

Grain-resolving simulations of submerged cohesive granular collapse

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1 Introduction

2 Computational model

3 Results

4 Conclusions



Granular collapse means a **granular column** is **suddenly** put into motion.

It is a kind of **unsteady granular flow**.



Pyroclastic flows



Landslides

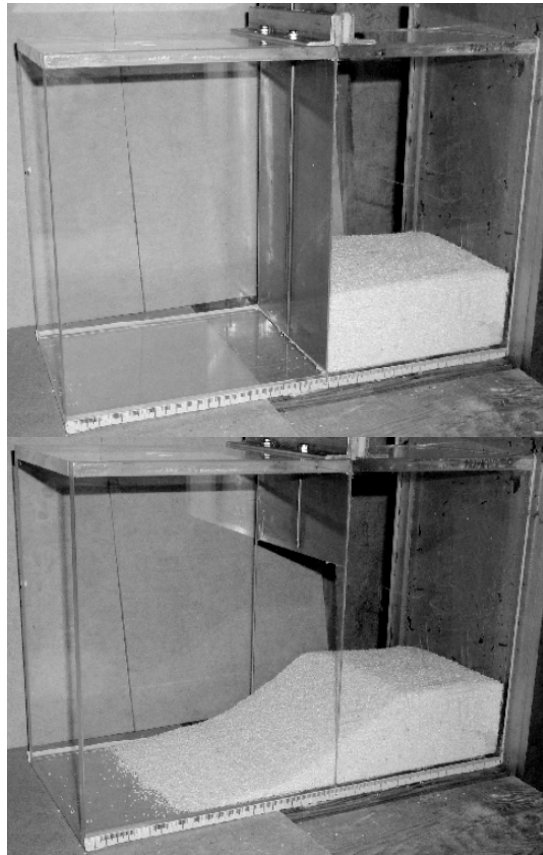
(Delannay et al. 2017)



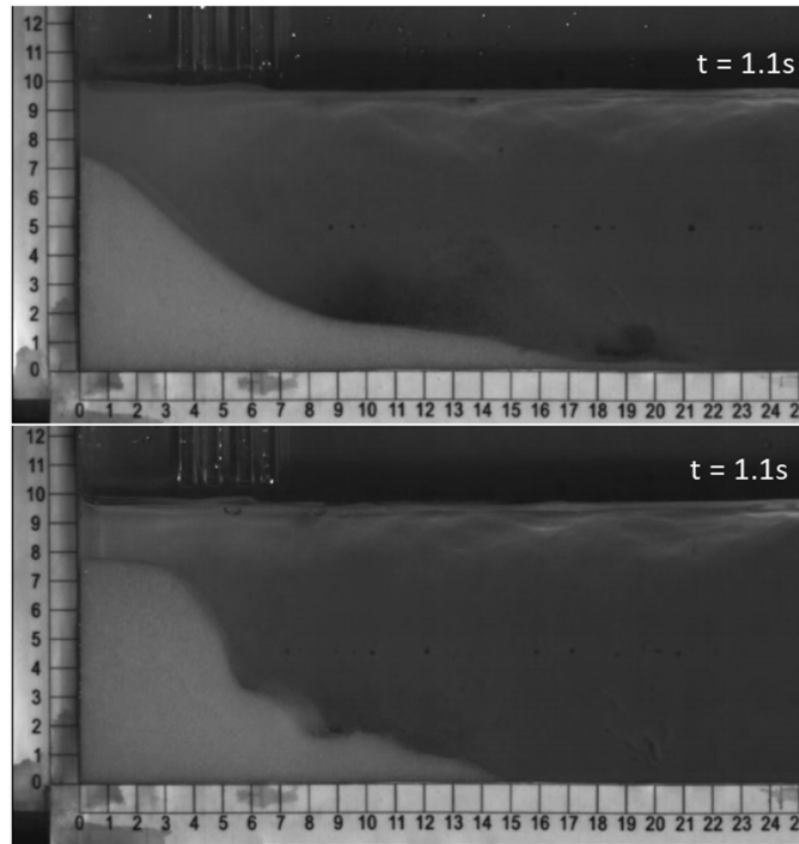
Submarine avalanches

Introduction——previous research

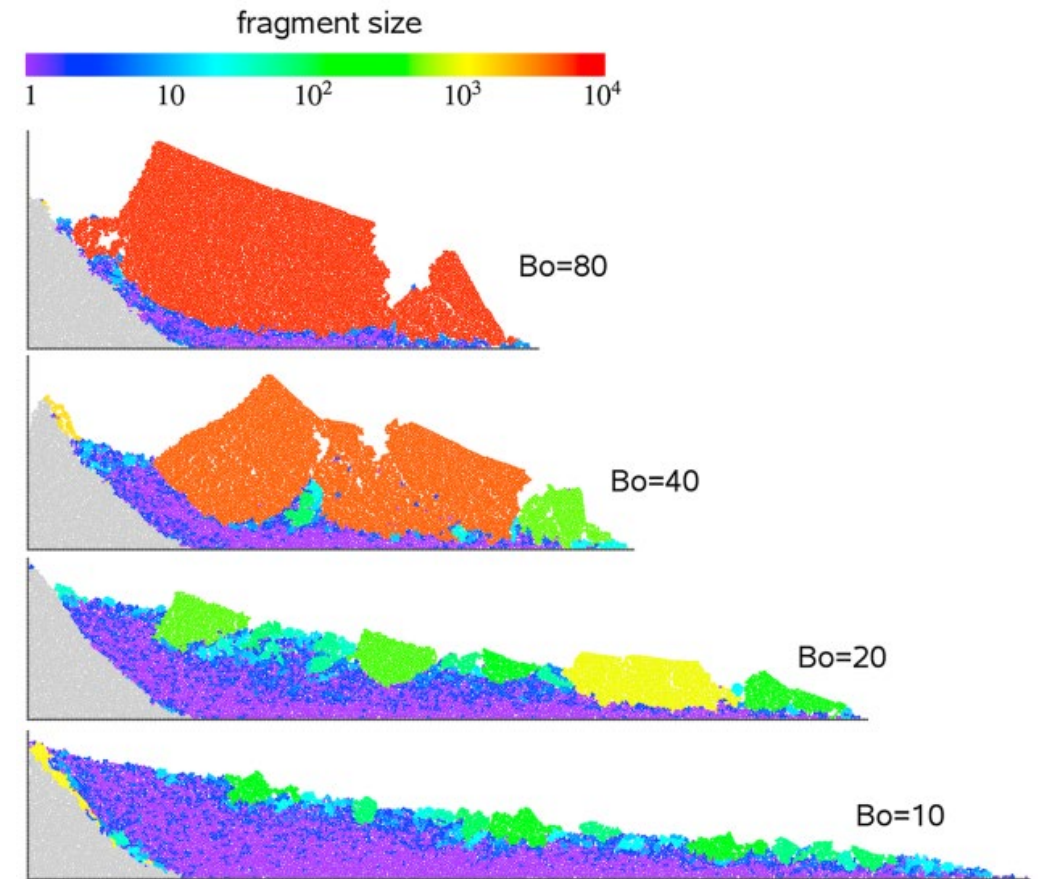
- Dry noncohesive granular collapse
- Submerged noncohesive granular collapse
- Dry cohesive granular collapse
- Submerged cohesive granular collapse ?



Balmforth & Kerswell 2005

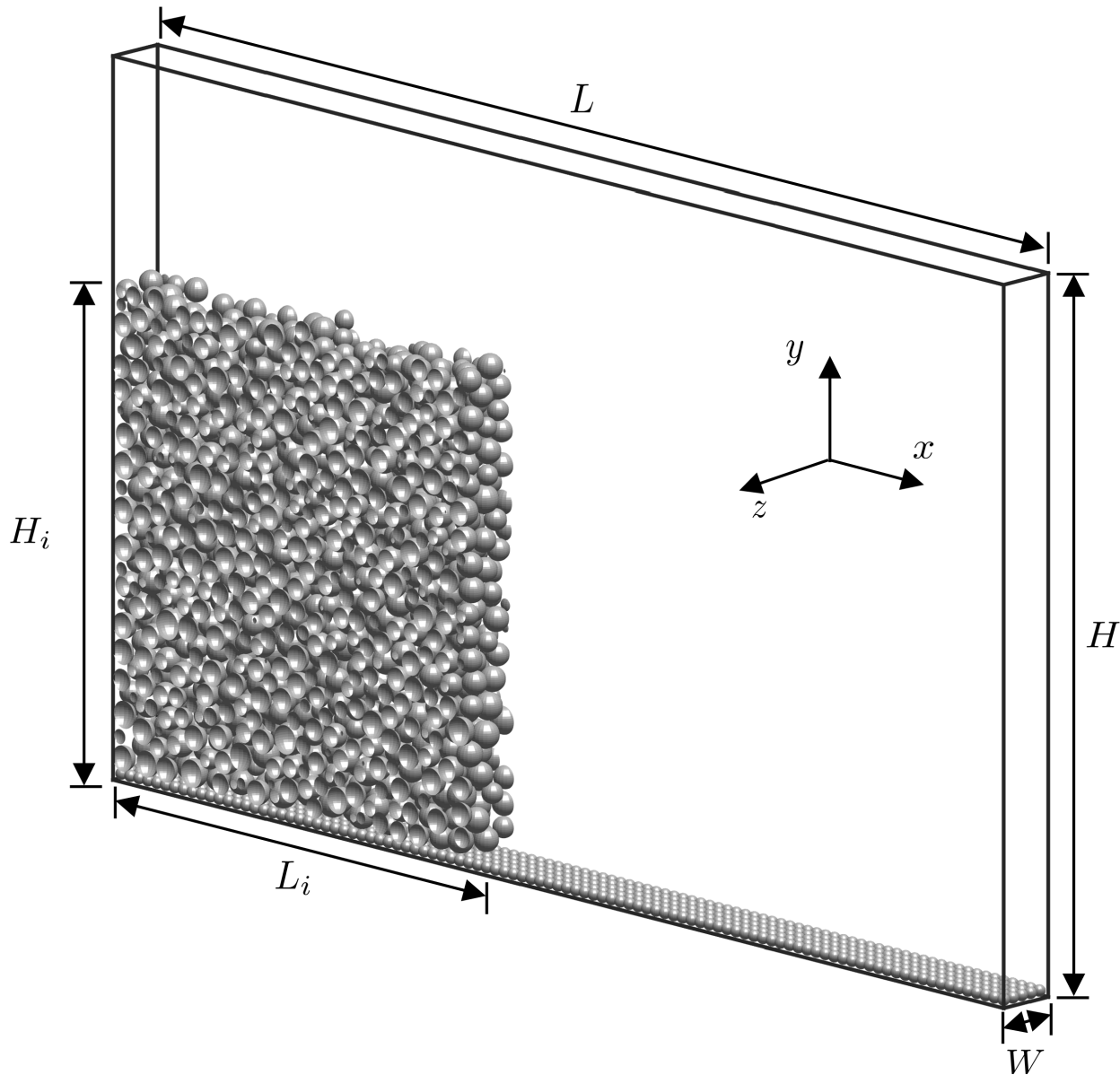


Wang et al. 2017



Langlois et al. 2015

Computational model——initial set-up



1. **Aspect ratios** of the initial column:

$$a = \frac{H_i}{L_i}$$

2. **Cohesive numbers:**

$$C_o = \frac{\max(||\mathbf{F}_{coh,50}||)}{m_{50}g'}$$

Computational model—governing equations

□ Fully coupled, grain-resolving direct numerical simulations.

□ Fluid:

The unsteady Navier-Stokes equations:

$$\frac{\partial \mathbf{u}}{\partial t} + \nabla \cdot (\mathbf{u}\mathbf{u}) = -\frac{1}{\rho_f} \nabla p + \nu_f \nabla^2 \mathbf{u} + \mathbf{f}_{IBM},$$

The continuity equation:

$$\nabla \cdot \mathbf{u} = 0,$$

□ Particle:

Within the framework of the IBM, we calculate the motion of each individual spherical particle.

➤ **Translational velocity:**

$$m_p \frac{d\mathbf{u}_p}{dt} = \mathbf{F}_{h,p} + \mathbf{F}_{g,p} + \mathbf{F}_{c,p},$$

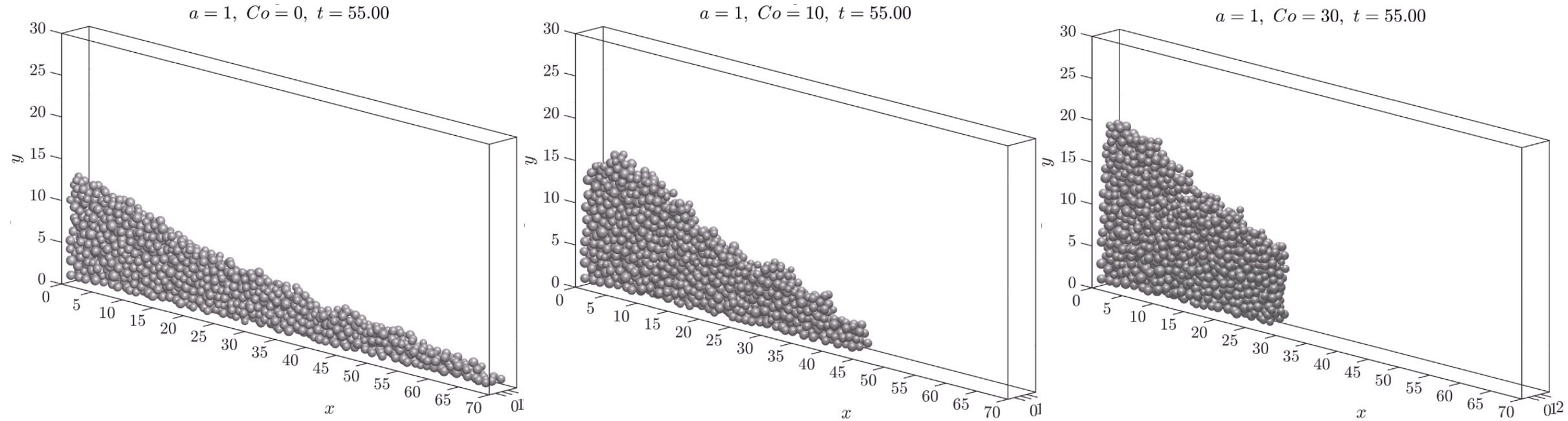
➤ **Angular velocity:**

$$I_p \frac{d\boldsymbol{\omega}_p}{dt} = \mathbf{T}_{h,p} + \mathbf{T}_{c,p},$$

➤ **$\mathbf{F}_{c,p}$: the force due to particle collisions:**

$$\mathbf{F}_{c,p} = \begin{cases} \mathbf{F}_l & \text{Lubrication force} \\ \mathbf{F}_n & \text{Normal contact force} \\ \mathbf{F}_t & \text{Tangential contact force} \\ \mathbf{F}_{coh} & \text{Cohesive force} \end{cases}$$

Observations — shallow columns



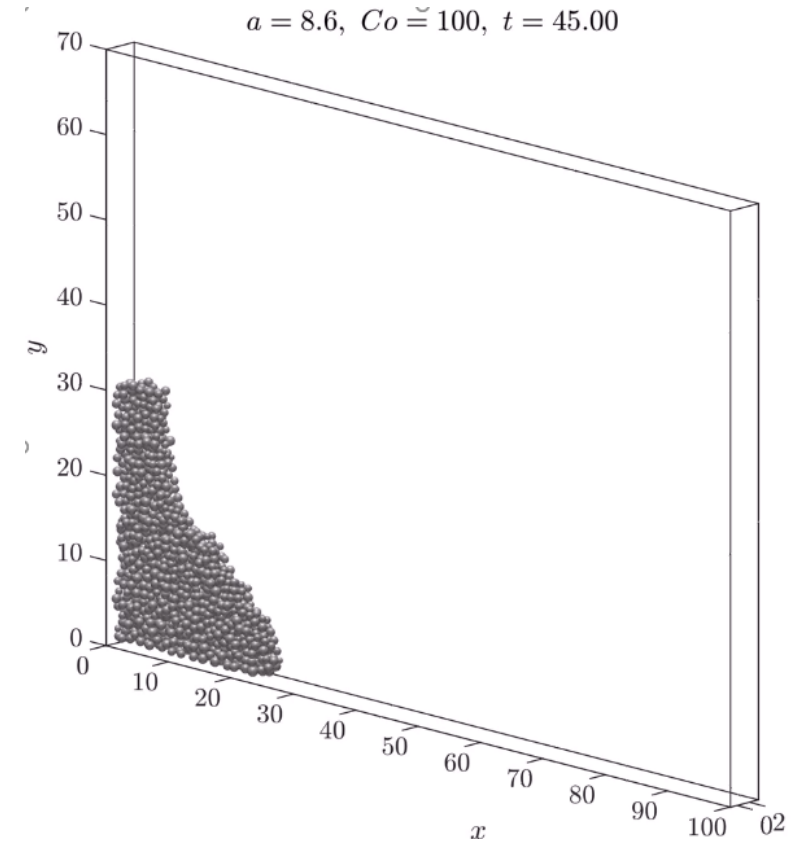
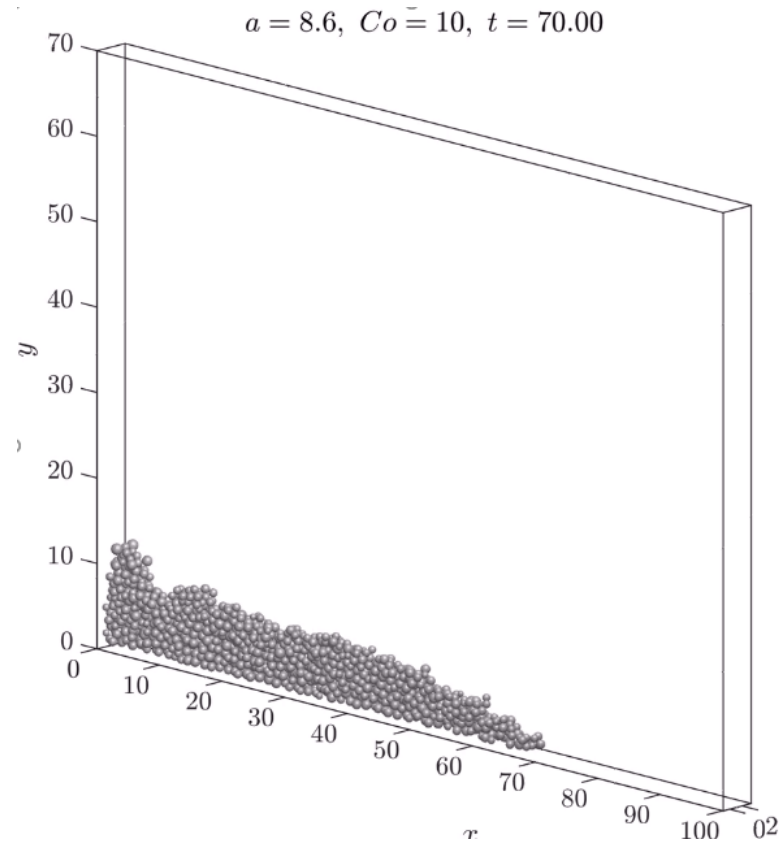
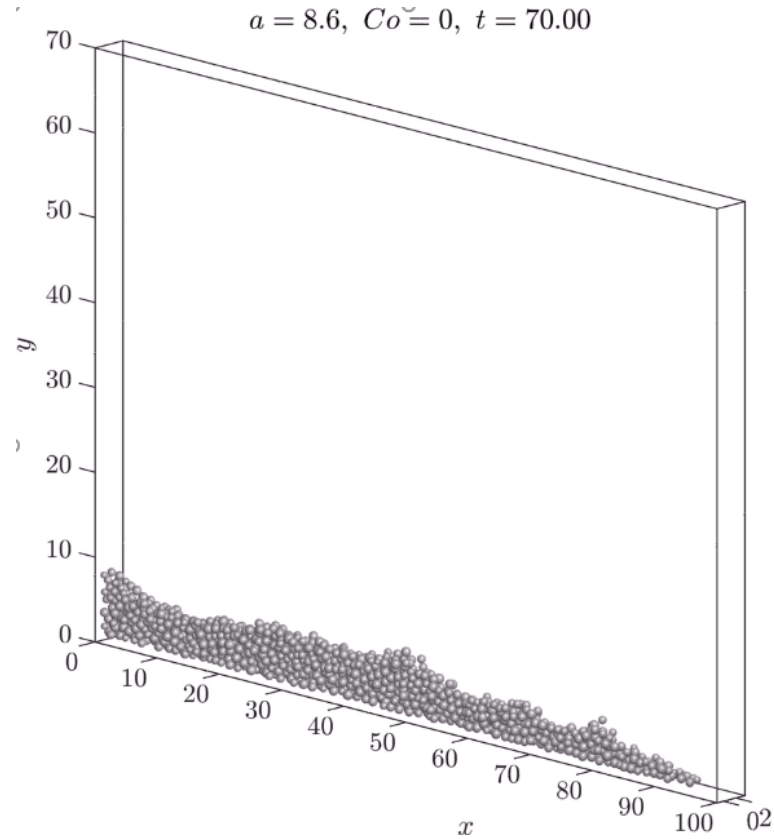
For $Co=0$:

1. The particles near the **upper right** part slide down along **an inclined shear surface**.
2. The counterflow **detaches some particles** from the main body.
3. The final deposit exhibits a **triangular shape**.

For $Co=10$ and 30 :

Particles spread **slower**. **No** particles are detached. The final deposit is **shorter** and **thicker**.

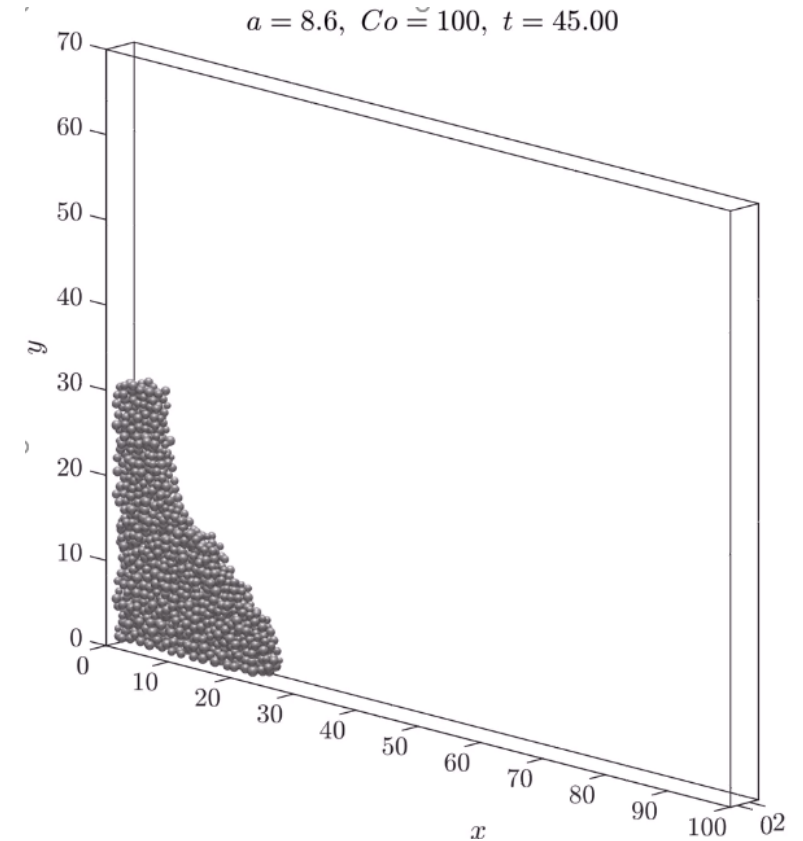
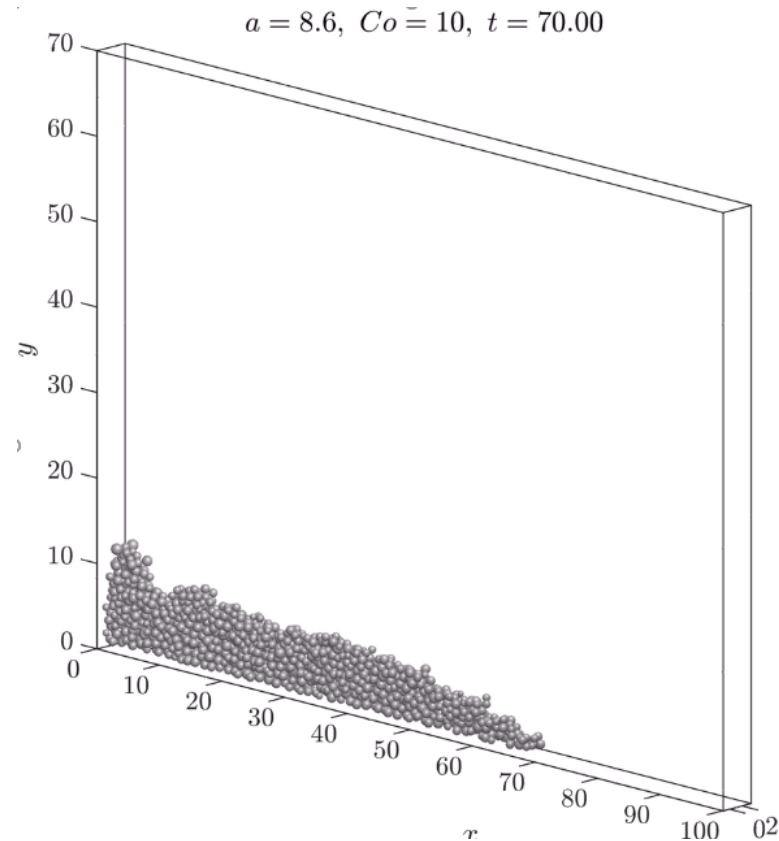
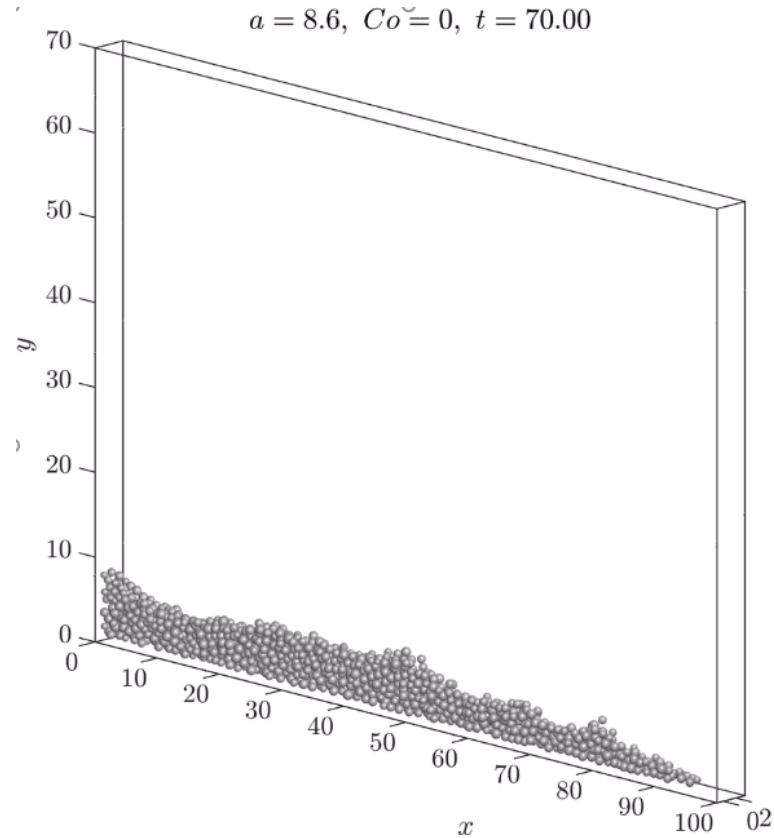
Observations — tall columns



For $Co=0$:

1. The **upper section freely falls** in the early stage.
2. Most of the **lower section moves to the right** and forms a flow head.
3. Many particles are **detached** from the collapsing column. They settle more slowly.

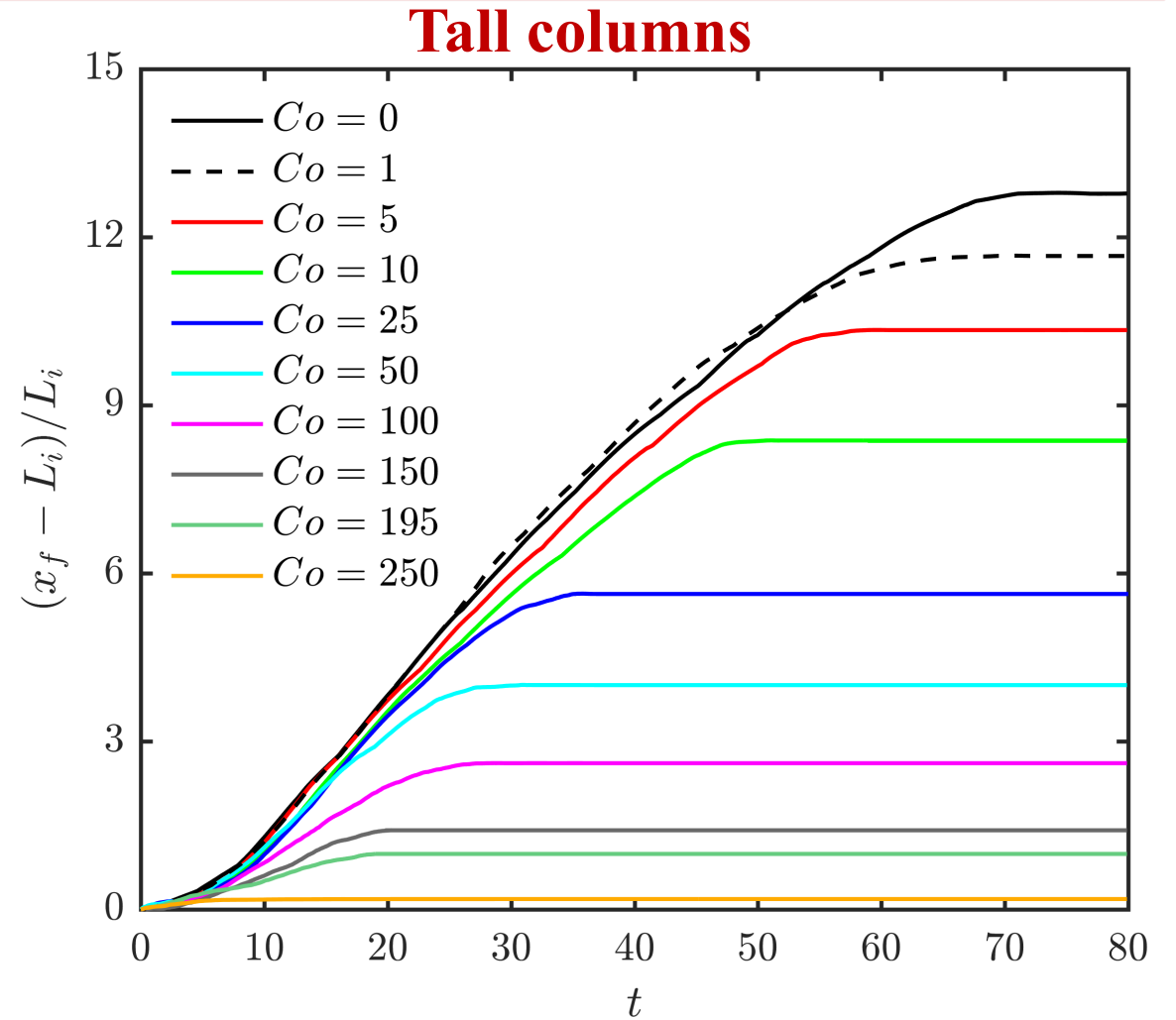
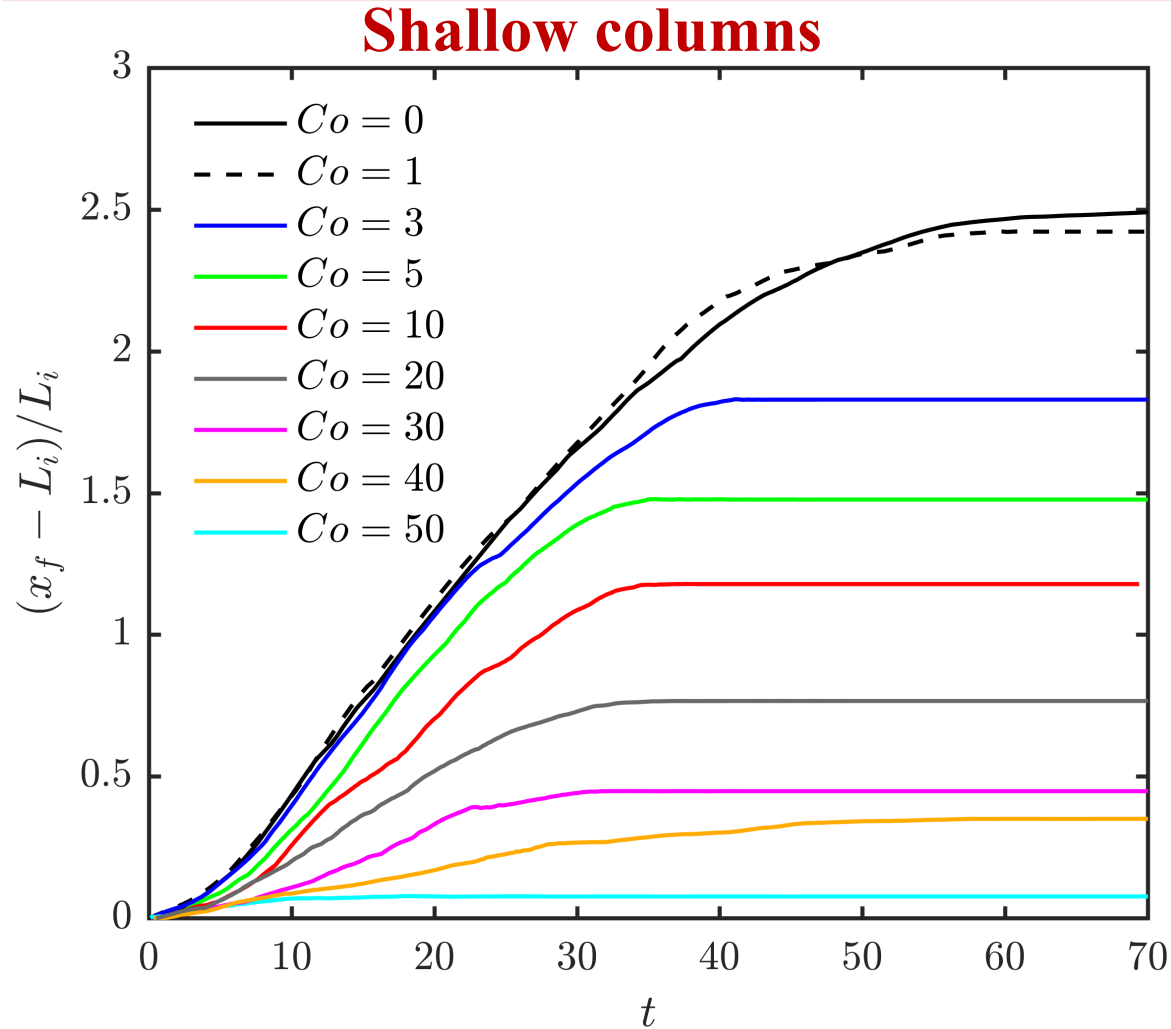
Observations — tall columns



For $Co=10$ and 100 :

1. Spreading slower and shorter runout distance.
2. The cohesive force **prevents the detachment** of individual particles from the collapsing column.
3. When $Co=100$ the cohesive force **slows down the horizontal spreading** of the particles near the bottom.

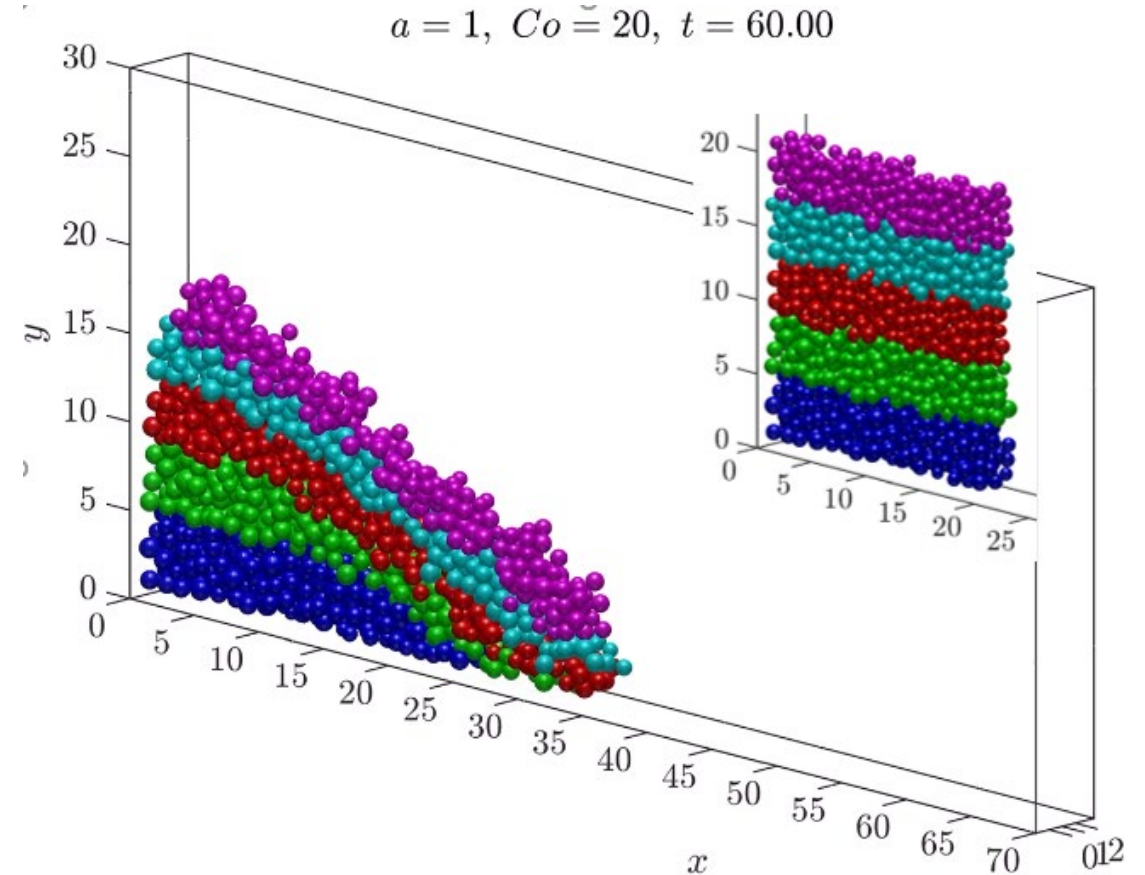
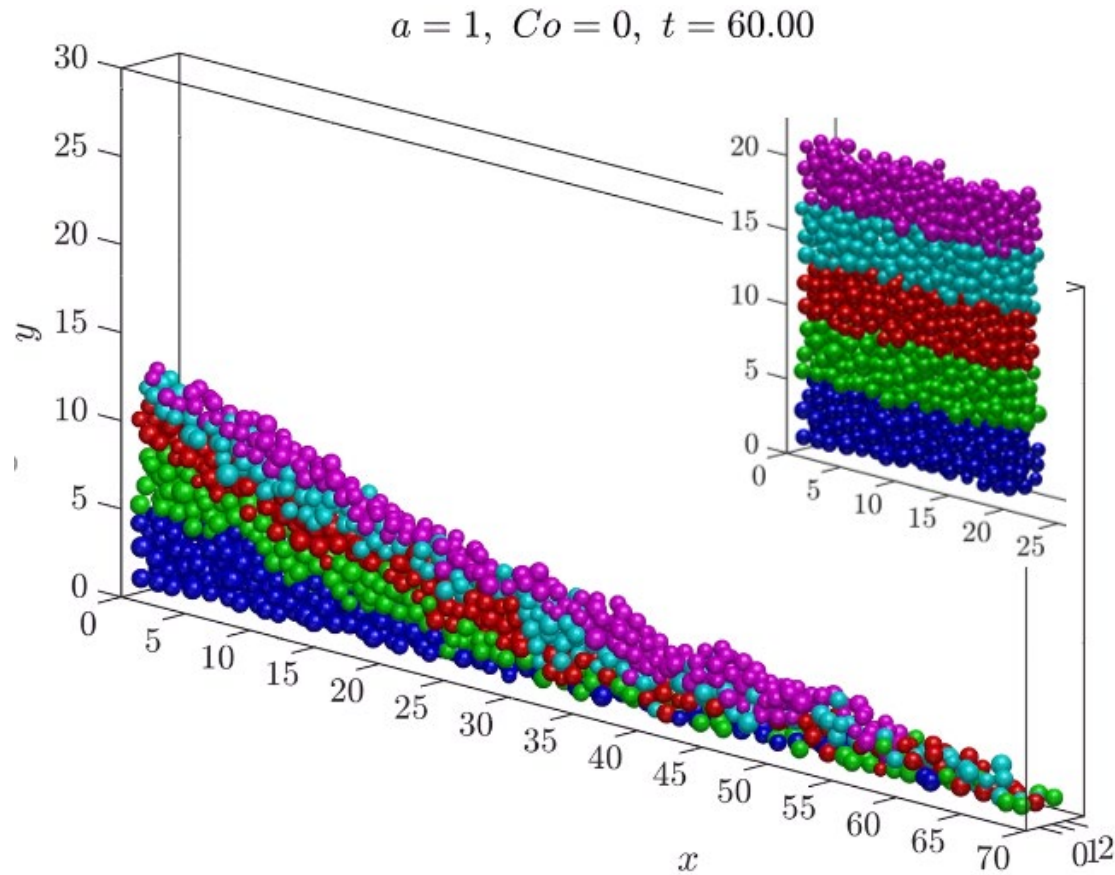
Front location



Acceleration stage → Constant front velocity stage → Deceleration stage (Meruane et al. 2010; Lee et al. 2018)

1. The **front location** decreases with increasing the **cohesive numbers Co** .
2. The **shallow and tall columns** no longer collapse for **$Co=50$** and **250** , respectively.

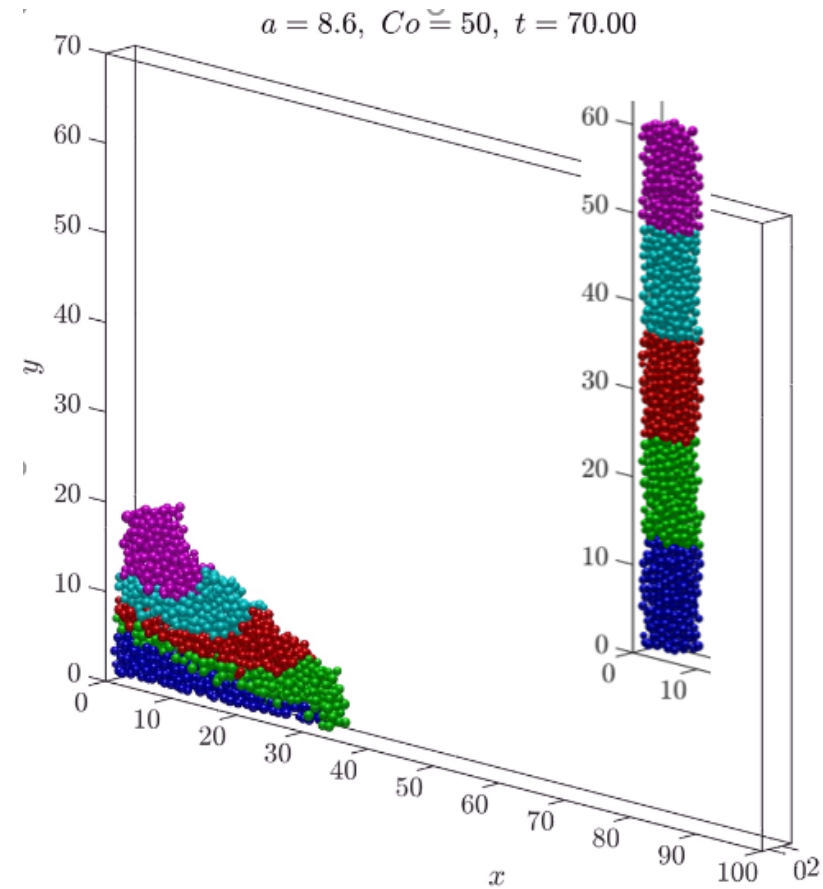
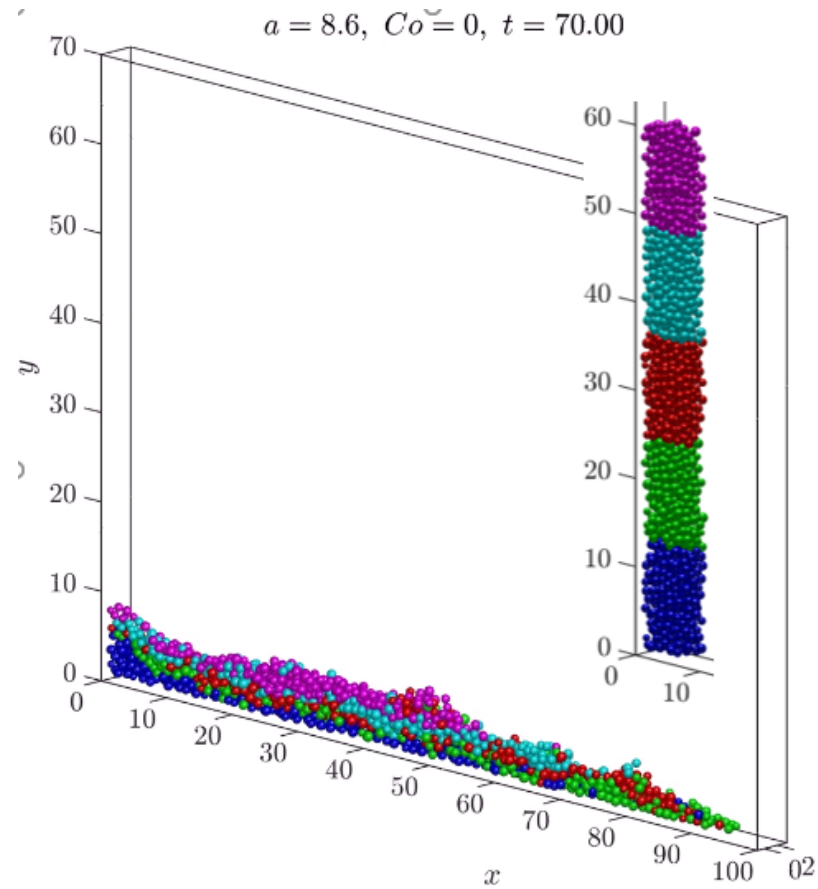
Internal structure of the deposit



For $Co=0$ and 20 :

1. The layers **stay mostly intact** towards the **left** of the final deposit.
2. There tends to be **more mixing** towards the **tip** of the deposit on the right.
3. The particles that **travel the farthest** are mostly from the **second and third-highest layers**.

Internal structure of the deposit



The particles that **travel the farthest** now originate in the **second layer from the bottom**.

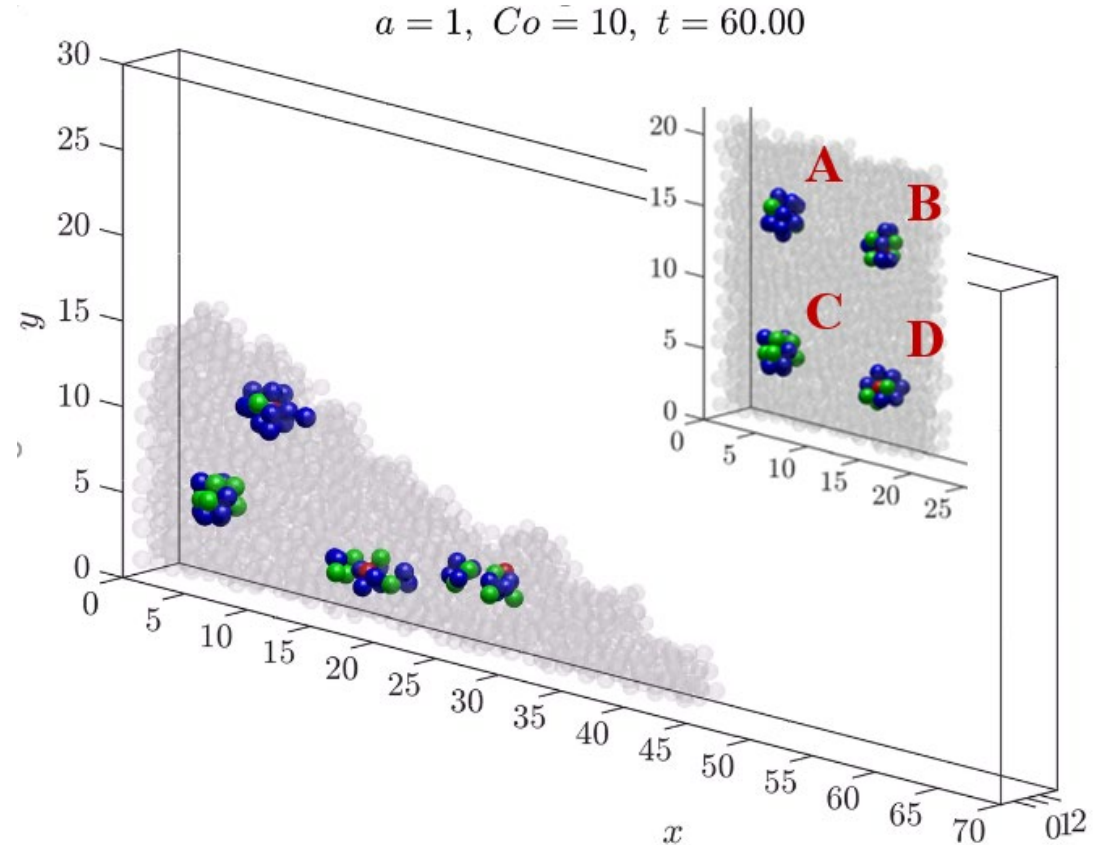
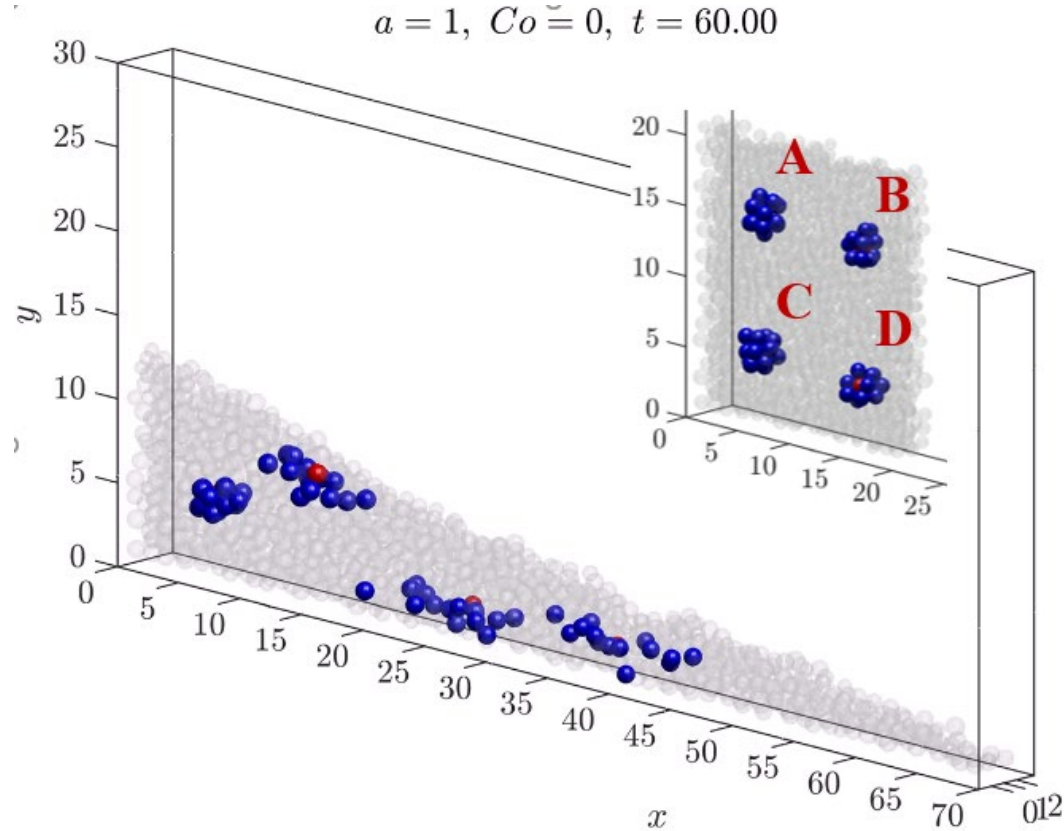
For $Co = 0$:

There is **more pronounced mixing** between the layers.

For $Co = 50$:

The layers tend to **retain their identities**.

Formation and persistence of flocs in the interior of the granular collapse



For $Co=0$ and 20:

Particles **A** and **C** and their neighbors travel a relatively **short distance**.

Particles **A** and **C** **remain close** to their initial neighbors.

Particles **B** and **D** travel much **larger distances**.

The **neighbors** of particles **B** and **D** become much **more separated** from each other for $Co = 0$ as compared to $Co = 10$.

Conclusions

1. The **cohesive force** acts to **prevent the detachment** of individual particles from the main body of the collapsing column, **reduce** its **front velocity**, and yield a **shorter** and **thicker final deposit**.
2. Computational particle tracking indicates that the **cohesive force reduces** the **mixing** of particles within the collapsing column, and it results in the **formation** and **persistence** of **flocs**

Thanks!