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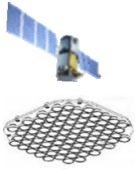
# **Laser GNSS Receiver for LEO POD, Laser Occultation and Time&Frequency Transfer of Optical Clocks in the Timing Labs**

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# Laser GNSS Receiver in LEO

GNSS  
SLR array



## Continuous-Wave (CW) Laser

- Stability  $< 5 \times 10^{-15}$  at 1 s
- Linewidth  $< 2$  Hz
- With cavity from NPL (LISA mission)
- Up to 1 kW power



No atmosphere in LEO  
One could improve LEO attitude  
(quadrant photodiode)

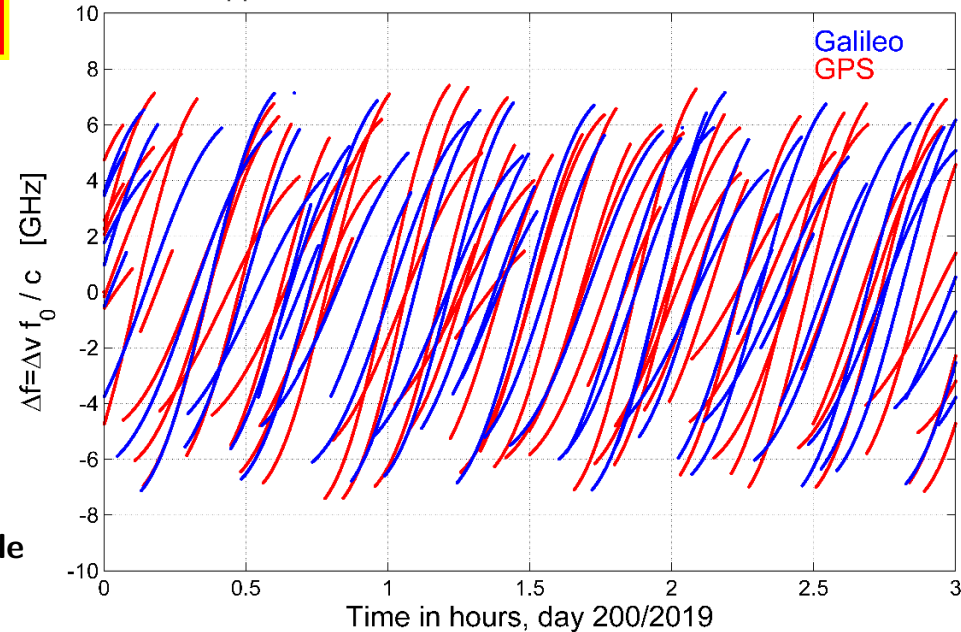
Carrier-phase  
on a continuous-  
wave (CW) laser

1064 nm

CW-laser

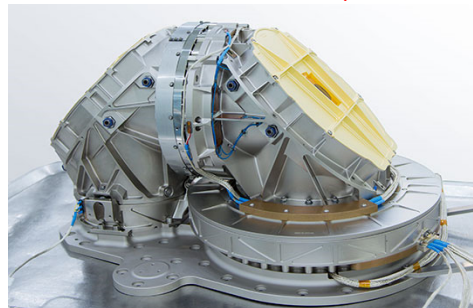


Doppler shift at 1064 nm, GRACE-FO - Galileo&GPS

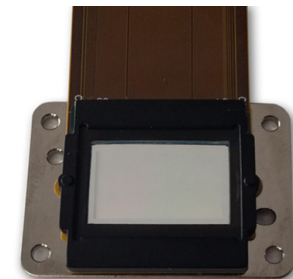


## Laser beam steering in LEO

“Sinopta” laser steering terminal for LEO  
steering over  $360^\circ/90^\circ$



Spatial Light Modulator  
steering over  $24^\circ$



# Link Budget for CW Laser: LEO, Ground, Moon

## Link Budget:

$$n_r = \frac{E_t \lambda}{hc} G_t \sigma_{ocs} \left( \frac{1}{4\pi R^2} \right)^2 A_r T_a^2 T_c^2 \eta_t \eta_r \eta_d$$

## Gaussian Beam:

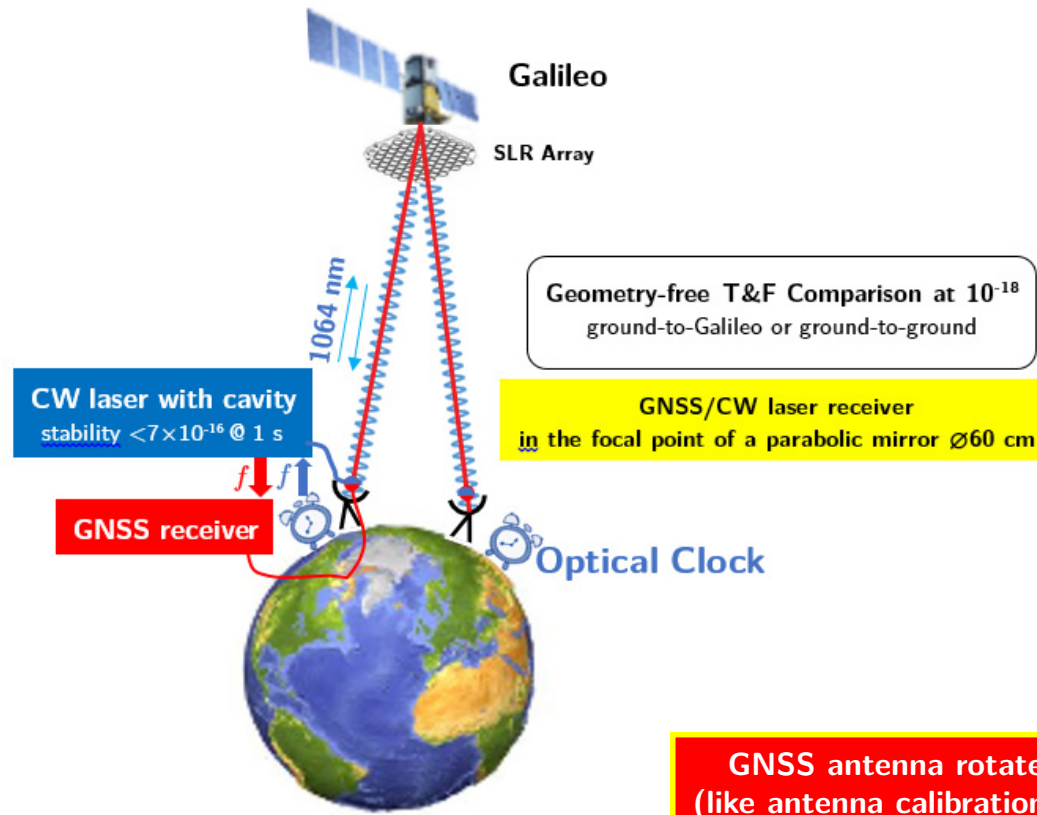
$$\theta_t = \frac{\lambda}{\pi \omega_0}$$

$$G_t(\theta) = \frac{8}{\theta_t^2} e^{-2\left(\frac{\theta}{\theta_t}\right)^2}$$

Link budget parameter	GRACE-FO – Galileo	Ground – Galileo	Ground – Moon
Transmitted power of a CW laser	200 W	1 kW	1 kW
$\sigma_{ocs}$ optical cross section for GNSS (Pearlman, 2008) and Moon (Arnold, NASA TN)	$45 - 80 \times 10^6 \text{ m}^2$	$45 - 80 \times 10^6 \text{ m}^2$	$1400 \times 10^6 \text{ m}^2$ 300 corner-cubes
Radius for the aperture area of the receiving optics $A_r$	0.15 m	0.30 m	0.50 m
Gaussian beam divergence half-angle, $\theta_t$	0.47°	0.23°	0.14°
Fine pointing error (steering in steps of 0.01°), $\theta$	0.20°	0.20°	0.10°
$G_t$ transmitter gain	$1.1 \times 10^{12}$	$1.4 \times 10^{12}$	$6.3 \times 10^{12}$
$T_a^2$ Two-way atmospheric transmission ( <a href="#">Degnan, 1993</a> ), ( <a href="#">Matthews 2020</a> )	1.0	0.5	0.5
$T_c^2$ Two-way cirrus transmission ( <a href="#">Degnan, 1993</a> )	1.0	0.5	0.8
$\eta_t$ efficiency of the transmitting optics ( <a href="#">Degnan, 1993</a> )	0.9	0.9	0.9
$\eta_r$ efficiency of the receiving optics ( <a href="#">Degnan, 1993</a> )	0.9	0.9	0.9
$\eta_d$ is the detector efficiency (photodiode sensitivity, Hamamatsu) see also ( <a href="#">Jennrich &amp; Heinzel, 2013</a> ),	$0.7 \frac{\text{A}}{\text{W}}$	$0.7 \frac{\text{A}}{\text{W}}$	$0.7 \frac{\text{A}}{\text{W}}$
Received laser power from a CW laser	16.5 pW	100.1 pW	507.2 fW (1.5 pW)

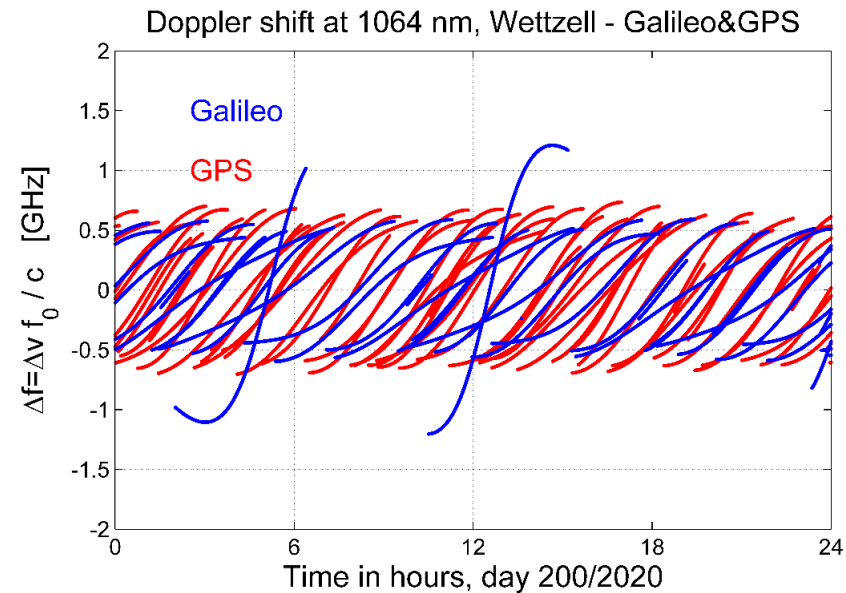
**Hamamatsu (Japan) confirmed  
InGaAs photodiode for all three cases**

# CW-Laser for T&F Transfer at $10^{-18}$ for Optical Clocks



Laser Cavity (NPL&PTB)  
can provide microwave GNSS  
frequency more stable than H-maser  
on GNSS (via frequency comb)

2D Antenna Rotation about Pivot point



With 'phase clock' approach for  
GPS we introduced in 2004 for  
LEOs, we managed to get T&F  
transfer down to  $10^{-16}$  level

# CW-Laser Onboard GNSS Satellites

**“two-way”**

**CW-laser in LEO, signal reflected from GNSS**

$$n_r = \left( E_t \frac{\lambda}{hc} \right) G_t \sigma_{ocs} \left( \frac{1}{4\pi R^2} \right)^2 A_r T_a^2 T_c^2 \eta_t \eta_r \eta_d$$

$$\sim \frac{1}{R^4}$$

**“one-way”**

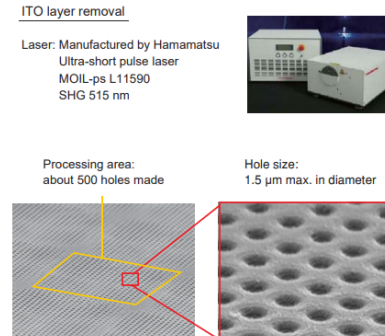
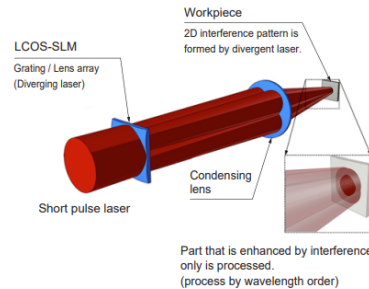
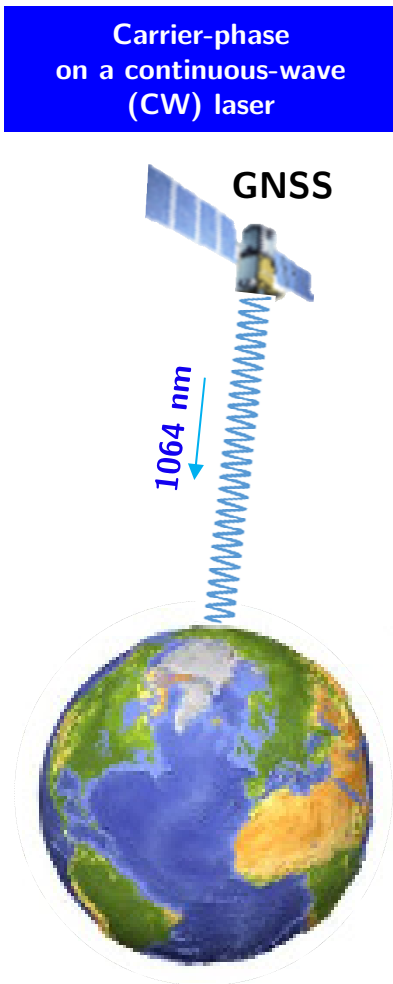
**CW-laser onboard GNSS satellites**

$$n_r = \left( E_t \frac{\lambda}{hc} \right) G_t \left( \frac{1}{4\pi R} \right)^2 A_r T_a^2 T_c^2 \eta_t \eta_r \eta_d$$

$$\sim \frac{1}{R^2}$$

Link budget	GRACE-FO – Galileo	Ground – Galileo	Ground – Moon
Two-way: Transmitted power of a CW laser	200 W	1 kW	1 kW
One-way: Transmitted power of a CW laser ( <b>1 nW received</b> )	20 mW	15 mW	20 mW

**We need 1-20 mW laser power transmitted from a GNSS to guarantee the received power of 1 nW (GRACE-FO) or 200 pW (LISA) with down to 10 cm collecting optics**



**With SLM, one can split laser beam into thousands of separate laser beams and steer them independently**