

# Regulation of Proton-Alpha Differential Flow by Compressive Fluctuations and Ion-scale Instabilities in the Solar Wind

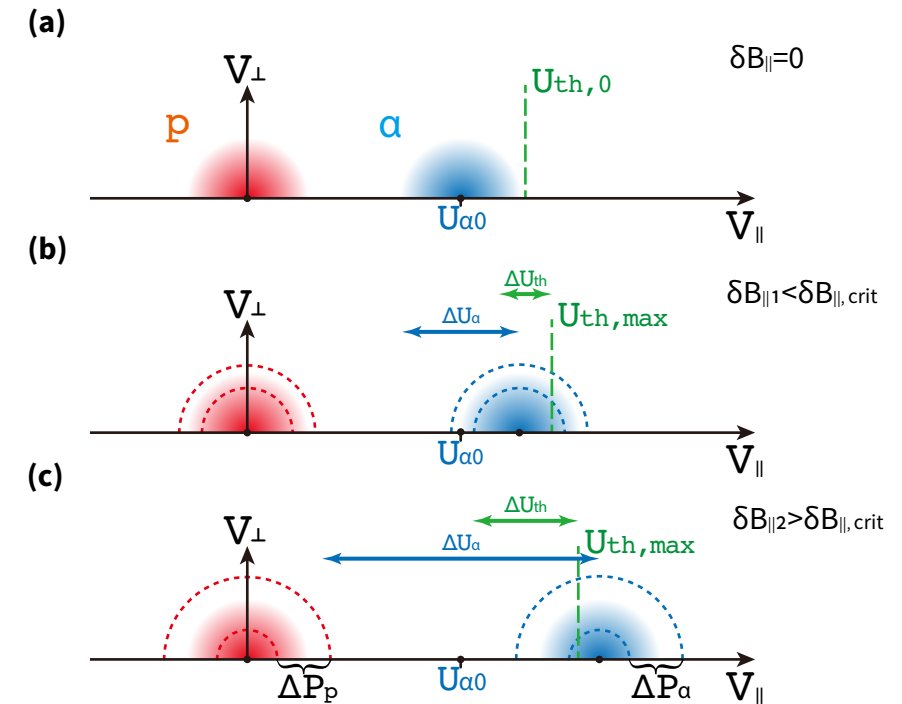
Xingyu Zhu<sup>1,2</sup>, Daniel Verscharen<sup>1</sup>, Jiansen He<sup>2</sup>, Benett A. Maruca<sup>3</sup>, Christopher J. Owen<sup>1</sup>

1. Mullard Space Science Laboratory, University College London, Dorking RH5 6NT, UK

2. School of Earth and Space Sciences, Peking University, Beijing 100871, China

3. Bartol Research Institute, Department of Physics and Astronomy, University of Delaware, Newark, DE 19716, USA

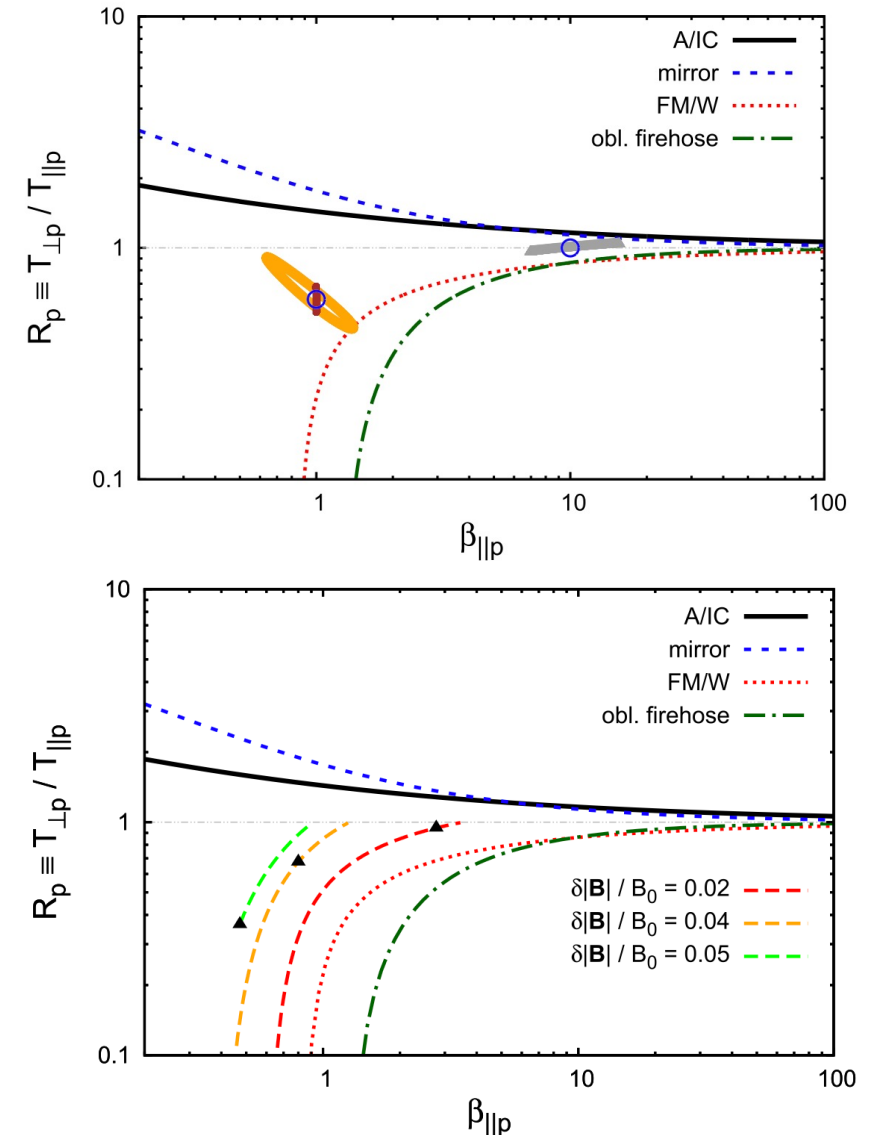
email: xingyu-zhu@ucl.ac.uk



# Fluctuating-anisotropy effect

- The slow-mode-like fluctuations can significantly modify  $n$ ,  $T$  and  $\mathbf{B}$ , when the amplitude is large. Therefore, it can transport the plasma in the  $\frac{T_{\perp p}}{T_{\parallel p}} - \beta_{p\parallel}$  plane around the background plasma parameter.
- Once the plasmas “touch” the instability threshold, The excited waves can scatter the protons and thereby raise up the average value of  $\frac{T_{\perp p}}{T_{\parallel p}}$ .
- This effect maintain the average value of  $\frac{T_{\perp p}}{T_{\parallel p}}$  away from the instability thresholds at a location in the parameter space depending on the wave-amplitude.

Verscharen et al. (2017)



# Motivation

- We extend the framework of the “fluctuating-anisotropy effect” to a broader “fluctuating-moment theory”, incorporating the differential flow between different species (like protons and  $\alpha$ -particles,  $U_\alpha$ ), which we also call “fluctuating-beam effect”.
- We aim to study the dependence of the effective threshold (maximum permitted value) of  $U_{\alpha 0}$  on  $\beta_{p\parallel}$  and  $\delta B_{\parallel}/B_0$  for Alfvén/ion-cyclotron (A/IC) and Fast-magnetosonic/whistler (FM/W) instabilities.
- We consider large-scale quasi-perpendicular slow waves using multi-fluid model considering isotropic temperature.

(Due to time limitation, we wouldn't present the results for FM/W instability, the thresholds of which is greater than A/IC instability.)

# Illustration of the Fluctuating-beam Effect

## (a). No slow mode ( $\delta B_{\parallel} = 0$ ):

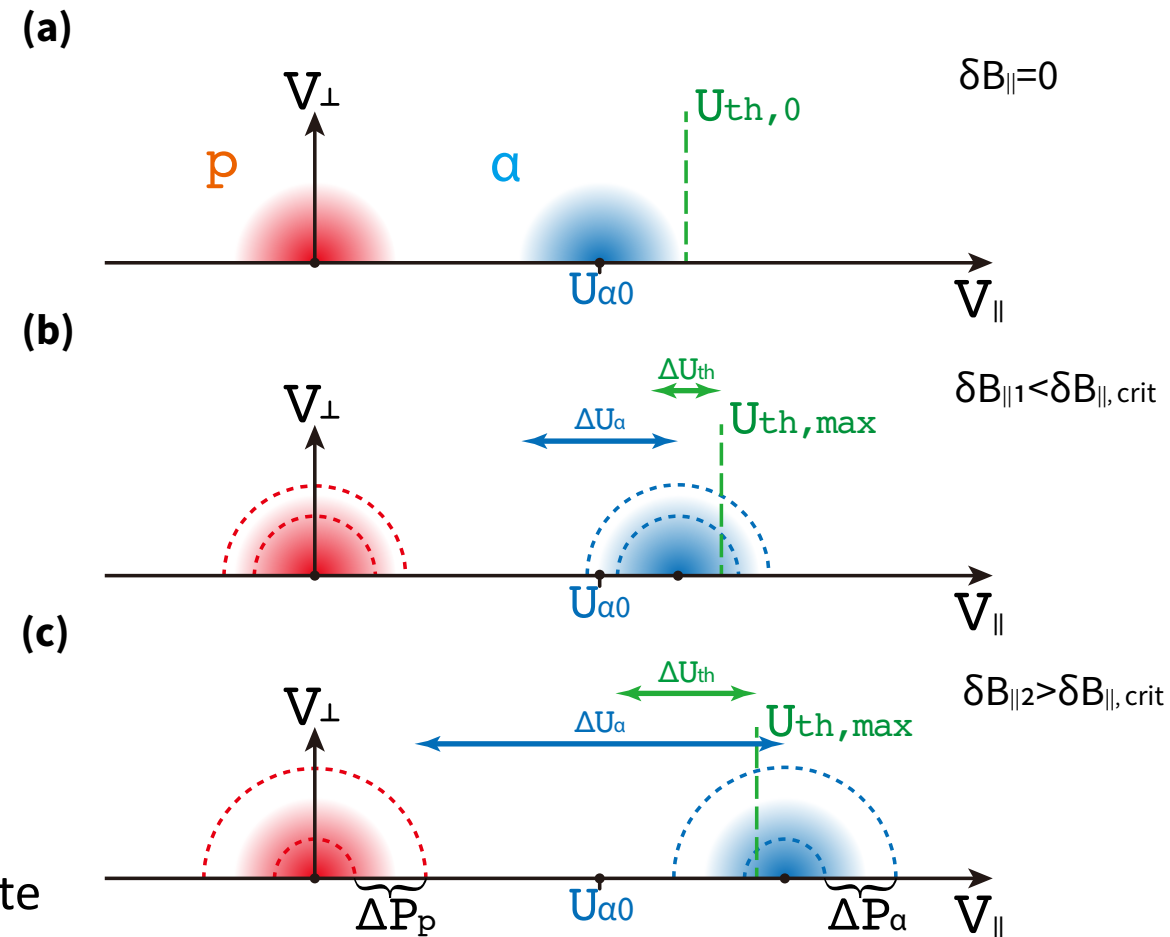
the  $\alpha$ -particle drift velocity ( $U_{\alpha 0}$ ) is below the instability threshold ( $U_{th,0}$ ) (stable configuration)

## (b). A small-amplitude slow wave ( $\delta B_{\parallel 1} < \delta B_{\parallel, \text{crit}}$ ):

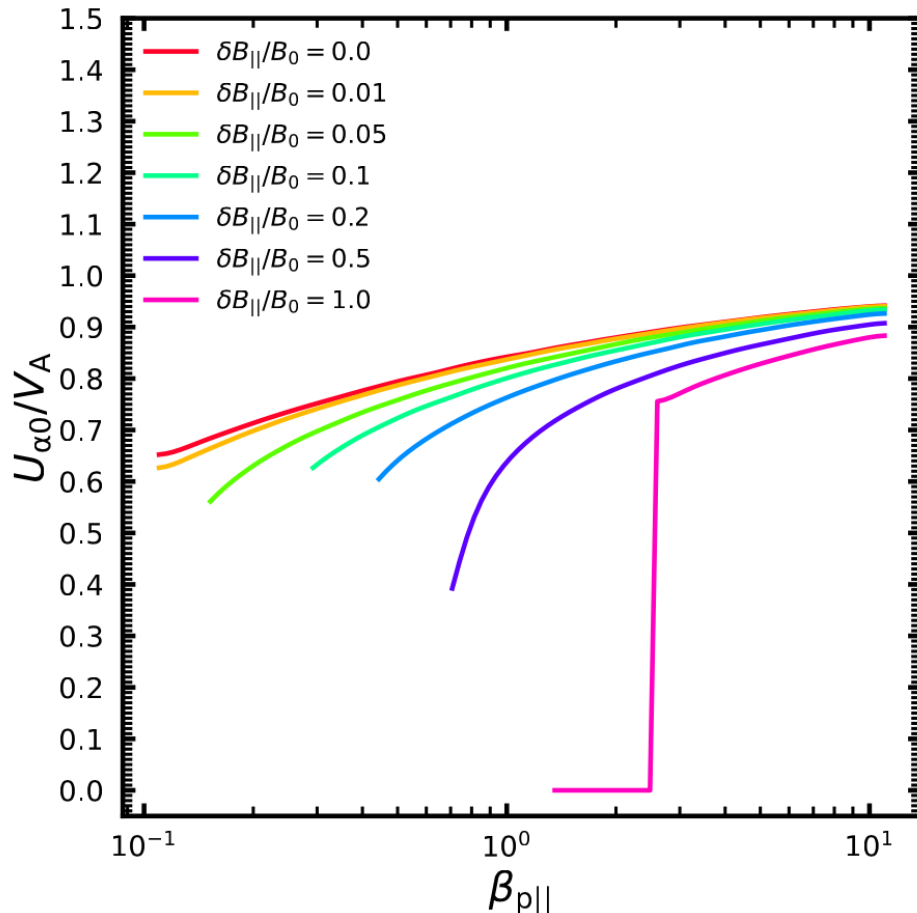
Plasma parameters and fields start to fluctuate. However, the amplitude is so small that  $U_{\alpha}(t) < U_{th}(t)$  throughout the evolution of slow-mode fluctuations. (stable configuration)

## (c). A large-amplitude slow wave ( $\delta B_{\parallel 2} > \delta B_{\parallel, \text{crit}}$ ):

As wave amplitude grows, the amplitude of beam velocity grows faster, it will exceed the instability threshold at a finite time interval during slow-mode wave evolution.



# Effective thresholds for A/IC Instability



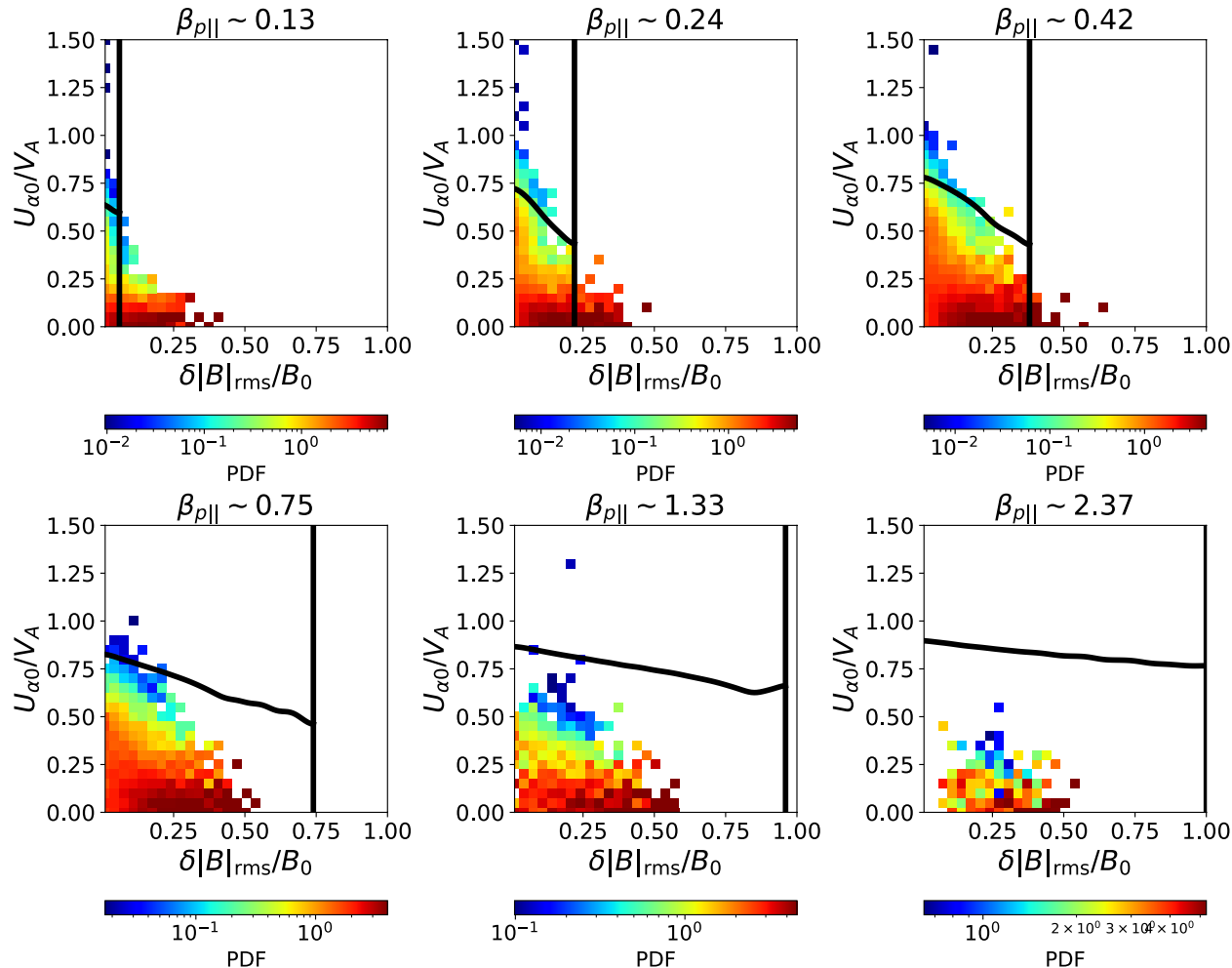
- Expression for A/IC instability threshold (Verscharen et al. 2013):

$$U_{\alpha} > \frac{T_{\perp\alpha} + T_{\parallel\alpha}}{4T_{\perp\alpha}} \left[ \sigma_1 w_{\parallel P} + V_A + \sqrt{(V_A + \sigma_1 w_{\parallel P})^2 - 2V_A^2} \right] - \sigma_1 w_{\parallel P},$$

- As  $\delta B_{||}/B_0$  becomes larger, the effective instability threshold  $U_{\alpha 0, \text{crit}}$  (maximum permitted  $U_{\alpha 0}$ ) is reduced (more significant when  $\beta < 1$ )
- Slow-mode-like fluctuations make it easier for plasmas to become unstable to the A/IC instability and thereby regulate  $U_{\alpha}$  through wave-particle interaction.

# Comparison with WIND Observations

■ 1-D probability distribution function of  $U_{\alpha 0}/V_A$  at different  $\delta|B|_{rms}/B_0$  at six  $\beta_{p||}$ :



- The effective A/IC instability thresholds (oblique black curves) decrease with the compression amplitude
- The linear assumption break-down amplitude increases with plasma beta (black vertical line).
- The effective instability thresholds restrict the data to the stable space.
- This finding is broadly consistent with our “fluctuating-beam effect” that the large-scale slow-mode-like fluctuations regulates the alpha drift velocity by kinetic instability in the solar wind .

# Conclusion

- We propose a new concept of **fluctuating-beam effect**:

The large-scale slow-mode-like fluctuations can reduce the effective instability thresholds depending on the compression amplitude and regulate  $\alpha$ -particle drift velocity through kinetic instability.

- The distributions of the measured  $\alpha$ -particle drift velocity have an inverse dependence on the compressive wave amplitude and are restricted by the effective A/IC instability thresholds.

