

Changes in pattern formation and behavior in rotating Rayleigh-Bénard convection due to inhomogeneous thermal insulation

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Introduction & Motivation

- Rayleigh-Bénard convection in molten planetary cores, affected by thermal conditions at the mantle (Mound & Davies, 2017)
- Explore the importance of thermal boundaries on fluid heat transport and temporal and spatial behaviour
- Can spatially inhomogeneous insulation cause time-independent behaviours and pin convection rolls in rapidly rotating system?

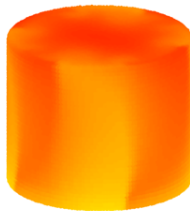
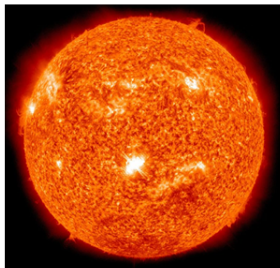


Figure: Image of the sun from National Geographic, a numerical simulation and a convection experiment at UCLA SpinLab.

Governing Equations

Boussinesq approximation applied to Navier Stokes equations:

$$\frac{D\mathbf{u}}{Dt} + 2\boldsymbol{\Omega} \times \mathbf{u} = -\frac{1}{\rho_0} \nabla p + \alpha T \mathbf{g} + \nu \nabla^2 \mathbf{u}, \quad (1)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (2)$$

$$\frac{DT}{Dt} = \kappa \nabla^2 T. \quad (3)$$

Obtain non-dimensional equations by apply scalings such that,

$$t = \hat{t} \frac{d^2}{\kappa} \quad \mathbf{x} = \hat{\mathbf{x}} d \quad \mathbf{u} = \hat{\mathbf{u}} \frac{\kappa}{d} \quad p = \hat{p} \frac{\kappa^2}{d^2} \quad T = \hat{T} \Delta T, \quad (4)$$

where d is the height of the system.

Governing Equations

Eqs (1)-(3) now become non-dimensional:

$$\frac{D\mathbf{u}}{Dt} = -\nabla p + Pr\nabla^2\mathbf{u} - RaPrT\tilde{\mathbf{z}} - \frac{Pr}{Ek}\tilde{\mathbf{z}} \times \mathbf{u}, \quad (5)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (6)$$

$$\frac{DT}{Dt} = \nabla^2 T. \quad (7)$$

With non-dimensional parameters Prandtl number, Rayleigh number, and Ekman number:

$$Pr = \frac{\nu}{\kappa} \quad Ra = \frac{\alpha\Delta T d^3 g}{\nu\kappa} \quad Ek = \frac{\nu}{2\Omega d^2} \quad (8)$$

Boundary Conditions

Top and Bottom ($z = z_1, z_0$):

- No-slip velocity: $\mathbf{u} = 0$
- Fixed temperature: $T(z_0, z_1) = 1, 0$

Sidewalls ($r = R$):

- No-slip velocity: $\mathbf{u} = 0$
- One of the following forms inhomogeneous heat flux:

① Positive/Negative:

$$\nabla T \cdot \mathbf{n} = \frac{\partial T}{\partial r} = \sum_{n=1}^{\infty} A_n \sin(f_n z) \sin(m_\theta \theta) \quad (9)$$

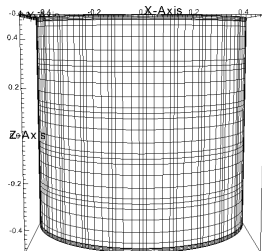
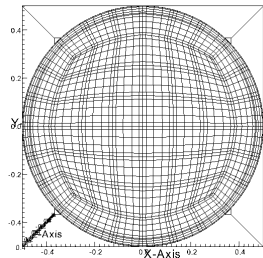
② Positive/Zero:

$$\frac{\partial T}{\partial r} = A_n \sin(\pi(z + 0.5))(1 + \sin(m_\theta \theta)) \quad (10)$$

A_n , f_n , and m_θ are constants

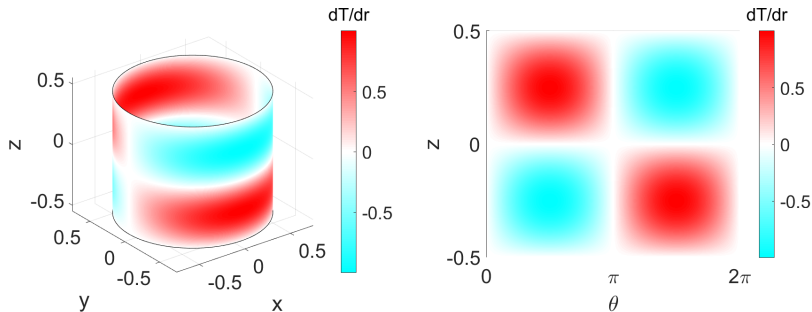
Numerical Simulations

- Nek5000: Spectral Element Method solver
- Cylindrical domain with aspect ratio $\Gamma = r/d = 0.7$
- Parameter range:
 $Pr = 0.7$
 $Ra \leq 10^8$
 $Ek \geq 10^{-5}$
- Initial condition: $\mathbf{u} = 0$ and $T = -z + 0.5$ with perturbation



Positive/Negative BC on sidewall

$$\nabla T \cdot \mathbf{n} = \frac{\partial T}{\partial r} = \sum_{n=1}^{\infty} A_n \sin(f_n z) \sin(m_\theta \theta)$$



Heat flux at $r = R$ when $A_n = 1$, $f_n = 1$, and $m_\theta = 1$

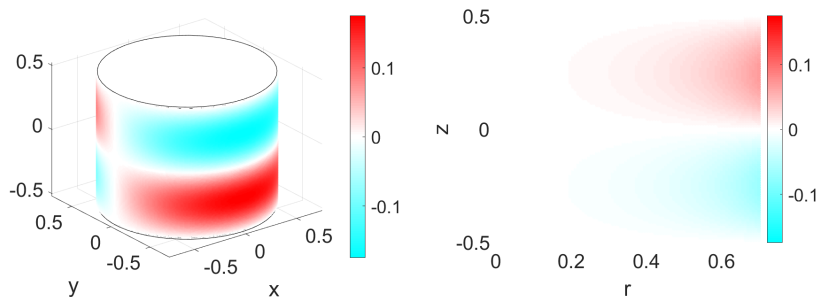
$f_n = 1$ for all cases to simplify vertical deviation

Conducting Solution with P/N BC on sidewall

Solving Eqs (5-7) with Eq (9) for a state where $\mathbf{u} = 0$ and $dT/dt = 0$,

$$T(z, \theta, r) = \bar{T}(z) + T'(z, \theta, r)$$

$$\bar{T} = \frac{z}{z_0 - z_1} - \frac{z_1}{z_0 - z_1}, \quad T' = \sum_{n=1}^{\infty} \frac{A_n}{\frac{dI_m(Rf_n)}{dr}} \sin(f_n z) \sin(m_\theta \theta) I_m(f_n r)$$



T' when $A_n = 1$, $f_n = 1$, and $m_\theta = 1$
Note I_m is a modified Bessel's function

Numerical Validation of P/N BC on sidewall

Use conducting solution (T_{cs}) and numerical simulation (T_{ns}) where $Ra < Ra_c$, so the system is conducting to validate numerical set up. Quantified with $T^\infty = |\max(T_{cs} - T_{ns})|$

Ek	Ra	m_θ	A_n	T^∞
∞	1	1	0.1	4×10^{-5}
			0.5	2×10^{-4}
		2	0.1	4×10^{-5}
10^{-3}	1	1	0.1	4×10^{-5}
10^{-4}	1	1	0.5	7×10^{-4}
			1	2×10^{-3}
		3	2	2×10^{-3}
10^{-5}	1	1	0.5	2×10^{-4}

T^∞ increases as A_n increases due to the dl_m/dr term in the conducting solution, which is calculated numerically.

So as dr decreases, T^∞ also decreases.

All results shown are calculated with $dr = 10^{-3}$.

Results of Positive/Negative BC on Sidewalls

$$\nabla T \cdot \mathbf{n} = \frac{\partial T}{\partial r} = \sum_{n=1}^{\infty} A_n \sin(f_n z) \sin(m_\theta \theta)$$

Azimuthal mode, m_θ is chosen in relation to the dominant mode, m_{UI} identified in the uniformly insulated (UI) system of the same Ek and Ra . m_θ generally is either:

- $= 1$, to apply a larger wavelength at the sidewall than in the UI system, more similar to planetary cores (Mound & Davies, 2017)
or
- $= m_{UI} > 1$, to investigate ability of lateral thermal boundary conditions to pin convection rolls (i.e. stop them from rotating about the domain)

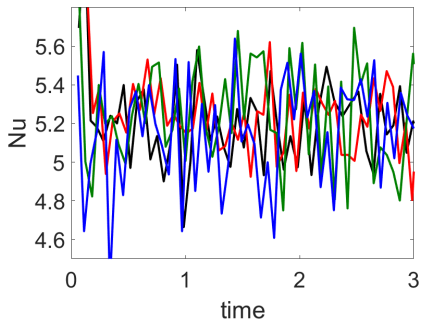
Amplitude, A_n , is increased from 0.5 until effects are observed

Heat transport is measured by Nusselt number, $Nu = \frac{\Delta T}{(dT/dr)|_{z=z_1} d}$.

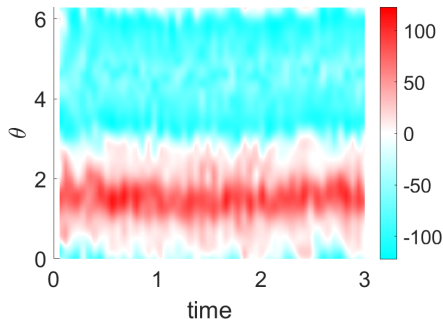
P/N BC on sidewall: Non-Rotating Example

$$Ek = \infty, Ra = 3 \times 10^5$$

UI system is chaotic with $m_{UI} = 1$. Apply $m_\theta = 1$ with $A_n = [0.5, 1, 2]$.



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



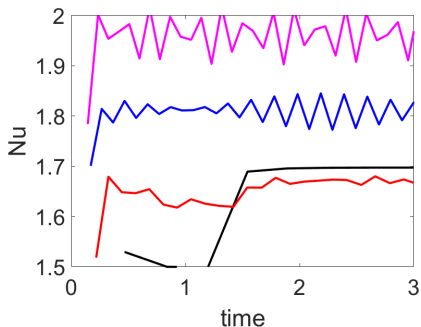
Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.65$

Temporal behaviour is chaotic and dominant mode is $m = 1$ for all A_n
No significant change in behaviour due to heterogeneous sidewall insulation

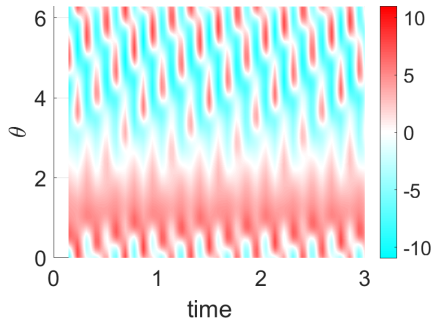
P/N BC on sidewall: Weakly Rotating Example

$$Ek = 10^{-3}, Ra = 4.8 \times 10^4$$

UI is time-independent with $m_{UI} = 3$. Apply $m_\theta = 1$ with $A_n = [0.5, 2, 3]$



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 3$



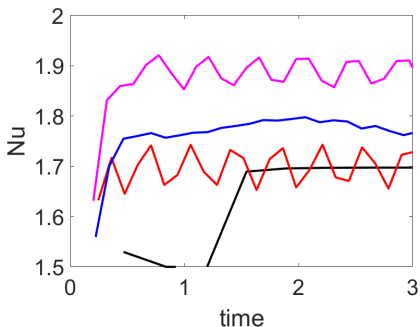
Hövmoller diagram of vertical velocity, w , for $A_n = 3$ taken at $z = 0.3$ and $r = 0.68$

$m_\theta = 1$ cause time-dependant behaviour and super-positions with $m_{UI} = 3$ mode, showing effects most strongly where applied $dT/dr > 0$.

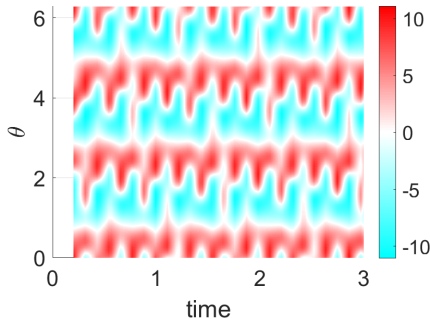
P/N BC on sidewall: Weakly Rotating Example

$$Ek = 10^{-3}, Ra = 4.8 \times 10^4$$

UI is time-independent with $m_{UI} = 3$. Apply $m_\theta = 3$ with $A_n = [0.5, 2, 3]$



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 3$



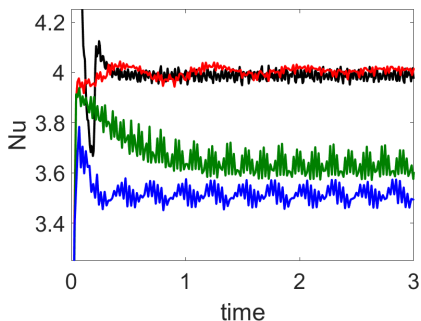
Hövmoller diagram of vertical velocity, w , for $A_n = 3$ taken at $z = 0.3$ and $r = 0.68$

$m_\theta = m_{UI} = 3$ causes oscillatory temporal behaviour. Convection rolls are pinned, but oscillate within range $\theta = 2\pi/m_\theta$.

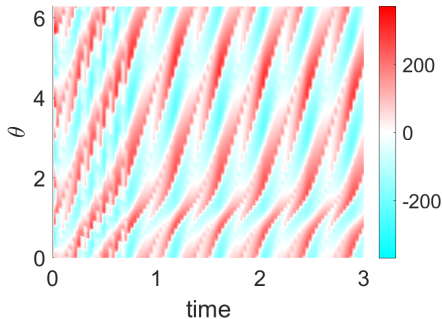
P/N BC on sidewall: Rapidly Rotating Example

$$Ek = 10^{-5}, Ra = 1.44 \times 10^7$$

UI is chaotic with small σ_{Nu} and $m_{UI} = 3$. Apply $m_\theta = 1$ & $A_n = [0.5, 1, 2]$



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



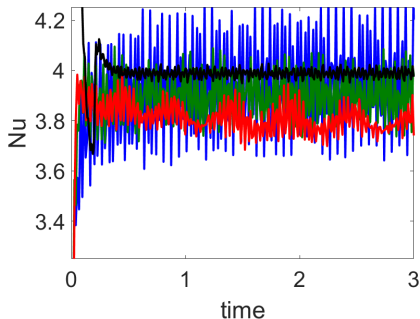
Hövmoller diagram of vertical velocity, w , for $A_n = 1$ taken at $z = 0.3$ and $r = 0.68$

As A_n increases, \overline{Nu} decreases, and quasi-oscillatory temporal behaviour begins. Fluid rotates slower than UI, and slows further where $dT/dr > 0$

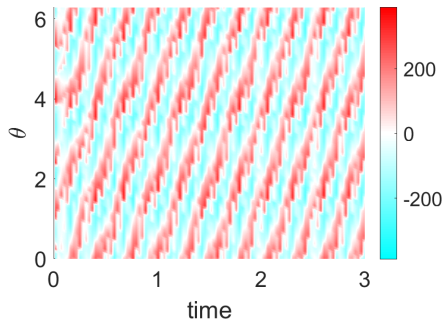
P/N BC on sidewall: Rapidly Rotating Example

$$Ek = 10^{-5}, Ra = 1.44 \times 10^7$$

UI is chaotic with small σ_{Nu} and $m_{UI} = 3$. Apply $m_\theta = 3$ & $A_n = [0.5, 1, 2]$



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.68$

As A_n increases, σ_{Nu} increases. For all A_n , fluid rotates slower than UI, upward and downward motion oscillate in strength 3 times in θ

P/N BC on sidewall: Discussion

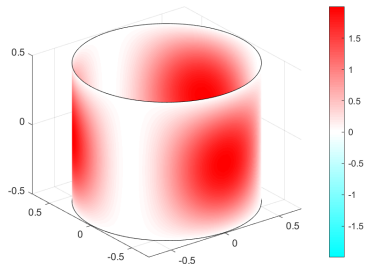
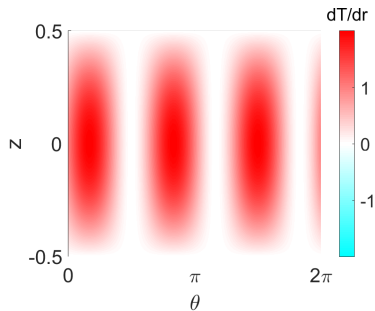
- Vary heat flux in both z and θ to investigate planetary-core-like system:

$$\nabla T \cdot \mathbf{n} = \frac{\partial T}{\partial r} = \sum_{n=1}^{\infty} A_n \sin(f_n z) \sin(m_\theta \theta)$$

- Choose either $m_\theta = 1$, to approximate planetary cores
 - Fluid rotates slower where $dT/dr > 0$or $m_\theta = m_{UI}$, the dominant mode of the uniformly insulated case (UI) to pin convection rolls
 - pinned rolls within $\theta = 2\pi/m_\theta$ only when weakly rotating
- Generally, non-rotating systems are not affected by the inhomogeneous sidewall insulation.
- In rotating systems, the sidewall heat flux variation is observed to cause the system to become more time-dependant and make the fluid rotate about the domain, regardless of the temporal and spatial behaviour of the uniformly insulated system

2. Positive/Zero BC on sidewall

$$\frac{\partial T}{\partial r} = A_n \sin(\pi(z + 0.5))(1 + \sin(m_\theta \theta))$$



Heat flux at $r = R$ when $A_n = 1$, and $m_\theta = 3$

P/Z BC on sidewall: Heat Transport

Previous Nu definition not applicable. Redefine from Kunnen et al, 2008, $Nu = 1 + \langle wT \rangle_V$, a volume average of the domain.

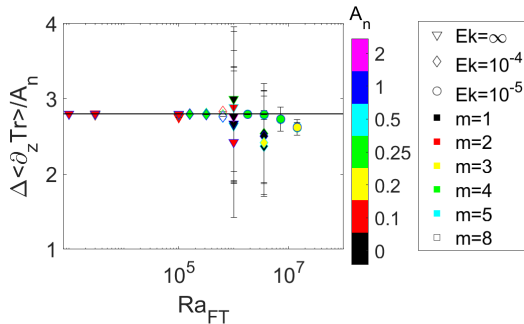
Following Hepworth, 2014, heat flux accounting is employed to validate numerical set-up. By integrating the heat equation, Eq(7), over the volume of the domain and applying the fixed temperature longitudinal and heterogeneous lateral (Eq (12)), we find that,

$$\int_0^R \int_0^{2\pi} \frac{\partial T}{\partial z} r d\theta dr|_{z=-Z} = \int_0^R \int_0^{2\pi} \frac{\partial T}{\partial z} r d\theta dr|_{z=Z} + 4RA_n \quad (12)$$

Thus the difference in heat flux from the bottom to top of the system should be $4RA_n = 2.8A_n$.

P/Z BC on sidewall: Numerical Validation

We simulated a range of $Ek - Ra - m_\theta - A_n$ systems and measured the difference in heat flux between the top and bottom, $\Delta \langle \partial_z Tr \rangle$, and divided by the A_n of each respective case.

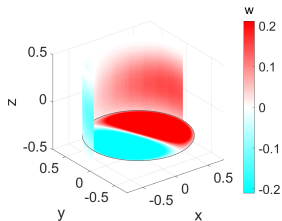


Increase in Ra and decrease Ek case variation away from 2.8, likely due to increased time-dependence causing less certainty in heat flux measurements.

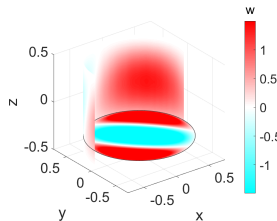
Ek , A_n , and m value are represented by the marker shape, edge colour, and fill colour, respectively. the solid line represents the expected value, 2.8. Error bars show the standard deviation.

P/Z BC on sidewall: Non-rotating Example

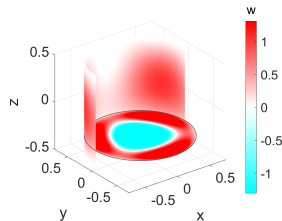
$Ek = \infty$, $Ra = 3 \times 10^3$ UI is time-dependant with $m_{UI} = 1$. However, Positive/Negative BC results found no effect due to $m_\theta = m_{UI} = 1$. We apply $m_\theta = [2, 3]$ with $A_n = 1$.



UI



$m_\theta = 2$



$m_\theta = 3$

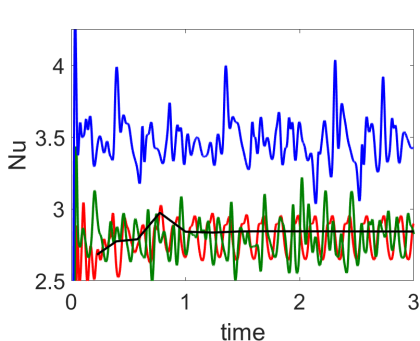
Vertical velocity, w , at $r = 0.69$, and $z = -0.3$

The dominant mode of the system becomes m_θ . The fluid motion in the bulk of the domain is affected, which was not observed in the P/N case.

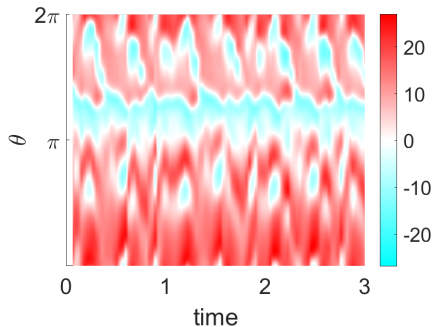
P/Z BC on sidewall: Weakly Rotating Example

$$Ek = 10^{-3}, Ra = 9.6 \times 10^4$$

UI is time-independent with $m_{UI} = 2$. We apply $m_\theta = 1$ & $A_n = [0.5, 1, 2]$



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



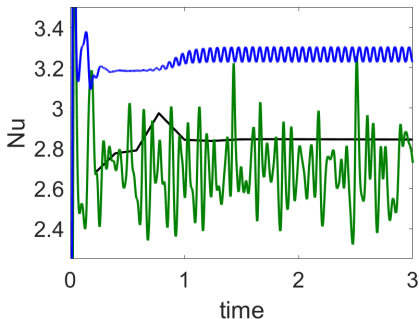
Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.68$

All systems become chaotic, but \overline{Nu} only increases when $A_n = 2$. One convection roll dominant, upward motion favoured where $dT/dr \neq 0$.

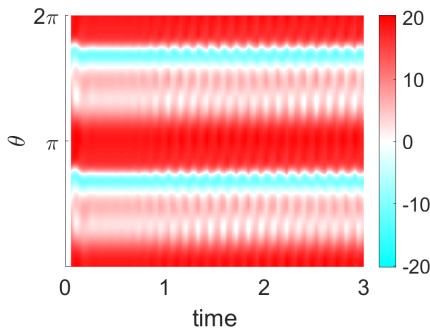
P/Z BC on sidewall: Weakly Rotating Example

$$Ek = 10^{-3}, Ra = 9.6 \times 10^4$$

UI is time-independent with $m_{UI} = 2$. We apply $m_\theta = 2$ & $A_n = [1, 2]$



UI; $A_n = 1$; $A_n = 2$



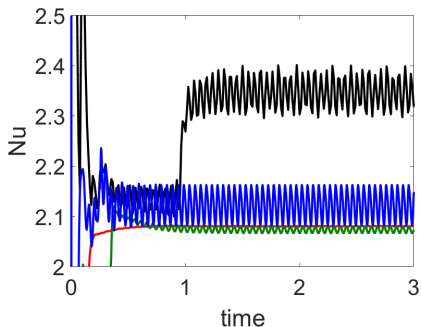
Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.68$

Transition from chaos to oscillation as A_n increases, \overline{Nu} also increases.
Upward motion strongly preferred where $dT/dr \neq 0$. Rolls pinned.

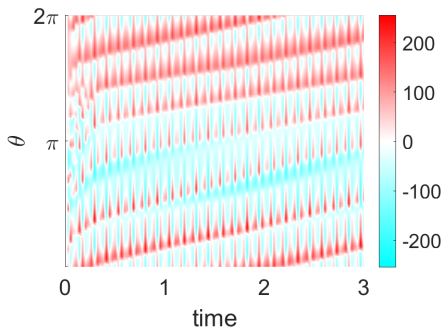
P/Z BC on sidewall: Rapidly Rotating Example

$$Ek = 10^{-5}, Ra = 7.2 \times 10^6$$

UI is oscillatory with $m_{UI} = 4$. We apply $m_\theta = 1$ & $A_n = [0.5, 1, 2]$.



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



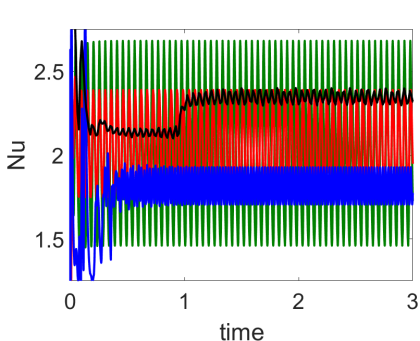
Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.68$

\overline{Nu} decreased with sidewall condition, amplitude of oscillation increases as A_n increases, strongest deformation of m_{UI} caused where $dT/dr \neq 0$

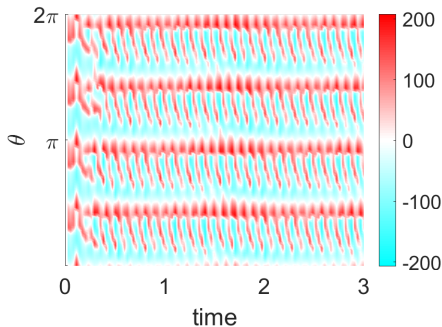
P/Z BC on sidewall: Rapidly Rotating Example

$$Ek = 10^{-5}, Ra = 7.2 \times 10^6$$

UI is oscillatory with $m_{UI} = 4$. We apply $m_\theta = 4$ & $A_n = [0.5, 1, 2]$.



UI; $A_n = 0.5$; $A_n = 1$; $A_n = 2$



Hövmoller diagram of vertical velocity, w , for $A_n = 2$ taken at $z = 0.3$ and $r = 0.68$

\overline{Nu} decreased as A_n increases. Convection rolls pinned where dT/dr strongest, with oscillations between $\theta = 2\pi/m_\theta$.

P/Z BC on sidewall: Discussion

- Define a thermal boundary condition which is $dT/dr > 0$ with a sinusoidal variation in z and θ :

$$\frac{\partial T}{\partial r} = A_n \sin(\pi(z + 0.5))(1 + \sin(m_\theta \theta))$$

- A Nu definition based on Kunnen, et al 2008 is defined
- Heat flux accounting following Hepworth, 2014 is followed to show that the difference in heat flux between the bottom and top of the system should be $2.8A_n$. Used to validate the numerical setup.
- Non-rotating: Convection rolls arrange into m_θ in the bulk of the fluid such that the centre of the system tends toward upward velocity
- Weakly rotating: $m_\theta = 1$ causes temporal chaos, $m_\theta = m_{UI}$ the mode is pinned, for sufficient A_n .
- Rapidly rotating: $m_\theta = 1$ shows strongest deviation from m_{UI} where $\max(dT/dr) = 1$; $m_\theta = m_{UI}$ convection pinned where dT/dr is strongest with oscillation between; increasing A_n decreases \overline{Nu} .

Conclusions and Future Work

● Positive/Negative Thermal BC

- Non-rotating systems not affected; Rotating systems become more time-dependant
- $m_\theta = 1$ causes slow of fluid rotation where $dT/dr > 0$
- $m_\theta = m_{UI}$ pins rolls but fluid oscillates within $\theta = 2\pi/m_\theta$

● Positive/Zero Thermal BC:

- Non-rotating systems form m_θ number of rolls in the bulk
- Rapidly rotating systems decrease in \overline{Nu} as A_n increases
- Rotating, for large A_n , pin rolls which oscillate between $\theta = 2\pi/m_\theta$
- Increasing the ratio between A_n and $\max(\mathbf{u})$ increases the effect on the system ($A_n \approx 2$ in Earth's core (Mound & Davies, 2017))
- Positive heat flux changes fluid behaviour more than negative heat flux variations - applicable to studies of planetary core dynamics, which have varying, positive, heat flux on boundaries
- In rotating systems, the application of $m_\theta = m_{UI}$ pins fluid everywhere- interesting potential experiment
- Future work: conduct experiments to verify results

References and Acknowledgements

- ① NEK5000 Version 19.0. (2019). Argonne National Laboratory, Illinois.
- ② Mound, J., and Davies, C. (2017). J. Fluid Mech.
- ③ Hepworth, B. (2014). PhD thesis. University of Leeds.

This work was undertaken on ARC3, part of the High Performance Computing facilities at the University of Leeds, UK.