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# A novel technique for studying volcanic gas chemistry and dispersion on short time scales



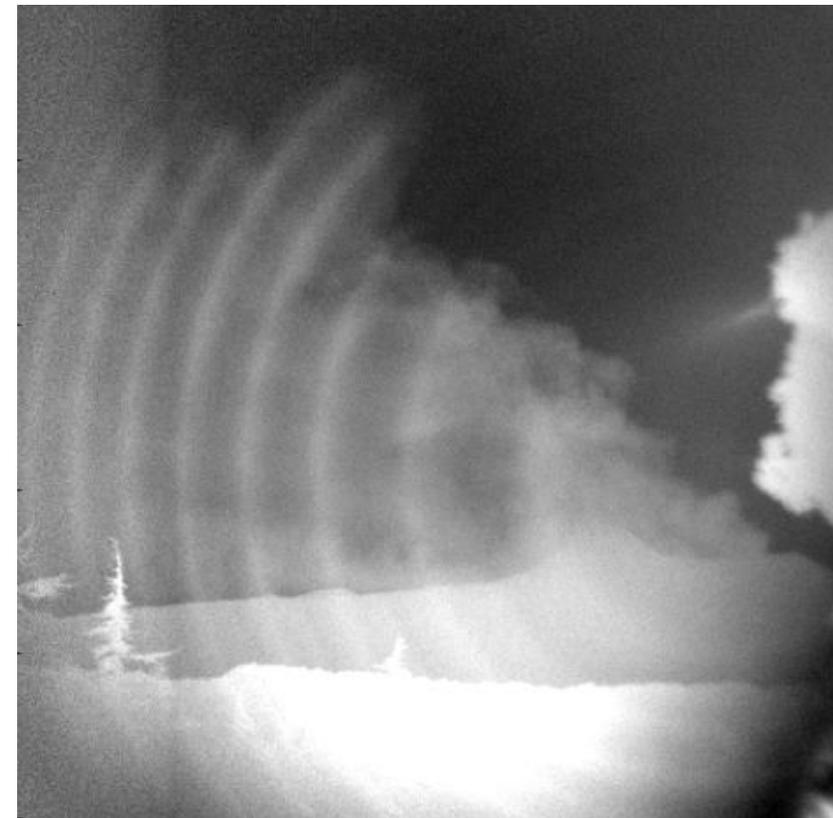
Christopher Fuchs<sup>1</sup>, Jonas Kuhn<sup>1,2</sup>, Nicole Bobrowski<sup>1,2</sup>, and Ulrich Platt<sup>1,2</sup>

<sup>1</sup> *Institute of Environmental Physics, Heidelberg, Germany*

<sup>2</sup> *Max Planck Institute for Chemistry, Mainz, Germany*

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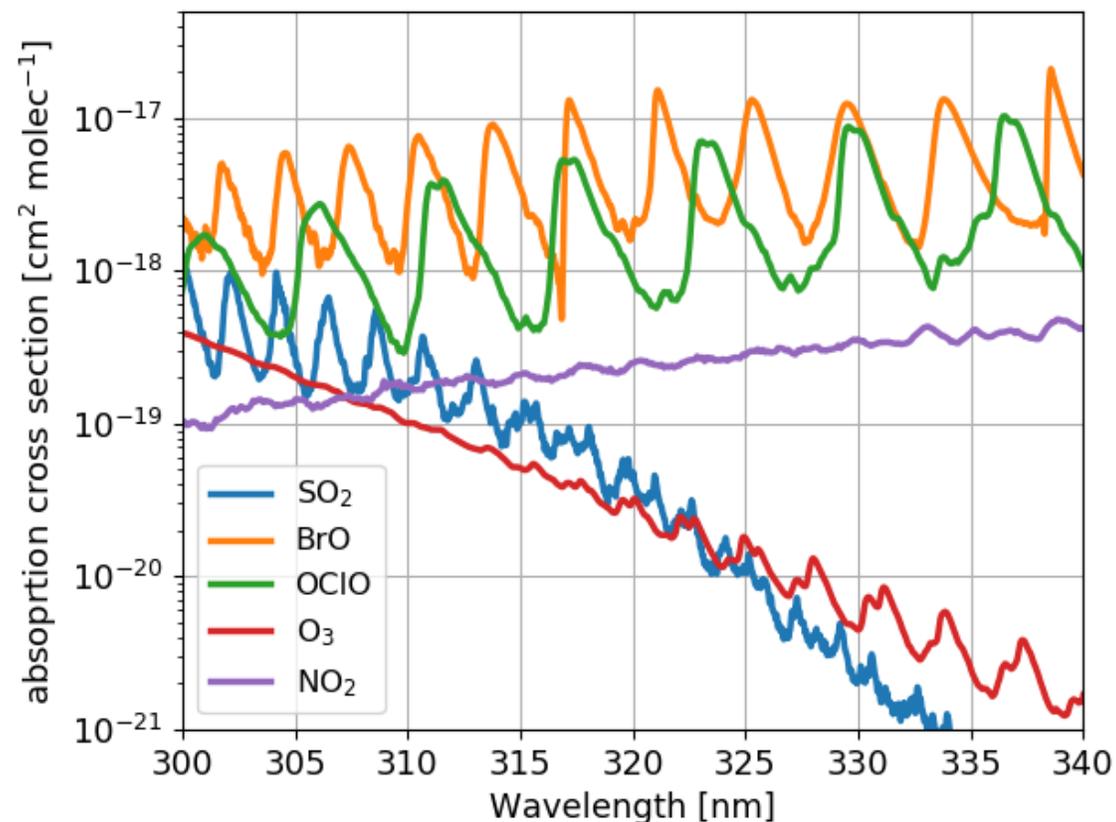
# Imaging of volcanic plume emissions

**Motivation:** resolve plume processes at their intrinsic time and spatial scale, e.g.:

- emission fluxes of trace gases
- investigation of halogen chemistry (e.g. bromine)

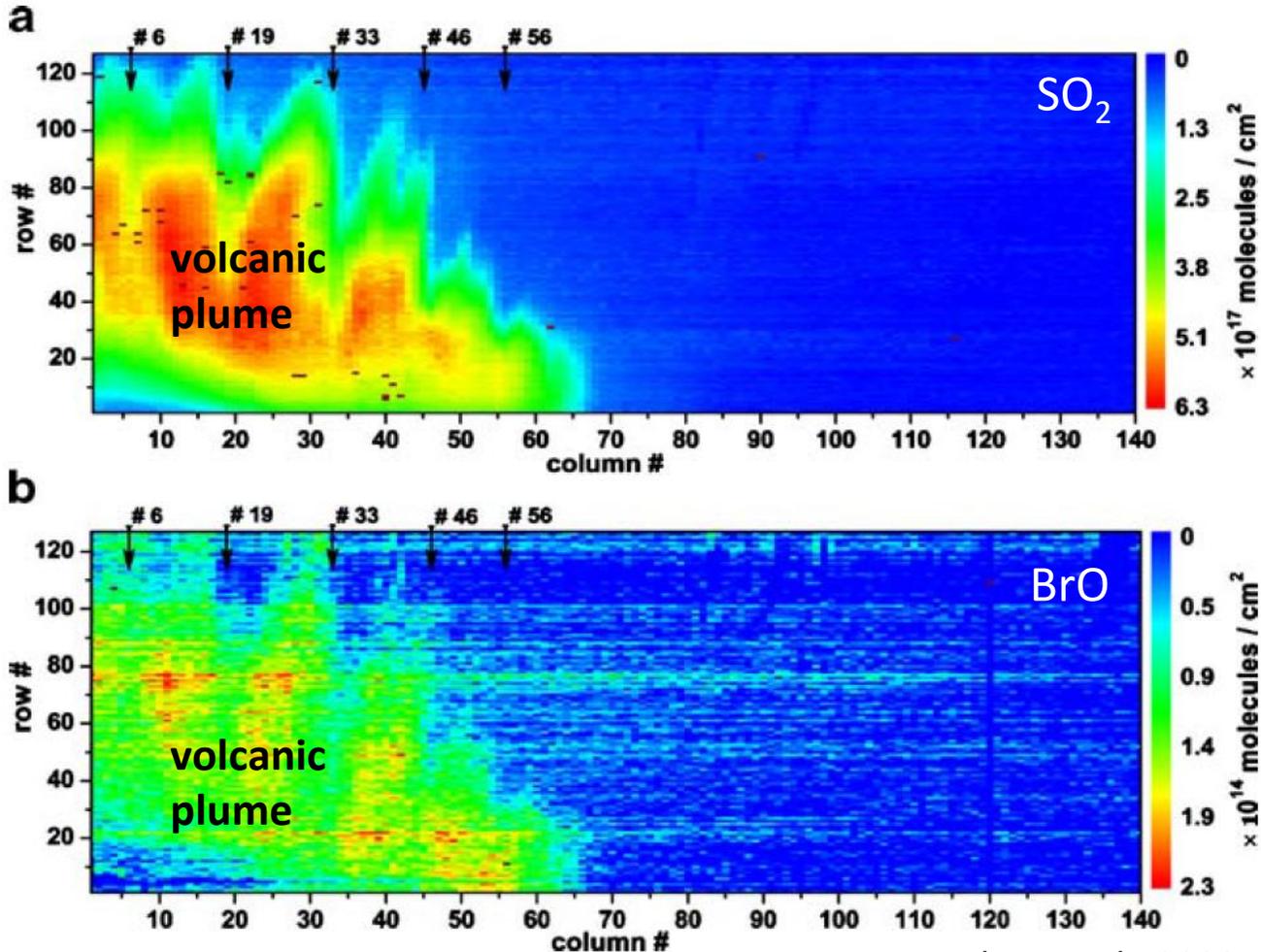
## Requirements:

- identification by **characteristic differential absorption structures** within a measured spectral optical density
- quantification of **column densities** (CDs, number density of trace gas along atmospheric light path)
- **'sub nm' spectral resolution** required for sufficient selectivity and sensitivity



# Imaging trace gases in volcanic emissions: Imaging DOAS

Imaging of volcanic SO<sub>2</sub> and BrO emissions (pushbroom approach, i.e. one column at a time)



- **High spectral resolution** → detection of various different trace gas species simultaneously

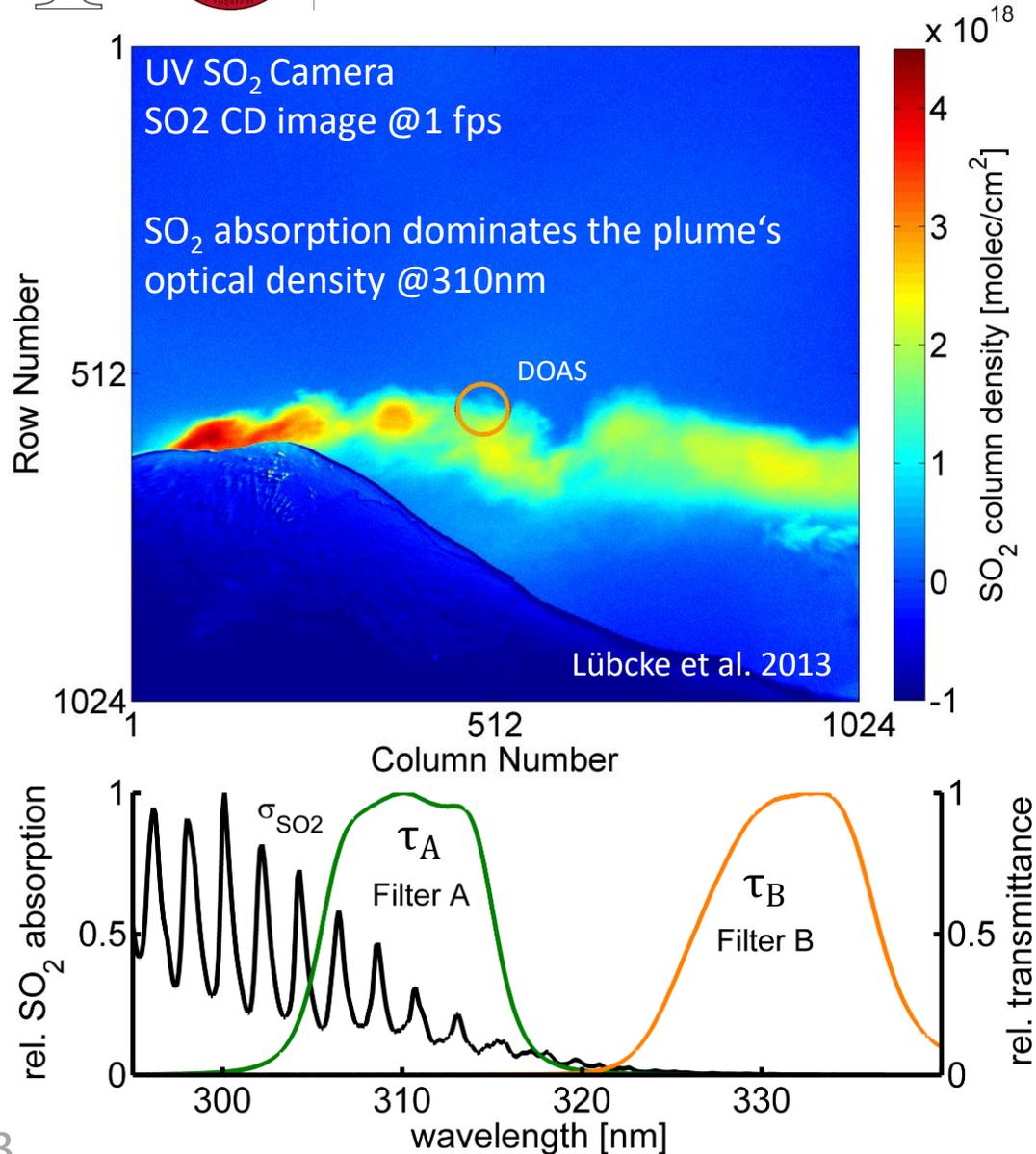
**BUT: too slow!**

imaging by spatial scanning:

- limited light throughput of the spectrometer
- scanning mechanisms or complex optics required



# Imaging trace gases in volcanic emissions: SO<sub>2</sub> Camera



high spatio-temporal resolution → flux monitoring

**BUT: limited selectivity & high cross interferences**

- interference to clouds, aerosol and O<sub>3</sub>
- Only works for high SO<sub>2</sub> column densities

Lambert-Beer's law:

$$dI(\lambda) = I_0(\lambda) \cdot e^{-\sigma(\lambda) \cdot S} d\lambda$$

$$= I_0(\lambda) \cdot e^{-\tau} d\lambda$$

$I_0$ : incident radiation  
 $\sigma$ : abs. cross section  
 $S$ : column density  
 $\tau$ : optical density

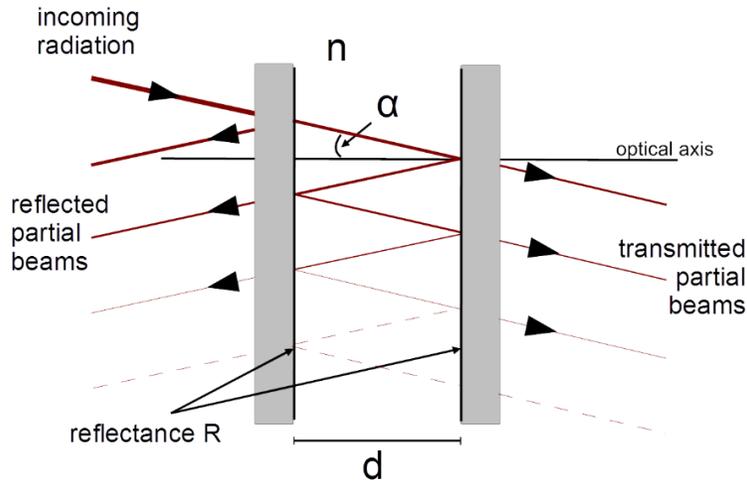
$$\tau_i = \ln \left( \frac{\int_{\lambda_{1,i}}^{\lambda_{2,i}} I_{0,i}(\lambda) d\lambda}{\int_{\lambda_{1,i}}^{\lambda_{2,i}} I_i(\lambda) d\lambda} \right) = \sigma_i \cdot S$$

$i \in A, B$  (Filter)

Apparent absorbance SO<sub>2</sub> Camera:

$$AA = \tau_A - \tau_B = k_{SO_2} \cdot S_{SO_2}$$

# Imaging Fabry-Perot Interferometer Correlation Spectroscopy (IFPICS)



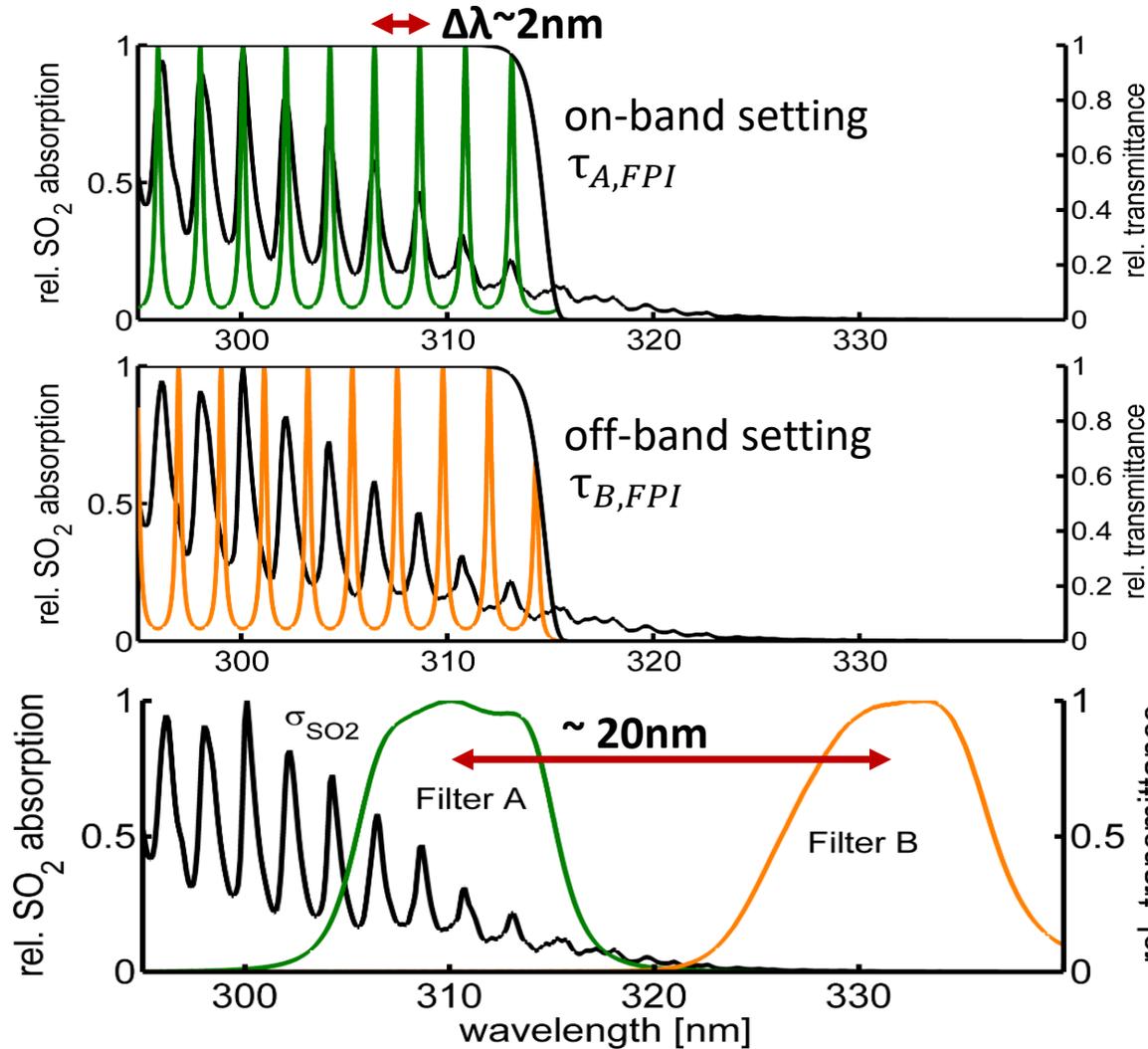
free spectral range:

$$\Delta\lambda \approx \frac{\lambda^2}{2 \cdot nd \cos(\alpha)}$$

Apparent absorbance FPI camera:

$$AA = \tau_{A,FPI} - \tau_{B,FPI} = k S_{SO_2}$$

- much **more specific** to the trace gas (narrow band structure)
- broad band **interferences reduced** (spectral variation 2nm vs. 20nm)



transmission at **maximum** SO<sub>2</sub> absorption  
→ **correlation**

transmission at **minimum** SO<sub>2</sub> absorption  
→ **anti-correlation**

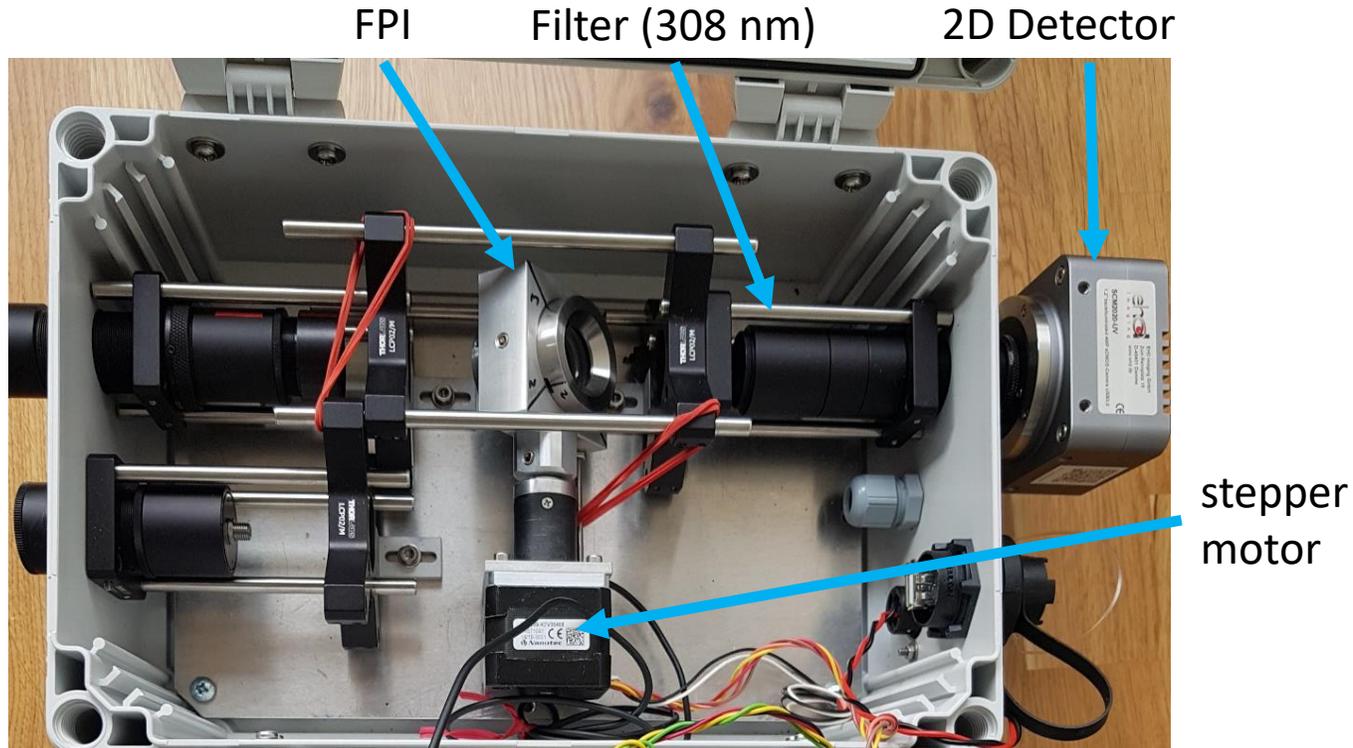
Filter based SO<sub>2</sub> Camera

**Comments:** The FPI transmission spectrum is matched to the SO<sub>2</sub> absorption spectrum in an on-band and off-band setting replacing the spectral filters used by SO<sub>2</sub> cameras. The spectral separation between the two spectral channels is thereby reduced by a factor of 10 (2 nm instead of 20 nm), thus reducing broad band interferences, e.g. to aerosols and increasing the specificity to the investigated trace gas. Tuning between on- and off-band setting is applied by tilting the FPI optical axis = variation of incidence angle  $\alpha$ .



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# Proof of concept study for SO<sub>2</sub> – Imaging Camera



Dimensions [mm]:  
200 × 350 × 130

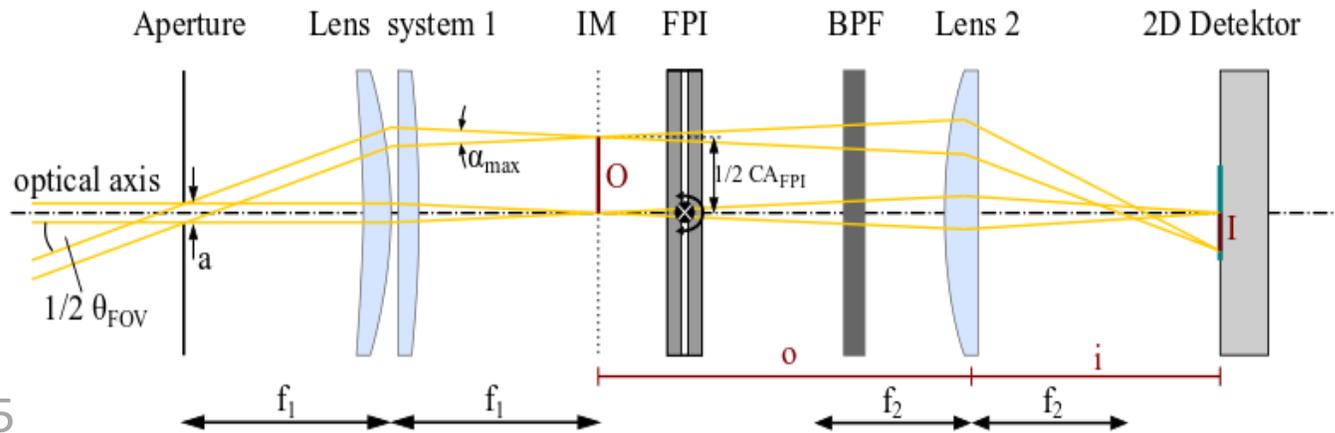
Weight: 4.8 kg

Power consumption: < 10W



### Comments:

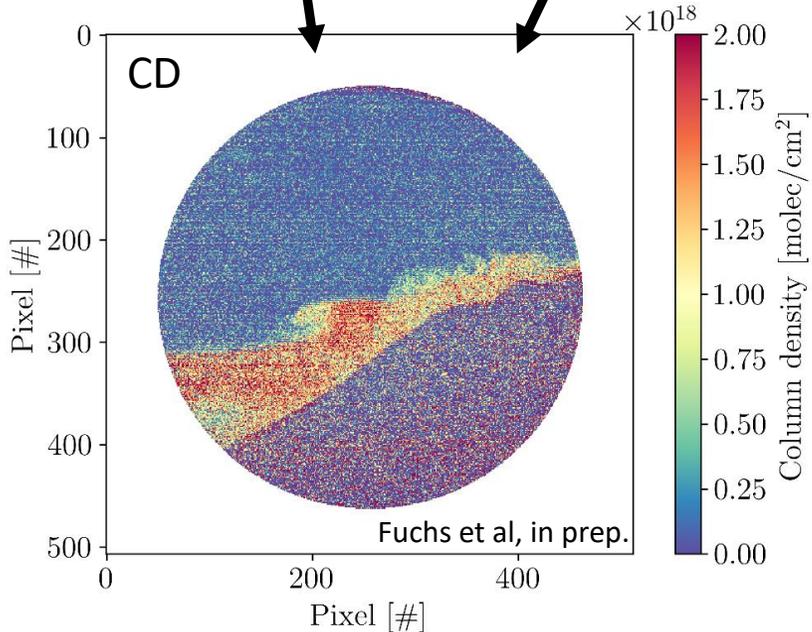
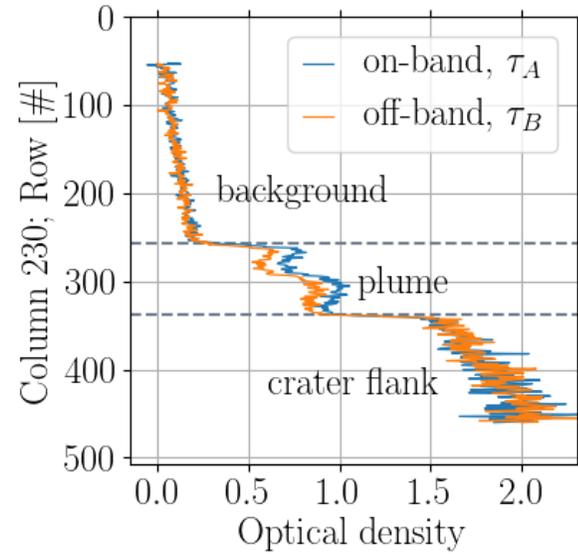
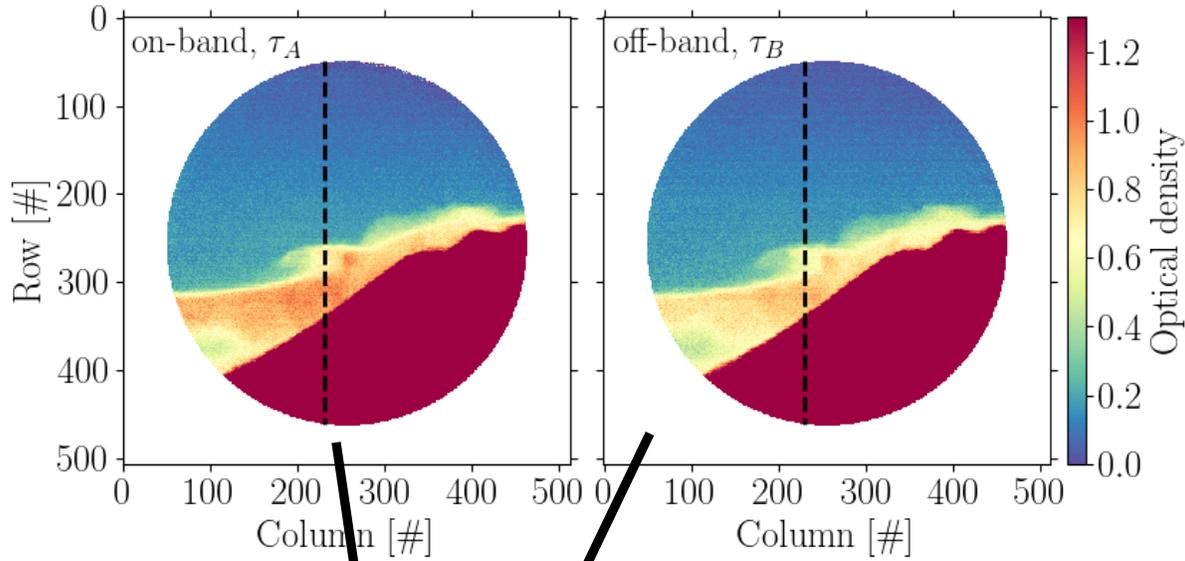
The IFPICS instrument uses an image-side telecentric optic setup (see schematic bottom left). Light enters from the left. Light is traversing an aperture before it gets parallelised by lens 1. The aperture entrance is thereby limiting the angle of divergence  $\alpha_{max}$ . The FPI is mounted on a motor that allows to tilt it around its optical axis to enable tuning the spectrum by varying the incidence angle onto the FPI. After traversing an band pass filter the light is imaged onto a 2D UV-sensitive detector.





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# Imaging of SO<sub>2</sub>: plume of SE-crater, Mt. Etna, Sicily



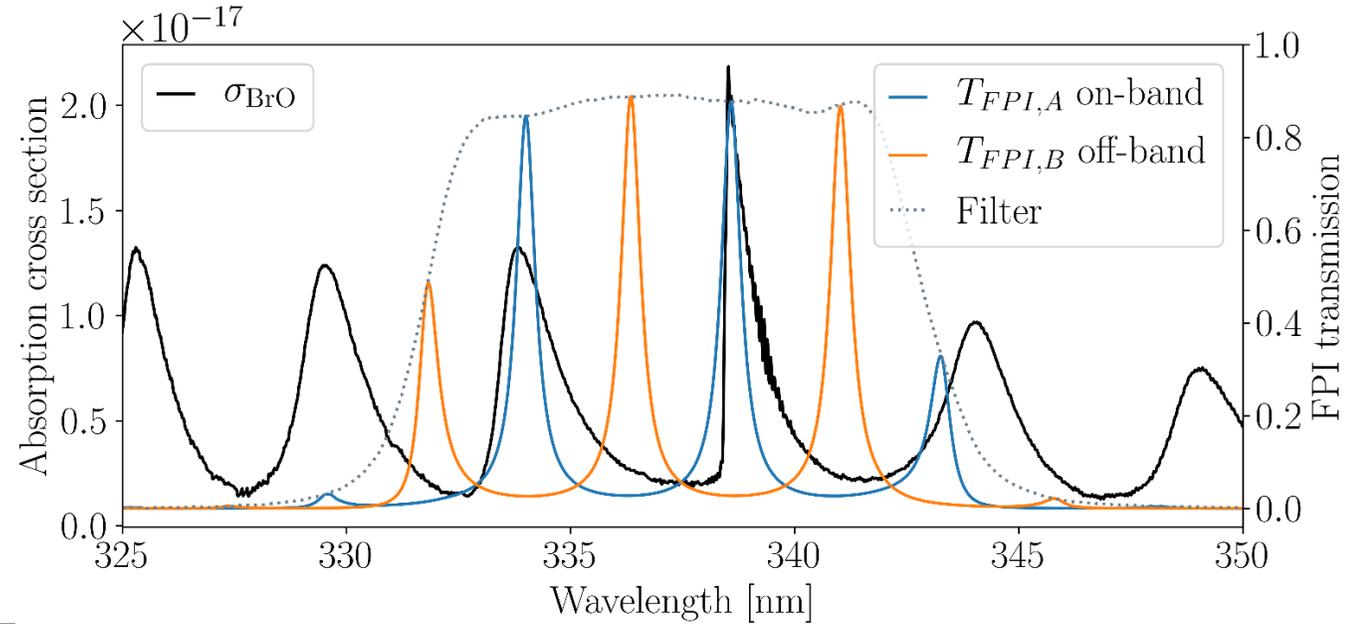
- Calibration via forward modelling the instrument response
- SO<sub>2</sub> sensitivity:  $AA = 10^{-19} \text{ cm}^2/\text{molec} \cdot S$
- Detect. limit  $\sim 4 \times 10^{17} \text{ molec cm}^{-2} \text{ s}^{-1/2}$

**Comments:**  
Optical density images shown in on-band  $\tau_A$  and off-band  $\tau_B$  setting (top left images). Plot of the optical densities along vertical dashed lines shows that the signal differ only in the plume region ( $\tau_A > \tau_B$ ), whereas the signal is equal in the background and crater flank area ( $\tau_A = \tau_B$ ) (top middle). Apparent absorbance AA given by  $AA = \tau_A - \tau_B$  is converted into SO<sub>2</sub> column densities (CD) (bottom left) via an instrument forward model, which was validated by gas cell measurements.

# IFPICS remote sensing of volcanic BrO – A model study

Simulated FPI transmission spectrum for atmospheric conditions (right)

Correlating and anti-correlating FPI transmission spectrum

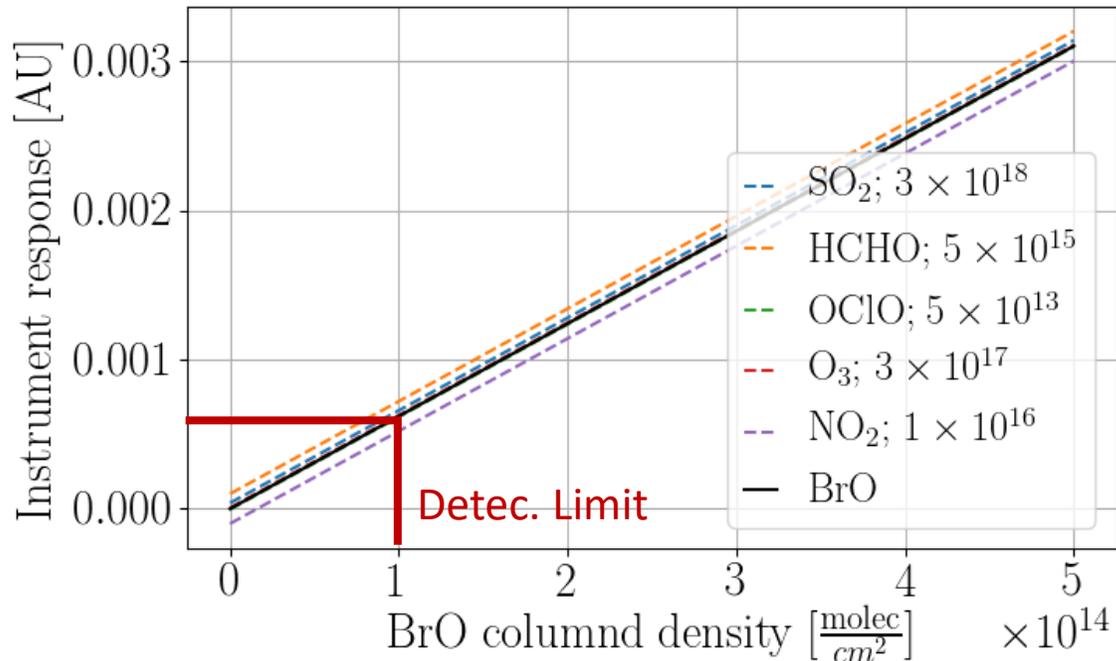


Cross interferences of BrO measurements to other trace gas species (left)

Sensitivity:  $AA = 6.5 \times 10^{-18} \text{ cm}^2/\text{molec} \cdot S$

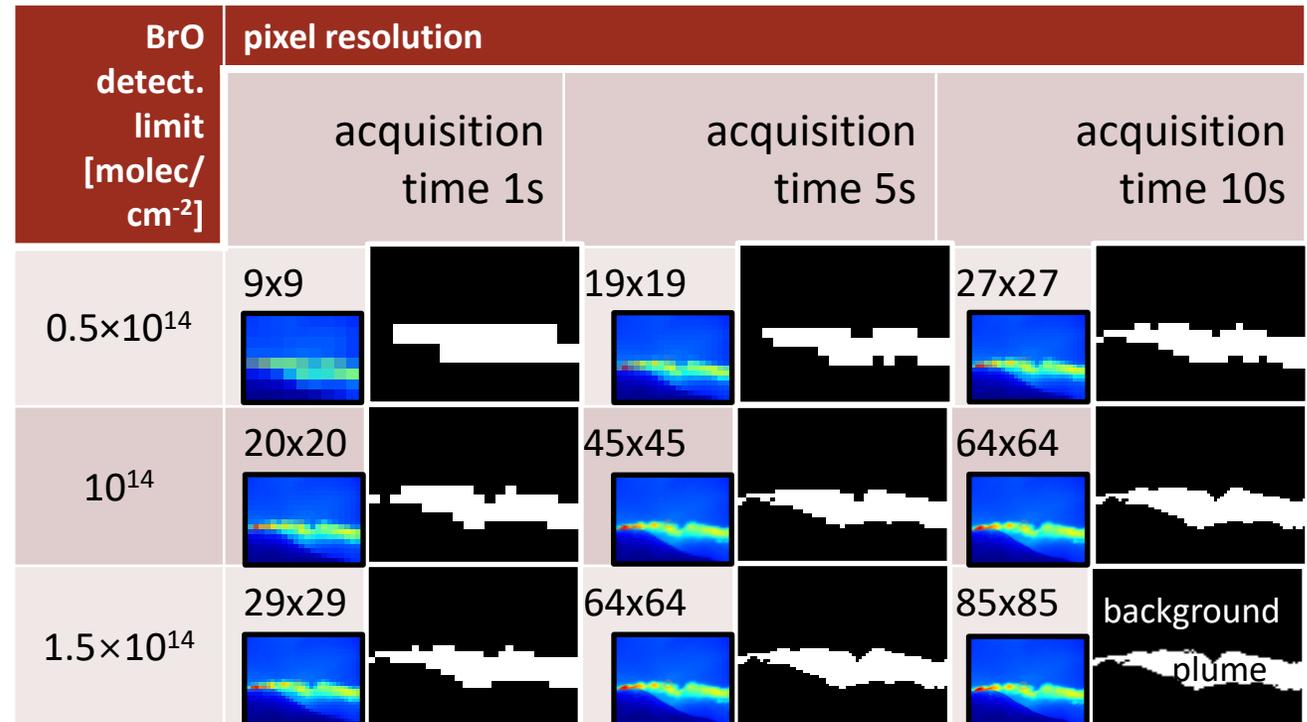
Intended detection limit for atmospheric  
Measurement conditions:  $\sim 10^{14} \text{ molec/cm}^2$   
( $\sim 10^{13} \text{ molec/cm}^2$  reached in laboratory)

→ **Low cross interferences!**



# IFPICS remote sensing of volcanic BrO – A model study

- Lower signal expected: apparent absorbance  $\sim 10^{-3}$  (for SO<sub>2</sub>  $\sim 10^{-1}$ )
- Required detection limit  $\sim 10^{14}$  molec/cm<sup>2</sup> → achieved with an one pixel prototype under laboratory measurement conditions
- Model study:  
Low photon budget - lower signal → trade spatial resolution against photons
- BrO detection limits for different degrees of binning and different acquisition times (right)



Calculations according to Kuhn et. al. 2019

## Conclusion:

### **SO<sub>2</sub>: First successful FPI imaging application on volcanic SO<sub>2</sub>**

- Prototype is robust and small
- Good sensitivity and low cross interferences
- Inherent calibration with forward model
- No background reference images required

### **BrO:**

- BrO model study yield promising spatio-temporal resolution
- Detection limit of  $\sim 10^{14}$  molec/cm<sup>2</sup> easily achieved with one-pixel prototype in laboratory

## Outlook:

- Field measurements of volcanic plume BrO
- Extend the technique to further trace gases
- Field measurements of weaker SO<sub>2</sub> point sources.

Correspondence:

[cfuchs@iup.uni-heidelberg.de](mailto:cfuchs@iup.uni-heidelberg.de)

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