



Concerning the Impact of Deorbiting Spacecraft to the Upper Atmosphere

Leonard Schulz and **Karl-Heinz Glassmeier**, EGU General Assembly 2020, Session PS2.2

Correspondence: l.schulz@tu-bs.de

Motivation/Problem

- Advancing commercialization of space industry
- Upcoming large satellite constellations (LC)
- Mitigation guidelines due to increasing space debris numbers

→ More and more human-made objects (satellites, rocket bodies, etc.) will reenter the atmosphere and disintegrate adding material to the upper atmosphere

Arising questions:

- How much mass enters annually?
- How much of this mass is injected into the atmosphere in what state?
- Which chemical composition does this material have?
- How will this change in the future (especially regarding LC)?
- How does this compare to the natural injection of meteoroids coming from asteroids, comets, etc.?
- Finally: How does this influence the environment/upper atmosphere? → Importance of metal injection

Annual Mass Input

Natural

- Differentiation between interplanetary dust particles (IDPs) and larger impactors ($>10^{-2}$ g)
 - IDP mass input: Many different studies with very different values:
 - Model of solar meteoroid flux (*Grün et al., 1985; Divine, 1993; Staubach et al., 1997; Dikarev et al., 2004*)
 - Visual, (radar), and satellite data (*Hughes, 1978*)
 - Impacts on satellite (*Love and Brownlee, 1993*)
 - Many other studies (see *Plane, 2012*; others)

→ About 20 kt/yr most probable
 - Larger impactor mass: about 2 kt/yr (after *Brown et al., 2002*)
- Around 22 kt impact Earth every year**

Anthropogenic

- Today about 190 t/yr, 60% satellites, 40% upper stages (combination of *Liou et al., 2018; McNair and Boykin, 1966*)
 - But: Core stages do reenter right after liftoff of a launch vehicle → Extra 0.5 kt per year
 - With future LC: 2 Scenarios:
 1. Most probable scenario: 14000 additional satellites and 235 upper stages within 5 years
→ 1 kt/yr (additional)
 2. Maximum scenario: 25000 additional satellites, 380 upper stages within 5 years
→ 1.7 kt/yr (additional)
- 2.4 kt/yr at maximum possible in the future**

Composition

Natural

- IDP composition after *Schramm et al., 1989; Arndt et al. 1996*
 - Largely silicates, but have a considerable amount of organic elements (51%) as they originate mostly from comets (*Jenniskens, 2015; Zolensky et al., 2006; Bardyn et al., 2017*)
- Large impactor composition is modelled after the meteorite composition found on Earth (*Grady, 2000; Lodders and Fegley, 1998; Mittlefehldt et al., 1998; Wasson, 1974; Demidova et al., 2007*)
 - Originate mostly from asteroids (*DeMeo and Binzel, 2008; Fernández et al., 2005; Bottke et al., 2002*)
 - Composition is less organic

→ 35 % of the mass are metal elements. Most abundant elements are O, Fe, Si, Mg, C, S each contributing more than 1kt/yr to the influx

Anthropogenic

- Hard to estimate as every object is different, but classes (e.g. rocket bodies) have similar composition
- Very high metal abundance (up to 95 % for upper stages), especially Al, Fe, and Ni coming from various alloys used for structural parts (*Henson, 2018*)
- Satellites have lower metal abundances, although they are still higher than for the natural material
- For satellites more diverse materials are used (*Finckenor, 2018*)

→ High metal abundance, non-metallic elements only minor. Main elements are Al, Fe, Ni

Atmospheric Processing upon Reentry

Highest uncertainties

Not all the mass is injected into the atmosphere, a portion hits the ground!

Different ablation products: Atoms, Dust (aerosols), Ground-reaching material (meteorites, debris)

Natural	Anthropogenic
<ul style="list-style-type: none">▪ IDPs:<ul style="list-style-type: none">• Atoms lost through sputtering (<i>Rogers et al., 2005</i>) at high altitudes and ablation (<i>e.g. Rietmeijer, 2002</i>) at lower altitudes• After that are slowed down → unmelted aerosols• All the mass is injected into the atmosphere!▪ Larger impactors (<i>Baldwin and Sheaffer, 1971; Klekociuk et al., 2005; Borovicka et al., 2019, others</i>):<ul style="list-style-type: none">• Negligible sputtering• Atoms and dust lost through ablation → Dust recondenses• Increasing aerosol mass with impactor size• Around 15 % of the mass reaches the ground	<ul style="list-style-type: none">▪ Different ablation behavior due to slower speeds, lower entry angle▪ Higher mass portion should survive, but measures to reduce risks on the ground are taken → complete demise is the goal → intentional breaking points, etc.▪ Our preliminary estimate: 30% gets to the ground, the rest is equally ablated in atomic and aerosol form

Summarizing Results

- Today, the natural atmosphere injection dominates, human-made objects contribute less than 1% of the mass
- With future LC this changes:
 1. Probable Scenario: 4% of the annually injected mass is human-made. Relative injection of metals is even higher (9%)
 2. Maximum Scenario: 7% of the annually injected mass is human-made. Relative metal injection up to 16%
- The anthropogenic injection of some metal elements might even prevail the natural injection: For example Al: 0.27 kt/yr natural, 0.37 or 0.62 kt/yr (Scenario 1 and 2 respectively).
- Disproportional increase of the injection of aerosols from human-made objects

→ With future satellite constellations, the anthropogenic injection becomes significant!

Problems and Outlook

- Although vast amounts of literature are available for some topics, often quite different values or results are given (e.g. natural mass input)
- Extensive study has to be performed (Paper is in preparation)
- Here given values might change marginally, but the overall significance of the anthropogenic injection remains

Possible effects on the environment?

- Climate (aerosols have a radiative forcing effect, see *Lawrence et al., 2018*)
 - Ionosphere (increase of the injection of atoms and ions)
 - Chemical processes in the atmosphere (maybe ozone?)
- Problems should be addressed in the future!

References I

Literature (in the order of appearance):

- Grün, E., Zook, H.A., Fechtig, H., Giese, R.H., 1985. Collisional balance of the meteoritic complex. *Icarus* 62, 244–272. doi:10.1016/0019-1035(85)90121-6.
- Divine, N., 1993. Five populations of interplanetary meteoroids. *Journal of Geophysical Research* 98, 17029–17048. doi:10.1029/93JE01203.
- Staubach, P., Grün, E., Jehn, R., 1997. The meteoroid environment near earth. *Advances in Space Research* 19, 301–308. doi:10.1016/S0273-1177(97)00017-3.
- Dikarev, V., Grün, E., Baggaley, J., Galligan, D., Landgraf, M., Jehn, R., 2004. Modeling the sporadic meteoroid background cloud. *Earth, Moon, and Planets* 95, 109–122. doi:10.1007/s11038-005-9017-y.
- Hughes, D.W., 1978. Meteors, in: McDonnell, J.A.M. (Ed.), *Cosmic Dust*. Wiley, pp. 123–185.
- Love, S.G., Brownlee, D.E., 1993. A direct measurement of the terrestrial mass accretion rate of cosmic dust. *Science* 262, 550–553. doi:10.1126/science.262.5133.550.
- Plane, J.M.C., 2012. Cosmic dust in the earth's atmosphere. *Chem. Soc. Rev.* 41, 6507–6518. doi:10.1039/C2CS35132C.
- Brown, P., Spalding, R.E., ReVelle, D.O., Tagliaferri, E., Worden, S.P., 2002. The flux of small near-earth objects colliding with the earth. *Nature* 420, 294–296. doi:10.1038/nature01238.
- Liou, J.C., Matney, M., Vavrin, A., Manis, A., Gates, D., 2018. NASA ODPO's Large Constellation Study. *Orbital Debris Quarterly News* 22, 4–7.
- Mc Nair, A., Boykin, E., 1966. Earth orbital lifetime prediction model and program. *NASA Technical Memorandum X-53385*.
- Schramm, L.S., Brownlee, D.E., Wheelock, M.M., 1989. Major element composition of stratospheric micrometeorites. *Meteoritics* 24, 99–112. doi:10.1111/j.1945-5100.1989.tb00950.x.
- Arndt, P., Bohsung, J., Maetz, M., Jessberger, E.K., 1996. The elemental abundances in interplanetary dust particles. *Meteoritics & Planetary Science* 31, 817–833. doi:10.1111/j.1945-5100.1996.tb02116.x.
- Jenniskens, P., 2015. Meteoroid Streams and the Zodiacal Cloud, in: Michel, P., DeMeo, F.E., Bottke, W.F. (Eds.), *Asteroids IV*, pp. 281–295. doi:10.2458/azu_uapress_9780816532131-ch015.
- Zolensky, M.E., Zega, T.J., Yano, H., Wirick, S., Westphal, A.J., Weisberg, M.K., Weber, I., Warren, J.L., Velbel, M.A., Tsuchiyama, A. et al., 2006b. Mineralogy and petrology of comet 81p/wild 2 nucleus samples. *Science* 314, 1735–1739. doi:10.1126/science.1135842.
- Bardyn, A., Baklouti, D., Cottin, H., Fray, N., Briois, C., Paquette, J., Stenzel, O., Engrand, C., Fischer, H., Hornung, K., et al., 2017. Carbon-rich dust in comet 67P/Churyumov-Gerasimenko measured by COSIMA/Rosetta. *Monthly Notices of the Royal Astronomical Society* 469, 712–722. doi:10.1093/mnras/stx2640.
- Grady, M.M., 2000. *Catalogue of Meteorites*. Cambridge University Press.
- Lodders, K., Fegley, B., 1998. *The Planetary Scientist's Companion*.
- Mittlefehldt, D.W., McCoy, T.J., Goodrich, C.A., Kracher, A., 1998. Non-chondritic meteorites from asteroidal bodies, in: Papike, J.J. (Ed.), *Planetary Materials*. Mineralogical Society of America. volume 36 of *Reviews in Mineralogy*, pp. 4–001–4–196.

References II

Literature (continued):

- Wasson, J., 1974. Meteorites: Classification and Properties. Minerals, Rocks and Mountains, Springer Berlin Heidelberg.
- Demidova, S.I., Nazarov, M.A., Lorenz, C.A., Kurat, G., Brandstätter, F., Ntaflos, T., 2007. Chemical composition of lunar meteorites and the lunar crust. *Petrology* 15, 386–407. doi:10.1134/S0869591107040042.
- DeMeo, F., Binzel, R.P., 2008. Comets in the near-earth object population. *Icarus* 194, 436 – 449. doi:https://doi.org/10.1016/j.icarus.2007.10.011.
- Fernández, Y.R., Jewitt, D.C., Sheppard, S.S., 2005. Albedos of asteroids in comet-like orbits. *The Astronomical Journal* 130, 308–318. doi:10.1086/430802.
- Bottke, W.F., Morbidelli, A., Jedicke, R., Petit, J.M., Levison, H.F., Michel, P., Metcalfe, T.S., 2002. Debaised orbital and absolute magnitude distribution of the near-earth objects. *Icarus* 156, 399–433. doi:https://doi.org/10.1006/icar.2001.6788.
- Henson, G., 2018. Materials for Launch Vehicle Structures, in: Bhat, B. N. (Ed.), *Aerospace Materials and Applications*. American Institute of Aeronautics and Astronautics, pp. 435-504.
- Finckenor, M. M., 2018. Materials for Spacecraft, in: Bhat, B. N. (Ed.), *Aerospace Materials and Applications*. American Institute of Aeronautics and Astronautics, pp. 403-434.
- Rogers, L., Hill, K., Hawkes, R., 2005. Mass loss due to sputtering and thermal processes in meteoroid ablation. *Planetary and Space Science* 53, 1341 – 1354. doi:https://doi.org/10.1016/j.pss. 2005.07.002.
- Rietmeijer, F.J.M., 2002. Collected Extraterrestrial Materials: Interplanetary Dust Particles Micrometeorites, Meteorites, and Meteoric Dust, in: Murad, E., Williams, I.P. (Eds.), *Meteors in the Earth's Atmosphere*, p. 215.
- Baldwin, B., Sheaffer, Y., 1971. Ablation and breakup of large meteoroids during atmospheric entry. *Journal of Geophysical Research (1896-1977)* 76, 4653–4668. doi:10.1029/JA076i019p04653.
- Klekociuk, A., Brown, P.G., Pack, D., ReVelle, D.O., Edwards, W.N., Spalding, R.E., Tagliaferri, E., Yoo, B.B., Zagari, J., 2005. Meteoric dust from the atmospheric disintegration of a large meteoroid. *Nature* 436, 1132–1135. doi:10.1038/nature03881.
- Borovicka, J., Popova, O., Spurný, P., 2019. The maribo cm2 meteorite fallsurvival of weak material at high entry speed. *Meteoritics & Planetary Science* 54, 1024–1041. doi:10.1111/maps.13259.
- Lawrence, M.G., Schäfer, S., Muri, H. et al., 2018. Evaluating climate geoengineering proposals in the context of the Paris Agreement temperature goals. *Nat Commun* 9, 3734. https://doi.org/10.1038/s41467-018-05938-3

Images:

- Earth space environment: NASA Orbital Debris Program Office. URL: <https://www.doncio.navy.mil/FileHandler.ashx?id=9327> [29.04.2020]
- ATV-1 Jules Verne reentry: NASA/ESA/Bill Moede and Jesse Carpenter. URL: https://www.nasa.gov/images/content/280802main_ATV_full.jpg [20.06.2019]
- Chelyabinks meteor trail: Nikita Plekhanov, 2013, unchanged. URL: <http://gallery.ru/watch?ph=z6Q-ewl8g> [29.04.2020]