

# Long-term MAX-DOAS Measurement of Aerosol and Trace Gases in the Environmental Research Station Schneefernerhaus (UFS), Germany

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#### Introduction

A MAX-DOAS Instrument has been working in the Environmental Research Station Schneefernerhaus (UFS) on Zugspitze, the highest mountain in Germany since 2011. The DSCDs (differential slant column densities) of  $O_4$ , NO<sub>2</sub>, HCHO, and SO<sub>2</sub> etc. are measured, and the vertical profiles of tropospheric aerosol, NO<sub>2</sub>, and HCHO, as well as the VCDs (vertical column densities) of stratospheric  $O_3$  and  $NO_2$  are retrieved.

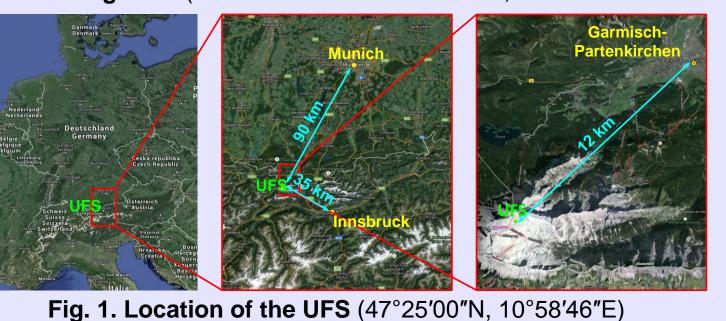
### **Observation Site and Instrument**

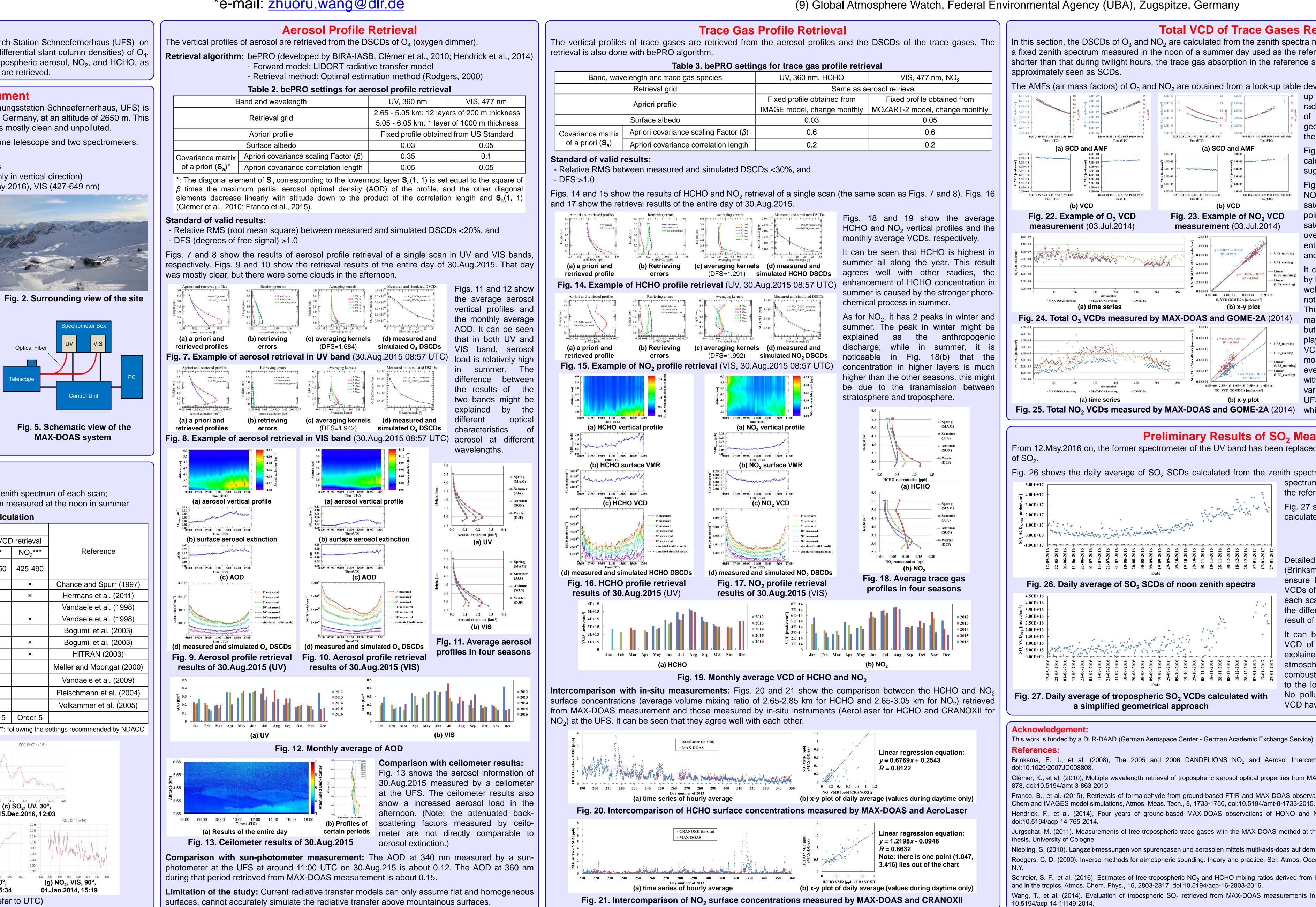
The Environmental Research Station Schneefernerhaus (Umwelt Forschungsstation Schneefernerhaus, UFS) is located under the summit of Zugspitze (2962 m), the highest mountain of Germany, at an altitude of 2650 m. This site is surrounded by the mountainous area of Alps, and the ambient air is mostly clean and unpolluted.

The MAX-DOAS instrument has been working since 2011. It consists of one telescope and two spectrometers.

**Telescope location:** on the terrace of the UFS Telescope direction: southwards, towards the mountainous area of Alps Vertical scanning sequence: 30°, 20°, 10°, 5°, 2°, 1°, -2°, 90° (scans only in vertical direction)

Spectral range: UV (320-478 nm before Feb 2016; 285-450 nm after May 2016), VIS (427-649 nm)





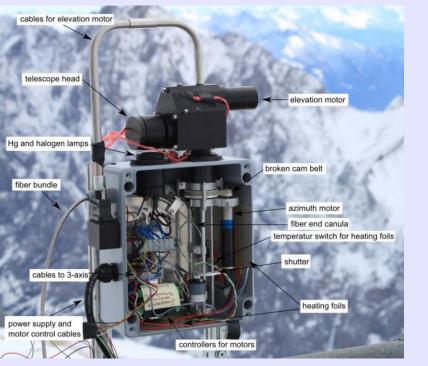


Fig. 3. Structure of the telescope (from Jurgschat, 2011)

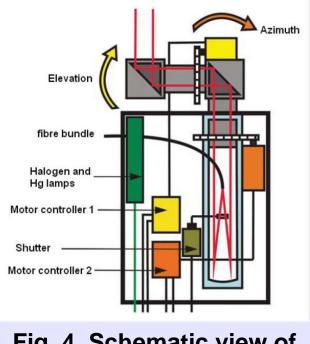
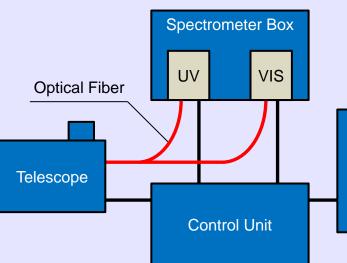


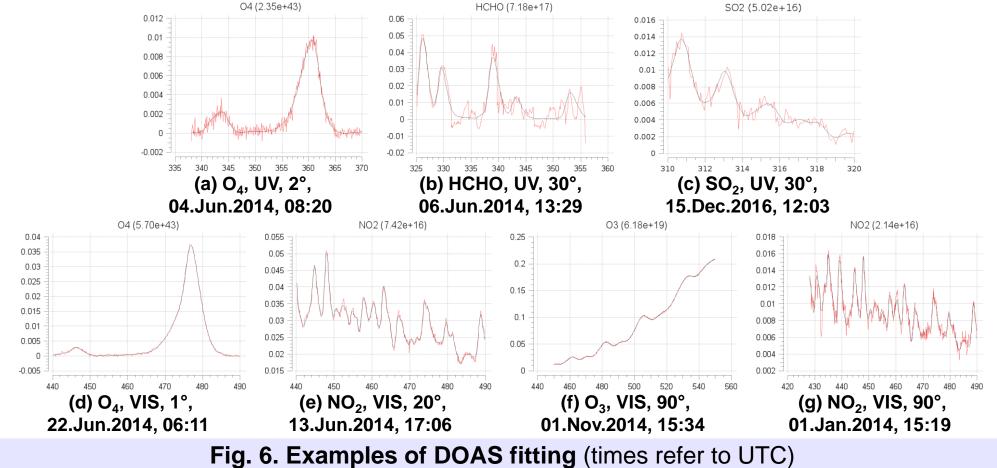
Fig. 4. Schematic view o the telescope (from Niebling, 2010)



**DSCD** Calculation **Calculation tool:** QDOAS (developed by BIRA-IASB) **Reference spectrum:** for retrieving aerosol and trace gas profiles, the zenith spectrum of each scan; for calculating total VCDs, a fixed zenith spectrum measured at the noon in summer Table 1. QDOAS settings for DSCD calculation

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Band	UV			VIS				
Data usage	profile retriev			al		total VCD retrieval		
Destination	O <sub>4</sub>	HCHO*	SO <sub>2</sub> **	O <sub>4</sub>	NO <sub>2</sub>	O <sub>3</sub> ***	NO <sub>2</sub> ***	Reference
Fitting window/nm	338-370	324.5-359	310-320	440-490	440-490	450-550	425-490	
Ring	×	×	×	×	×	×	×	Chance and Spurr (199
O <sub>4</sub>	×	×	×	×	×	×	×	Hermans et al. (2011)
NO <sub>2</sub> 294 K	×	×	×	×	×			Vandaele et al. (1998
NO <sub>2</sub> 220 K	×	×		×		×	×	Vandaele et al. (1998
O <sub>3</sub> 293K	×	×	×	×	×			Bogumil et al. (2003)
O <sub>3</sub> 223K	×	×	×	×	×	×	×	Bogumil et al. (2003)
H <sub>2</sub> O	×	×		×	×	×	×	HITRAN (2003)
НСНО	×	×	×					Meller and Moortgat (20
SO <sub>2</sub>			×					Vandaele et al. (2009
BrO			×					Fleischmann et al. (200
СНОСНО				×	×			Volkammer et al. (200
Debreensiel	Ordor 1			Order 1	Ordor 1			

Polynomial | Order 4 | Order 5 | Order 5 | Order 4 | Order 4 | Order 5 | Order 5 \*: following the QA4EVC HCHO DSCD calculation setting \*\*: only after 12.May.2016 \*\*\*: following the settings recommended by NDACC



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erman Aerospace Center (DLR), Oberpfaffenhofen, Germany eering (BGU), Technical University of Munich (TUM), Munich, Germany ation of Meteorological Satellites (EUMETSAT), Darmstadt, Germany SB), Brussels, Belgium rsity of Heidelberg, Heidelberg, Germany eißenberg, Germany niversity of Munich (LMU), Munich, Germany ), Technical University of Munich (TUM), Freising, Germany ntal Agency (UBA), Zugspitze, Germany	<image/> <text><text><image/></text></text>
<b>Total VCD of Trace Gases Retrieval</b> ection, the DSCDs of $O_3$ and $NO_2$ are calculated from the zenith spectra measured during twilight enith spectrum measured in the noon of a summer day used as the reference spectrum. As the han that during twilight hours, the trace gas absorption in the reference spectrum can be neglect nately seen as SCDs.	light path at noon is much
Fs (air mass factors) of $O_3$ and $NO_2$ are obtained from a look-up table developed by BIRA-IASB.	The AMF data in the look-

(air mass factors) of $O_3$ and	$NO_2$ are obtained from a look-up table				
18 1.4E+20 18   16 1.2E+20 1.4E+20   12 1.0E+20 1.0E+20   10 1.0E+20 0   6 0.0E+10 0   2 0.0E+00 18:40 18:45 18:50 18:55 19:00 19:05   (UTC) Time (UTC)	1.0E+17 25 1.0E+17 25   1.0E+17 25 1.0E+17 25   1.0E+17 1.0E+17 25 20   1.0E+16 1.0E+16 1.0E+16 1.0E+16   1.0E+16 1.0E+16 1.0E+16 1.0E+16   1.0E+10 1.0E+16 1.0E+16 1.0E+16   0.0E+00 3:25 3:30 3:35 3:40 3:45 3:50 3:55 4:00 1.8:40 18:45 18:50 18:55 19:00 19:05 19:10 19:15   Time (UTC) Time (UTC) Time (UTC)				
(a) SCD and AMF	(a) SCD and AMF				
9.0E+18 9.0E+18 7.0E+18 7.0E+18 6.0E+18 5.0E+18 0.0E+08 0.0E+00 18:40 18:45 18:50 18:55 19:00 19:05 Time (UTC) (b) VCD	5.0E+15 1.0E+15 2.0E+15 0.0E+00 3:25 3:30 3:35 3:40 3:45 3:50 3:55 4:00 Time (UTC) (b) VCD				
2. Example of O <sub>3</sub> VCD surement (03.Jul.2014)	Fig. 23. Example of NO <sub>2</sub> VCD measurement (03.Jul.2014)				
50 100 150 200 250 30 day number MAX-DOAS morning MAX-DOAS evening GOME-2A (a) time series	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$				
Total O <sub>3</sub> VCDs measured by	MAX-DOAS and GOME-2A (2014)				
, ,	1.0E+16				
	$\begin{array}{c c} \hline \\ \hline $				
	$\begin{array}{c c c c c c c c c c c c c c c c c c c $				
50 100 150 200 250 30 day number	$\begin{array}{c} \hline \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\$				
MAX-DOAS morning * MAX-DOAS evening * GOME-2A	NO <sub>2</sub> VCD GOME-2A [molec/cm <sup>2</sup> ]				
(a) time series	(b) x-y plot				
otal NO <sub>2</sub> VCDs measured b	v MAX-DOAS and GOME-2A (2014)				

the look up table are calculated using the UVSPEC/DISORT radiative transfer model which includes a treatment of the multiple scattering in a pseudo-spherical geometry. The VCDs can then be derived by dividing the SCDs by the corresponding AMFs.

Figs. 22 and 23 show the SCDs, AMFs and calculated VCDs of one day. The stable VCD value suggests that the AMF is reasonable.

Figs. 24 and 25 compare the total VCDs of  $O_3$  and NO<sub>2</sub> measured by MAX-DOAS and GOME-2A atellite. The satellite data are chosen from the data points within 150 km from the UFS. (The GOME-2A satellite is in a sun-synchronous orbit, it always pass over a place at 9:30 local time.) The data of the entire year of 2014 are shown in both time series and x-y plots.

It can be seen that the total VCDs of  $O_3$  measured by MAX-DOAS and by GOME-2A satellite agree very well with each other. As for NO<sub>2</sub>, the correlation is not so good as  $O_3$ , but still in a reasonable range. This might be explained that the total VCD of  $O_3$ mainly exists in the Ozone layer in the stratosphere; but for  $NO_2$ , the existence in the troposphere also plays an important role. It is also noticeable that the VCDs of NO<sub>2</sub> measured by MAX-DOAS in the morning are lower than those measured in the evening, and the evening results agree much better with satellite data. This can be explained by the daily variation of NO<sub>2</sub>. Moreover, the MAX-DOAS in the UFS measures almost only the free troposphere, while the satellite data cover some polluted areas.

## **Preliminary Results of SO<sub>2</sub> Measurement**

From 12.May.2016 on, the former spectrometer of the UV band has been replaced by a new one which can cover the fitting window

Fig. 26 shows the daily average of SO<sub>2</sub> SCDs calculated from the zenith spectra in the noon (170°<SAA<190°), with the zenith

spectrum measured at the noon of 20.Jun.2016 used as the reference spectrum.

Fig. 27 shows the daily average of tropospheric SO<sub>2</sub> VCD calculated with a simplified geometrical approach:

$$D_{trop} = \frac{DSCD}{\frac{1}{\sin(elev)} - 1}$$

Detailed explanation of this approach can be seen (Brinksma et al., 2008) and (Jurgschat, 2011). In order to ensure the validity of such assumptions, the geometric VCDs of the highest two elevation angles (20° and 30°) of each scan are calculated and then compared, and only if the difference between the two results is below 30%, the result of this scan is retained.

It can be seen that both zenith SCD and tropospheric VCD of SO<sub>2</sub> are relatively high in winter. This can be explained by the higher SO<sub>2</sub> concentration in the atmosphere in winter due to the increased coal combustion. Moreover, the longer average light path due to the lower SZA in winter can also increases the SCD. No pollution episodes with significantly enhanced  $SO_2$ VCD have been measured during this period.

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