

Laboratoire de Climatologie et Topoclimatologie



Comparison of the surface mass and energy balance of CESM and MAR forced by CESM over Greenland: present and future

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• Prior to this study:

GCMs/ESMs used to force the lateral boundaries of an RCM (Regional Climate Model, here MAR) couldn't simulate the surface mass balance (SMB) of an ice sheet The potential added value of an RCM could only be evaluated by comparing atmospheric variables in both models

• In this study:

Land/snow module of CESM simulating the surface mass and energy balance at the surface of the snow pack allows us for the first time to directly compare the SMB and its components in both models as well as the sensitivity of the snow module of both models to the projected temperature increase

Simulations setup

Period: 1981 – 2100 Scenario: SSP585 Model 1: CESM2 (Community Earth System Model), same version as in CMIP6 Model 2: MAR (Modèle Atmosphérique Régional) v3.11 forced by CESM

Aim of the study is to show the sensitivity of the snow module of both models to the projected climate change, not an *who's the best* contest



No correction at the MAR lateral boundaries in this study (e.g. to match MAR Greenland reference run (15km forced by ER5))

Initially two spatial resolutions for MAR runs (15 and 50 km) to highlight potential differences due to spatial resolution rather than used model



Differences between future anomalies: MAR vs MAR << MAR vs CESM

Only the MAR 15km simulation is used afterwards



2071 – 2100 mean SMB anomaly



2071 - 2100 mean SMB anomaly



Most of the SMB anomaly between MAR and CESM comes from the runoff anomaly (total amount and pattern)

	MB ± SD
SMB	-238±328
RU	272±322
SF+RF (=P)	10±63
	R ²
SMB vs RU	0.96
SMB vs P	0.02

250

200

150 100

50 0

-50

-100

-150

-200-250



SMB components anomalies: temporal evolution



Difference (Gt yr-1)*		
SMB	-350	
Runoff	380 25	
Sublimation	25)/15
Snowfall	10	
Rainfall	0	

* MAR–CESM anomaly difference averaged over 2091 – 2100 Factor >10 between RU MAR-CESM difference and SU difference (next largest difference) at the end of the century

SMB components anomalies vs temperature anomalies



1/ For the same temperature increase, SMB/RU decreases/ increases more in MAR than in CESM
2/ Temperature increases more in MAR than in CESM

Half of SMB/RU anomaly at the end of the century comes from sensitivity to temperature increase Half of SMB/RU anomaly at the end of the century comes from larger temperature increase in MAR

R² of quadratic fit

	CESM	MAR
SMB	0.91	0.95
RU	0.94	0.96
RF	0.84	0.85
SF	0.03	0.001
SU	0.77	0.87

SMB and its components anomaly vs

Energy balance components anomalies: temporal evolution

-0.5

4

60 -60 • NET Ñ Δ=+12 E fluxes anom. JJA (W m^{-2}) 50 -50 B • SWA \mathbb{N} • LWD 40 40 Up. longwave anom. JJA • LWU 30 -30 • SHF $\Delta = +12$ • LHF 20 --20 — MAR 10 -10-- CESM -10+ 10 2020 2000 2040 2060 2080 1980 2100 Difference (W m⁻²)* 12 Net radiation (NET) 12 Absorbed shortwave (SWA) **NET MAR-CESM difference at the end of the** Down. longwave (LWD) -1 century mostly driven by SWA differences, and, Up. longwave (LWU) -3 to a lesser extent, by LHF and LWU

Temporal evolution of the anomaly of energy balance components over 1981 – 2100 (10-yr running mean)

- SHF difference is negligible
- LWD difference reduced towards the end of the century
- NET = SWA + LWD LWU + SHF + LHF SWA = SWD (1-a) SWD = incoming solar radiation at the surface a = albedo

* MAR-CESM anomaly difference averaged

Sensible heat flux (SHF)

Latent heat flux (LHF)

over 2091 - 2100

Energy balance components anomalies vs temperature anomalies



Sensitivity to temperature change or larger temperature increase?

NET	40/60
SWA	70/30
LWD	Cancel out
LWU	T increase
SHF	Cancel out
LHF	70/30

R² of quadratic fit

	CESM	MAR
NET	0.97	0.97
SWA	0.66	0.88
LWD*	0.94	0.96
LWU*	0.998	0.998
SHF	0.82	0.85
LHF	0.79	0.91
		* Linear fit

Absorbed shortwave radiation: albedo vs SWD



R² of quadratic fit

SWA_A = Absorbed SW with incoming solar radiation kept constant in time. Shows influence of albedo changes on SWA SWA_S = Absorbed SW with albedo kept constant in time. Shows the influence of incoming shortwave radiation changes on SWA

Larger sensitivity of SWA to temperature change in MAR driven by larger sensitivity of albedo to temperature change

Absorbed shortwave radiation: albedo vs SWD

Fraction of the GrIS in the ablation zone vs temperature anomaly



Conclusions and perspective

- 1. First time that we can compare the sensitivities of the snow modules of MAR and the model used to force the atmosphere at the MAR domain boundaries
- 2. SMB and runoff simulated by MAR appear to be more sensitive than CESM's to the projected temperature increase
- 3. Main cause is the higher sensitivity of the MAR albedo (snow density increase) to temperature increase through its effect on the absorbed incoming radiation
- 4. Other causes for more minor different sensitivities in MAR and CESM might include:
 - Differences in parametrisations for sensible and latent heat fluxes
 - Effect of cloud cover on incoming shortwave and long wave radiation

Next step: try and determine which processes/parametrisations cause the sensitivity differences in both models

Supplementary material





Cloud cover fraction JJA

-0.05

1980

2000

2020

2040

2071 - 2100 JJA cloud cover fraction (0 - 1)



Temporal evolution of the anomaly of cloud cover fraction 1981 – 2100 (10-yr running mean) • MAR • CESM

2060

2080

2100