

Improving the Empirical Solar Wind Forecast

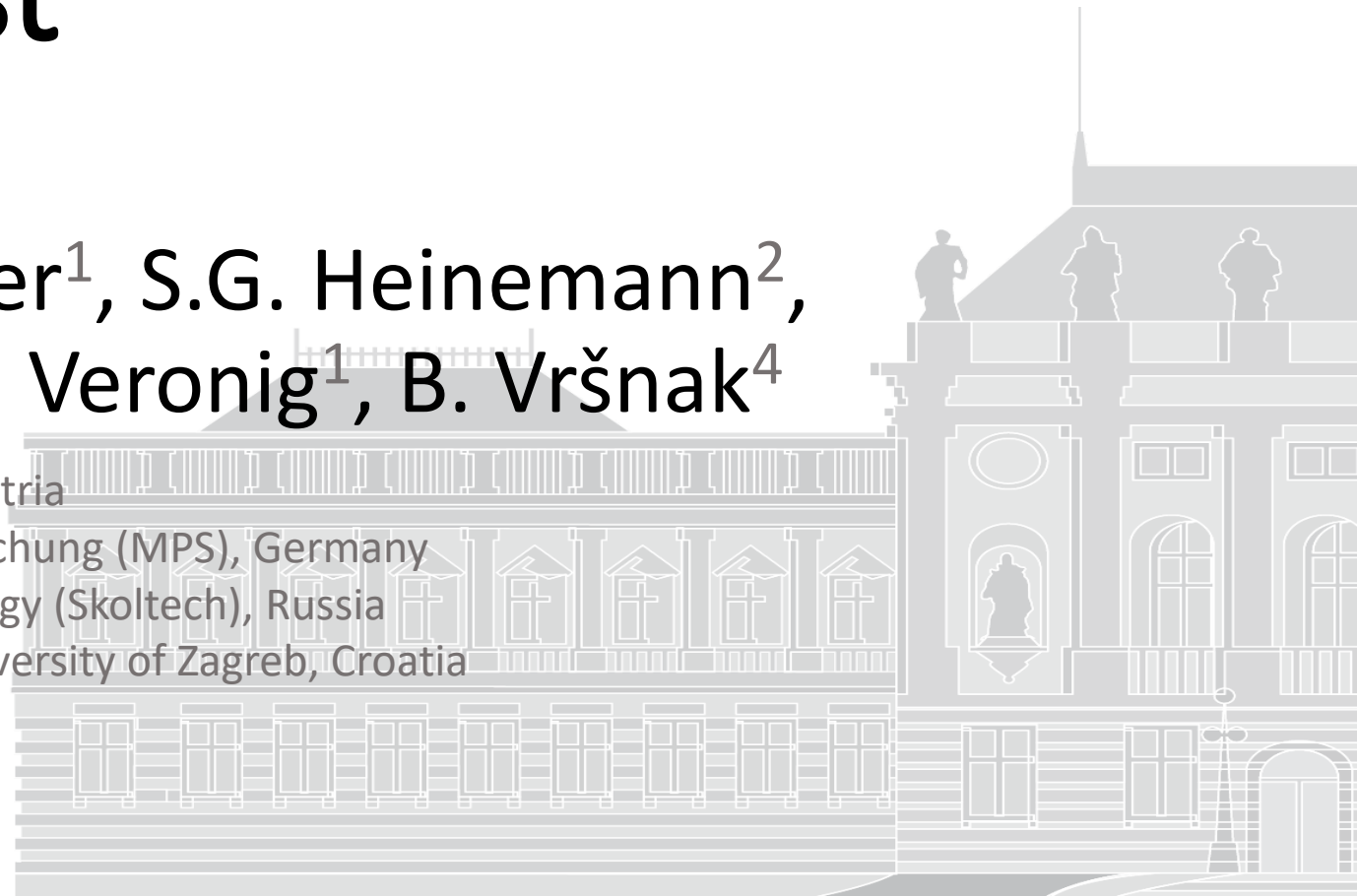
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ESWF:

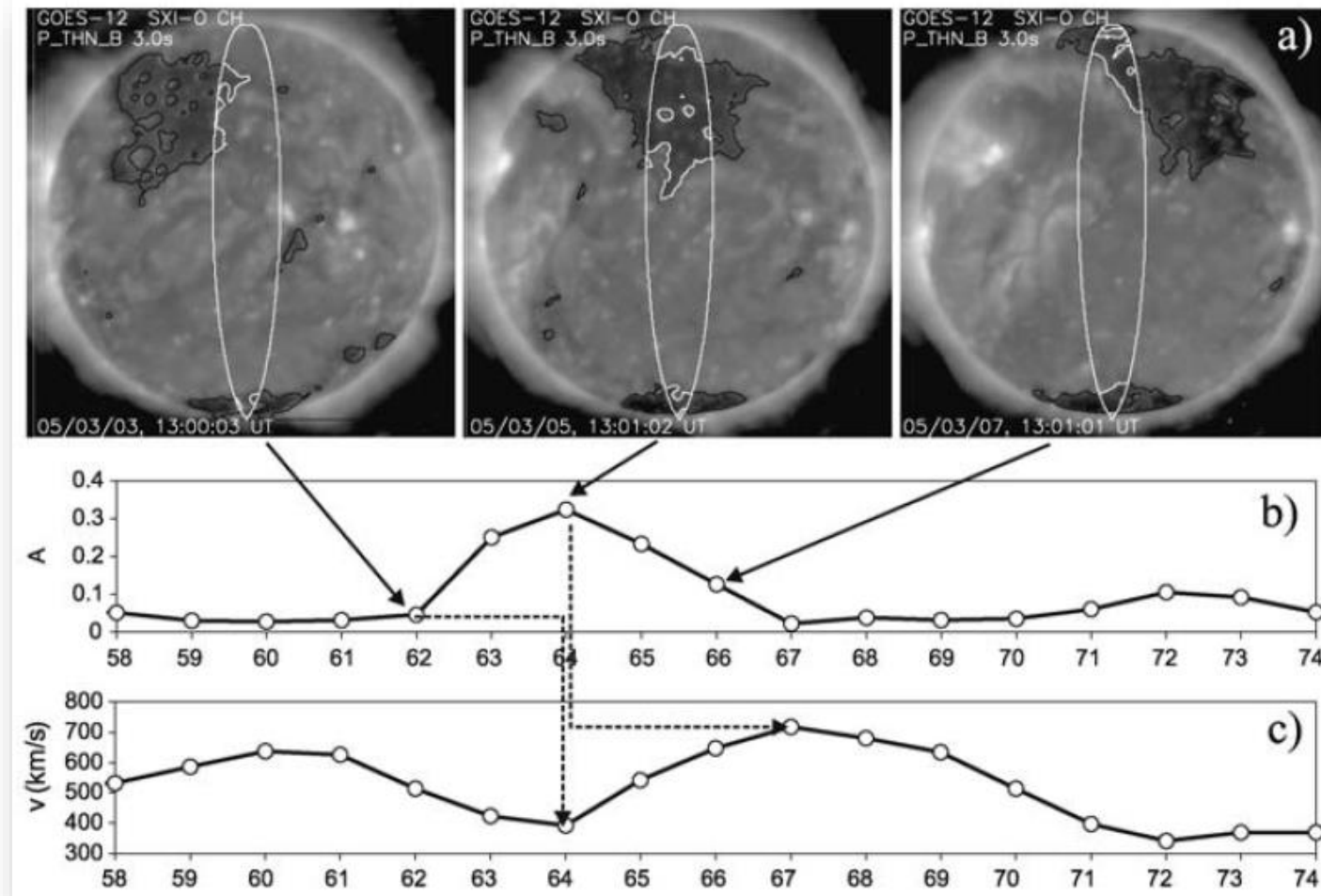
Empirical model for forecasting solar wind speed from relative Coronal Hole area (Vršnak et al., 2007; Rotter et al., 2015; Reiss et al., 2016)

$$v(t) = c_0 + c_1 * A(t')$$
$$t' = t + 4d$$

<https://swe.uni-graz.at/index.php/services/solar-wind-forecast>

Forecasting High Speed Streams (HSS):

measured profile: asymmetrical
model profile: symmetrical



Vršnak et al.
(2007)

Improvement Steps

- ESWF 1.0

- Constant parameters c_0, c_1
- $t' - t = 4d$

– Vršnak et al., 2007

- ESWF 2.0

- Parameters c_0, c_1 re-evaluated every CR
- $t' - t = 4d$

– Rotter et al., 2015

- ESWF 3.0

- PROP – algorithm:

$$t' - t = \frac{1AU}{v_{ESWF}(t)}$$

Improvement Steps

- ESWF 3.1

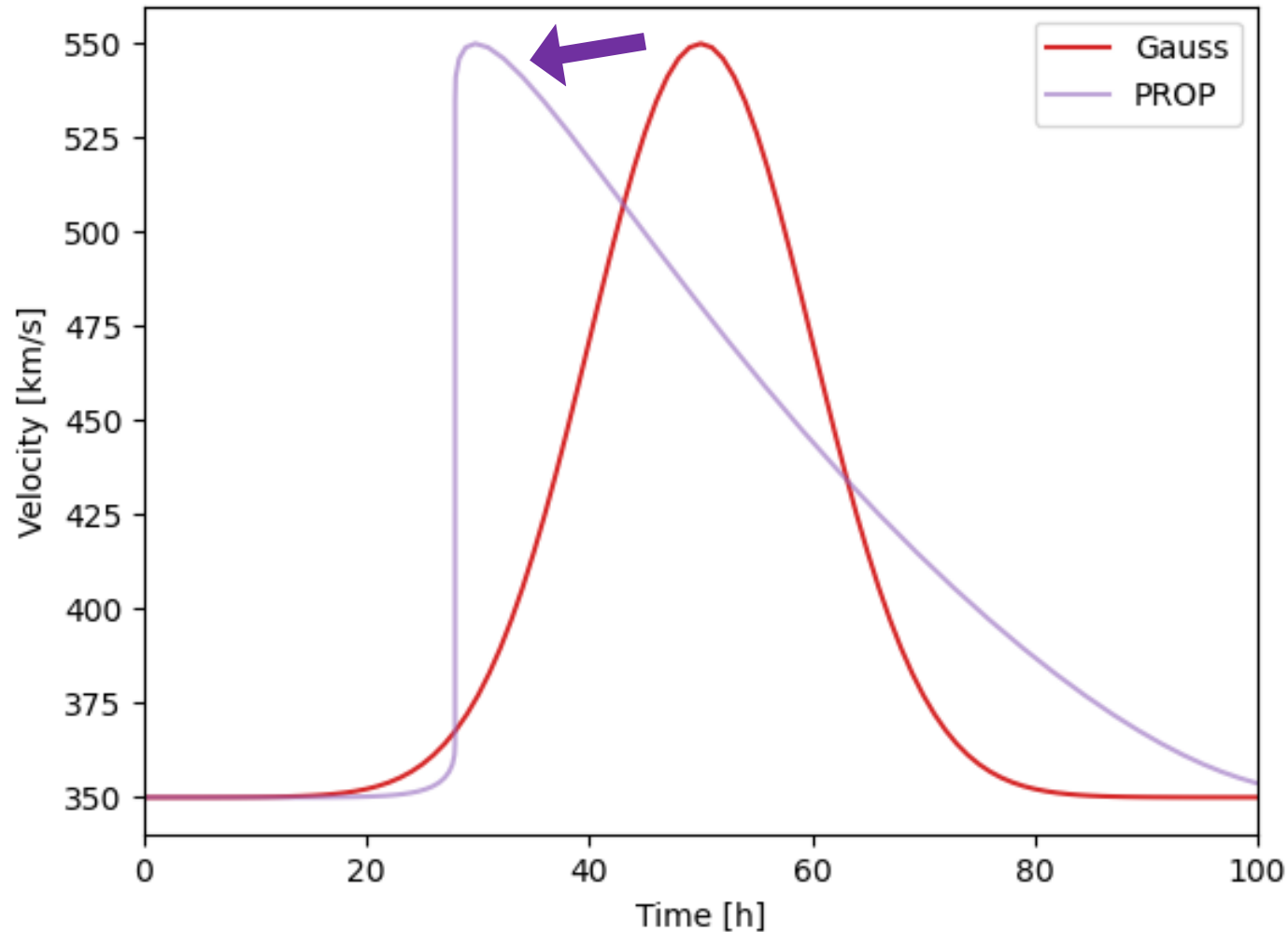
- PILE – algorithm:

$$t' - t = \frac{1AU}{(v_{ESWF}(t) + v_{AMB})/2}$$

- ESWF 3.2

- PILE – algorithm
 - Dynamic Thresholding (Heinemann et al., 2018)
 - Co-latitude information

PROP – Algorithm



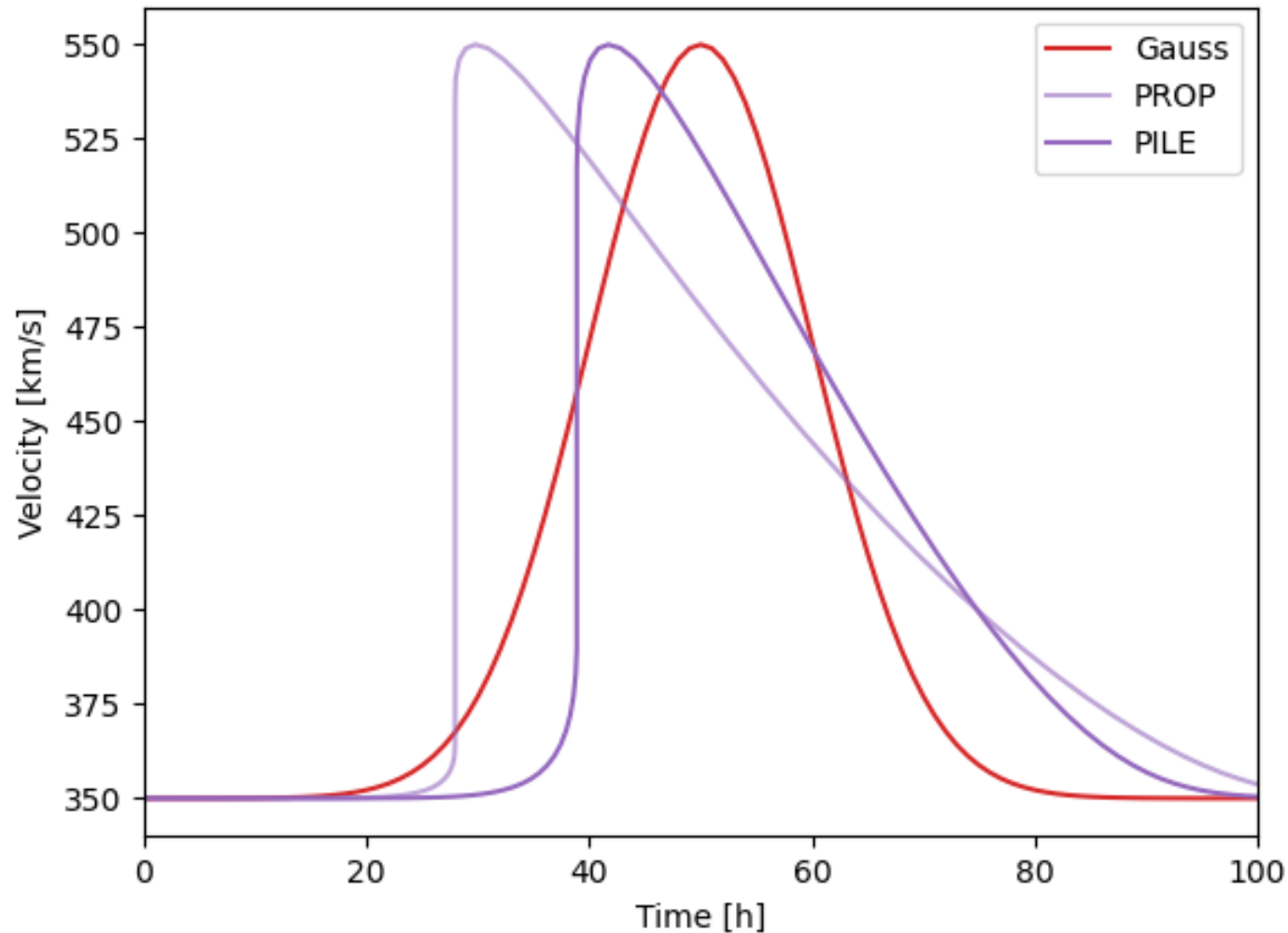
Propagation of
velocities from the sun
to 1AU

$$t' - t = \frac{1AU}{v_{ESWF}}$$

-> Asymmetric Profile

PILE – Algorithm

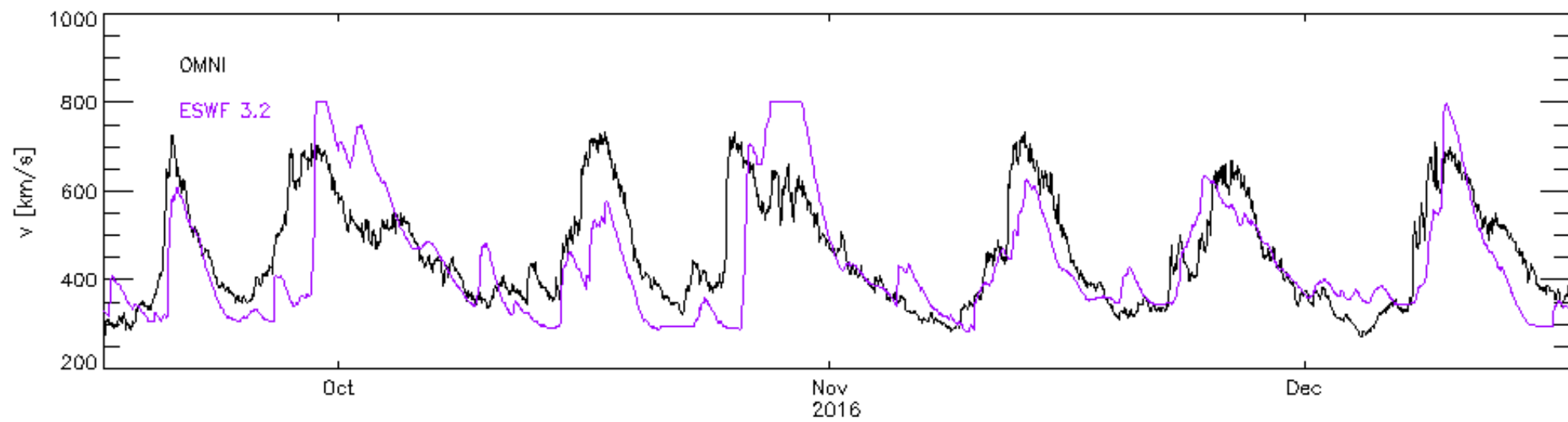
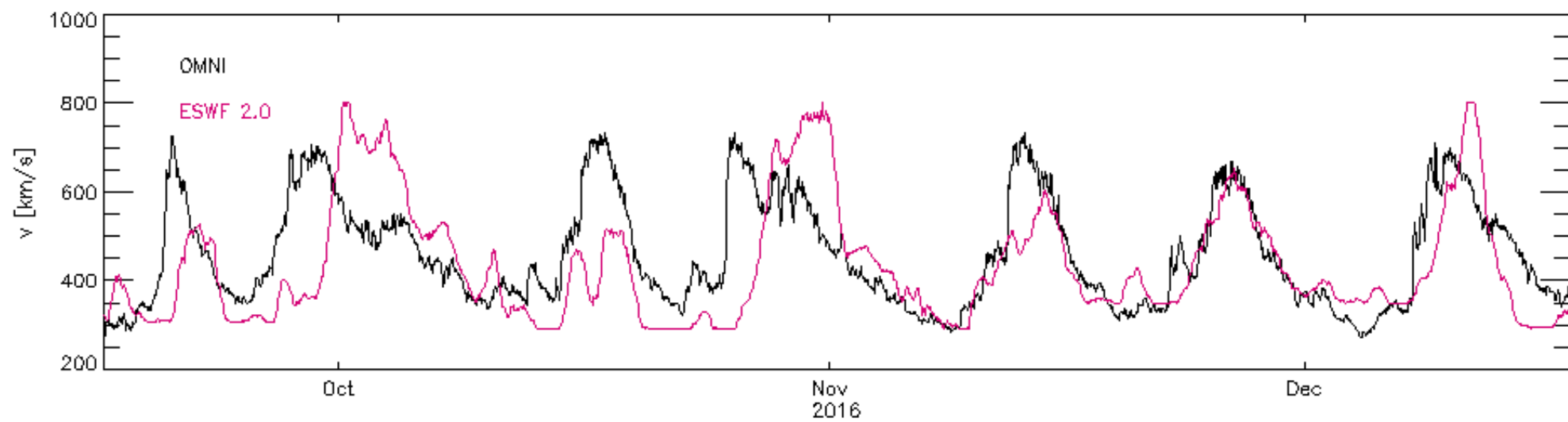
Less aggressive!



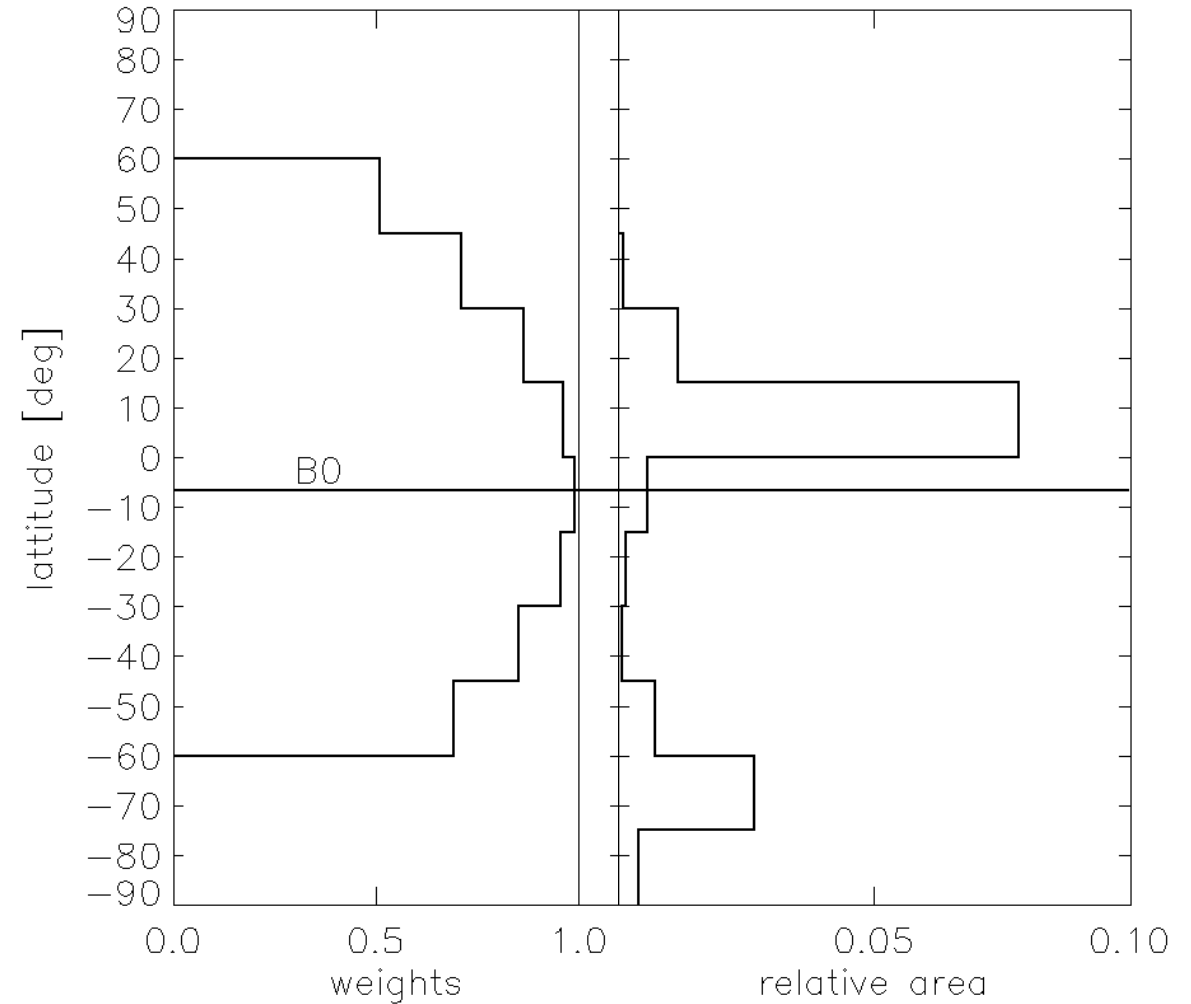
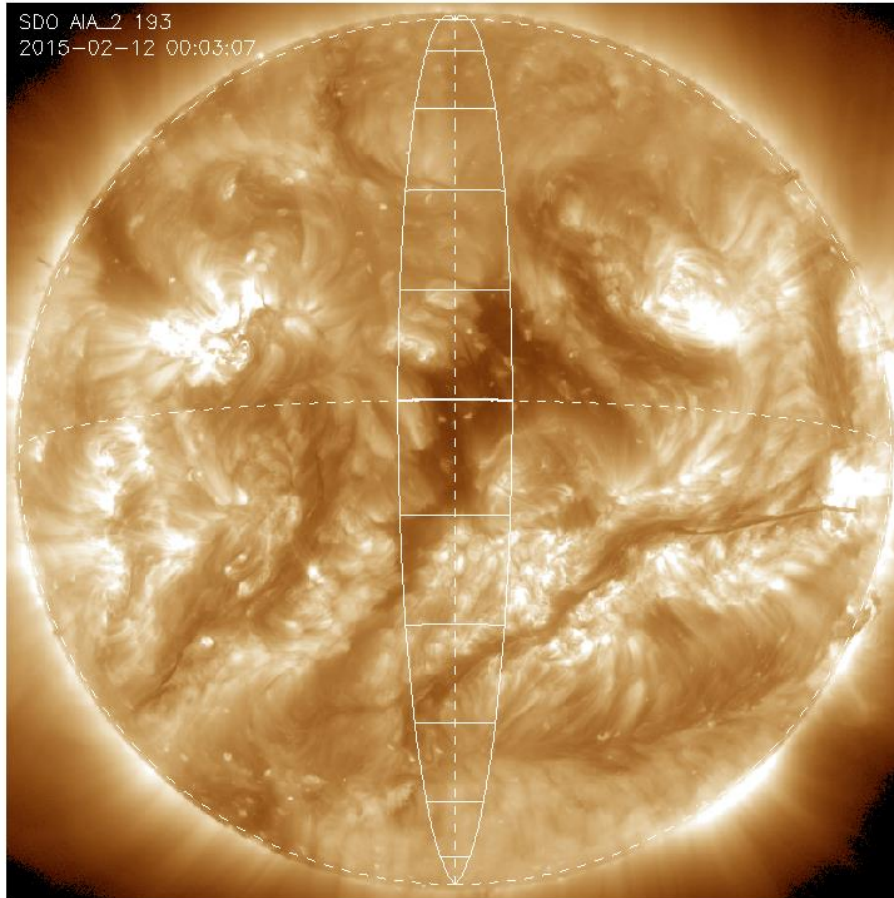
Propagation of
velocities from the sun
to 1AU

$$t' - t = \frac{1AU}{(v_{ESWF}(t) + v_{AMB})/2}$$

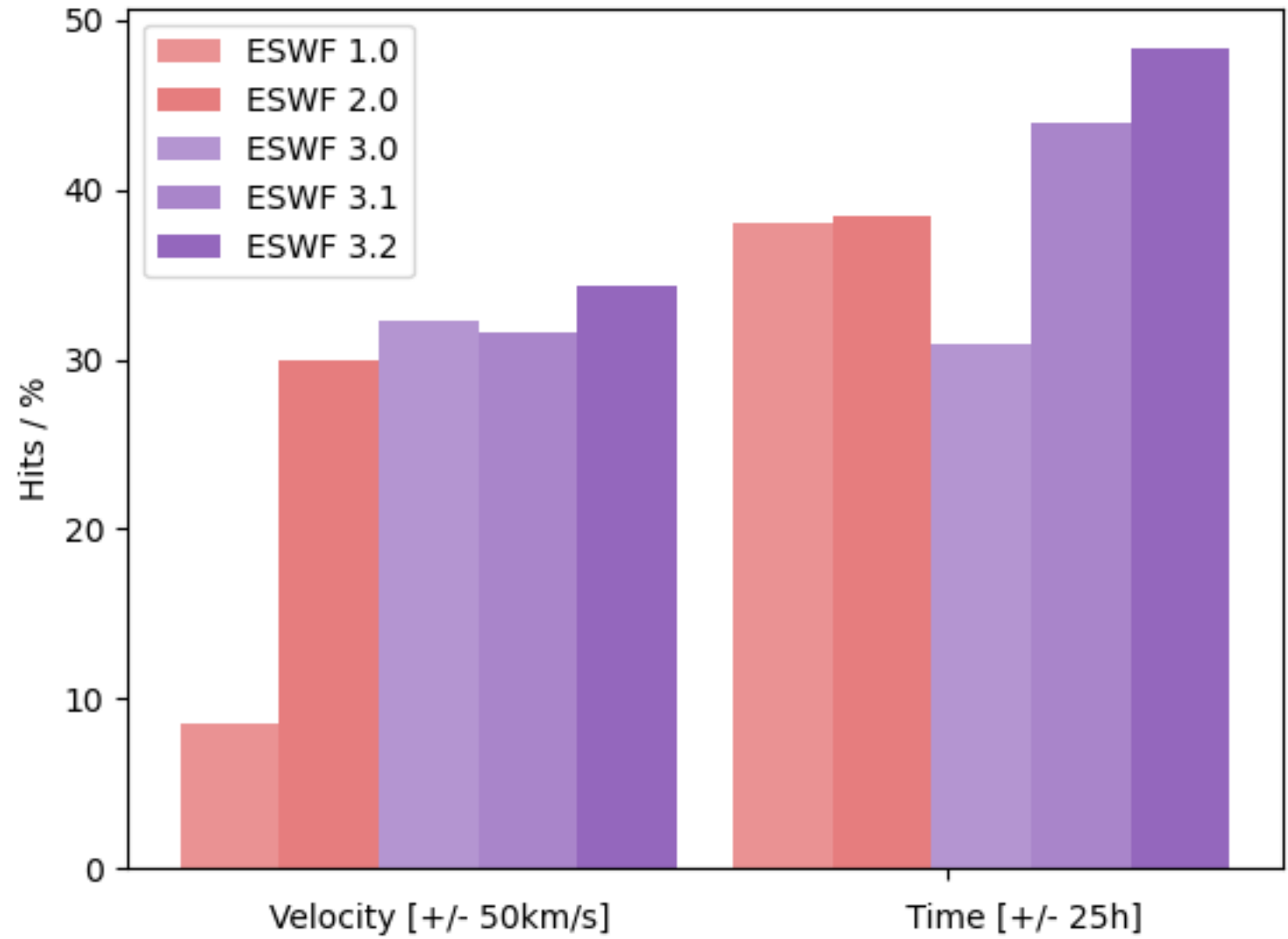
-> Slowed down by
ambient SW



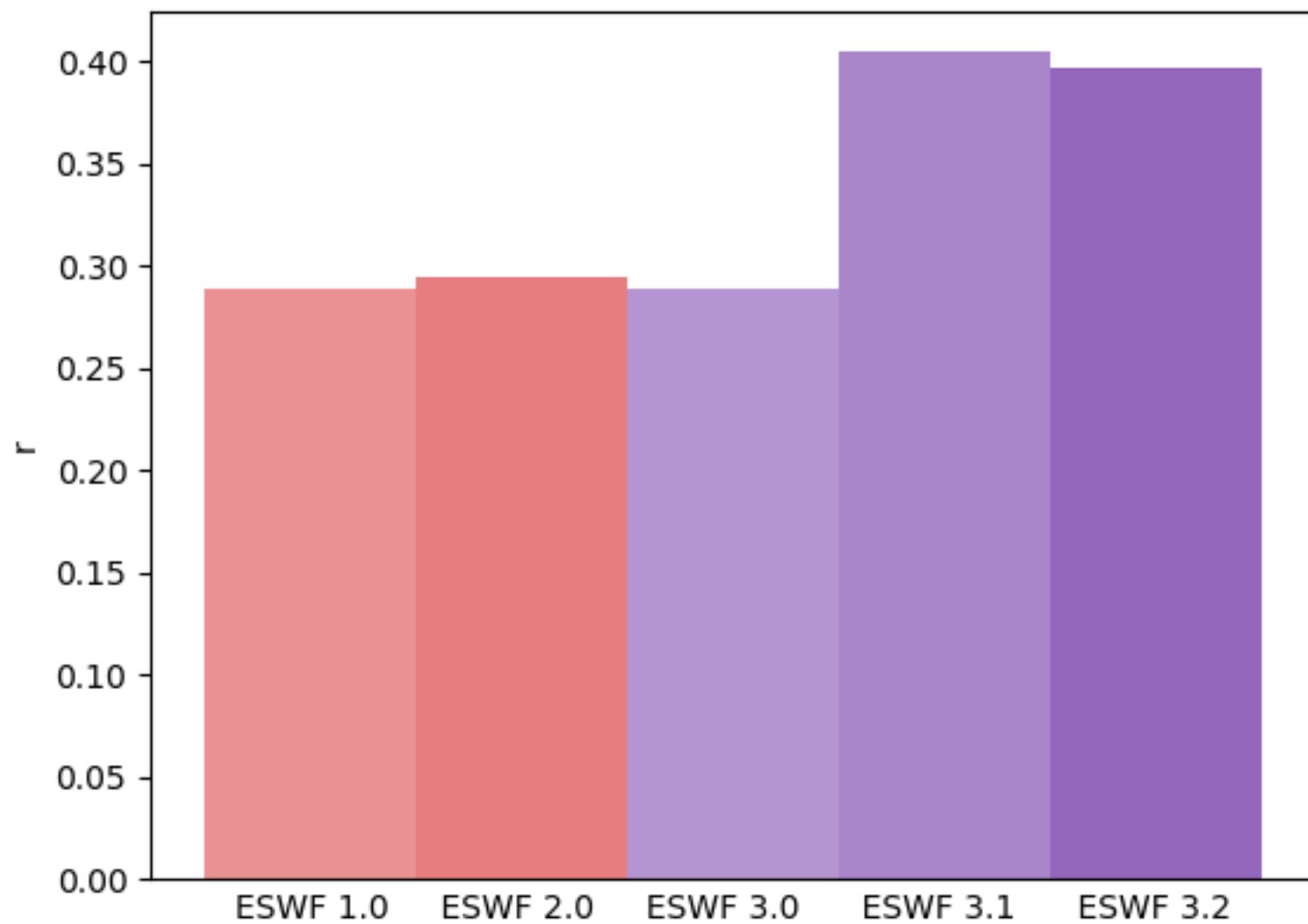
Co-latitude Information



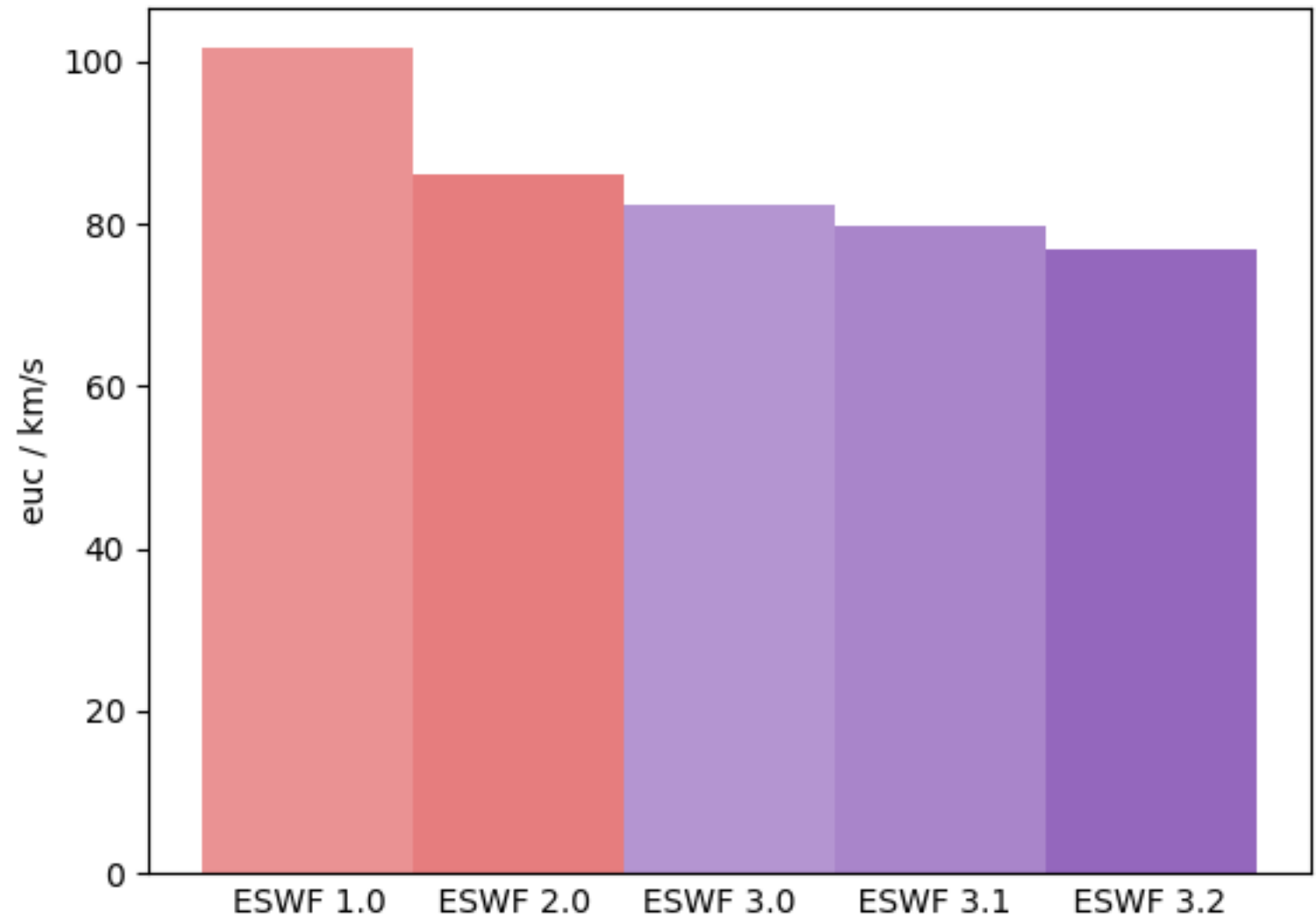
Hit & Miss



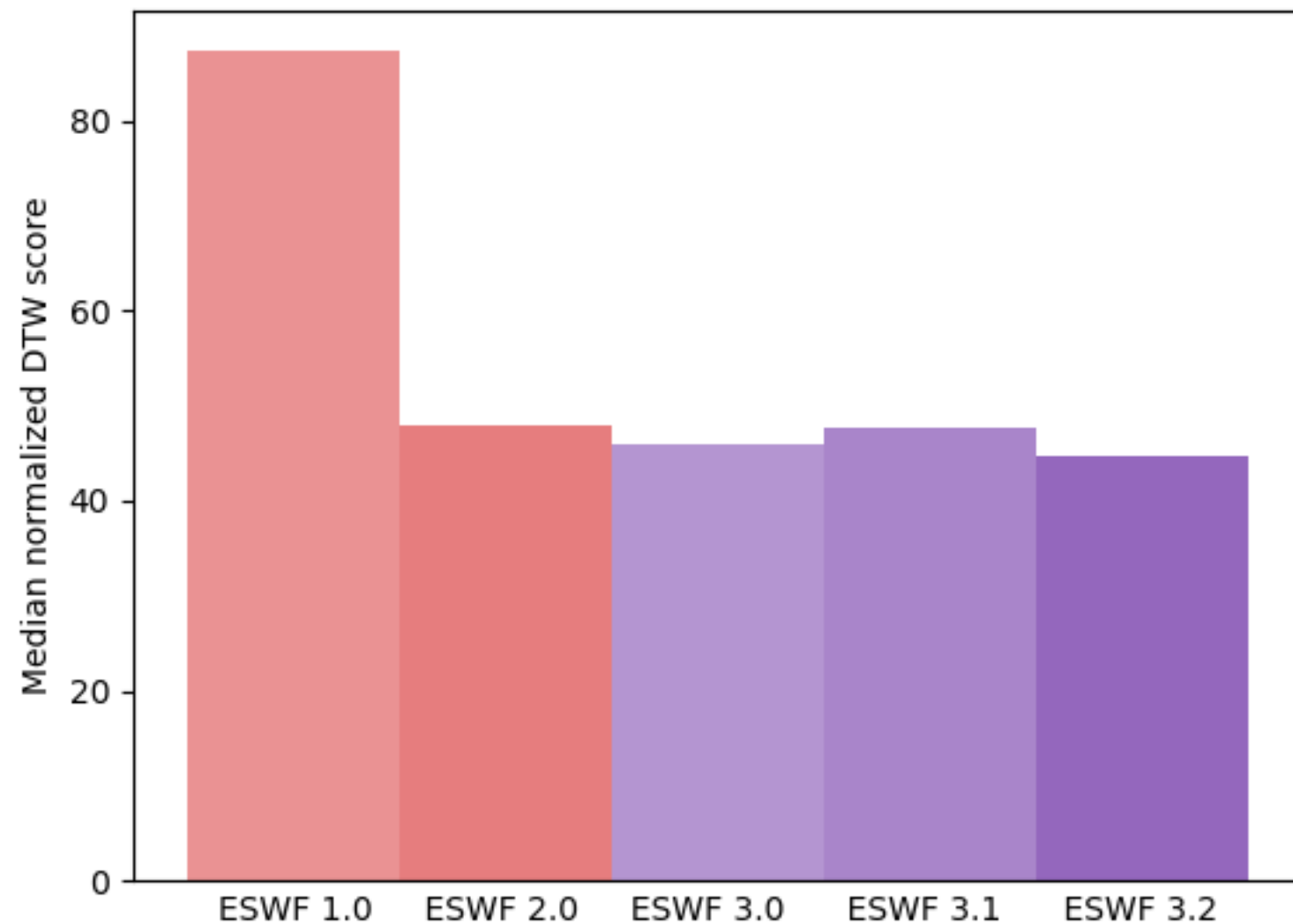
Pearson Correlation



Euclidian Distance

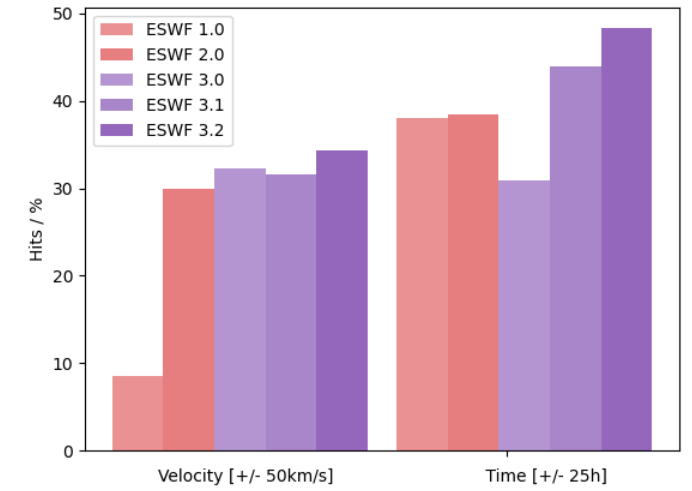
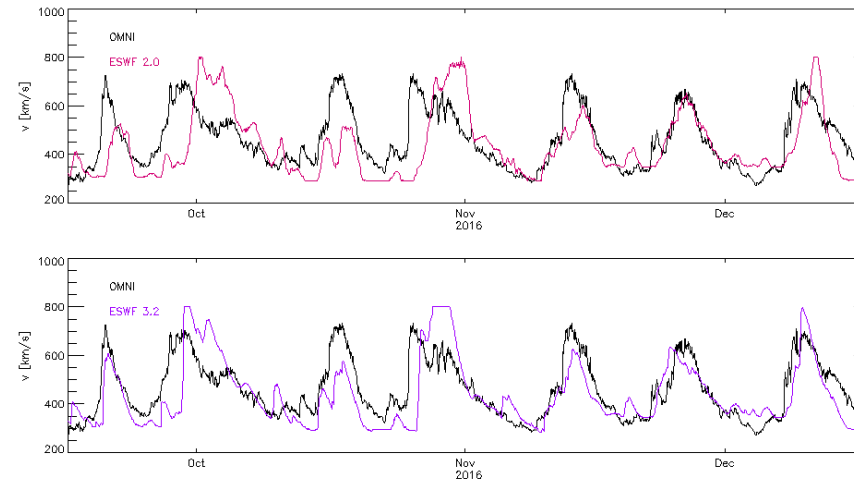
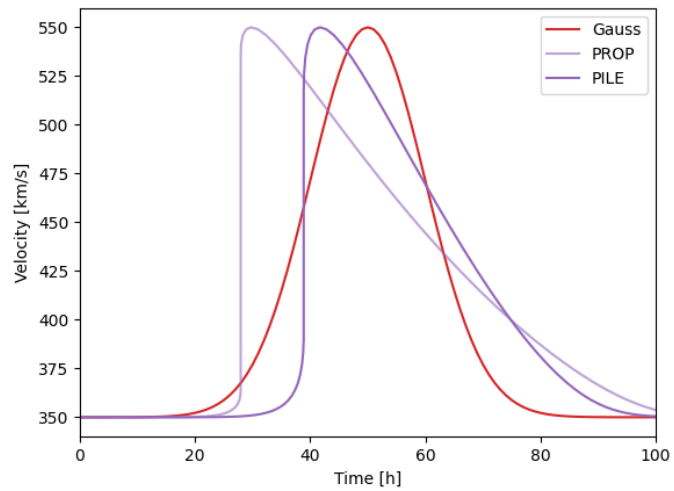


Dynamic Time Warping Score



References

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Thank You!

