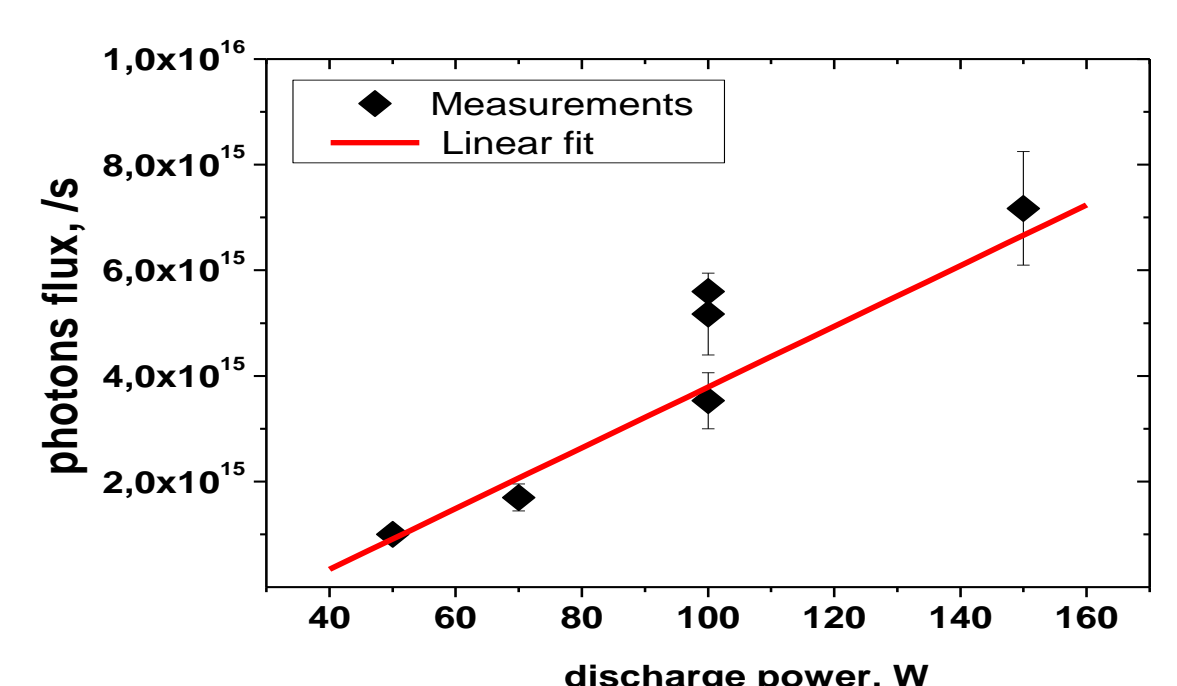
Fig. 7: integrated line intensity of (a) H (Lyman alpha) and (b) H<sub>2</sub> (B→X) as a function of %Ar added in H<sub>2</sub>/He. (Three series of measurements in the same plasma conditions)

➔ The addition of Ar in the H<sub>2</sub>/He gas mixture induces a maximum increase of the H (Lyman alpha) line intensity of 80% relative to the total emission flux.

Indeed, H atoms in the excited state are generated through energy transfer collisions between Ar(<sup>3</sup>P<sub>2</sub>) metastable and H<sub>2</sub>:  $\text{Ar}(\text{}^3\text{P}_2) + \text{H}_2 \rightarrow \text{Ar} + \text{H} + \text{H}$  and  $\text{Ar}(\text{}^3\text{P}_2) + \text{H}_2 \rightarrow \text{ArH}^* + \text{H}$ ; In parallel,  $\text{Ar}^* + \text{H}_2 \rightarrow \text{ArH}^* + \text{H}$ . Therefore, H (Lyman alpha) emission is enhanced through:  
 $\text{Ar}(\text{}^3\text{P}_2) + \text{H} \rightarrow \text{ArH}^* \rightarrow \text{ArH}^* + \text{H} (n=2)$   
 $\text{H} (n=2) \rightarrow \text{H} (n=1) + (\text{Lyman } \alpha = 121.6 \text{ nm})$   
For high Ar amount added, the H (Lyman alpha) intensity drops and the H<sub>2</sub>(B→X) intensity increases due to quenching of Ar(<sup>3</sup>P<sub>2</sub>) metastable by H<sub>2</sub>.

#### 3. Photon Flux at Lyman alpha (121.6 nm)

Fig. 8: Dependence of the photon flux on the discharge power. The experimental conditions are: flow rate=15 sccm, 2% H<sub>2</sub>/He+30%Ar, p=3.5 Torr. The photon flux was determined through Eq. 1.



#### 4. CH<sub>4</sub> Photolysis at Lyman alpha (121.6 nm)

Methane photolysis at Lyman alpha is studied in the frame of a program dedicated to simulations of Titan's atmosphere [2]. Irradiations have been performed using a microwave He/H<sub>2</sub> (98/2) discharge lamp. The chemical evolution of the gas mixture resulting from the irradiations is monitored by FTIR. The determination of the chemical mechanism involved requires the comparison with a dedicated OD kinetic model in which photolysis rates are fundamental parameters. Their values depend on the emission spectrum of the source.

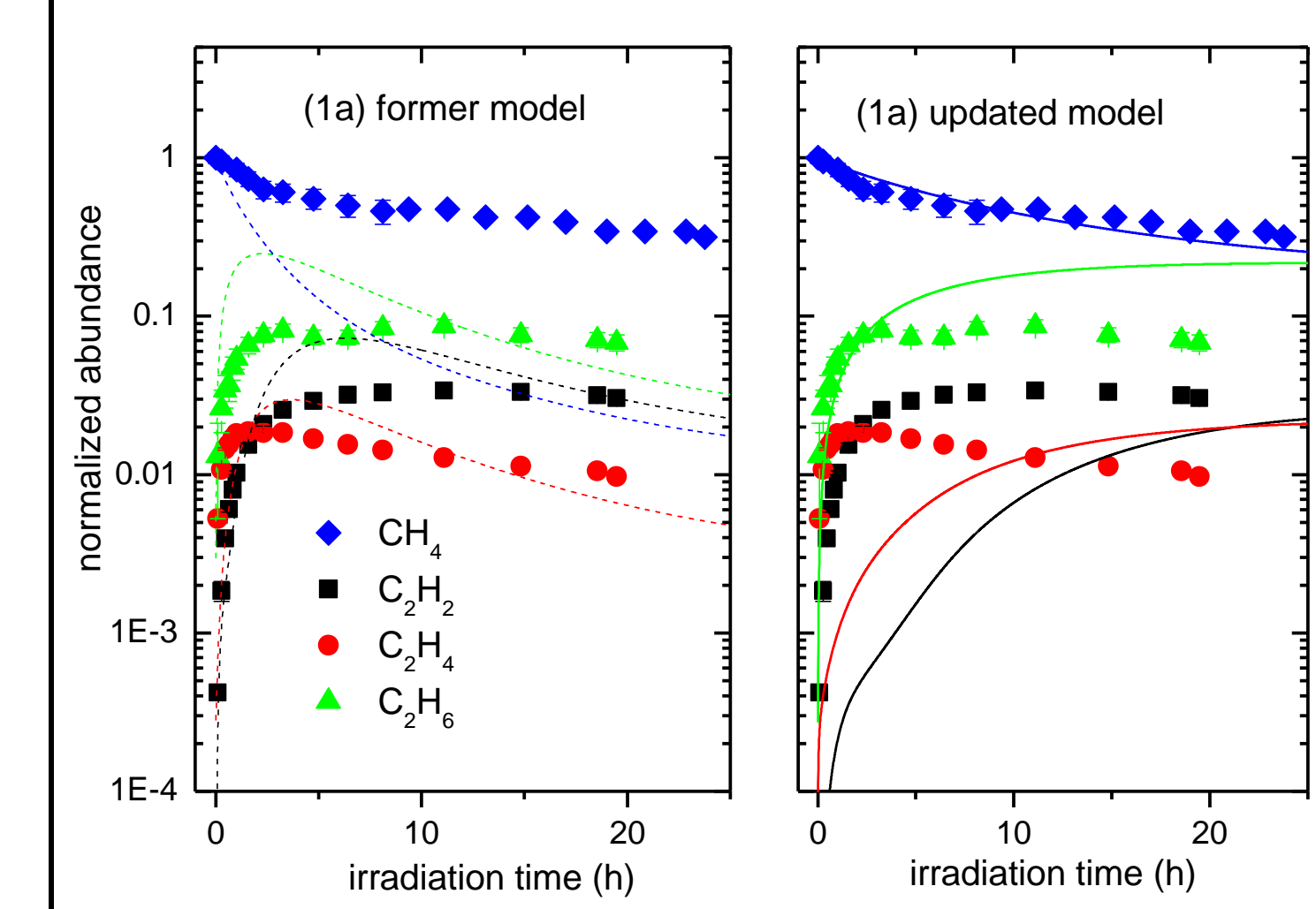


Fig. 9: Comparison between experimental abundances and calculated abundances obtained with the former model (left panel), and the improved model (right panel).

In a former model [2], the comparison of experimental and theoretical results was not entirely satisfactory since the destruction of methane was overestimated by the model as well as the production of ethane.

Recently, the fact that lamp is not monochromatic at Lyman alpha, as shown above, has been taken into account. In the updated model, new values for the photon flux,  $\Phi_{120-180}^A$  and the photolysis rates,  $J_{120-180}^A$ , have been used since CO<sub>2</sub> as well as CH<sub>4</sub> and C<sub>2</sub> compounds absorb in the emission domain of the lamp (120-180 nm):

$$d[CO]/dt = \int_{120}^{180} \Phi_{\lambda} \cdot \eta \cdot \beta \cdot d\lambda$$

$$\Phi_{120-180}^A = \int_{120}^{180} \Phi_{\lambda} \cdot d\lambda$$

$$\text{and } J_{120-180}^{CH_4} = \int_{120}^{180} (\sigma_{\lambda}^{CH_4} \cdot \eta_{\lambda}^{CH_4} \cdot \Phi_{\lambda} / S) \cdot d\lambda \quad S: \text{irradiated surface (cm}^2\text{)}$$

➔ This improvement allows a better agreement between experimental and theoretical data concerning CH<sub>4</sub> and C<sub>2</sub>H<sub>6</sub>. Reliable comparison can only be achieved when the irradiation source is well characterised in terms of flux and wavelength dependency.

### REFERENCES

[1] Jacquinet-Husson, N et al, The (2003) GEISA/IASI spectroscopic database. *Journal of Quantitative Spectroscopy and Radiative Transfer* 95 (4),429–467.  
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[3] Yoshino, K., Esmond, J.R., Sun, Y., Parkinson, W.H., Ito, K., Matsui, T., (1996). *Journal of Quantitative Spectroscopy and Radiative Transfer* 55 (1), 53–60.

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