





Determination of CME orientation and consequences for their propagation

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- Derive the inclination of (I)CME from remote and in-situ data
- > To study whether the inclination has an effect on the propagation => non-radial flows (NRF) in the sheath region

DATA

- 63 associated CME-ICME
- 2008-2016

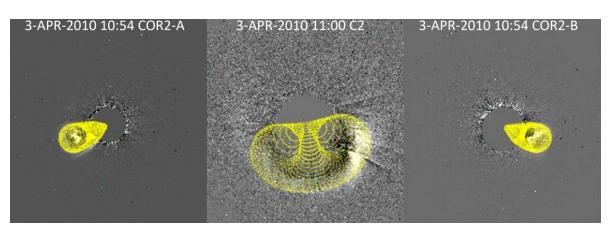
- Palmerio+2018
- ➤ Maricic+2020
- > Temmer+2021
- Nitta&Mulligan+2017

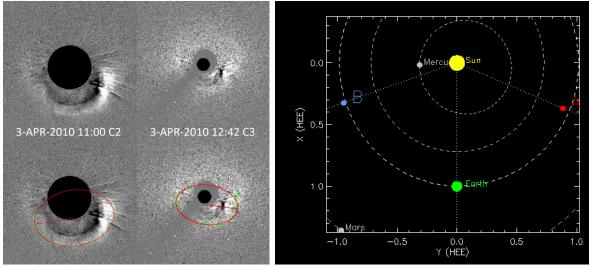
- MC in-situ signatures
- ▶ GCS reconstraction,≥ 2 vantage point

22 associated CME-ICME

Remote observations

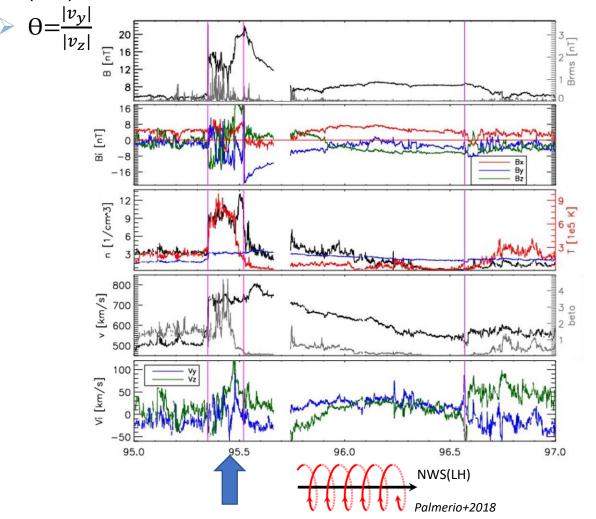
- Graduated Cylindricall Shell (GCS) 3D (croissant) geometry
- Ellipse fit 2D (ellipse) geometry
- Inclination of CME axis to solar equator as measured from solar west





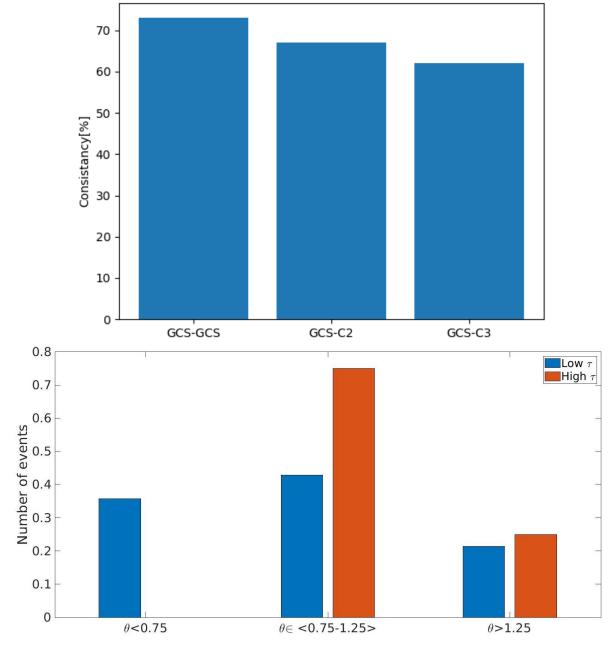
In-situ observations

- Manual boundary detection and FR type determination=> inclination
- Calculation of ratio of NRF in the sheath region in y- and z- direction (GSE)



RESULTS

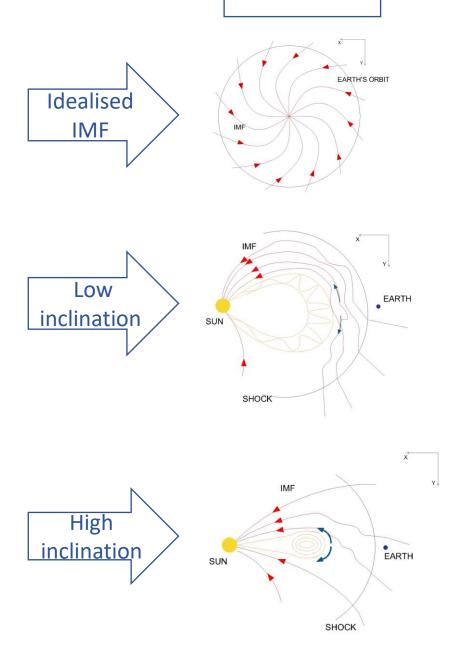
- GCS and ellipse fit are equally good in determining the *dominant inclination* of the event, high or low inclination
- The majority of events (66%) have consistent dominant inclination in the near Sun environment (GCS) and in the near Earth environment (in-situ)
- The majority of events have a Θ ratio close to 1, regardless of their dominant inclination
- There is no high inclination event with an Θ ratio smaller than 0.75. Also, the number of events with a high Θ ratio is higher for high inclination events than for low inclination events.
- ➤ High inclination events have more pronounced NRF in the y-direction, while low inclination events have more pronounced NRF in the z-direction



relative number of events (i.e., occurrence frequency) separately for high (orange) and low(blue) inclination events, with respect to the calculated θ ratios.

INTERPRETATION

Ecliptic plane



- > Axial curvature due to the rooting of the footpoints of the CME at the Sun => toroidal extent
- Cross-sectional curvature due to internal magnetic field structure => cross-sectional extent
- Low inclination CME: ambient plasma to be more easily deflected in the north-south (z-) direction where the CME extent is smaller, that is we might expect Θ <1 ($v_z > v_y$)
- \succ High inclination CME: ambient plasma to be more easily deflected in the east-west (y-) direction where the CME size is smaller, that is we might expect $\Theta>1$ ($v_{v}>v_{z}$)







Thank you for your attention!









CONCLUSION

- The majority of events have consistent tilt as observed remotely and in-situ
- >NRF for low inclination events are more pronounced in the N-S direction
- >NRF for high inclination events are more pronounced in the E-S direction
- > There might be a relation between the inclination and propagation
- ➤GCS and ellipse fit methods give equally good tilt estimations which will allow us to improve the sample size when considering the whole Soho era

TO DO:

- Increase the event sample (SOHO era, ellipse fit)
- > Check the transit time and gamma parameter

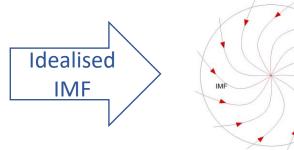
INTERPRETATION

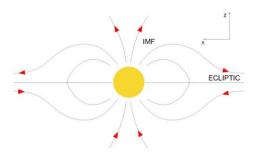
Ecliptic plane

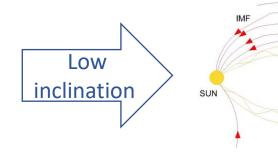
EARTH'S ORBIT

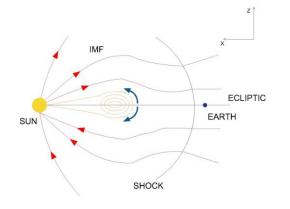
EARTH

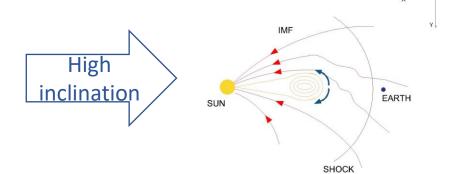
Meridional plane



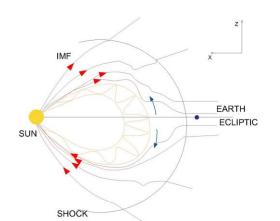








SHOCK



- > Axial curvature due to the rooting of the footpoints of the CME at the Sun
- Cross-sectional curvature due to internal magnetic field structure.
- Low inclination CME: ambient plasma to be more easily deflected in the northsouth direction where the CME extent is smaller, that is we might expect Θ < 1 (vz>vy)
- High inclination CME: ambient plasma to be more easily deflected in the eastwest direction where the CME size is smaller, that is we might expect Θ > 1 (vz<vy)