

Extraterrestrial ^3He shows that Mesozoic marl-limestone alternations are mainly driven by CaCO_3 variations at the astronomical timescale

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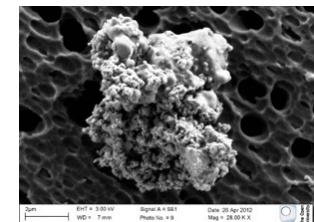
1 CRPG, CNRS, Université de Lorraine, Nancy, France, 2 Aix Marseille Université, CNRS, IRD, INRAE, Collège de France, CEREGE, Aix-en-Provence, France, 3 LGLTPE UMR CNRS 5276, Université de Lyon 1, France

^3He
ET



A new look with Extra Terrestrial ^3He !

Interplanetary Dust

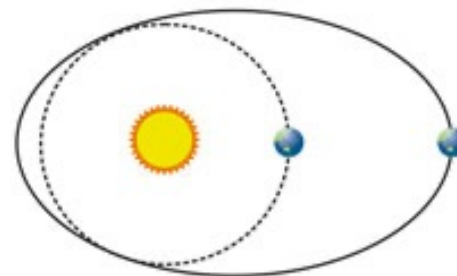


1 – Scientific question: why are they marl-limestone alternations?



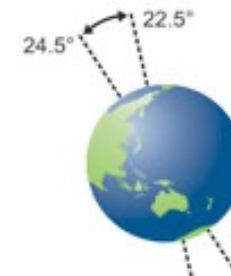
Drôme, France, Hauterivian

Milankovitch Cycles



Eccentricity

100 ka
400 ka



Obliquity

41 ka



Precession

19 ka
23 ka

The drivers are the astronomical cycles, but what is the internal mechanism?

The image shows a 12x3 grid of squares. The pattern of gray and white squares is as follows:

Gray	Gray	Gray
White	White	White
White	White	White
Gray	White	White
White	White	White
White	White	White
Gray	White	White
White	White	White
White	White	White
Gray	White	White
White	White	White
White	White	White

Silico-clastic sedimentation

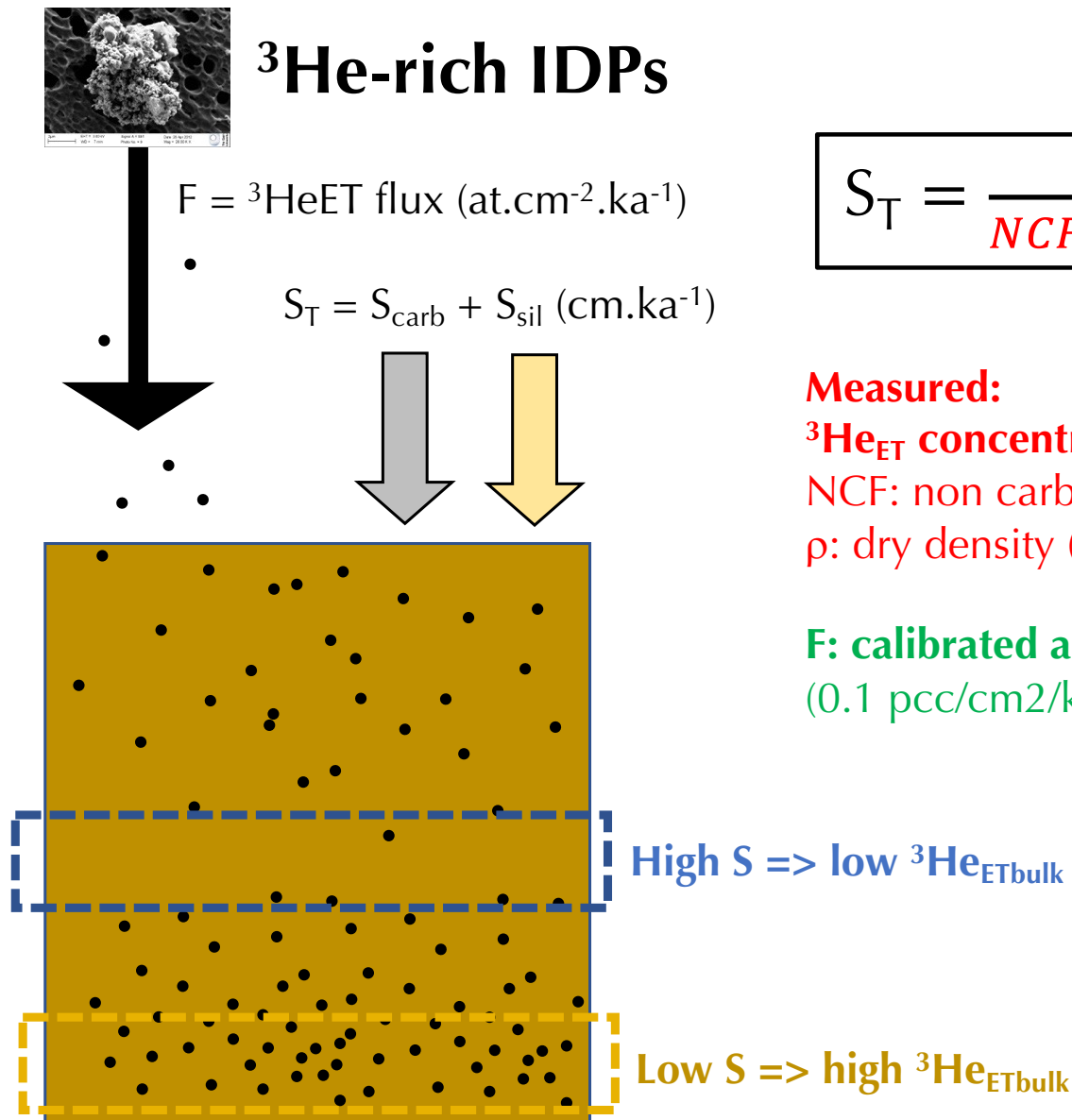
**Variable
CaCO₃
sedimentation**

CaCO₃ sedimentation

Variable Silico-clastic sedimentation

Einsele and Seilacher, 1982

2 – Method: extraterrestrial ^3He to introduce time into space



$$S_{\text{T}} = \frac{F}{NCF \times \rho \times ^3\text{He}_{\text{ET}}}$$

Measured:

$^3\text{He}_{\text{ET}}$ concentration (by mass spectrometry).

NCF: non carbonate fraction (by acid digestion)

ρ : dry density (g.cm^{-3}).

F: calibrated and assumed constant
(0.1 pcc/cm²/ka, Farley et al., 2012)

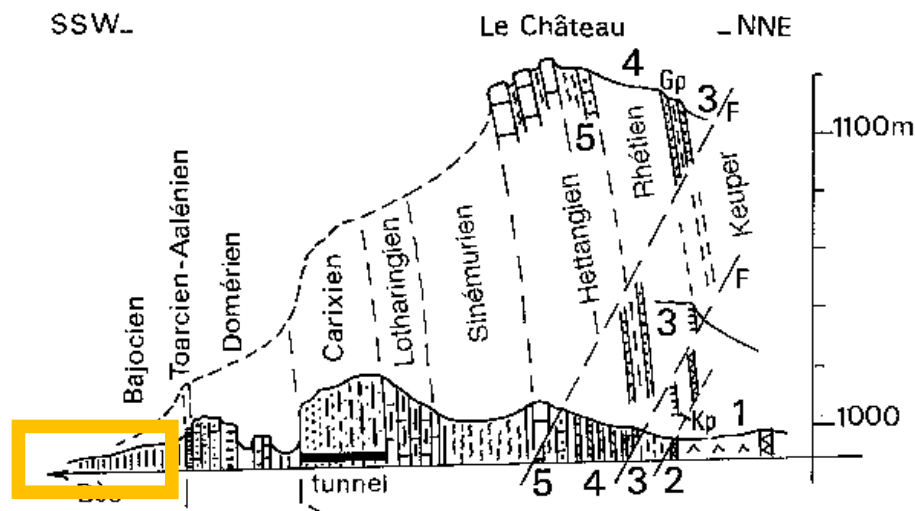


2 – Geological settings: 2 sampled sections: Bajocian and Valanginian



2 sections in the
Southern French Alps:

Bajocian (170 Ma)
Valanginian (135 Ma)



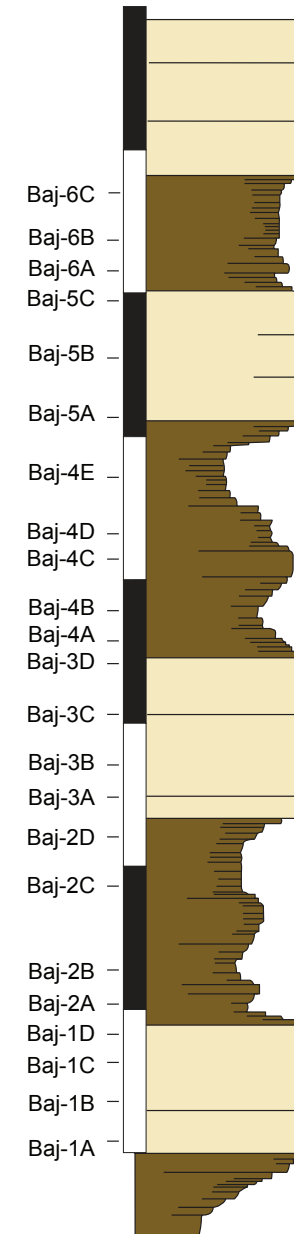
3 – Sedimentation rates results: Bajocian

Bajocian (170 Ma) :

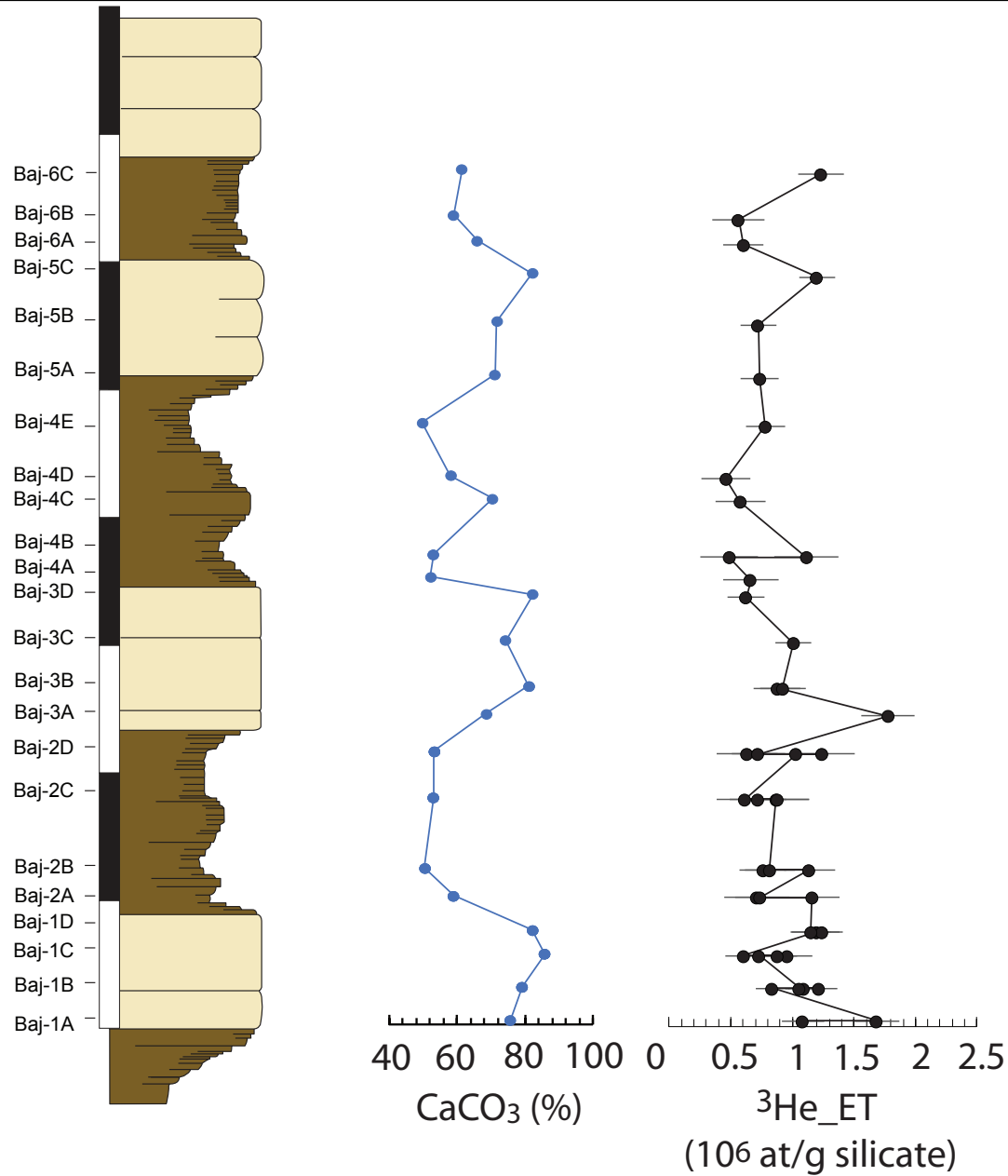
- 3.5 m section
- 3 marl limestones doublets
- 23 samples



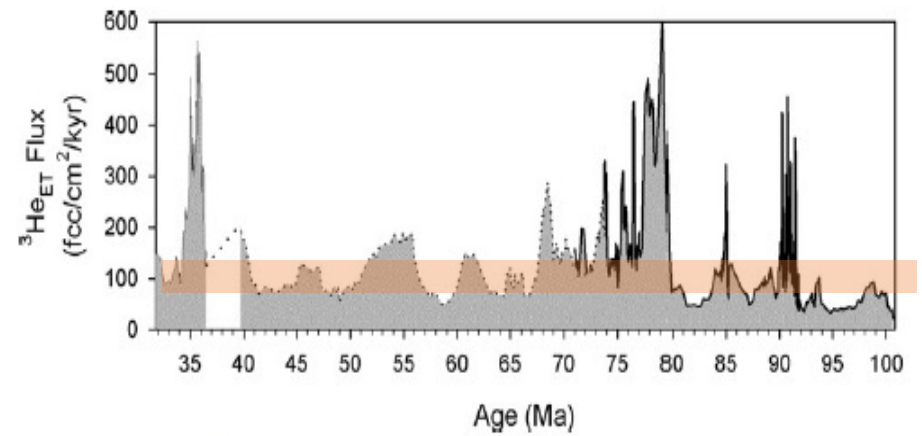
Blard et al., in prep



3 – Sedimentation rates results: Bajocian



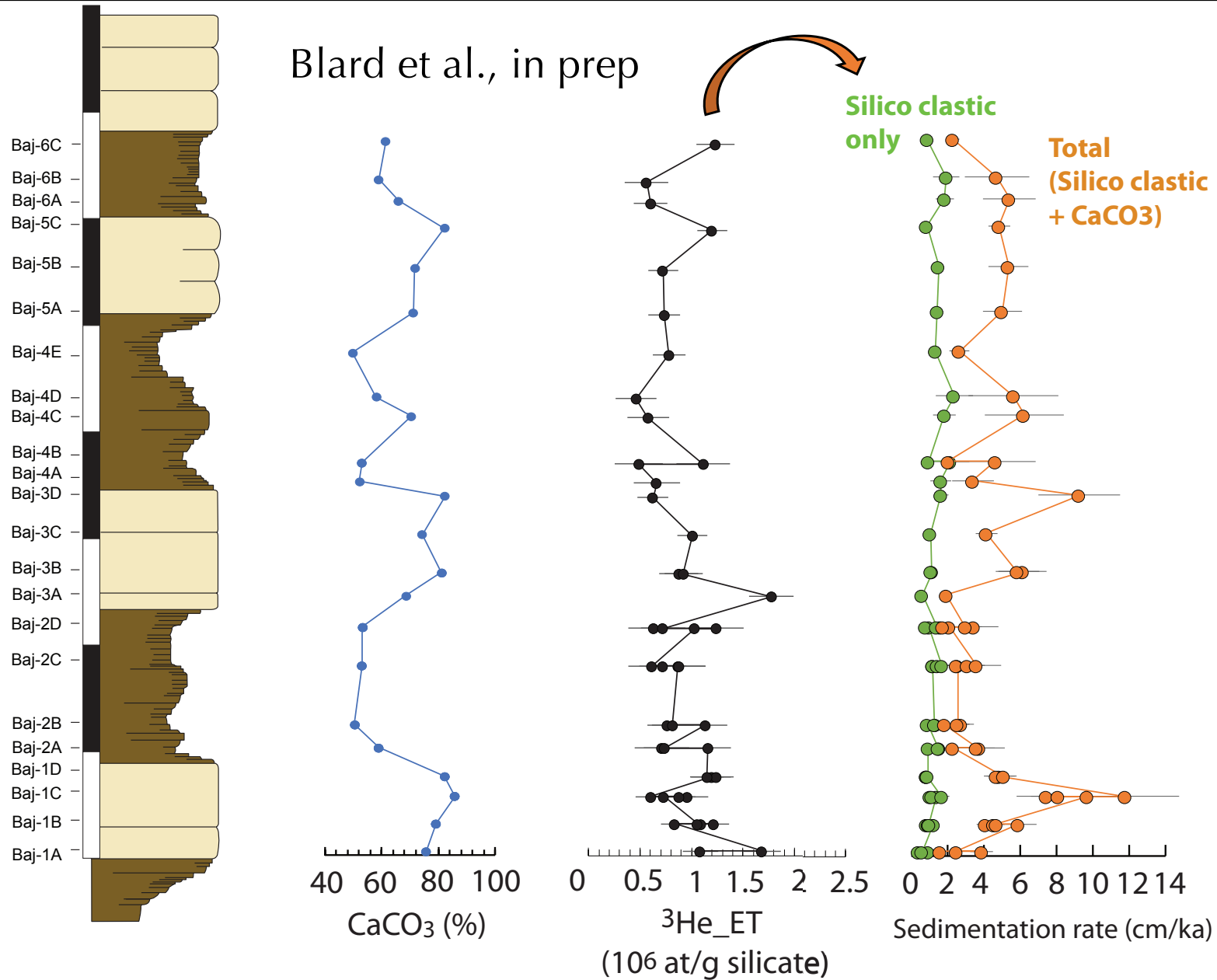
$$F = {}^3\text{He}_{\text{ET}} \text{ flux} = 0.1 \text{ pcc.cm}^{-2}.\text{ka}^{-1}$$



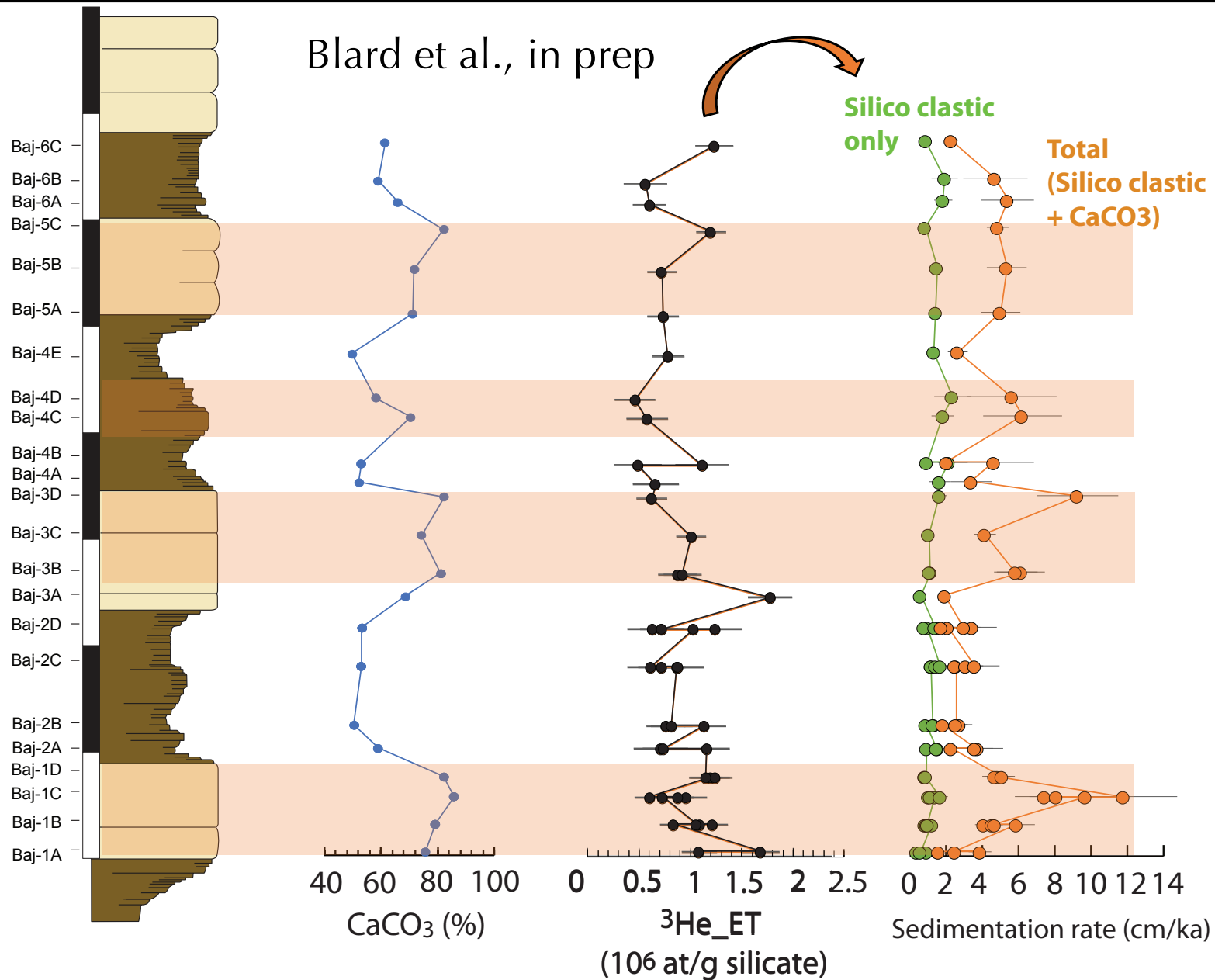
Farley et al., 2012

Blard et al., in prep

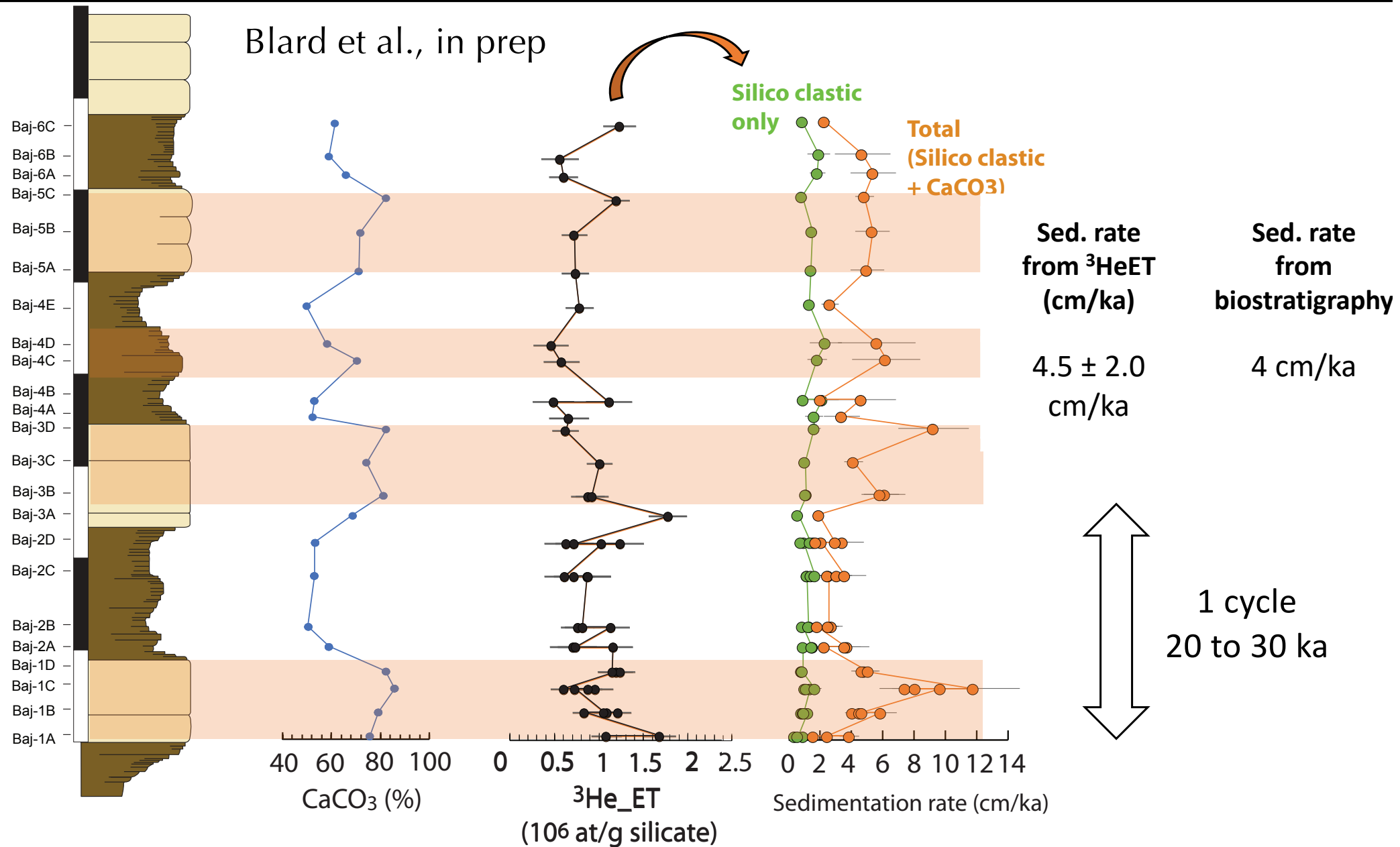
3 – Sedimentation rates results: Bajocian



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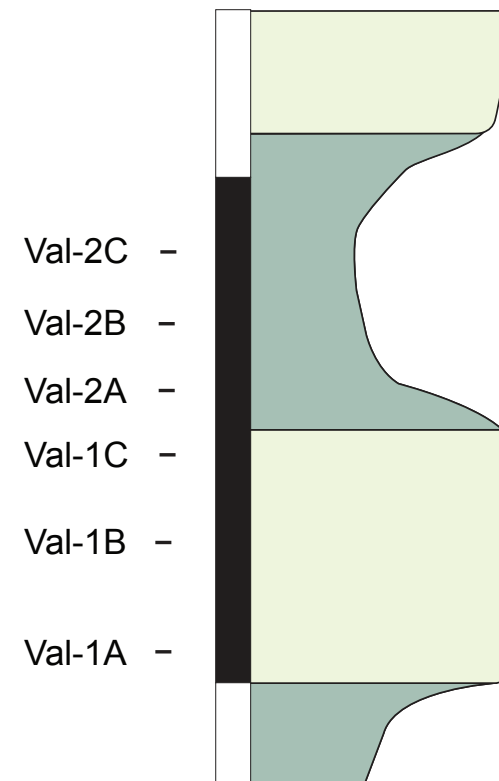


3 – Sedimentation rates results: Valanginian

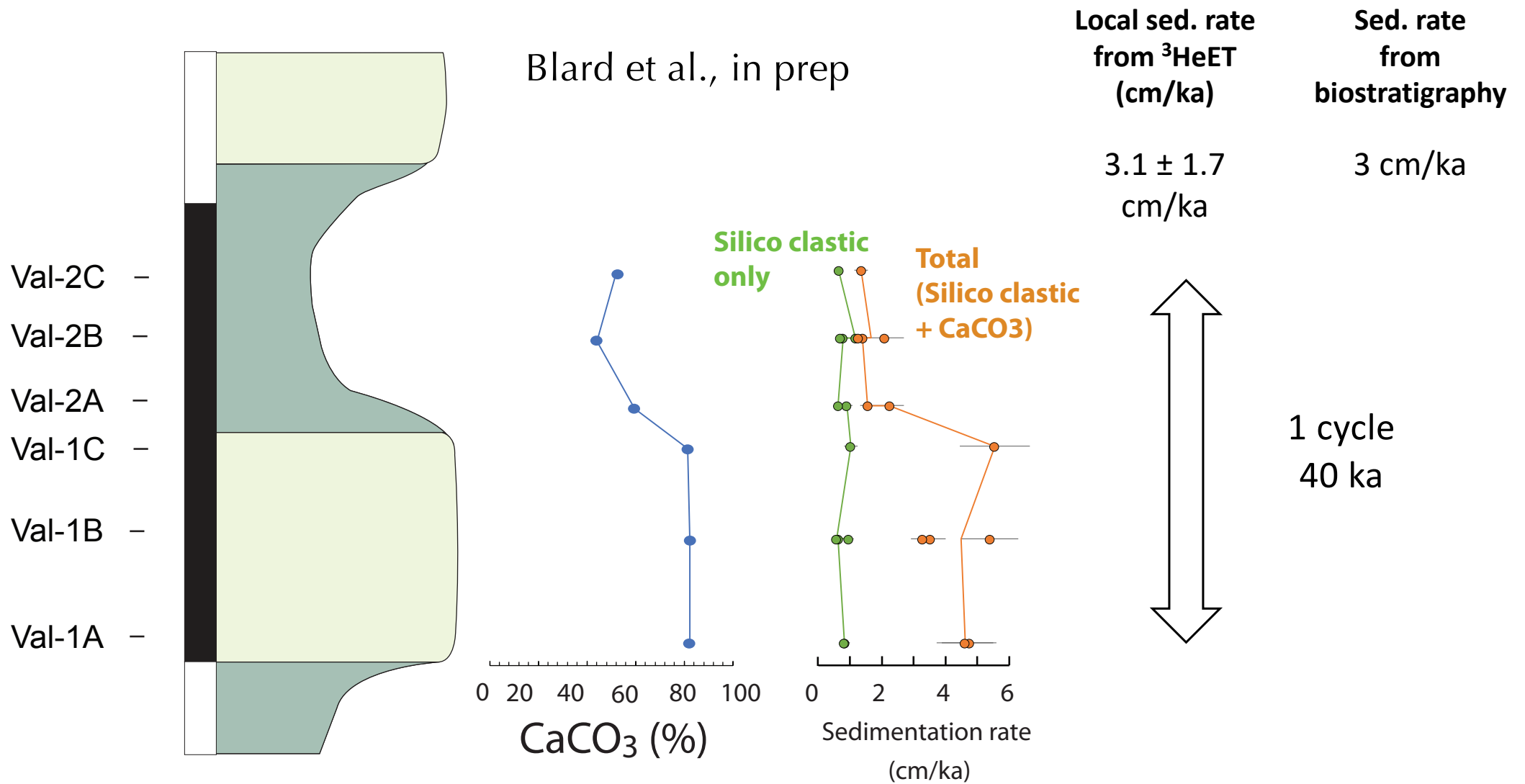
Valanginian (135 Ma) :

- **0.9 m section**
- **1 marl limestones doublets**
- **6 samples**

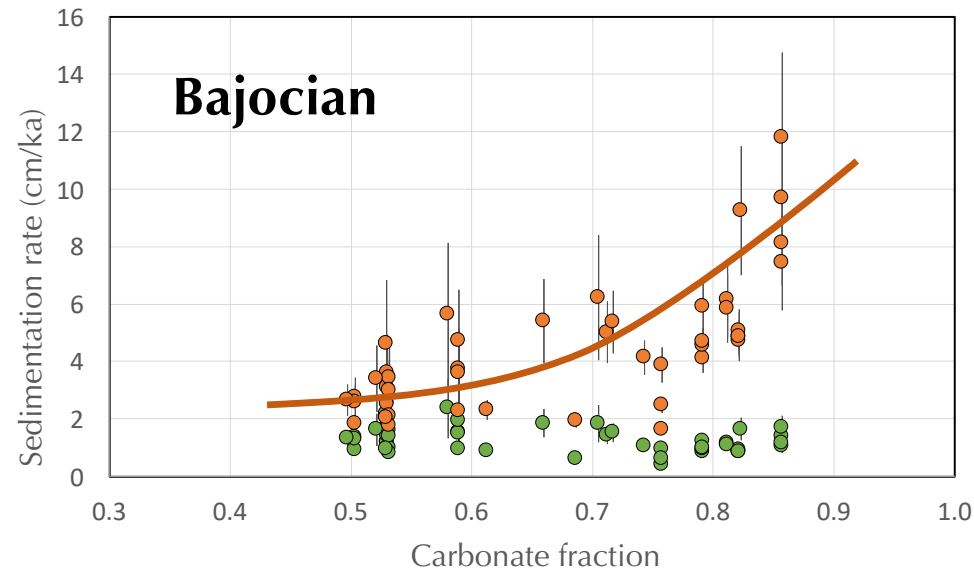
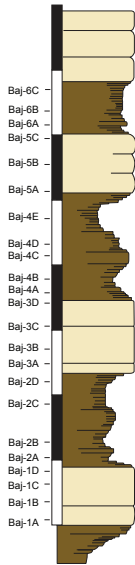
Blard et al., in prep



3 – Sedimentation rates results: Valanginian



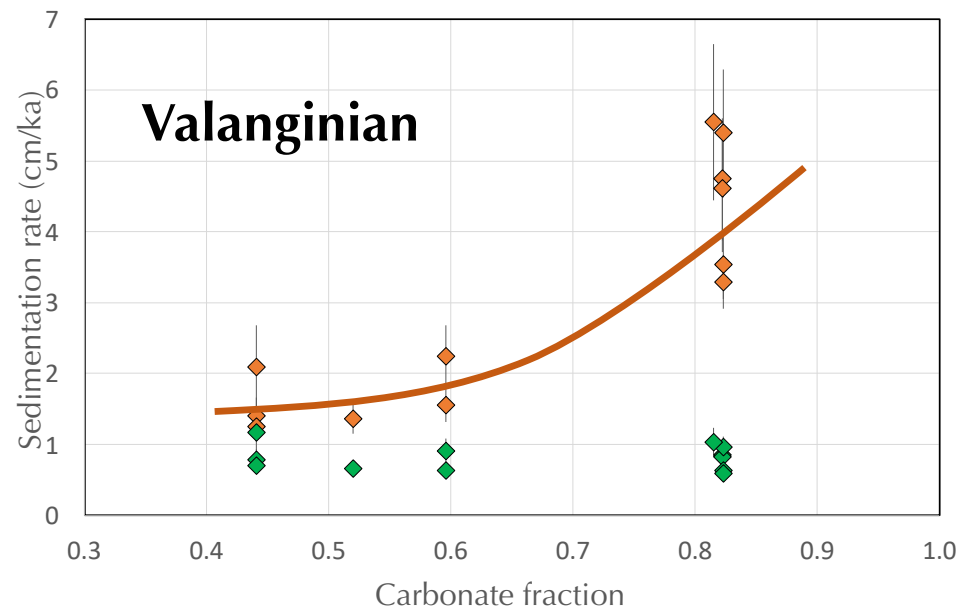
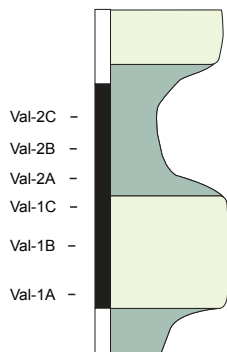
3 – Sedimentation rates summary



Blard et al., in prep

S_T = Total sed. rate

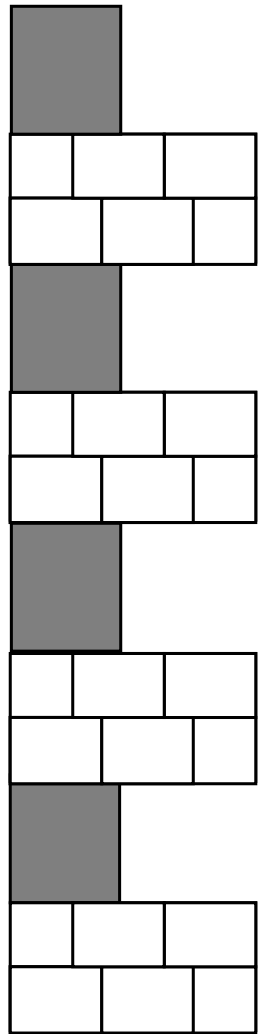
S_{sil} = Silico clastic sed. rate



Equation of a dilution by CaCO_3 (Ricken, 1993):

$$S_T = \frac{S_{sil}}{1 - NCF}$$

4 – Conclusion: ^3He ET shows constant terrigenous flux



Hypothesis 1

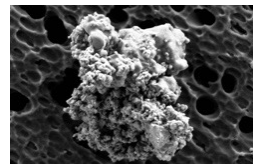
Silico-clastic sedimentation

Variable
 CaCO_3
sedimentation

Haute Provence

Blard et al.,
in prep.

3
He



Hypothesis 2

CaCO_3
sedimentation

?

Variable
Silico-clastic
sedimentation

Now hiring a new PhD at CRPG about
 ^3He ET in the Cretaceous marl-limestone alternations
Application deadline: June 15th 2022
Contact: blard@crpg.cnrs-nancy.fr

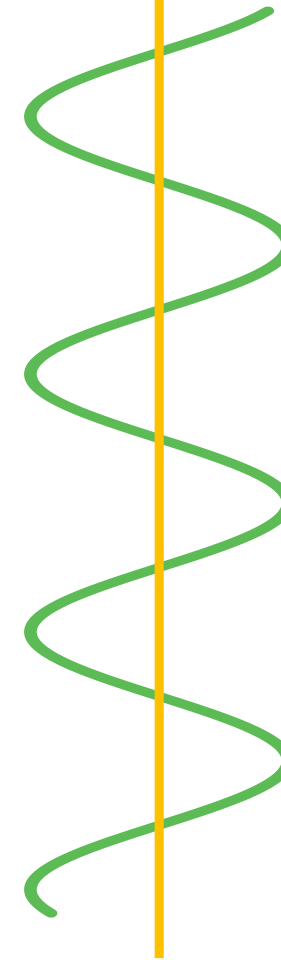


Cretaceous
Blakey, 2018

Hypothesis 2

CaCO_3
sedimentation

Variable
Silico-clastic
sedimentation



Locklair et Sageman, 2008

A scenic view of a beach with clear blue water, rocks, and a green hillside in the background. The water is a vibrant turquoise color, and the beach is sandy with some people and umbrellas. The hillside is covered in lush green vegetation.

Thank you!!

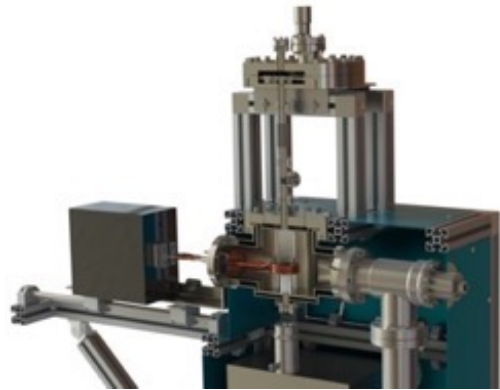
Merci !!

Grazie !!

Danke !!

¡¡ Gracias !!

2 – Method: sample preparation



CRPG induction furnace
Zimmermann et al., 2018

1 – Gentle crushing

2 - Weighting

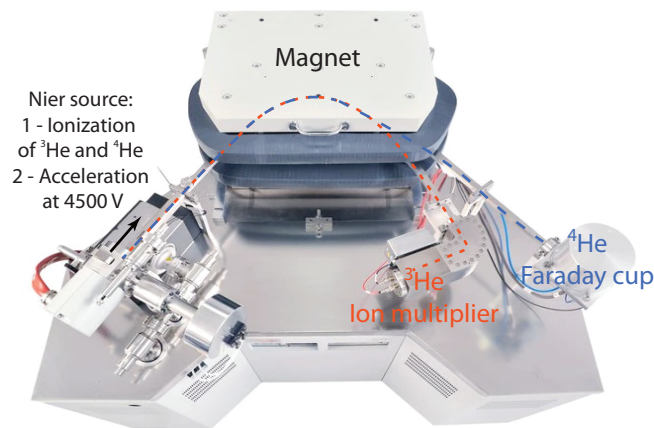
3 – 48 h in HCl 10%



4 – Drying (60°C) and weighting => CaCO₃ fraction

5 – Heating in high vacuum furnace 15 min 1400°C

6 – ³He ⁴He measurements with the SFT at CRPG



CRPG SFT

3 – Results: calculation of the IDP-Terrigenous mixing

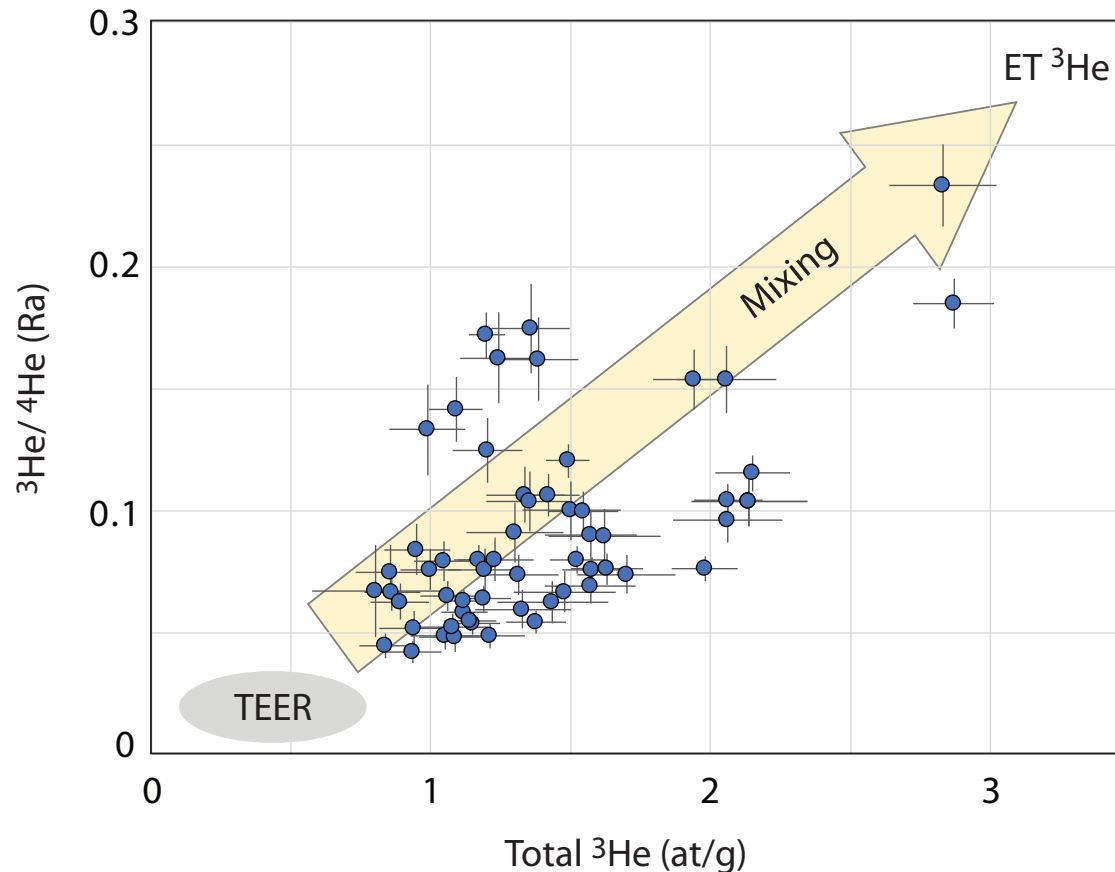
Calculation of $^3\text{He}_{ET}$ concentrations : 2 component mixing

$$[^3\text{He}_{ET}] = \left(\frac{1 - \frac{^3\text{He}/^4\text{He}_{TERR}}{^3\text{He}/^4\text{He}_{meas}}}{1 - \frac{^3\text{He}/^4\text{He}_{TERR}}{^3\text{He}/^4\text{He}_{IDP}}} \right) \cdot [^3\text{He}_{meas}]$$

Endmembers

$^3\text{He}/^4\text{He}_{IDP} = 300 \text{ Ra} \pm 50\%$

$^3\text{He}/^4\text{He}_{TERR} = 0.02 \text{ Ra} \pm 50\%$



**$\Rightarrow$ 60 to 90% of
 ^3He is ET in
these samples**

Blard et al., in prep