

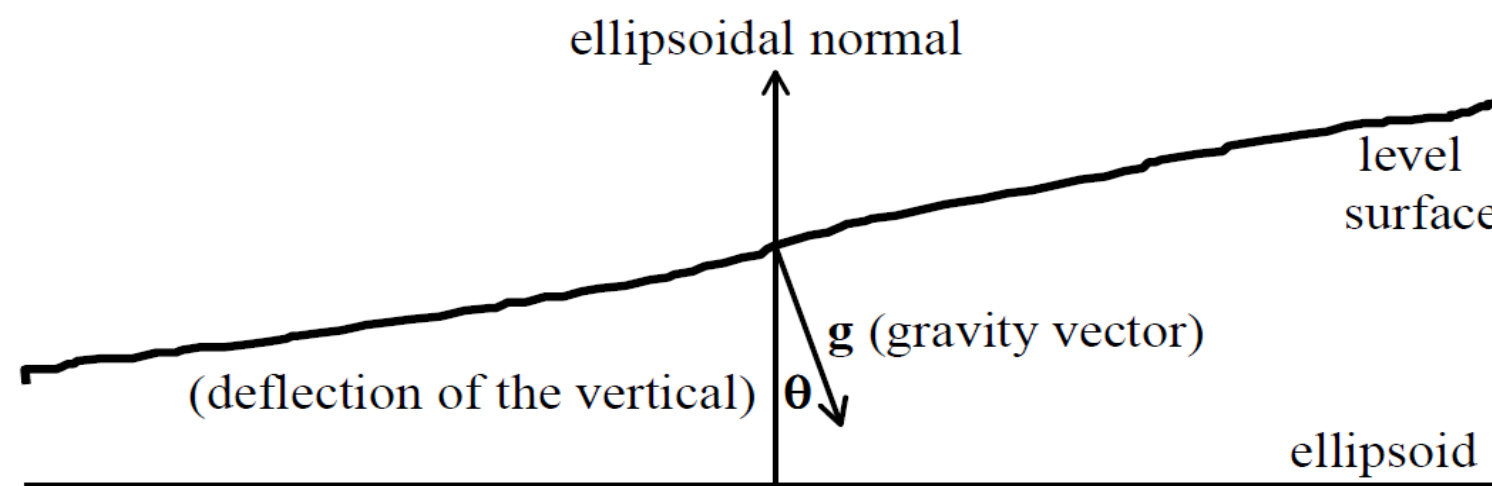
# The digital zenith camera as an additional technique for quasi-geoid model determination of Latvia

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Reiner Jaeger

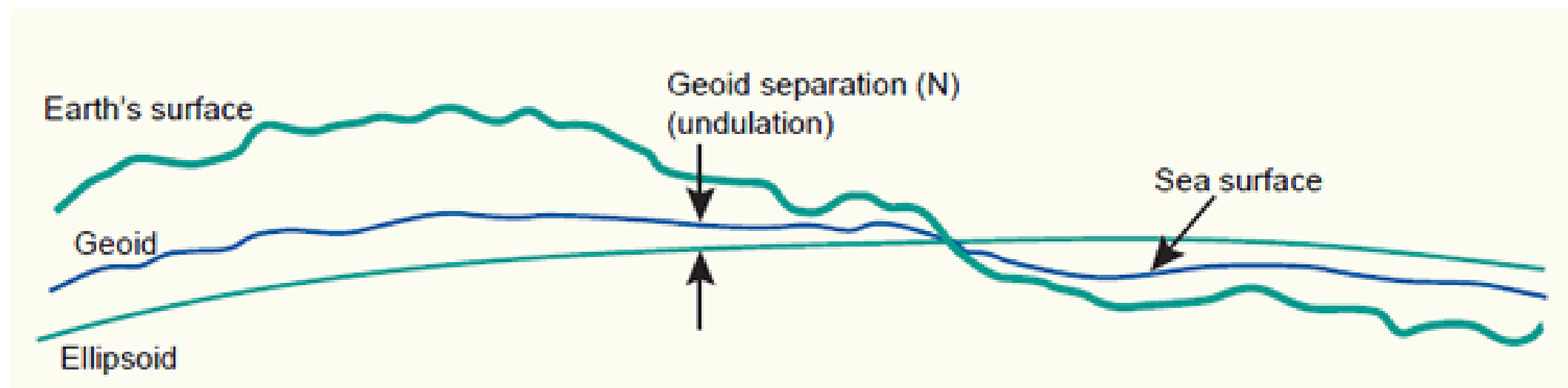
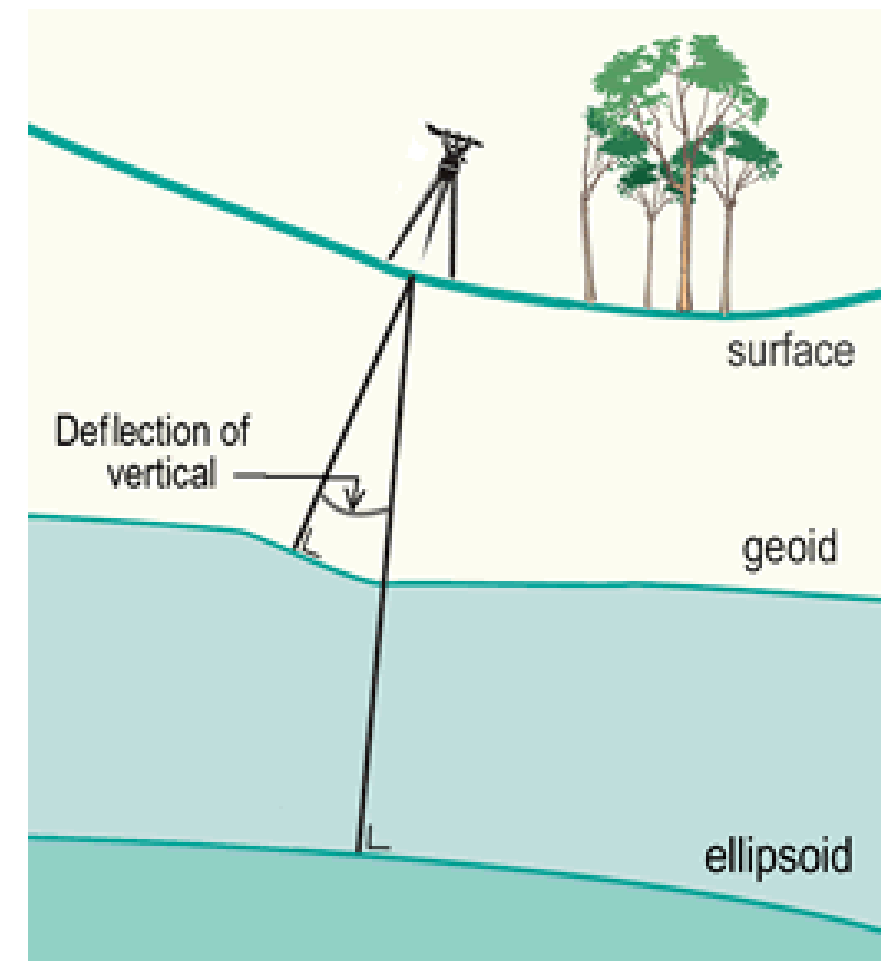
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# Definition

- Instrument for measurements of vertical deflection - difference between direction of the plumb line (the normal to the geoid) and the vertical direction on the ellipsoid (the normal to the ellipsoid).



- Vertical deflection represents inclination of geoid surface to reference ellipsoid surface.
- Amplitude of vertical deflection is typically about 10 arcseconds, can reach arcminute in mountain regions.



*Two main reference surfaces are used to approximate the shape of the Earth: geoid and ellipsoid.*

# Digital Zenith Camera

Portable astrometric instrument for vertical deflection measurements - between gravity field direction and normal to reference ellipsoid

## Data can be used for:

- Local quasi-geoid precision improvement;
- Earth crust movement monitoring;
- Local geological structure qualities determination.



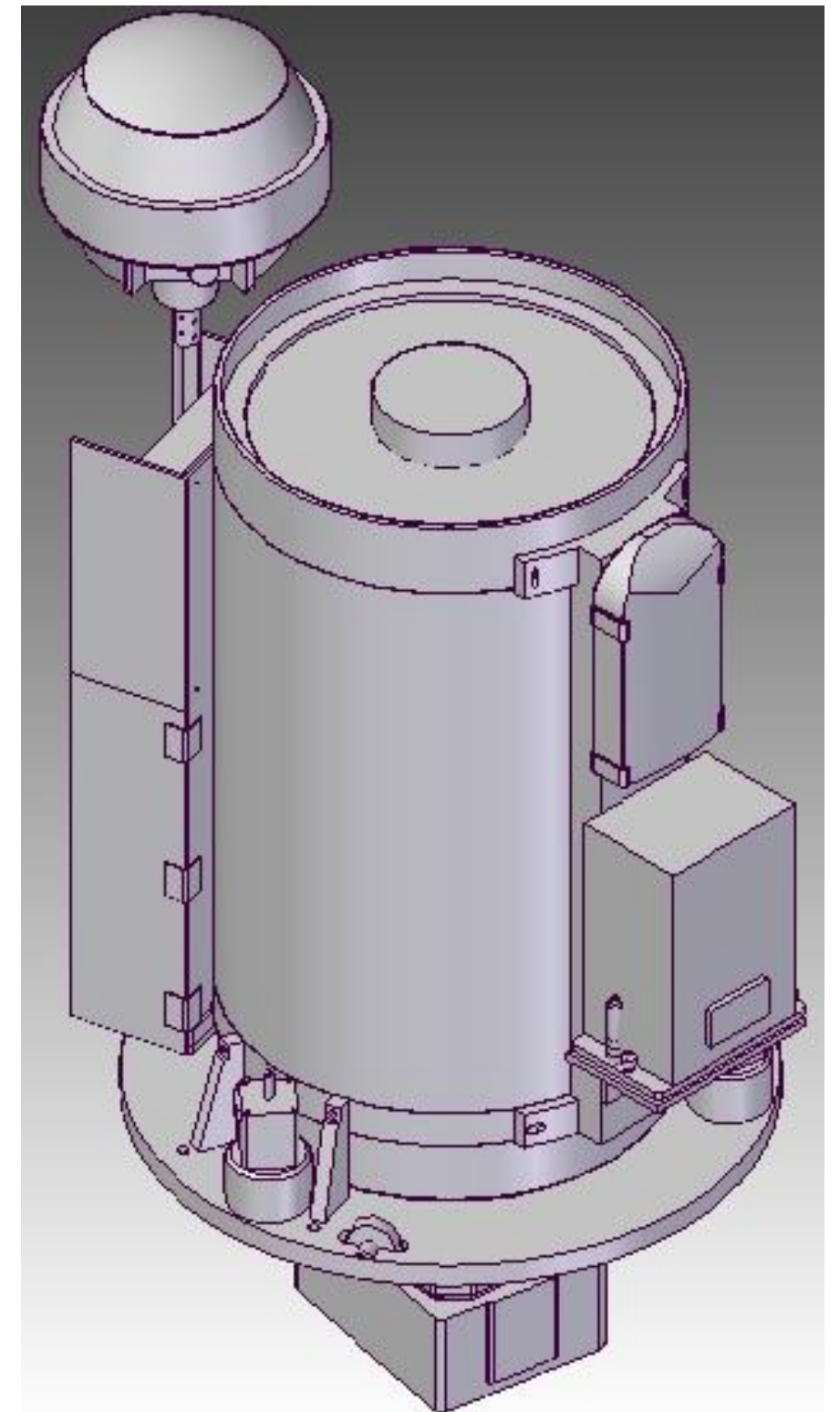
Vertical deflection measurement accuracy  $\sim 0.1$  arcsecond.  
Tripod with detachable rotating part:

astrometric quality coma free 8-inch catadioptric telescope;  
two-axis electronic HRTM tiltmeter (resolution  $\sim 0.02$  arc milliseconds);  
GNSS receiver (plane coordinate accuracy  $< 0.6$  m with SBAS);  
CCD matrix 8.3 MPx,  
linear actuators with resolution of  $< 0.01$   $\mu$ m;  
on-board computer MinnowBoard;  
WLAN for communication with remote laptop;  
Accurate levelling, setting of azimuth and prescribed schedule of observations are done automatically.

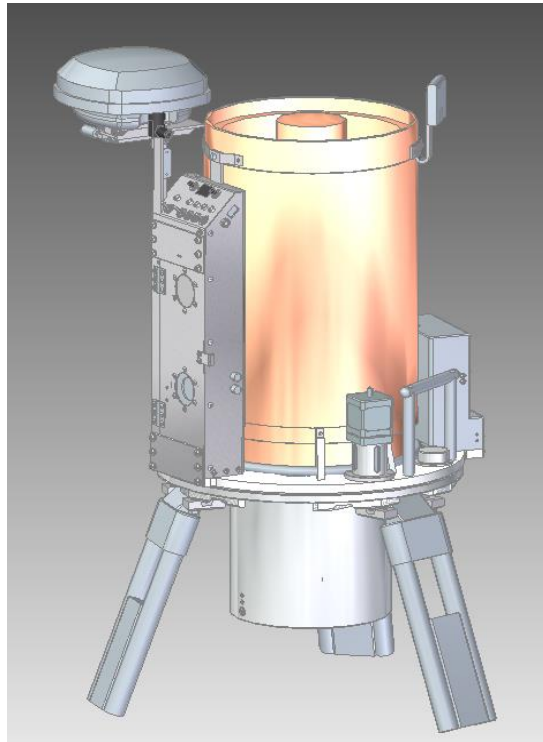


## Our design of instrument takes into account:

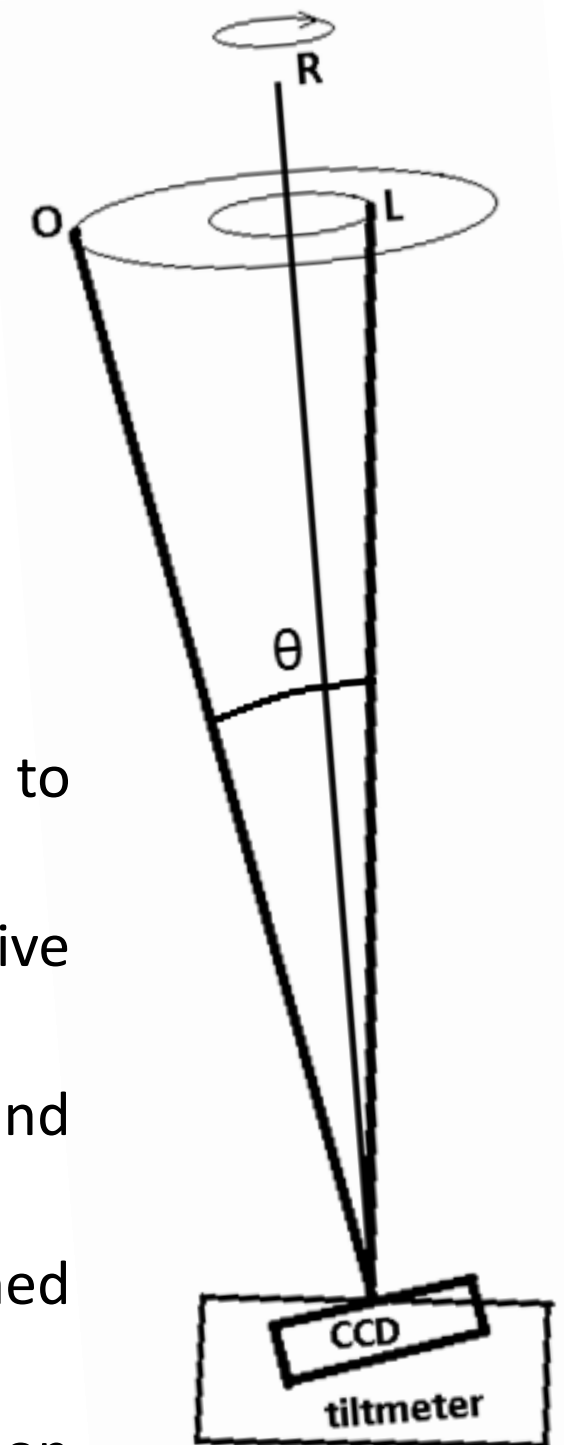
- remote control of all basic operations is necessary to avoid vibrations, convection, mechanical influence of cabling;
- leveling in all measurement positions is necessary to minimize effects of tiltmeter scale uncertainty, axes misalignment;
- tiltmeter position adjustment mechanism allows simple and independent adjustment related to main axis for both measurement planes;
- rotation mechanism supports any rotation angle.



# Measurement method

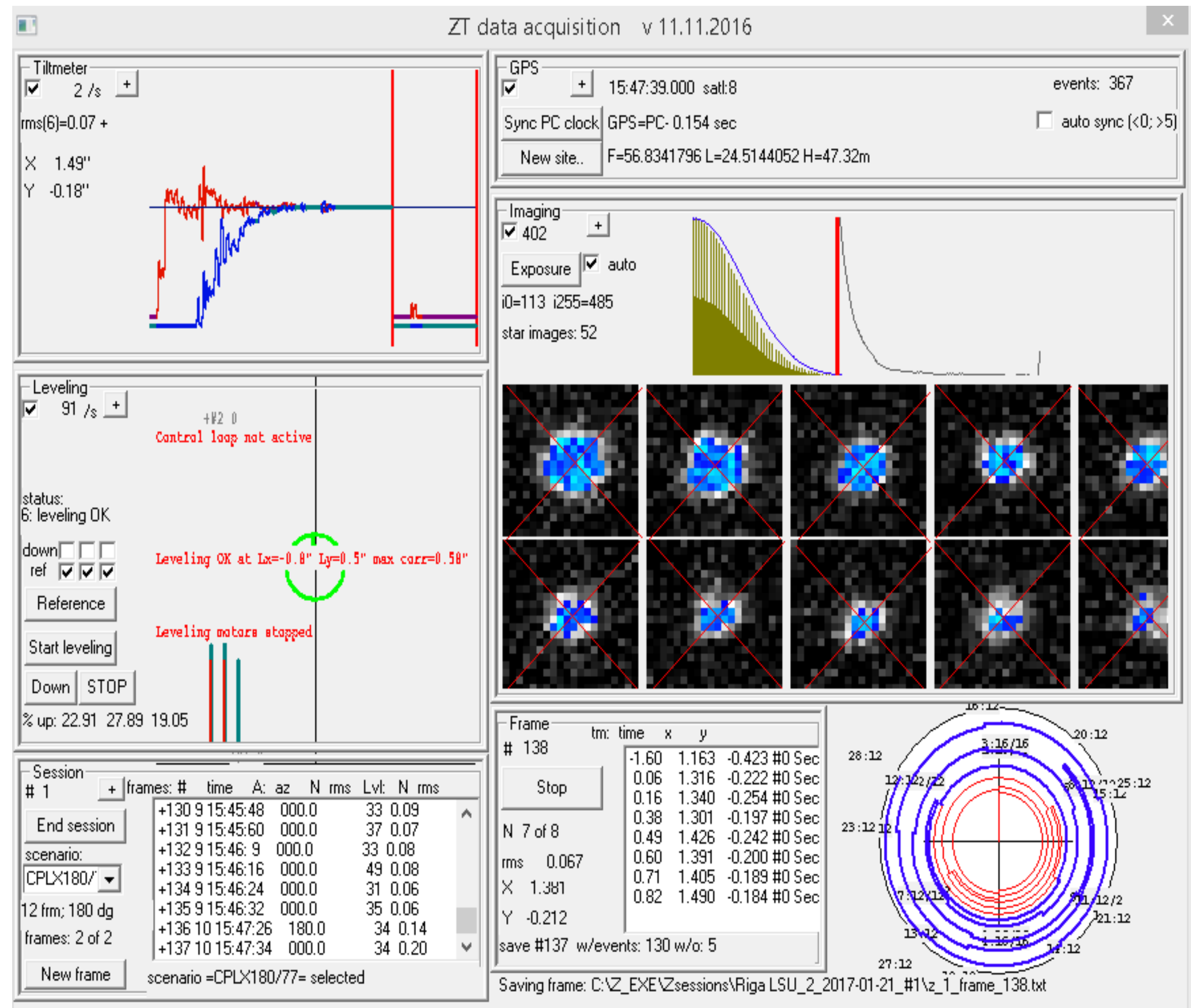


- Instrument consists of vertically oriented telescope, equipped with sensitive tiltmeter; assembly can be rotated around vertical axis.
- GNSS provides geocentric coordinates (15-20 cm) of instrument site in precise (10 mks) UTC time so that direction of reference ellipsoid's normal in astrometric coordinate system can be calculated.
- Reference star observations provide position of instrument main axis relative to reference ellipsoid's normal.
- Tiltmeter readings provide inclination corrections for the instrument main axis relative to plumb line.
- Difference of these positions would make vertical deflection, if all alignments and tiltmeter zero-points would be accurate.
- In reality, a number of measurements in different rotation positions must be performed to exclude alignment and zero-point errors.
- Directions of plumb line (L) and ellipsoid's normal (O) describe conuses around rotation axis (R) in rotating coordinate system of instrument main axis; their difference also is a conus, width and phase of it gives vertical deflection value.



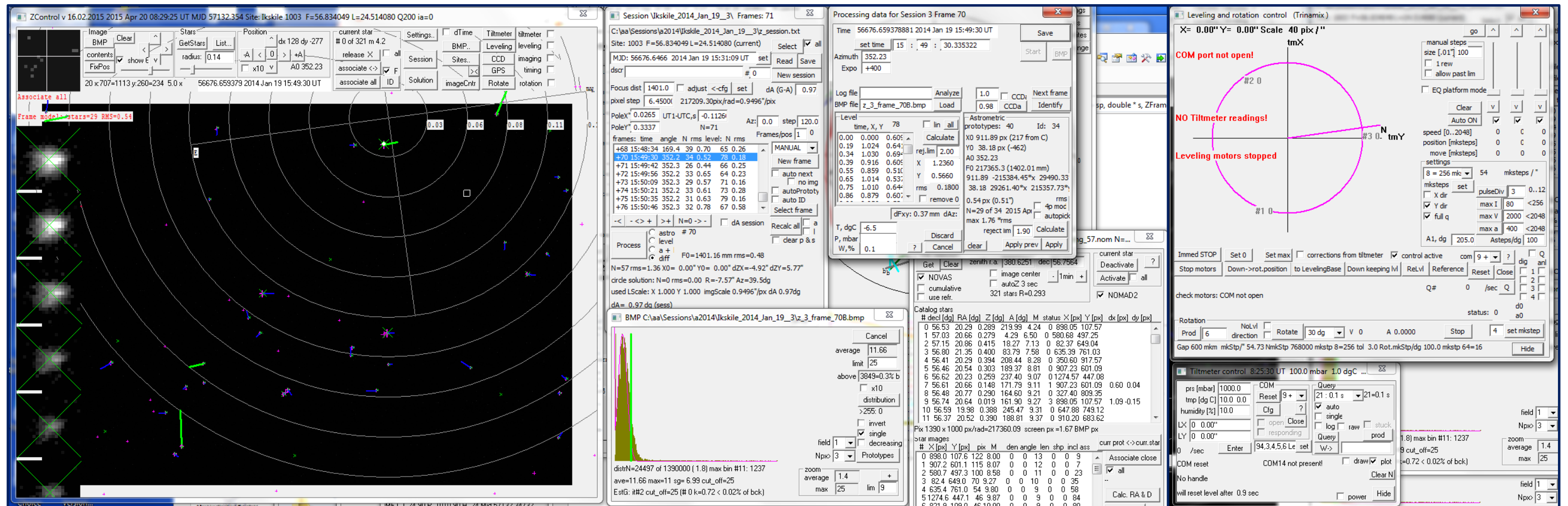
# Measurement Control

- All observation functionality within single interface window,
- Automatic mode supported,
- Measurement sequence specified in scenario script,
- 16 bit image intensity preserved,
- Support of uneven background intensity distribution analysis, improving twilight performance,
- GPS coordinates recorded for each session





# Data processing Software

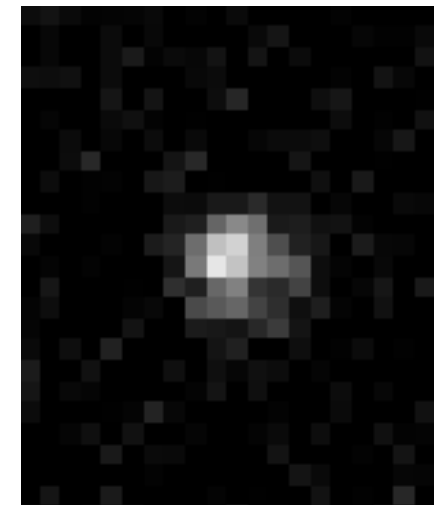
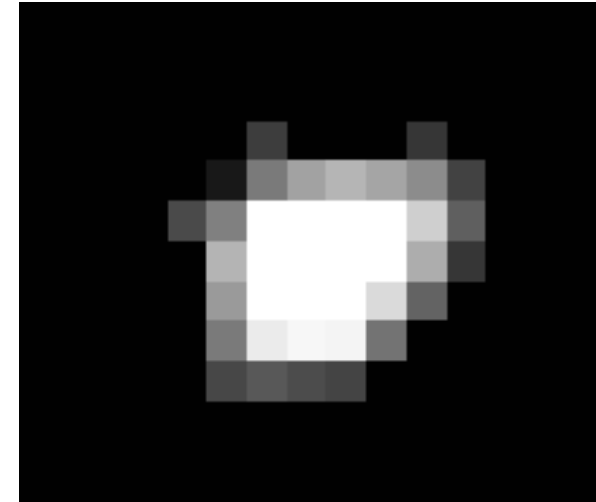


- C++ program, executing in Windows 8 environment.
- A subset of NOMAD (Naval Observatory Merged Astrometric Dataset) star catalog is used.
- NOVAS (Naval Observatory Vector Astrometry Software) package used for astrometric apparent places calculation.
- Automatic mode option is implemented for both measurements and data processing.
- Automatic reference star identification.
- In case of high vibrations data collecting is automatically interrupted.

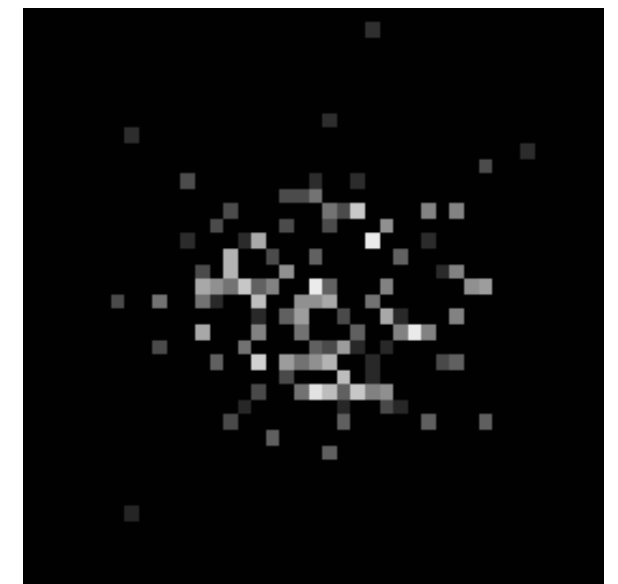
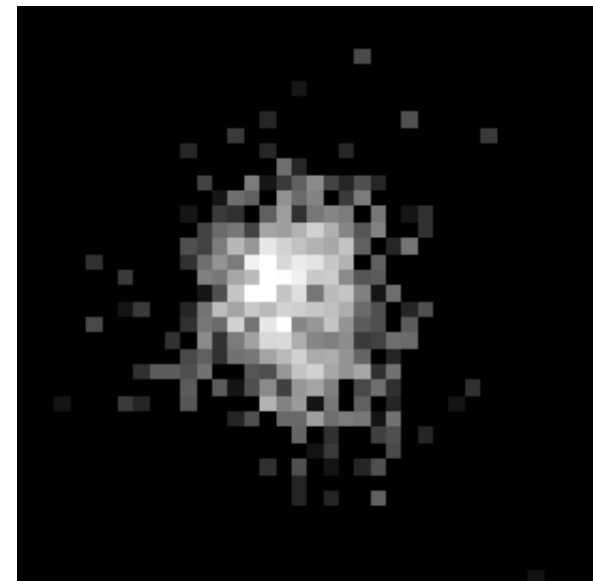
# Star images

➤ Aim: to obtain symmetrical image of good quality

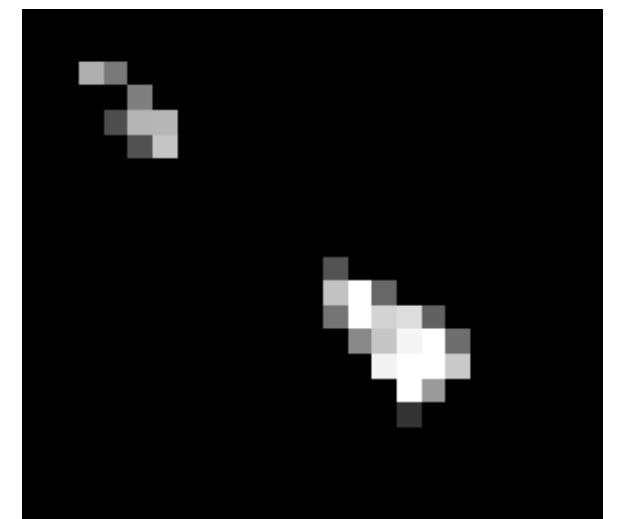
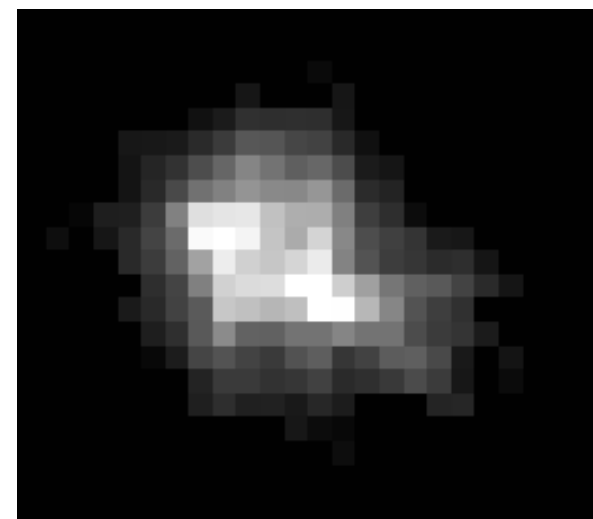
*Good conditions*



*Convection*

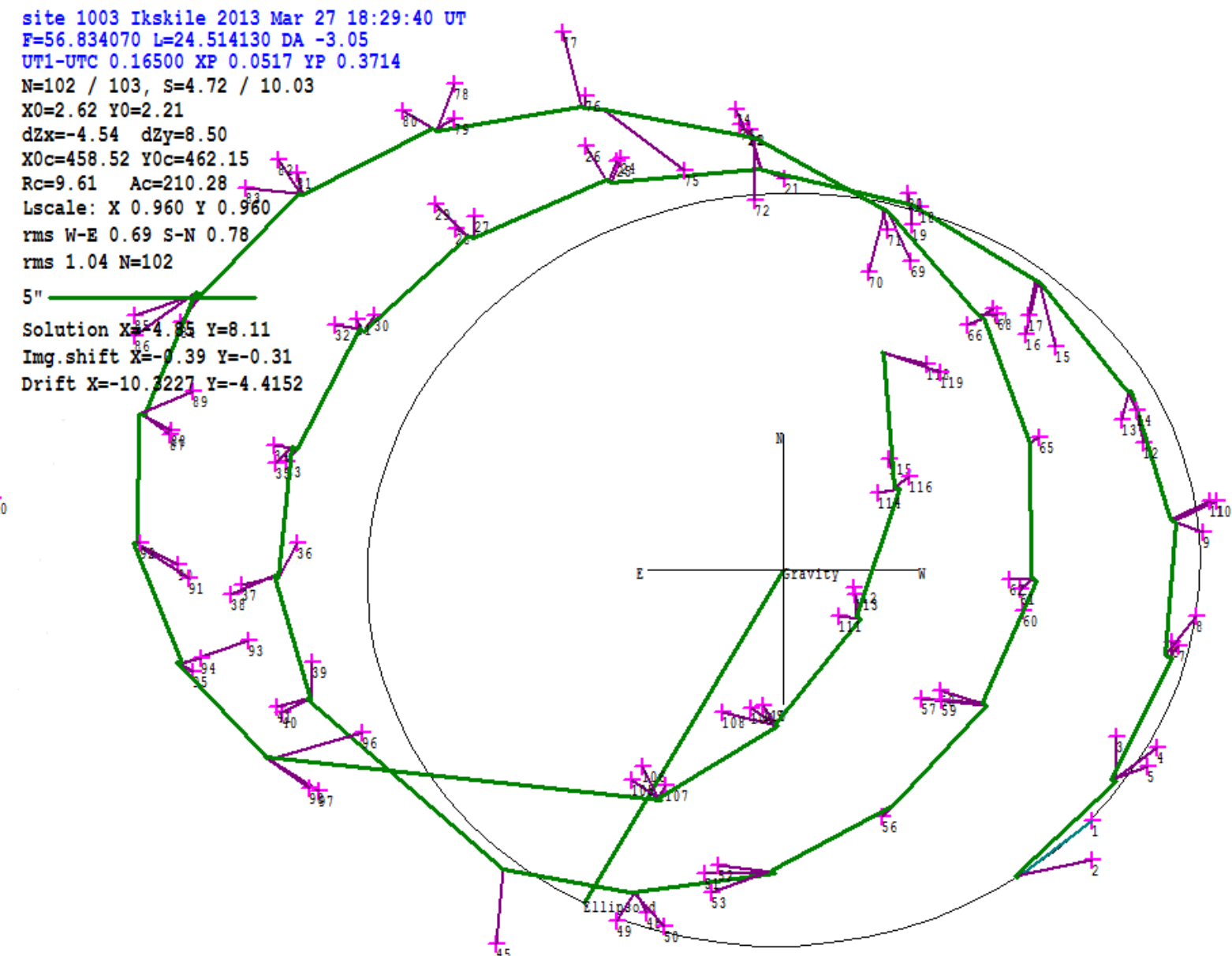
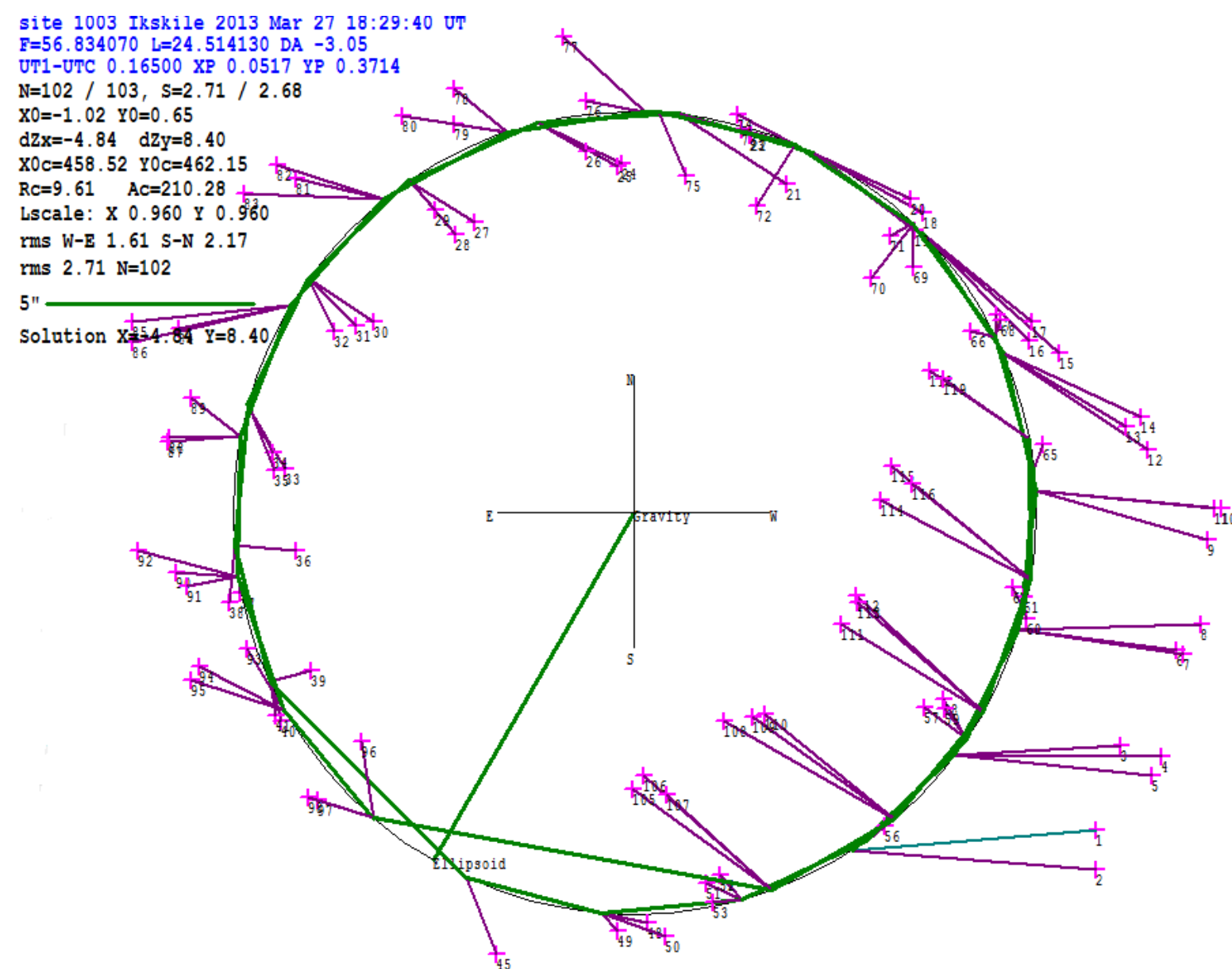


*Vibrations*





# The impact of linear drift of position ( $\sim 10''/h$ )



- In ideal conditions measurements should make **circle with radius of vertical deflection value**.
- In reality, thermal deformations change tiltmeter axis direction relative to optical system, resulting in spiralling trajectory.
- If dependence of deformations on time is close to linear, evaluation of simple compensating drift model is possible.

*Difference between directions to reference ellipsoid normal and plumb line in rotating imager coordinate system, where drift of relative component orientation has been almost linear during 40 minutes.*



# DZC-2019 (improvements)

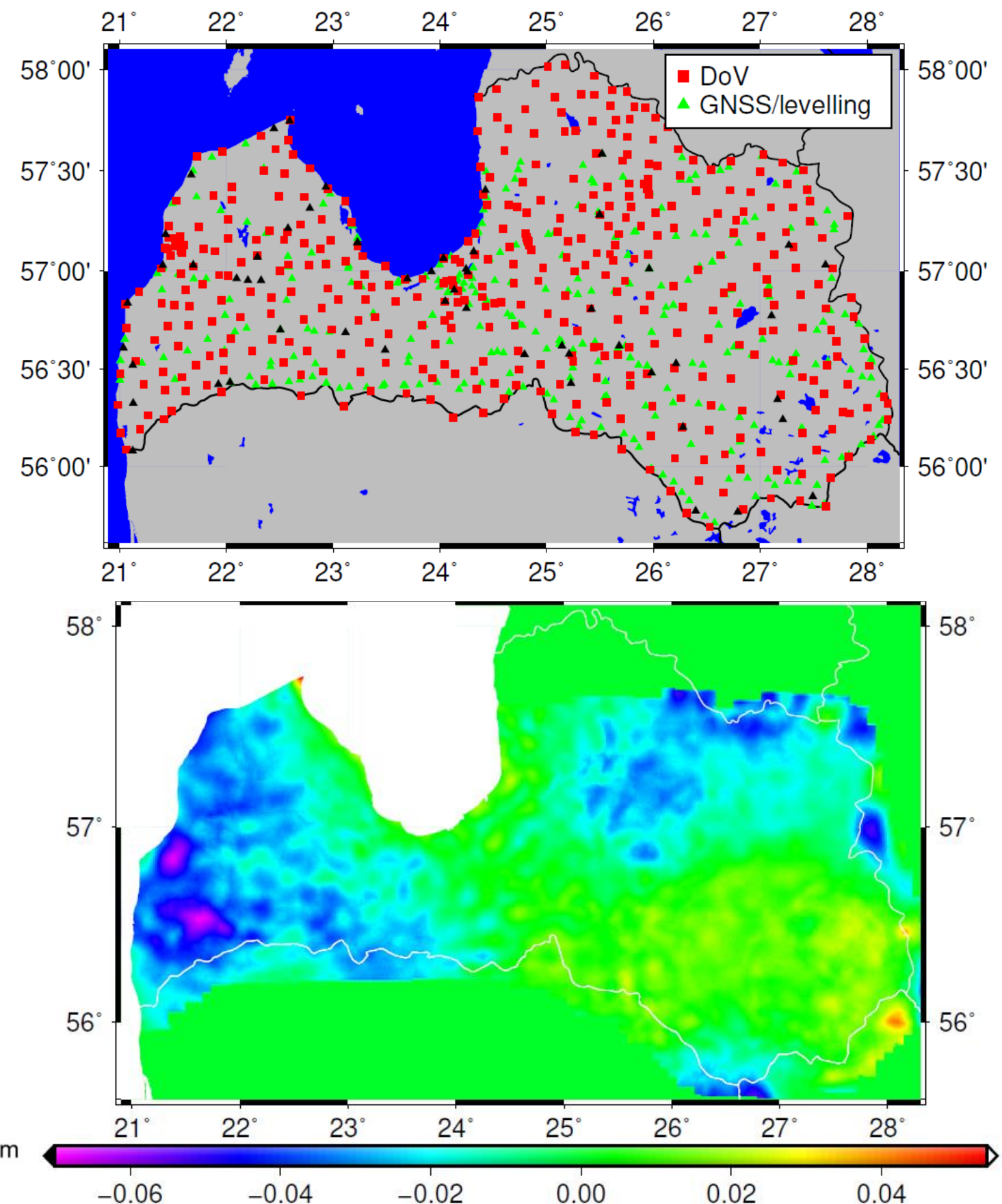
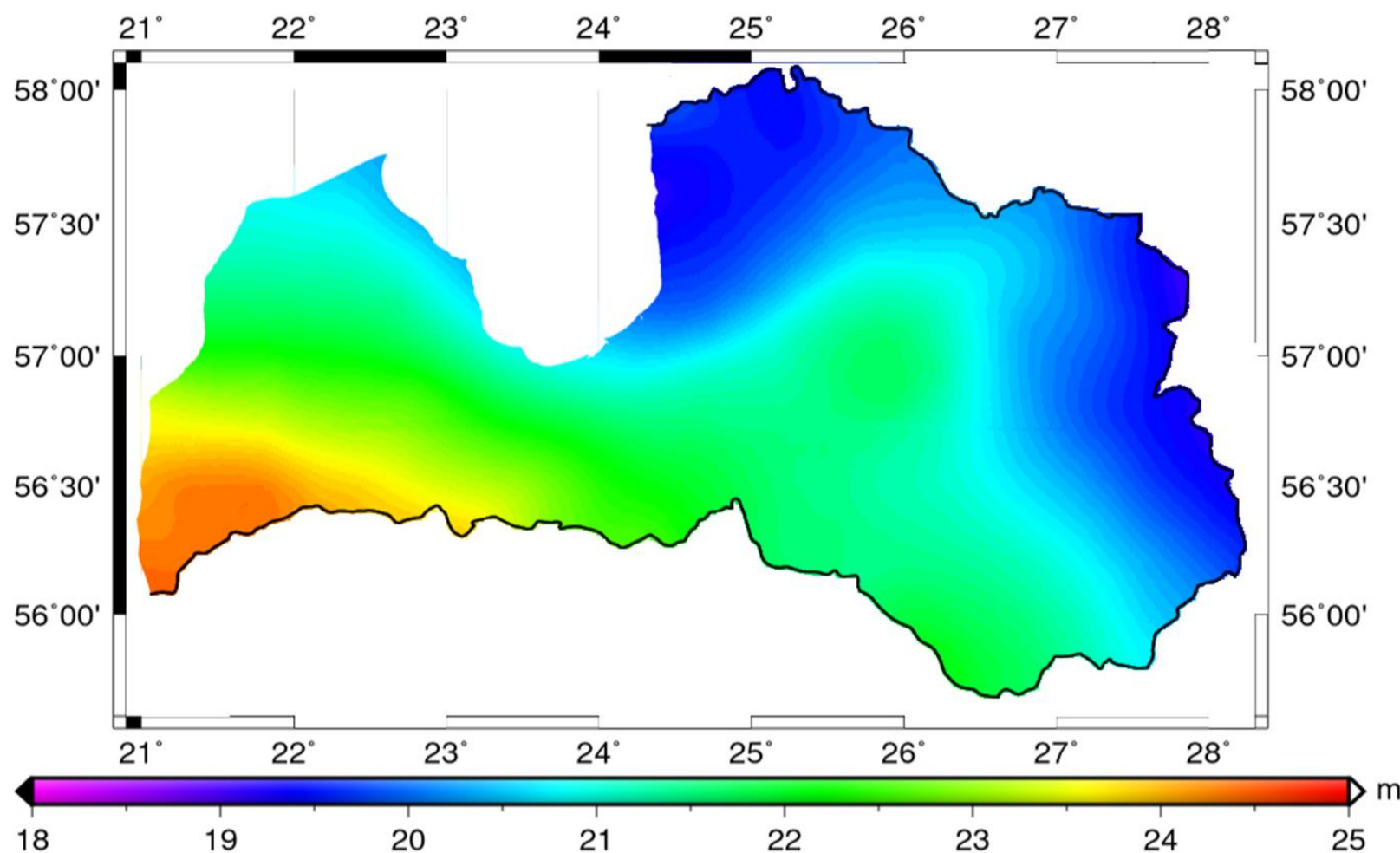
- Elaboration of measurement methodology,
- Measurement control software corrections and complements,
- Data processing improvements and automation,
- Transition to GAIA data release 2 star catalog,
- Revision of mechanical design, development of a new model,
- Currently 4 cameras are manufactured
- Commercial agreement with Louisiana State University





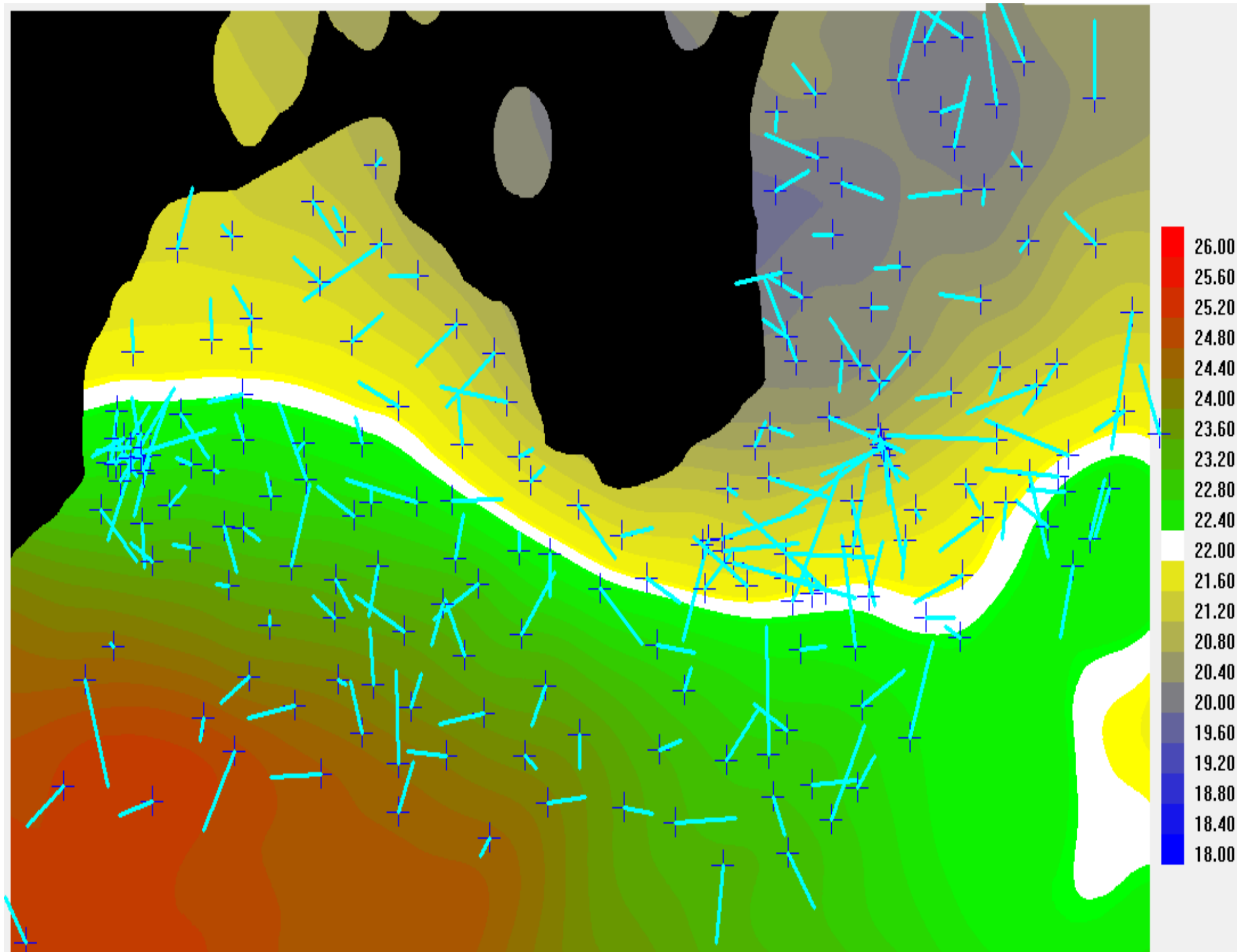
# High precision quasi-geoid model for the territory of Latvia

- 414 points (2017-2020)
- 1 – 1.5 hours one session
- max 6 points per night
- accuracy  $\sim 0.1''$



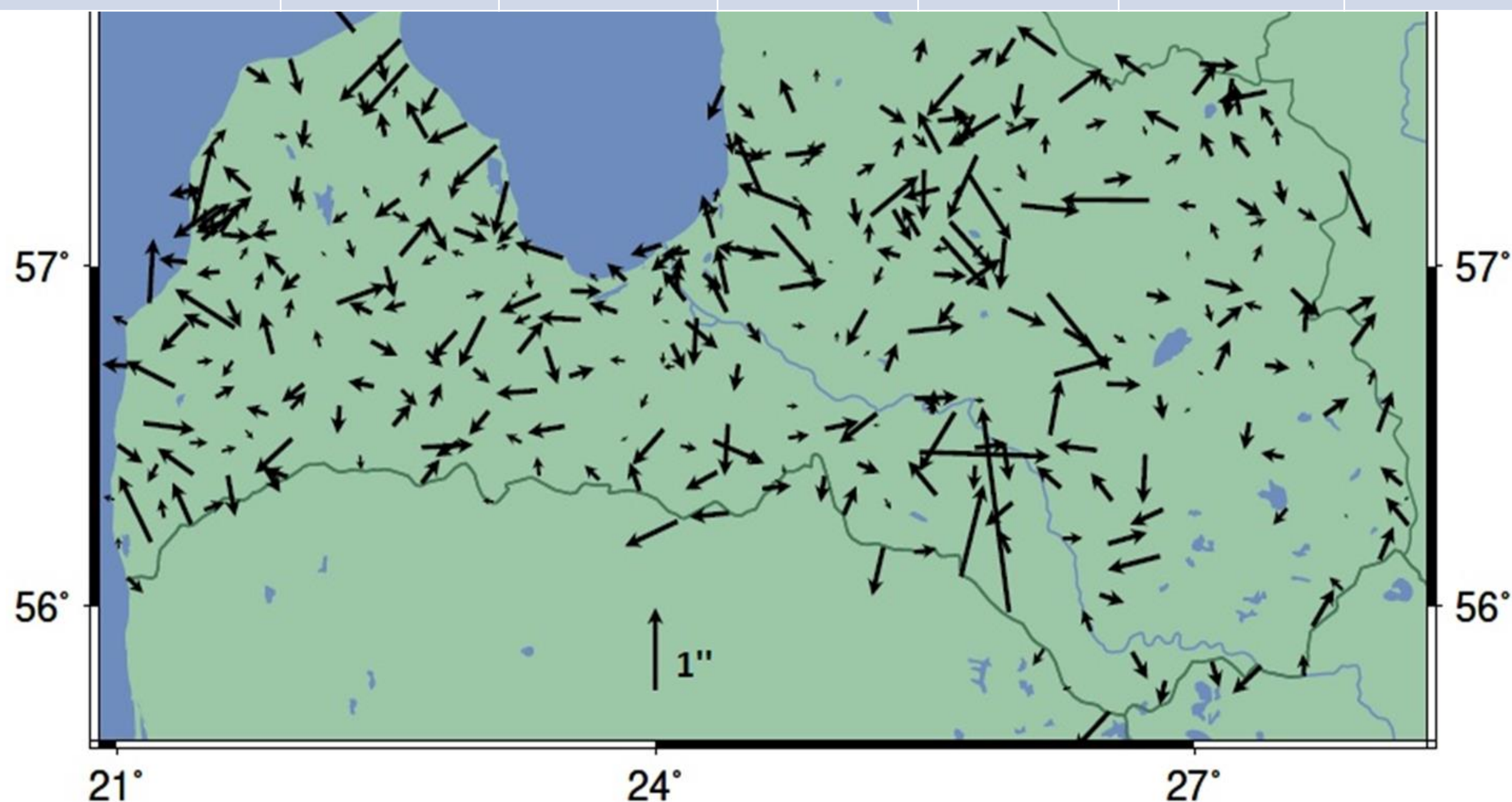


# Comparison of terrestrial VD and global GGM+ model



- Comparison with GGM+:
  - average 0.02"
  - rms 0.4"
  - Amplitude up to ~1"
- Can be used for:
  - Local geoid models,
  - Geological structure qualities,
  - Alternative to levelling.

	Min		Max		Avg	
	Xi	Eta	Xi	Eta	Xi	Eta
DFHRS	-0.190	-0.348	0.162	0.216	0.033	0.041
EGM2008	-1.128	-0.939	1.669	2.491	0.256	0.252



# Thank you!

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