







Revisiting the wave telescope for larger numbers of spacecraft

Leonard Schulz ¹, Karl-Heinz Glassmeier ^{1,2}, Ferdinand Plaschke ¹, Uwe Motschmann ^{3,4}

EGU General Assembly 2022, Session ST 2.4

Correspondence: l.schulz@tu-bs.de

¹Institute of Geophysics and Extraterrestrial Physics, Technische Universität Braunschweig, Germany

² Max Planck Institute for Solar System Research, Göttingen, Germany

³ Institute of Theoretical Physics, Technische Universität Braunschweig, Germany

⁴German Aerospace Center, Institute of Planetary Research, Berlin, Germany

Motivation/Problem

Wave telescope (e.g. Motschmann et al. 1996, JGR) uses few spatial points (spacecraft) to estimate spatial fourier transform

→ Able to find wave vectors in magnetic field data

Currently 4 S/C (MMS, Cluster)

→ But can be used with larger numbers of S/C! (e.g. HelioSwarm)

Some arising questions:

- What are the advantages of more satellites? → Better resolution, increased number of waves resolved
- What is the Nyquist limit for larger numbers of S/C?
- What happens to the spatial blindness for more than 4 S/C?

→ Simulations









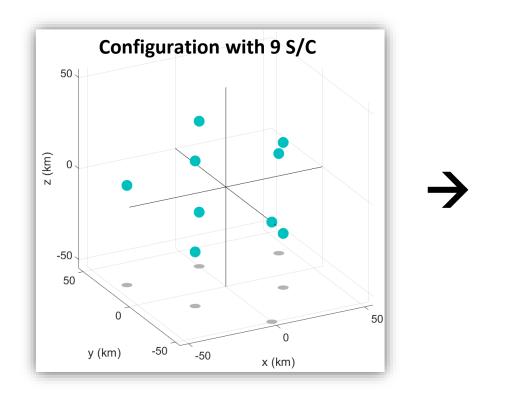
Number of detectable waves

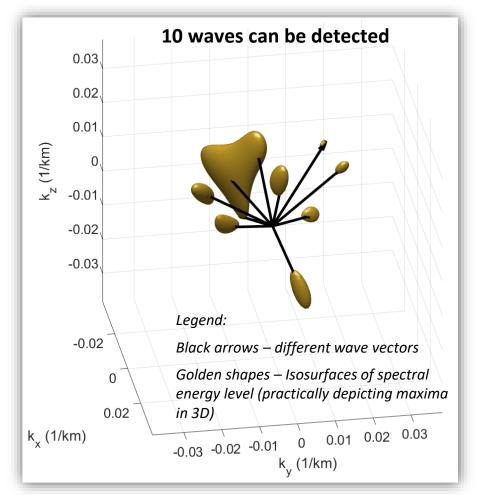
Theoretical limit of 3*n-1 waves (e.g. Tjulin et al. 2005) is never reached (e.g. Motschmann et al. 1996, JGR → 7 waves, 4 S/C;

own simulations)

Higher number of S/C can increase the number of detected waves!

For example:











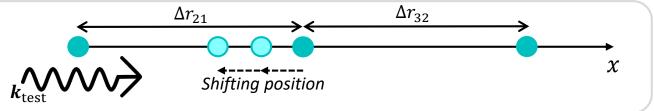




Nyquist limit in 1D

3 S/C configuration:

Center S/C position shifted in every simulation

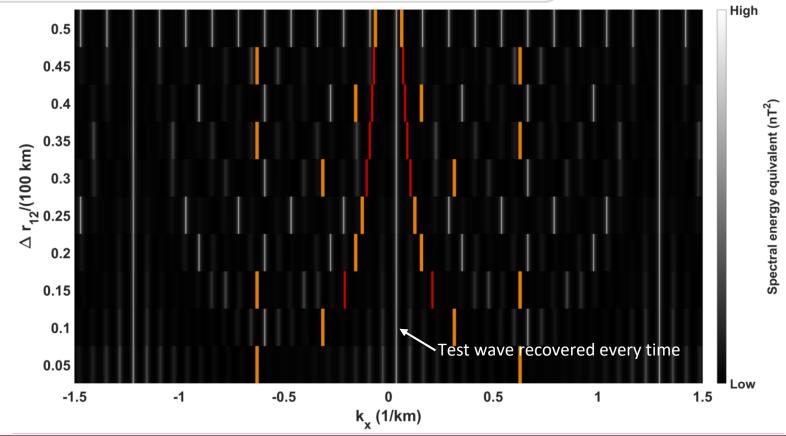


Nyquist wave vector defined by spacecraft distances:

$$2 \text{ S/C: } \frac{k_{\text{Ny}}}{\Delta r_{21}}$$

3 s/c:
$$k_{\text{Ny}} = \frac{\pi}{\gcd(\Delta r_{21}, \Delta r_{32})}$$

Implications for 2D and 3D case!











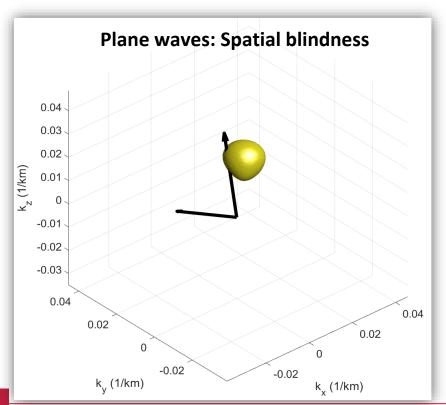
Spatial blindness

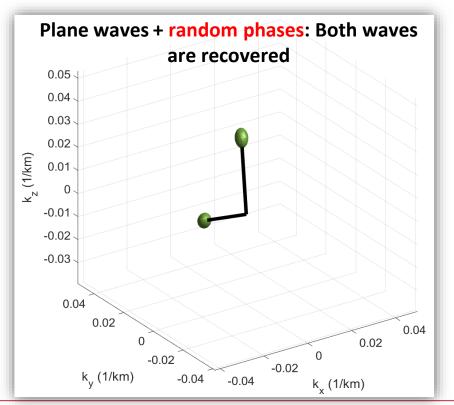
Using only plane waves: $\mathbf{B} = \mathbf{B_0} \exp(i(\omega_0 t - \mathbf{k}\mathbf{x})) \Rightarrow$ wave telescope suffers from spatial blindness (can't resolve waves with same ω_0 but different wave vectors \mathbf{k})

Introducing random phases ϕ_q : $\mathbf{B} = \mathbf{B_0} \exp(i(\omega_0 t - \mathbf{k}\mathbf{x})) exp(i\phi_q) \rightarrow$ wave telescope can separate and recover waves with different \mathbf{k}

→ Essentially introduction of phase jumps during the measurement, elimination of coherence

We found: spatial blindness is also present for more than 4 S/C!













Summary

The wave telescope is able to work with more than 4 S/C

Larger numbers of detection points (spacecraft) lead to:

- Increased resolution in k-space
- Larger numbers of waves that can be separated
- The Nyquist limit in k-space (defining the analysis limit) gets larger and more complex in 1D,
 even more so in 2D and 3D
- Spatial blindness is still present
 - → Our research in all of those areas is ongoing.







