

# The baroclinic instability of an initially stratified fluid layer

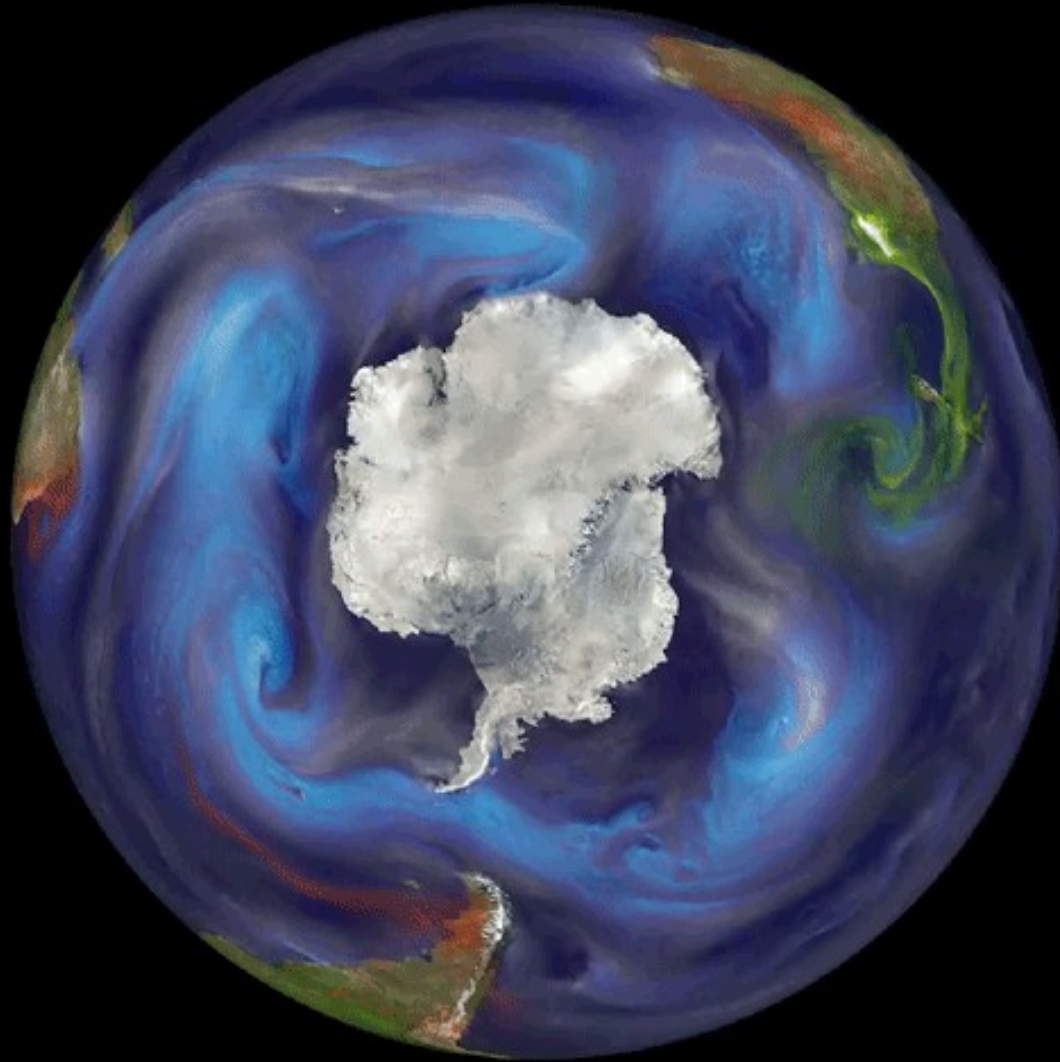
Patrice Le Gal, Miklos Vincze and Uwe Harlander

Institut de Recherche sur les Phénomènes Hors Equilibre  
CNRS, Aix-Marseille Université, France

Hungarian Academy of Sciences-Eötvös University  
Theoretical Physics Research Group, Budapest, Hungary

Department of Aerodynamics and Fluid  
Mechanics Brandenburg University of Technology, Cottbus, Germany

# General motivation



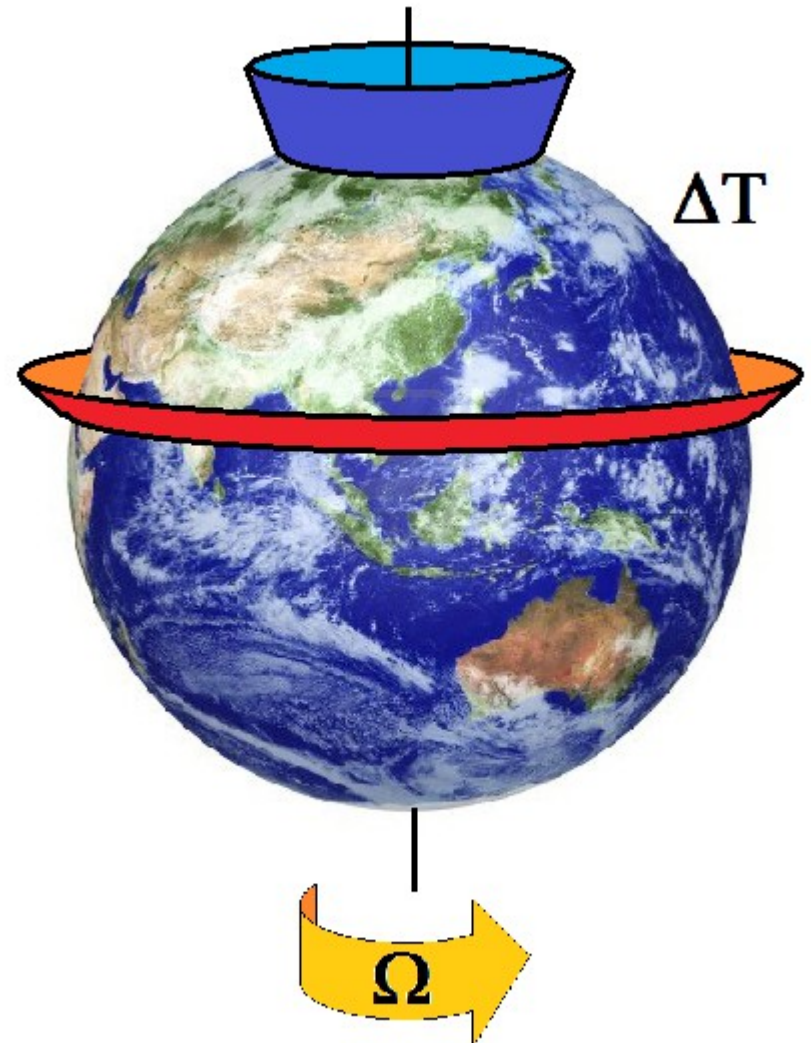
*Credit NASA: Aerosol transport patterns from the GOES-5 satellite data.*



<http://gmao.gsfc.nasa.gov/research/aerosol/>

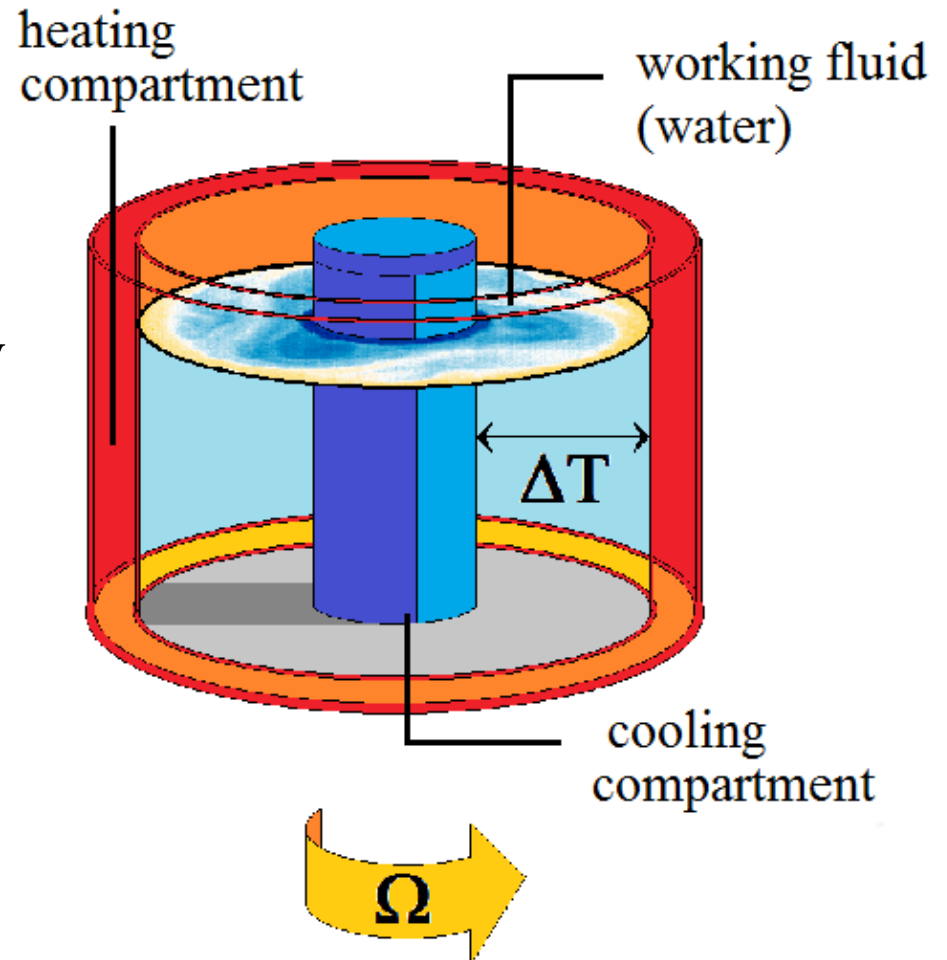
# A minimal model of weather-like dynamics

- A large variety of the typical atmospheric phenomena of the mid-latitudes are primarily driven by **two factors** only.
- *Rotation + meridional temperature difference  $\approx$  weather*



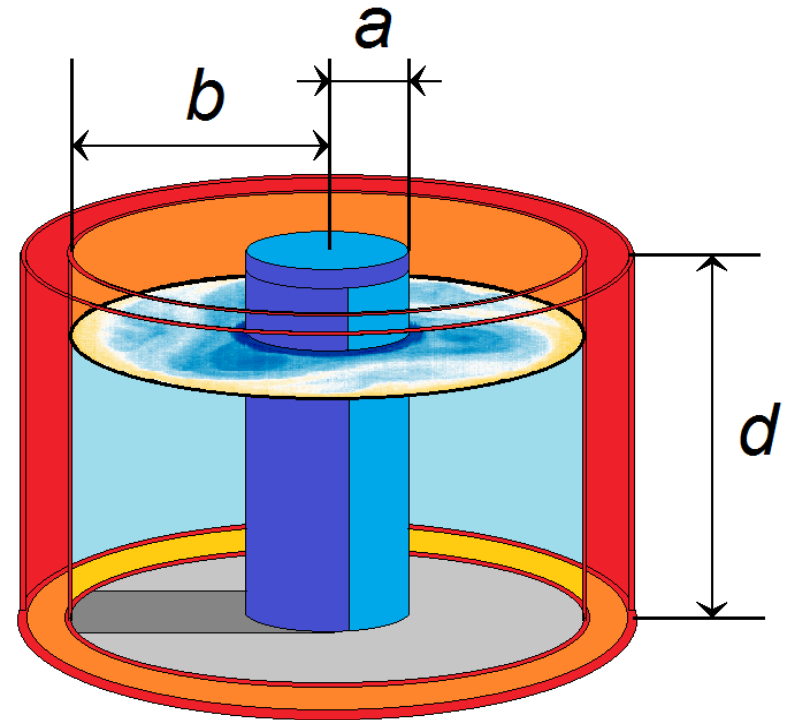
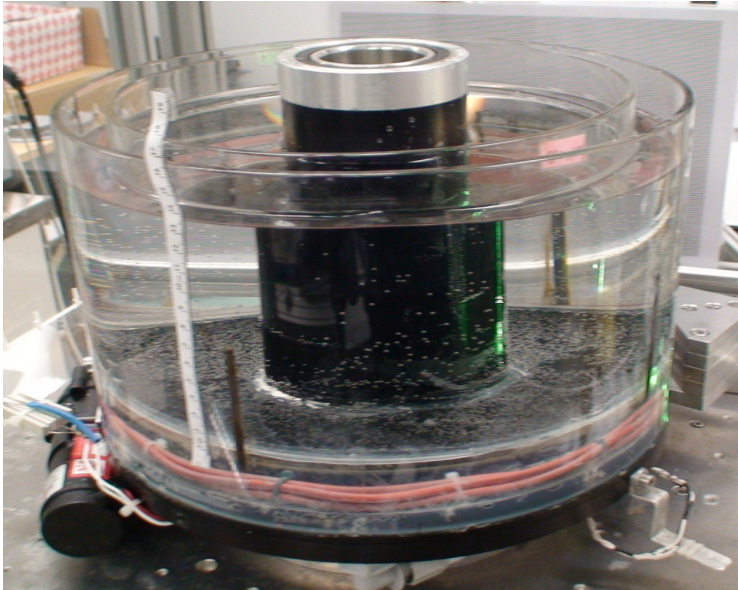
# A minimal model of weather-like dynamics

Laboratory experiments  
On the baroclinic instability



# CoGeoF1 platform

## Brandenburg University of Technology (BTU) Cottbus



### Geometric parameters:

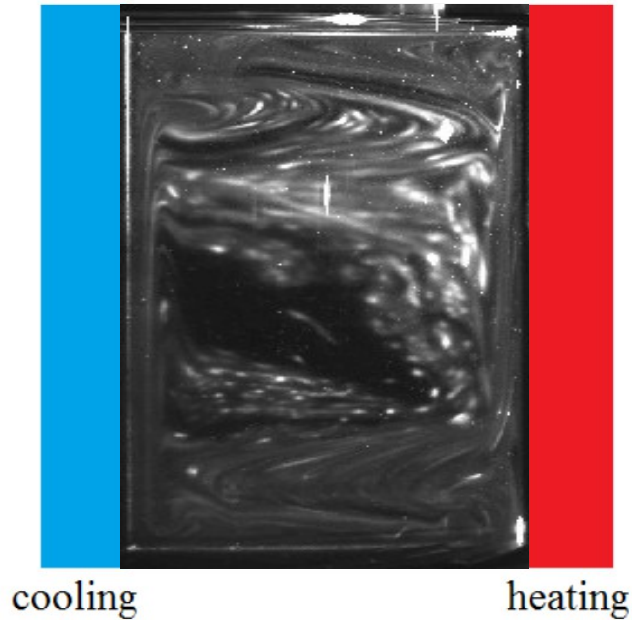
$$a = 45 \text{ mm}$$

$$b = 120 \text{ mm}$$

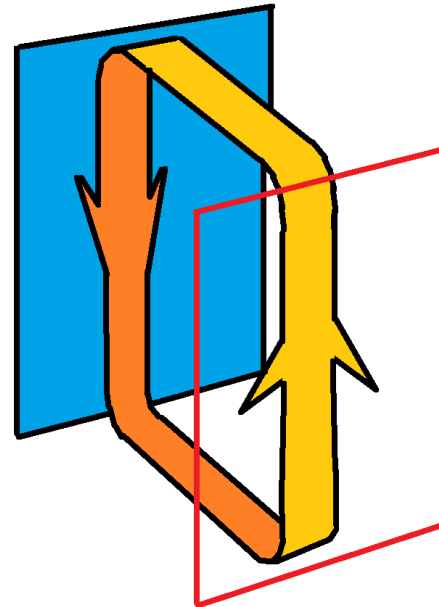
$$d = 135 \text{ mm}$$



# Basics: baroclinic instability



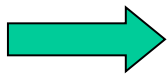
No Rotation



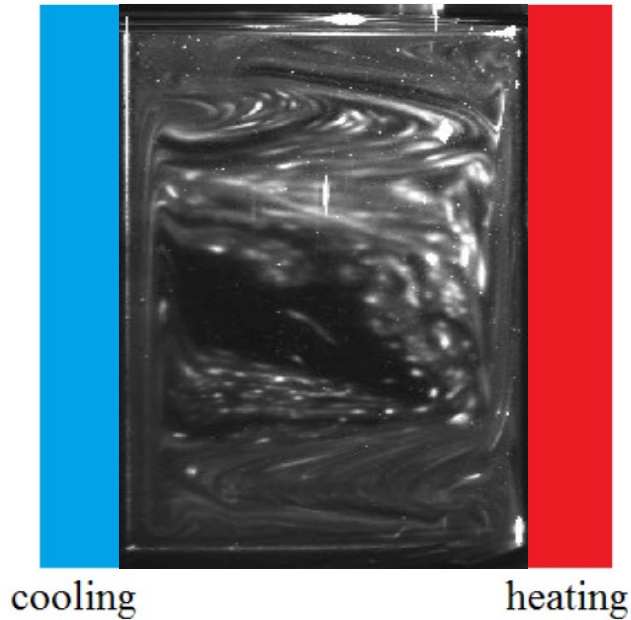
**“Sideways convection”** – no threshold  
in  $\Delta T$   
(i.e. No ‘critical Rayleigh number’)

***Any*** temperature difference can initiate  
the flow





# Baroclinic instability



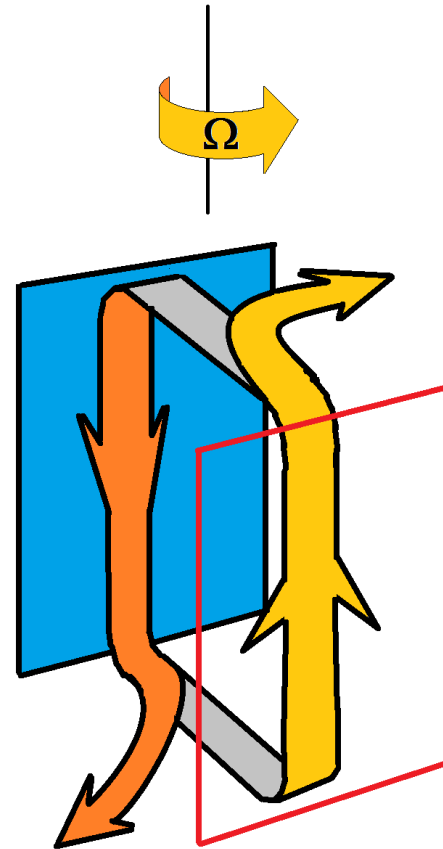
**Rotation!**



Zonal flow  
(thermal wind)

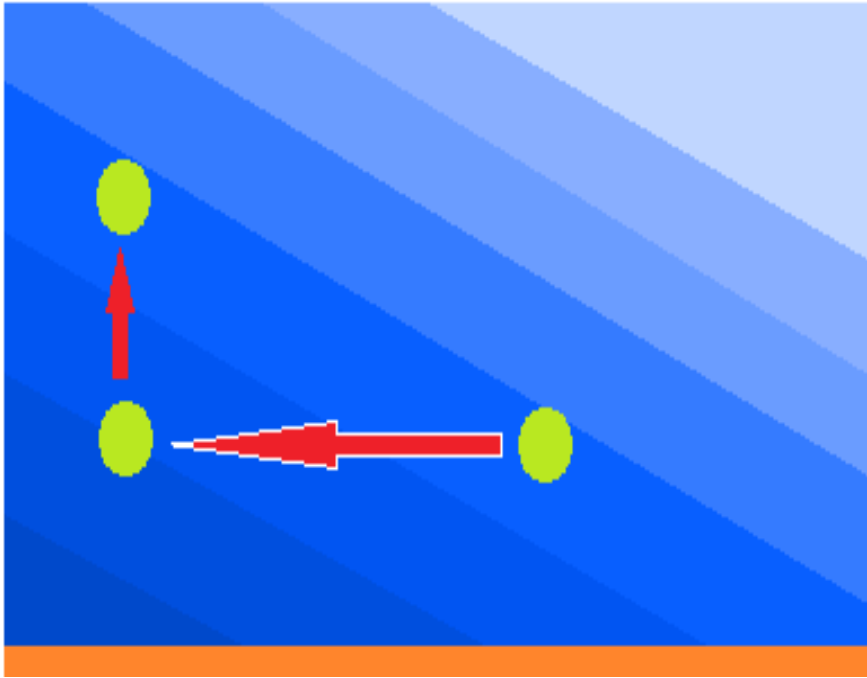


Geostrophic theory:  
**Tilted density  
surfaces**



$$-2\Omega \vec{e}_z \times \vec{u}$$

# Baroclinic instability



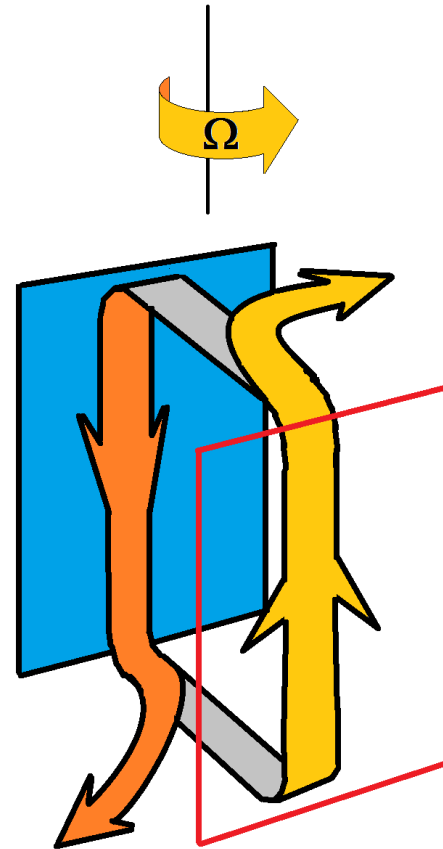
**Rotation!**



**Zonal flow**  
(thermal wind)



**Geostrophic theory:**  
**Tilted density**  
**surfaces**



$$-2\Omega \vec{e}_z \times \vec{u}$$



**INSTABILITY**





# Experimental baroclinic instability set-up mimics: Convective cells, jets and zonal winds, Rossby waves of planetary atmospheres

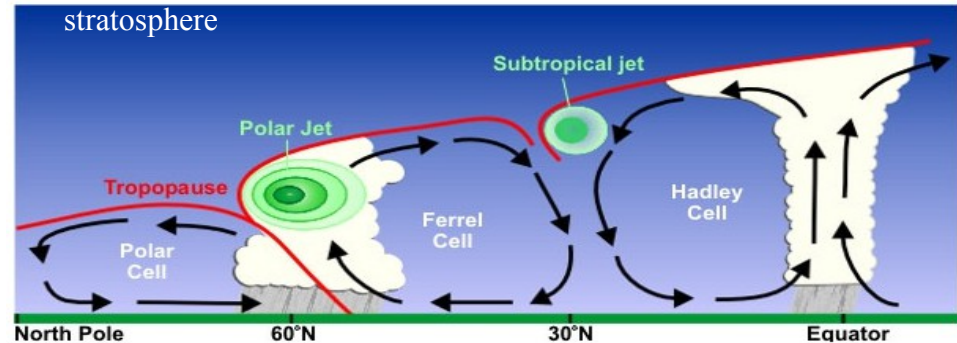
## - control parameters:

rotation rate, radial  
temperature difference

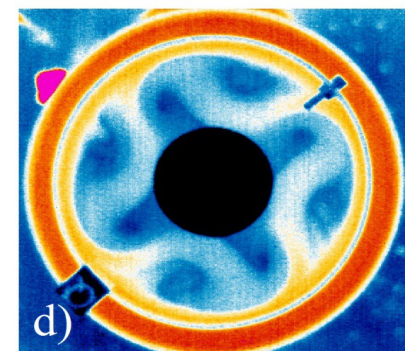
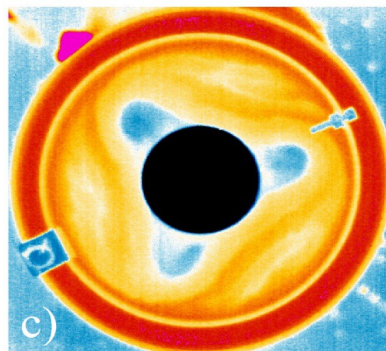
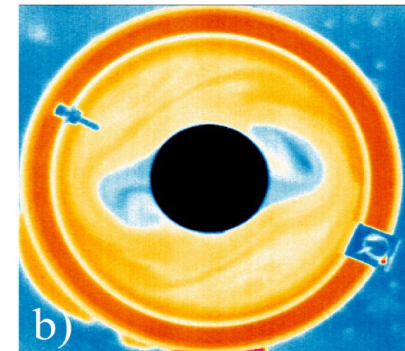
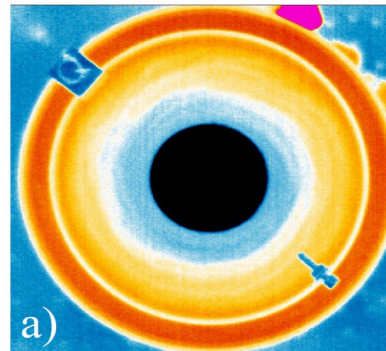
- Different planetary  
atmospheres can be  
modelled

Venus: slow rotation, zonal  
flow

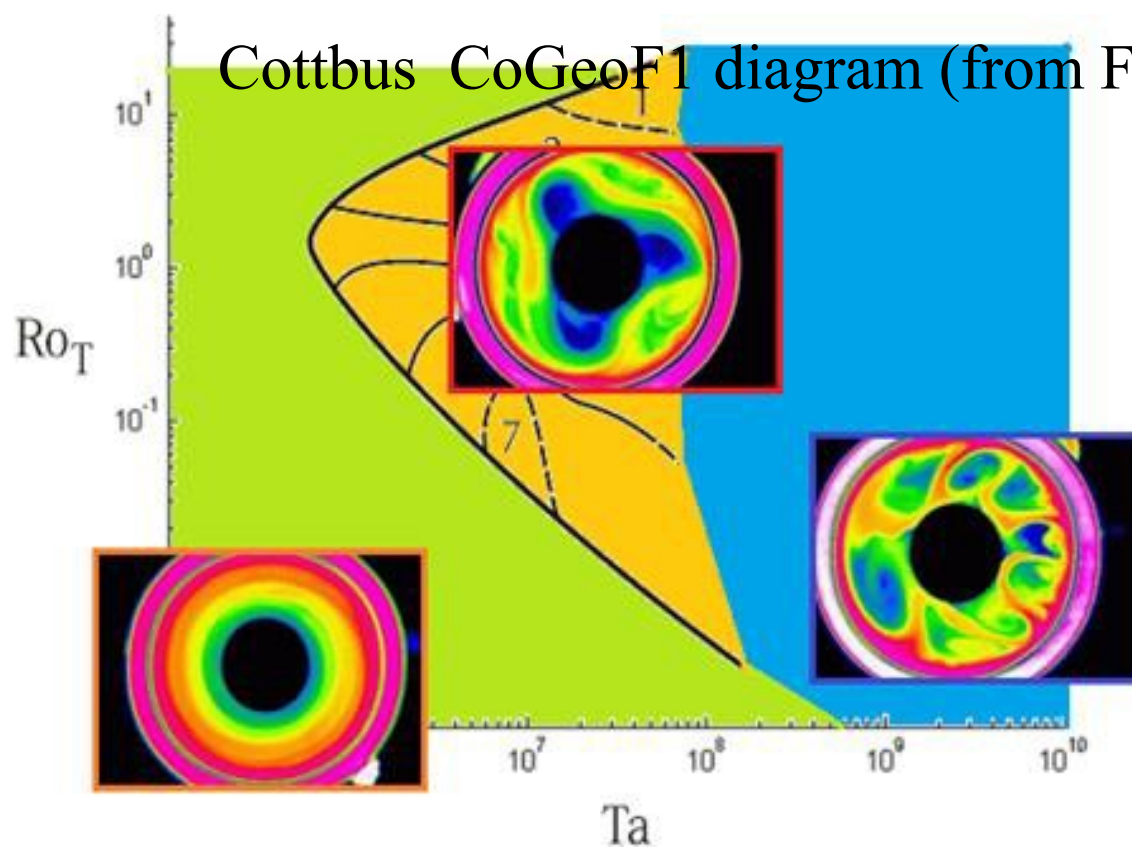
Earth: fast rotation →  
Coriolis effect → vortices  
("weather")



*Credit : wikipedia*



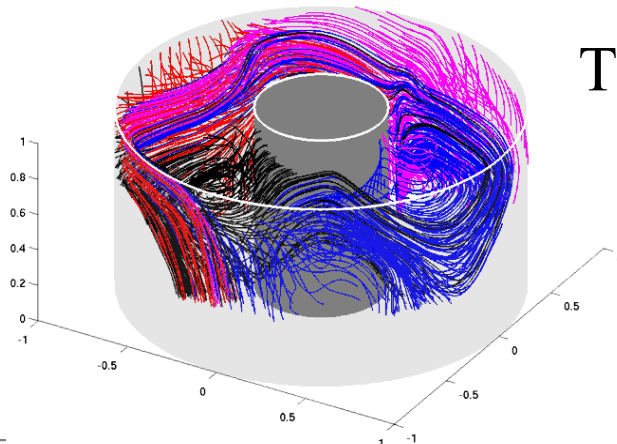
# Cottbus CoGeoF1 diagram (from Fultz):



$$Ta = \frac{4 \cdot \Omega^2 \cdot (b - a)^5}{\nu^2 \cdot d}$$

$$Ro = \frac{g \cdot d \cdot \alpha \Delta T}{\Omega^2 \cdot (b - a)^2}$$

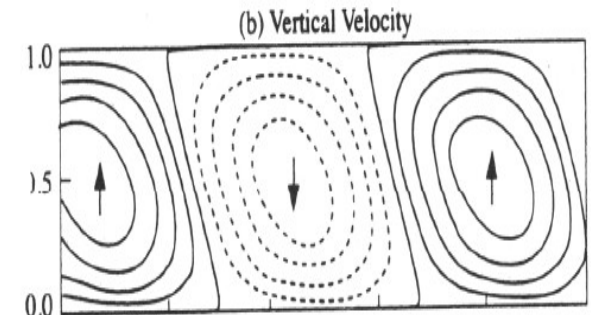
Tilted columnar vortices



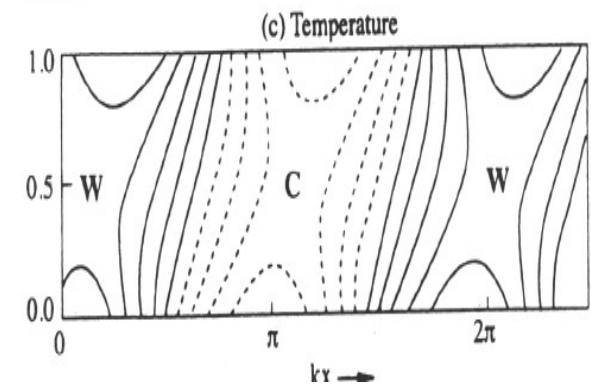
3D flow field from simultaneous PIV

thermography observations

$z/d$



$z/d$



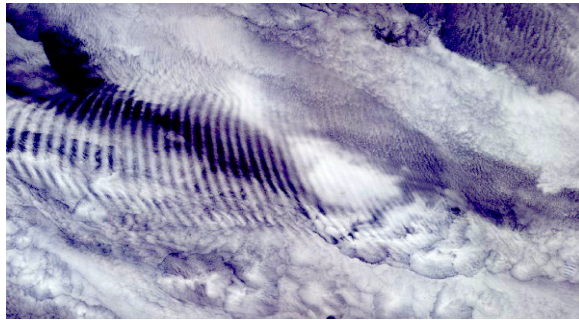
Credit : [www.atmosp.physics.utoronto.ca](http://www.atmosp.physics.utoronto.ca)

# Can we add a stratosphere ?

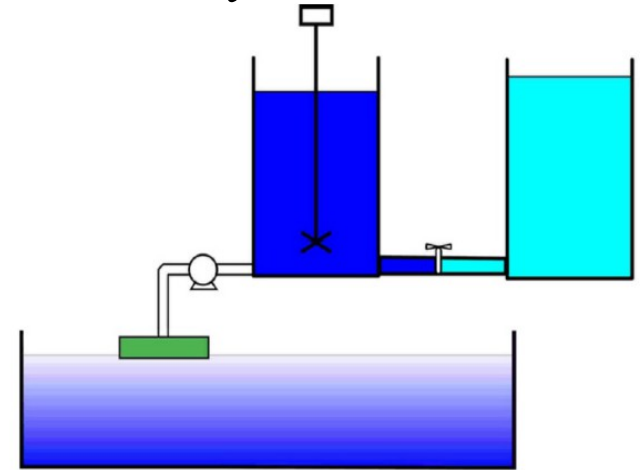
- To get the « right » horizontal stress free Boundary Condition
- To generate « pancake » vortices in rotating and stratified flows : Jupiter Great Red Spot
- To excite Gravity Waves from eventual unbalanced dynamics



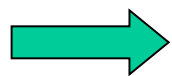
*Credit: NASA*



*Credit: <http://www.planetpals.com>*



**Salinity stratification:**  
**“Double-bucket” method**



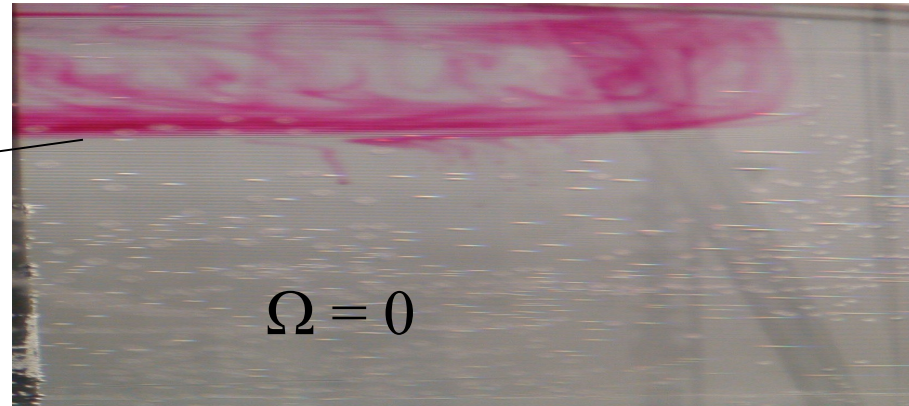
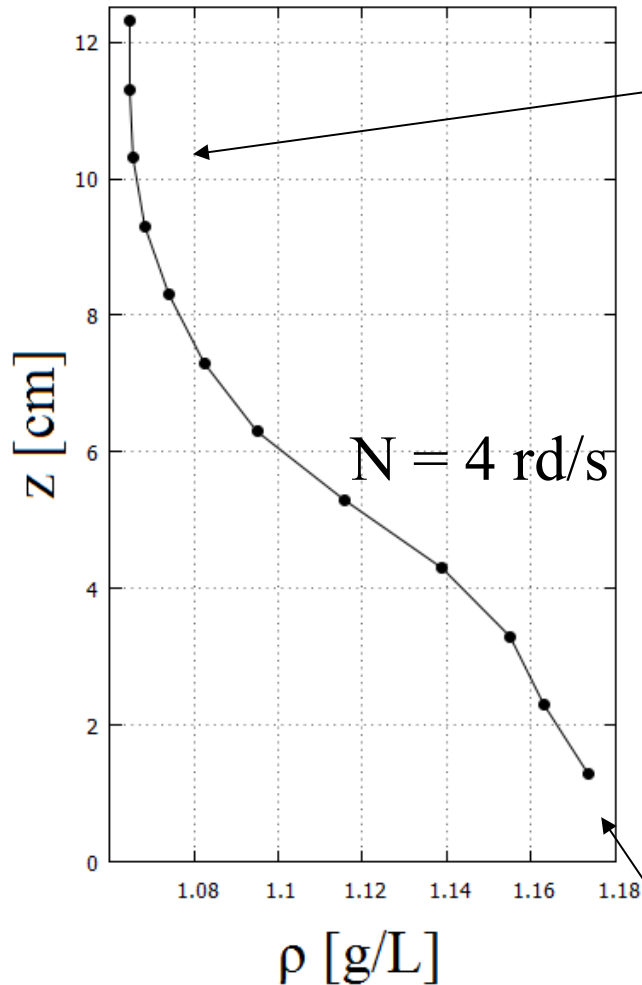
The baroclinic instability of an initially stratified fluid layer :

**The BAROSTRAT Instability**

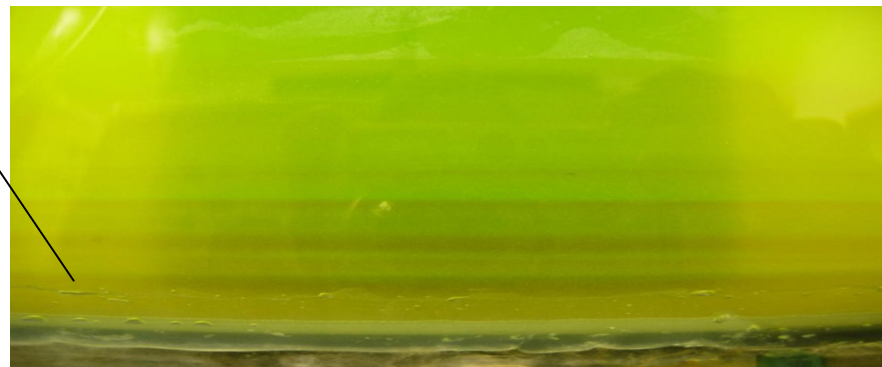




# Stratification will oppose to convection !

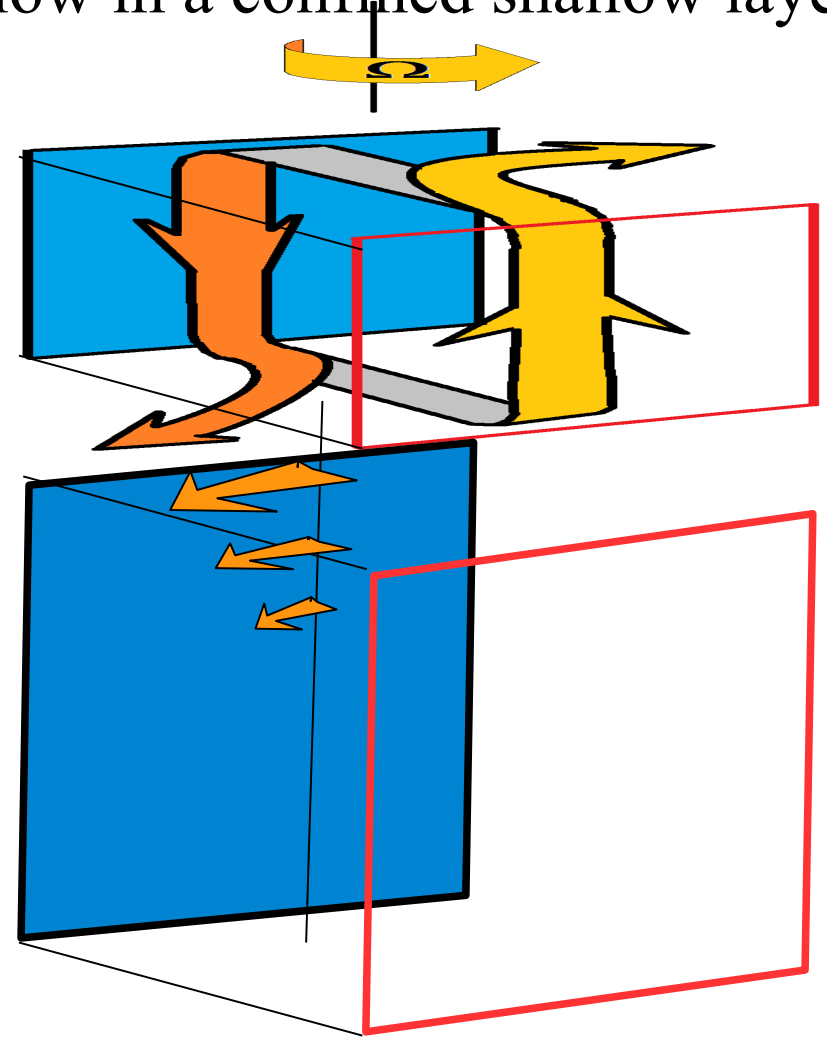
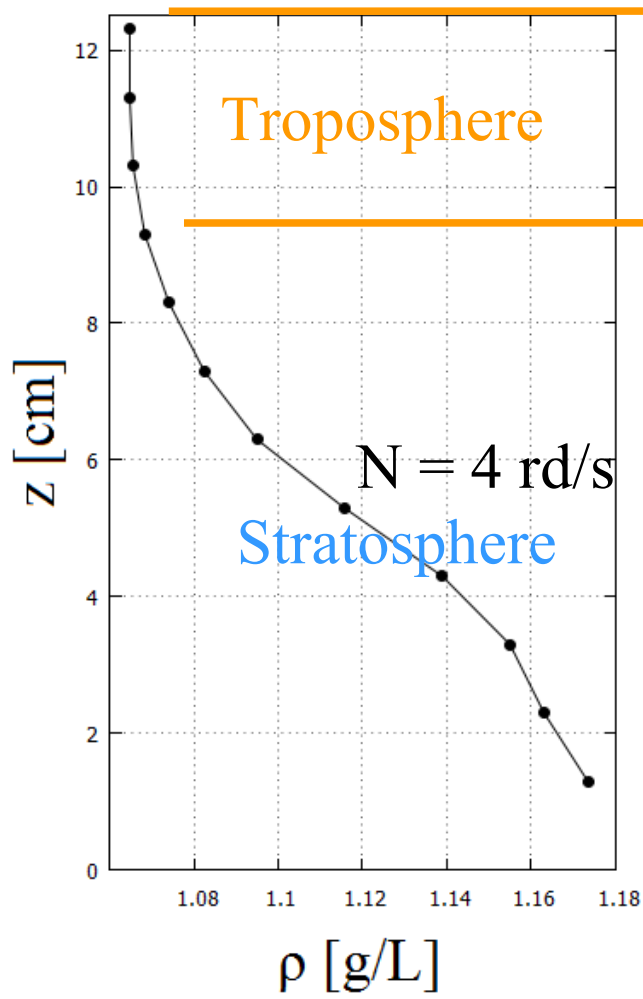


Above a temperature threshold,  
convection confined in weakly  
stratified shallow layers  
(Lee, Kang, Son 1999)



Double diffusive staircase cells at bottom

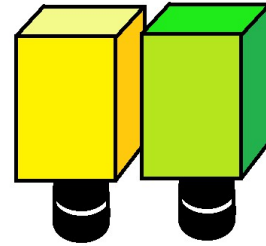
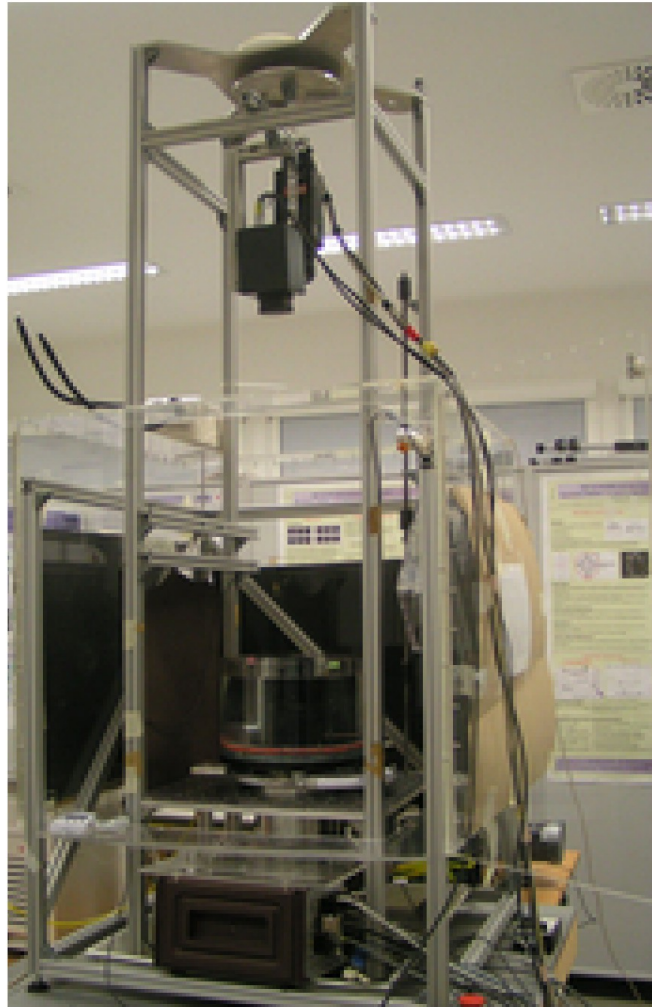
➡ Rotation will initiate zonal flow in a confined shallow layer !



$f \sim 0.4$  rd/s

➡ Above a critical  $\Omega$  : Baroclinic Instability in a shallow layer  
above a stratified layer

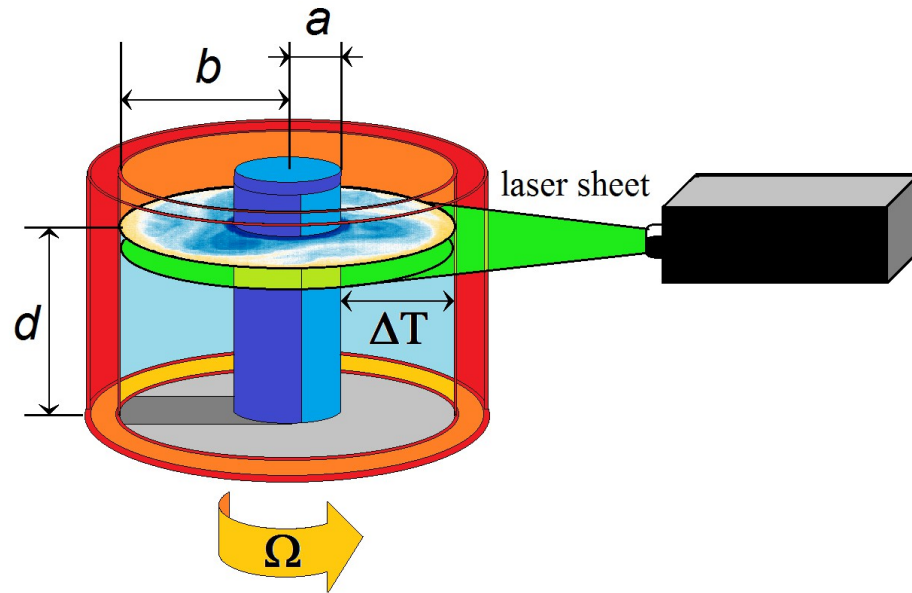
# Measurement techniques

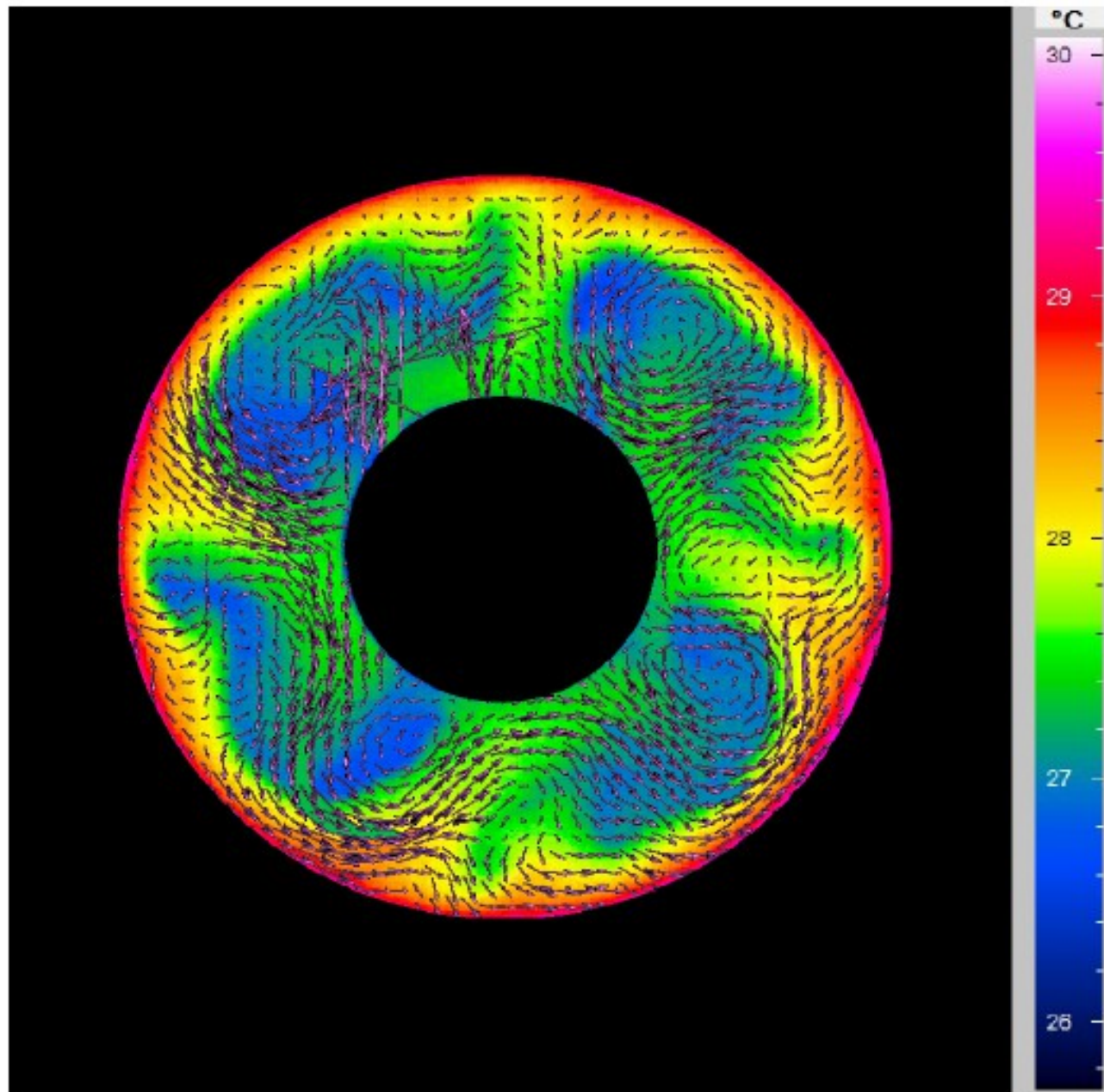


IR camera

PIV camera

**Co-rotating simultaneous  
IR thermography and PIV**

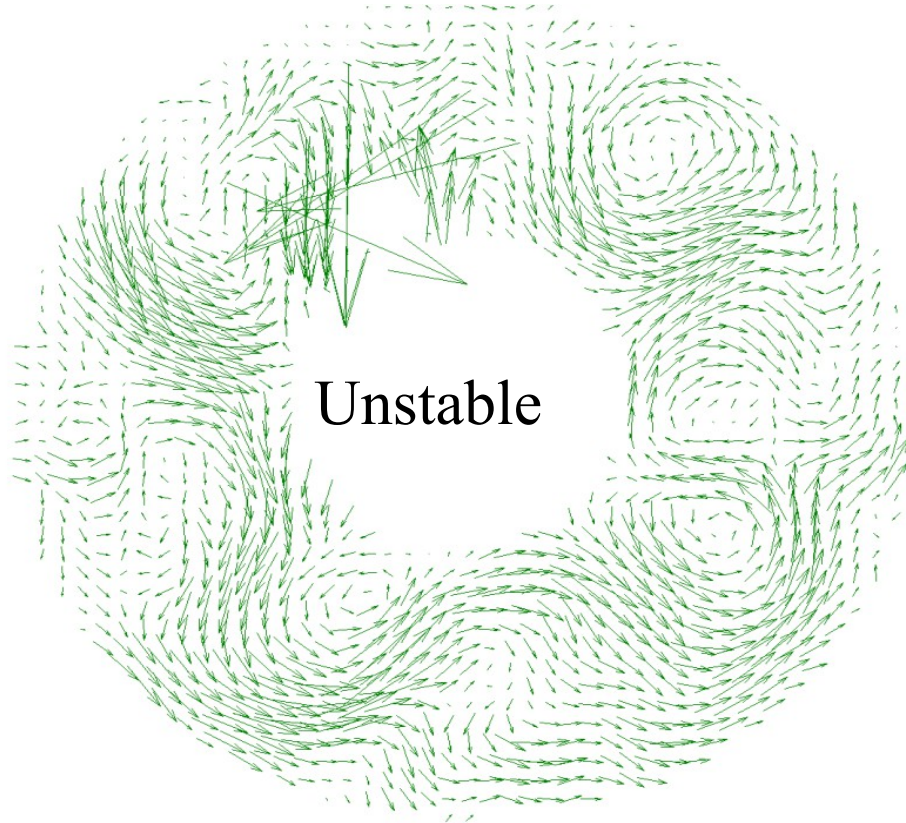




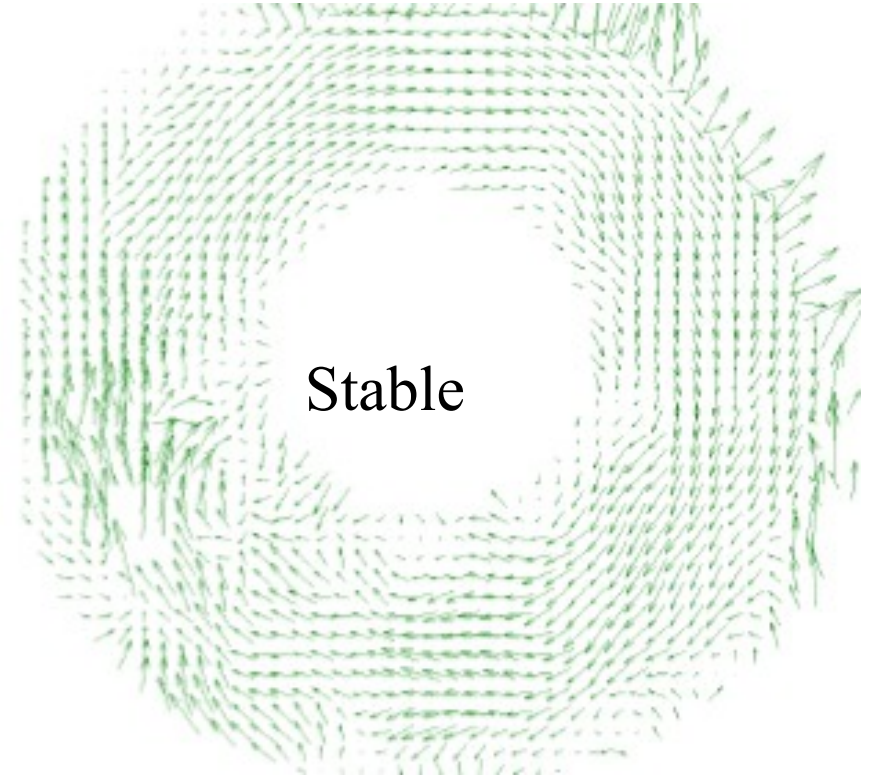
$\Delta T = 6 \text{ K}$   
 $\Omega = 2.7 \text{ rpm}$



PIV field near  
surface



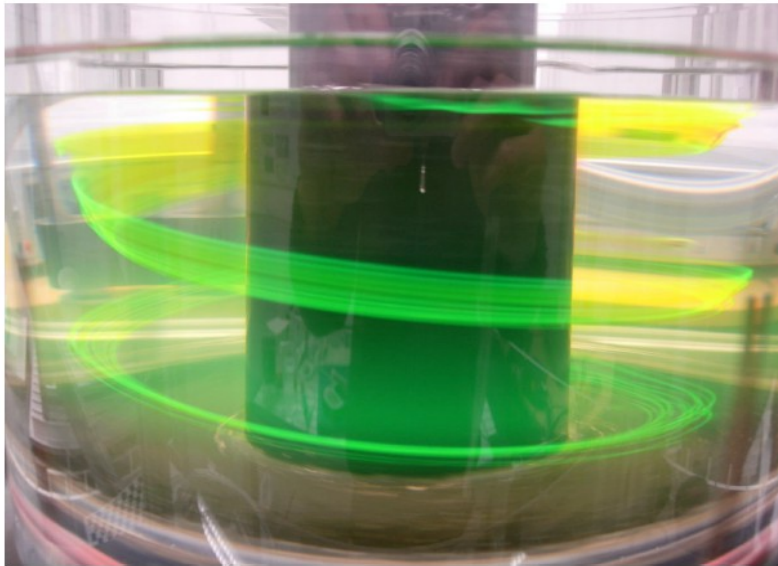
PIV field near  
bottom



Zonal flow only



Fluoresceine visualization  
of thermal wind

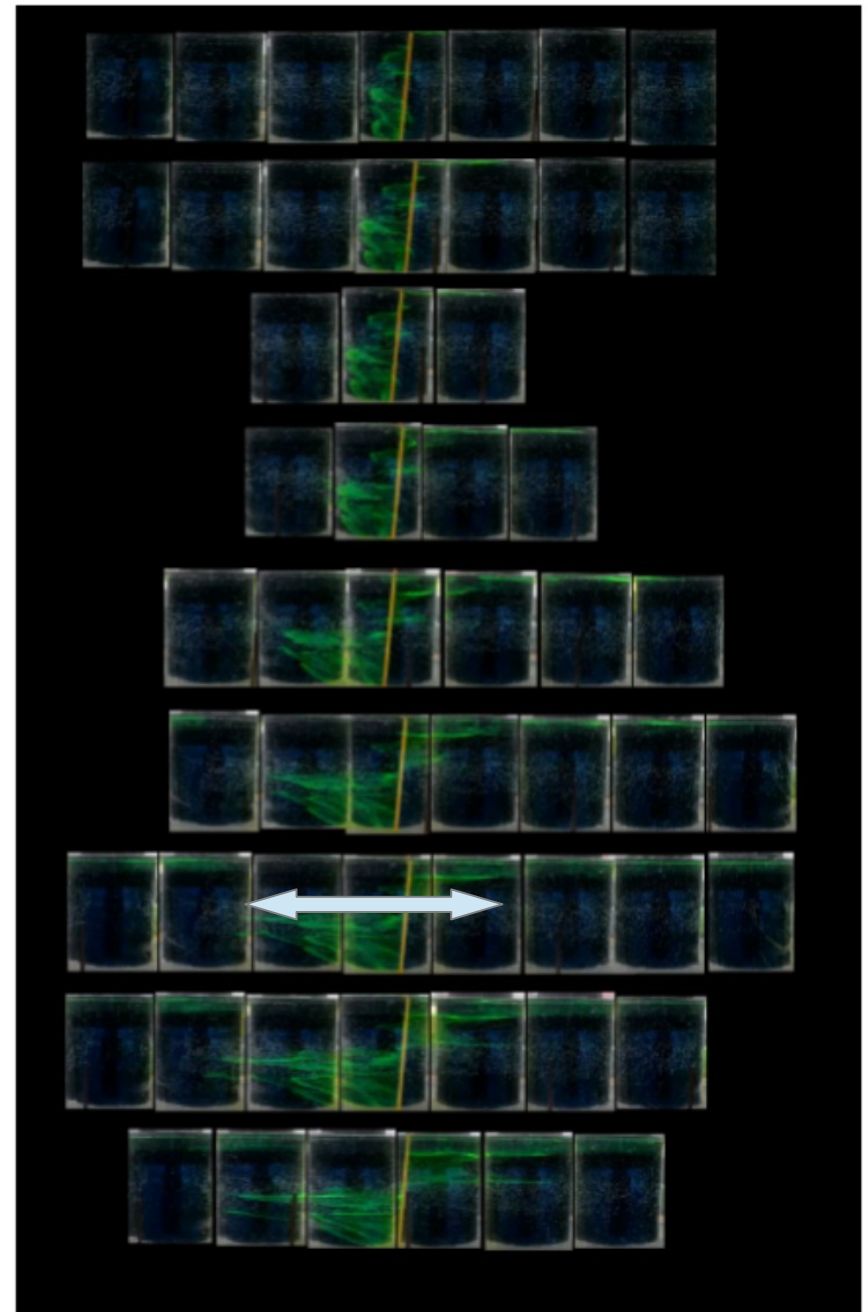


Note the stretching of dye  
filaments between top and  
bottom of the convective cell

$t=0$

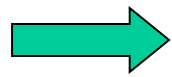
$\vdots$

$t=180s$

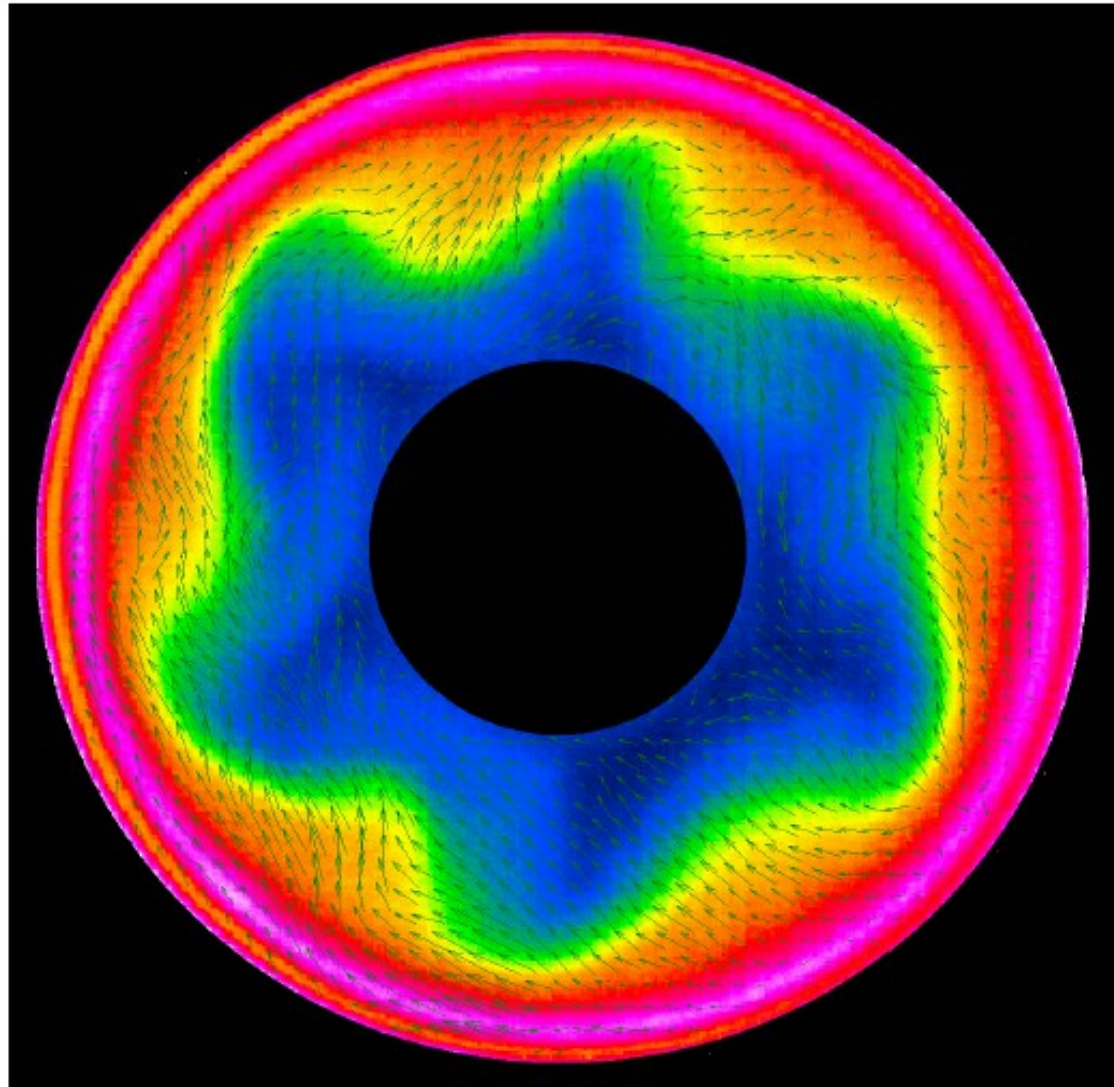


azimuthal direction





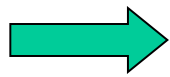
Sucking top mixed layer : decreases its thickness and changes azimuthal wave number



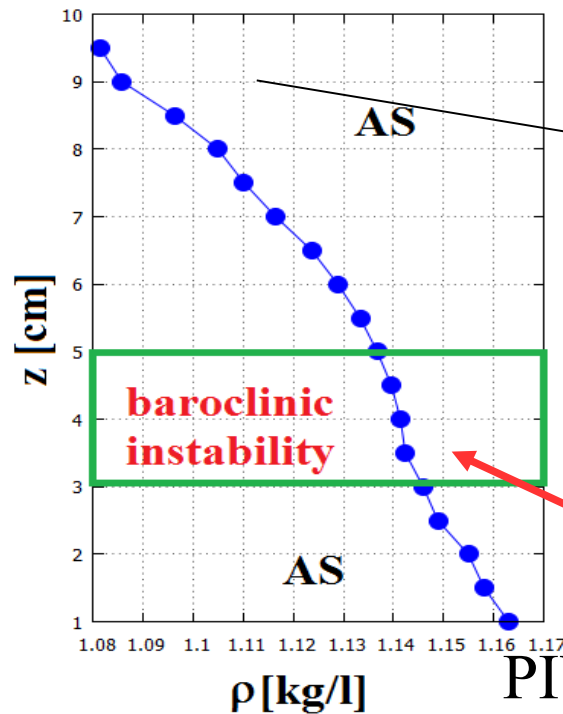
$$\Delta T = 6 \text{ K}$$

$$\Omega = 1.7 \text{ rpm}$$

*Combined thermographic/PIV image (surface temperature + PIV data from  $h = 12 \text{ cm}$ )*



Sucking out completely top mixed layer on a vertical  
Density profile with a zero gradient inner layer



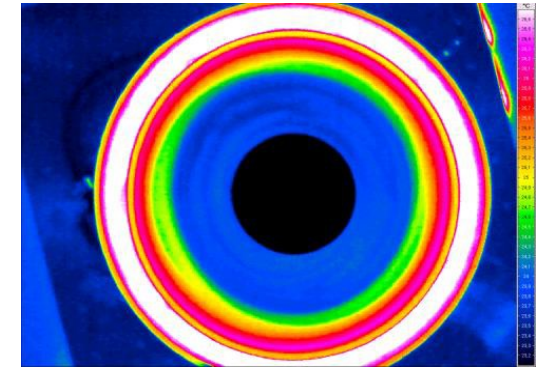
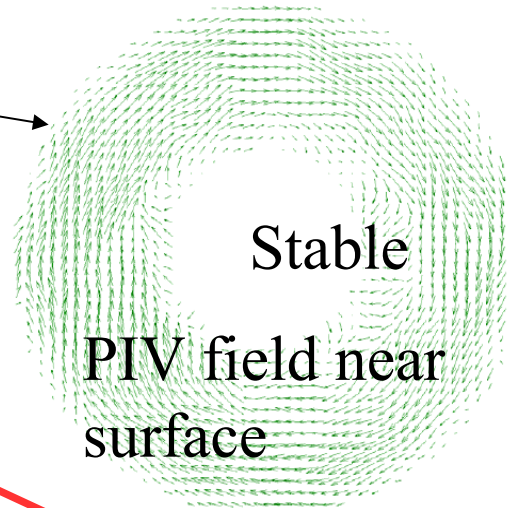
AS

baroclinic  
instability

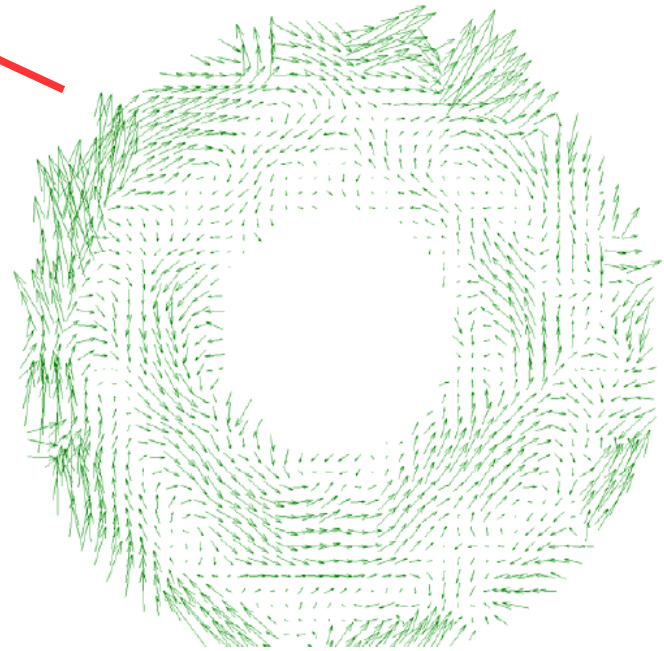
AS

$\rho$  [kg/l]

$z$  [cm]



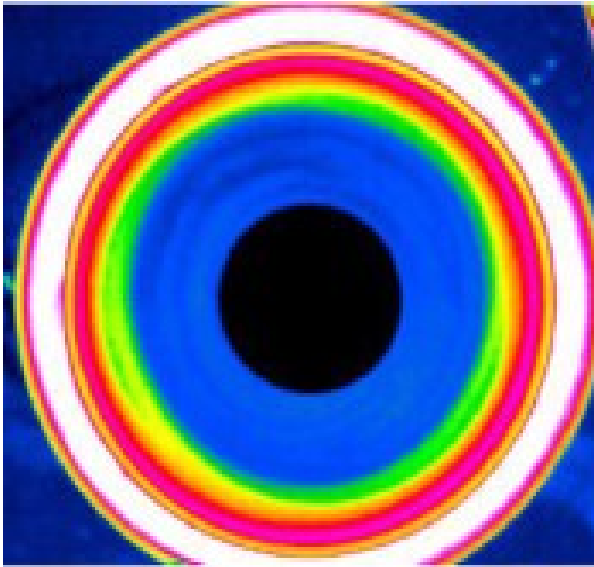
PIV field in  
inner layer



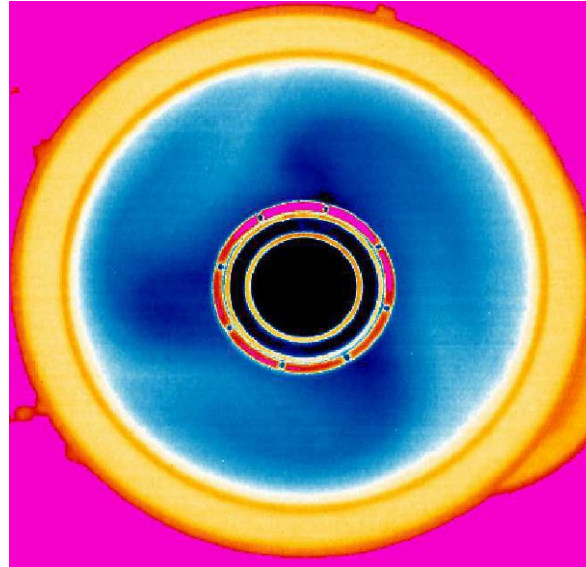
Unstable in a confined inner layer :  
Chain of pancake vortices between  
two stratified regions

# Zoology of different patterns of the « Barostrat instability »

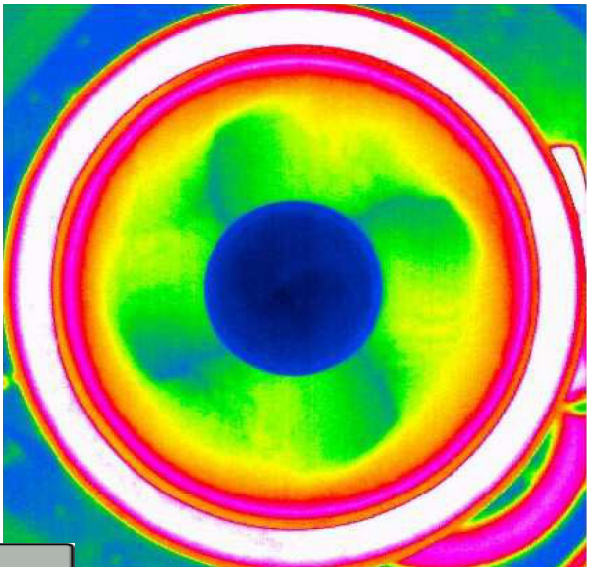
$m = 0$



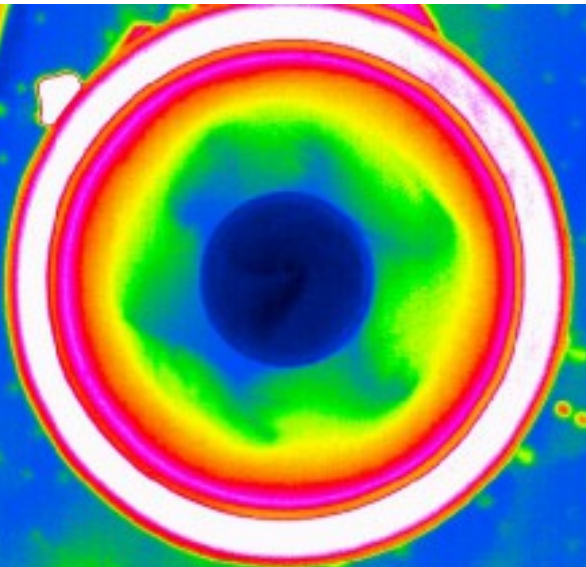
$m = 3$



$m = 4$



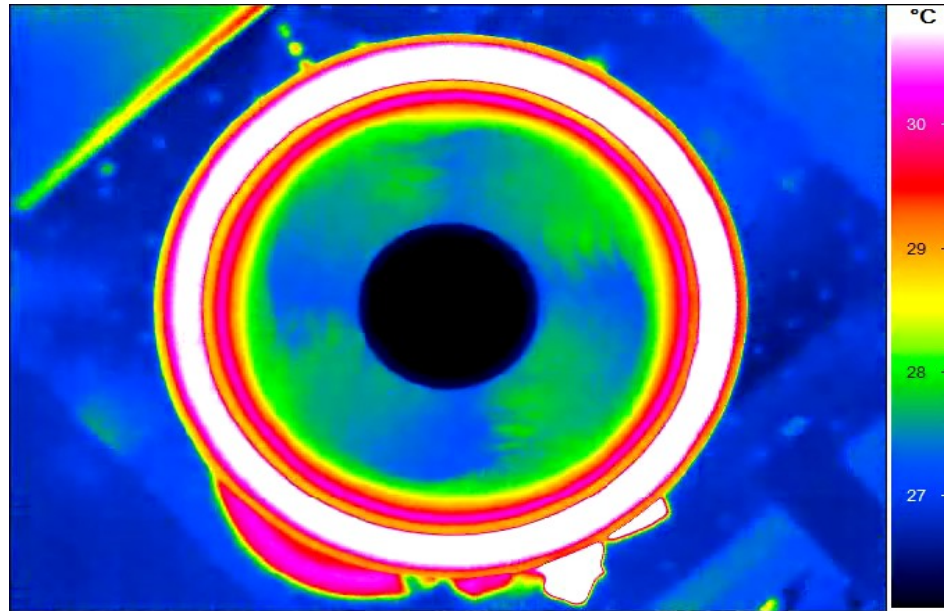
$m = 6$







Study unbalanced dynamics  
(spontaneous emission of internal waves)



$$\Delta T \approx 6.5 \text{ K}$$
$$\Omega = 2.5 \text{ rpm}$$

*Waves have been observed in a 2 immiscible fluid  
layers experiment (Williams, Haine, Read, 2008)*

