A low-cost miniaturised detector for environmental radioactivity measurements Karen Aplin¹, Aaron Briggs¹, Peter Hastings¹, Giles Harrison², Graeme Marlton² and Adam Baird¹ ¹ Physics Department, University of Oxford, Keble Road, Oxford OX1 3RH UK UNIVERSITY OF ² Department of Meteorology, University of Reading, Earley Gate, Reading RG6 6BB UK

1. Abstract

We have developed an inexpensive, low-power (20mA), low-mass detector (30g) for environmental radioactivity measurements, using scintillator and solid state technology. The detector can measure energy and therefore has the capability to distinguish different types of energetic particle. Results from recent tests, when our detector was integrated with a meteorological radiosonde system, and flown on a balloon up to ~25km, identified the transition region between energetic particles near the surface dominated by terrestrial gamma emissions, and higher-energy particles in the free troposphere from cosmic rays. The detector can also be used with Bluetooth for remote monitoring, which is particularly useful for hazardous areas. It is small and cheap enough to be used in sensor networks for a wide range of applications, from atmospheric science to disaster monitoring.

2. Device

The radiation detector uses a 1 x 1 x 0.8 cm³ CsI:TI scintillator with a integral photodiode. Energetic particles cause a flash of light in the scintillator, which is then converted to a charge pulse by the photodiode. Amplification, signal conditioning and a trigger circuit allow a microcontroller to count and measure the height of each pulse. The instrument can be carried on and interfaced with a meteorological radiosonde, as shown in Figs 1-3. Alternatively, data can be output via USB, Bluetooth or on request over the serial bus for a hand-held detector or remotely piloted vehicle.



Fig 1 Block diagram of the instrument operation showing the signal flow. In this case the output data is sent to a radiosonde, which carries it aloft. The PANDORA is a data acquisition system for a radiosonde allowing interfacing of custom instrumentation in addition to the standard quantities measured (Harrison et al, 2012)



Fig 2 Detector board (top) ready for radiosonde integration. The green board carries the PANDORA radiosonde data acquisition system.



Fig 3 Meteorological radiosonde (top) ready for deployment with the boxed detectors from Fig 2 (bottom)

3. Calibration

The radiation detector was calibrated with laboratory gamma sources providing energies from 511 – 1332 keV. More energetic peaks from natural radioactivity at 1460 and 2610 keV (from ⁴⁰K and ²⁰⁸TI respectively) were identified and used as additional calibration points (Fig 4). The instrument detects the "photopeak" where all the gamma energy is deposited in the scintillator.

Fig 4 Pulse size against particle energy. The sensitivity is ~0.177 mV/keV with an offset of 160 mV.



Based on our device specification, and the response of similar instruments calibrated with more energetic particles (Viesti et al, 1986), we anticipate that our detector will respond to gammas of energy up to ~ 17 MeV.

Separate laboratory tests have also established that: The resolution for gammas is ~100 keV

- The counting error is $\sim 10\%$
- The detector also responds to cosmic ray muons

4. Radiosonde test flights

Test flights were made from Reading University Atmospheric Observatory on 16th August and 6th October 2016. The final destination of each radiosonde is shown in Fig 5. Space and meteorological weather conditions were calm for both test flights. The radiosondes ascended to 29km before the balloons burst. (The devices were not retrieved.)



Fig 5 Map of the southern UK, showing the launch site in Reading, with red symbols indicating the final destination of each radiosonde. The August flight is indicated with a square, October a triangle. For reference, the distance between Reading and and London ~60 km.

5. Results

Pre-flight tests at Reading University Observatory and Meteorology Building (located ~500m apart) indicated similar count rates and gamma ray spectra, as expected for natural radioactivity dominated by surface geology (Figs 6-7). Count rates were also consistent with the background laboratory tests at Oxford.





Fig 8. Pulse height in mV calibrated to gamma energy (right hand axis) for the boundary layer (BL) height determined from the measured temperature profiles for the launch, and in the free troposphere for (a) August flight and (b) October flight.

6. Conclusions

This novel radioactivity detector:

- Resolves gamma energy from ~150 keV 17 MeV
- Is inexpensive, light weight and low power
- Offers a variety of connectivity capabilities

Acknowledgements

UK Science and Technology Facilities Council (Impact Accelerator). References Science & Technology Facilities Council

Aplin et al, <u>Space Weather</u>, 15 (2017) Harrison et al Rev. Sci. Instrum. 83, 036106 (2012) Viesti et al Nuc. Inst. Meth. Phys. Res. A252 75-59 (1986)



Determines count rates for a range of energetic particle types A more detailed account can be found in Aplin *et al* (2017)