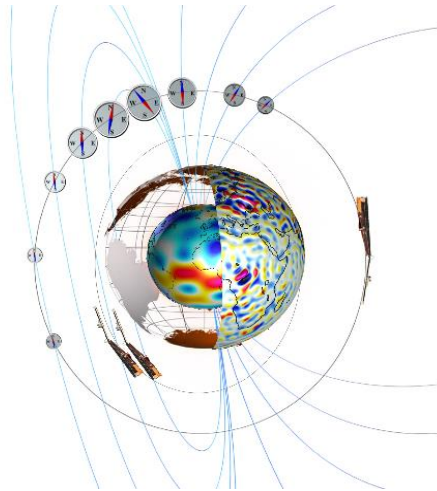


# TIRO

## Topside Ionosphere Radio Observations from multiple LEO-missions

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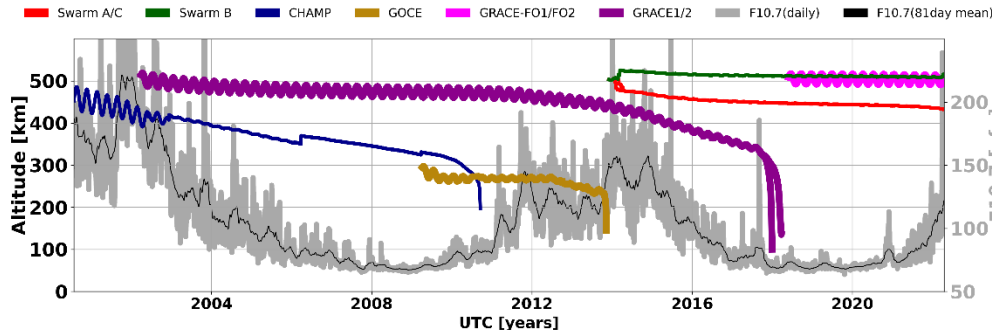


- 1: GFZ German Research Centre for Geosciences, Potsdam, Germany
- 2: Leibniz Institute for Atmospheric Physics, Kühlungsborn, Germany
- 3: Faculty of Aerospace Engineering, Delft University of Technology, Delft, Netherlands
- 4: Astronomical Institute, University of Bern, Bern, Switzerland
- 5: Wuhan University, Wuhan, China

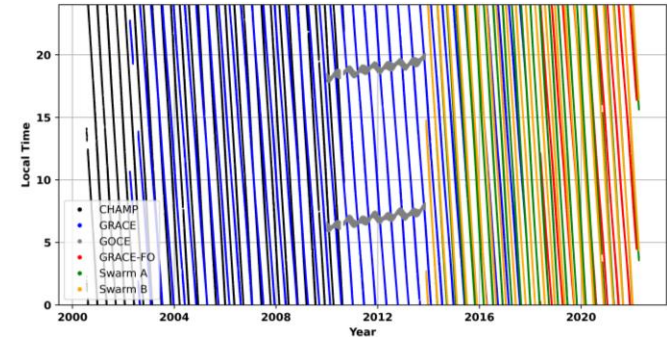
# Orbit characteristics

Extending Swarm topside observations with:

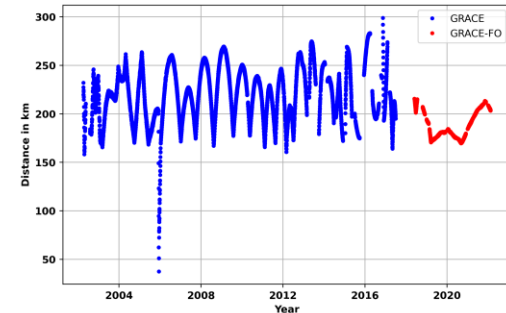
- GPS derived Total Electron Content (TEC)
  - CHAMP
  - GRACE A/B
  - GRACE-FO 1/2
- Electron density from inter satellite K-band
  - GRACE
  - GRACE-FO



Local time evolution



GRACE(-FO) inter-satellite distance



# TEC computation from dual-frequency GPS

Slant TEC is derived using the GPS carrier phase observations ( $L_1, L_2$ ) and calibrated using code observations ( $P_1, P_2$ ). Differential code biases (DCB) and multi-path corrections are applied to  $P_2 - P_1$

$$L_{gf} = L_1 - L_2$$

$$P_{gf}^{cal} = P_2 - P_1 + DCB_T + DCB_R + MP$$

Absolute slant TEC:

$$sTEC = K \cdot (L_{gf} + mn_{arc}(P_{gf}^{cal} - L_{gf}))$$

$$K \approx 40.3 \text{ m}^3 \text{ s}^{-2}$$

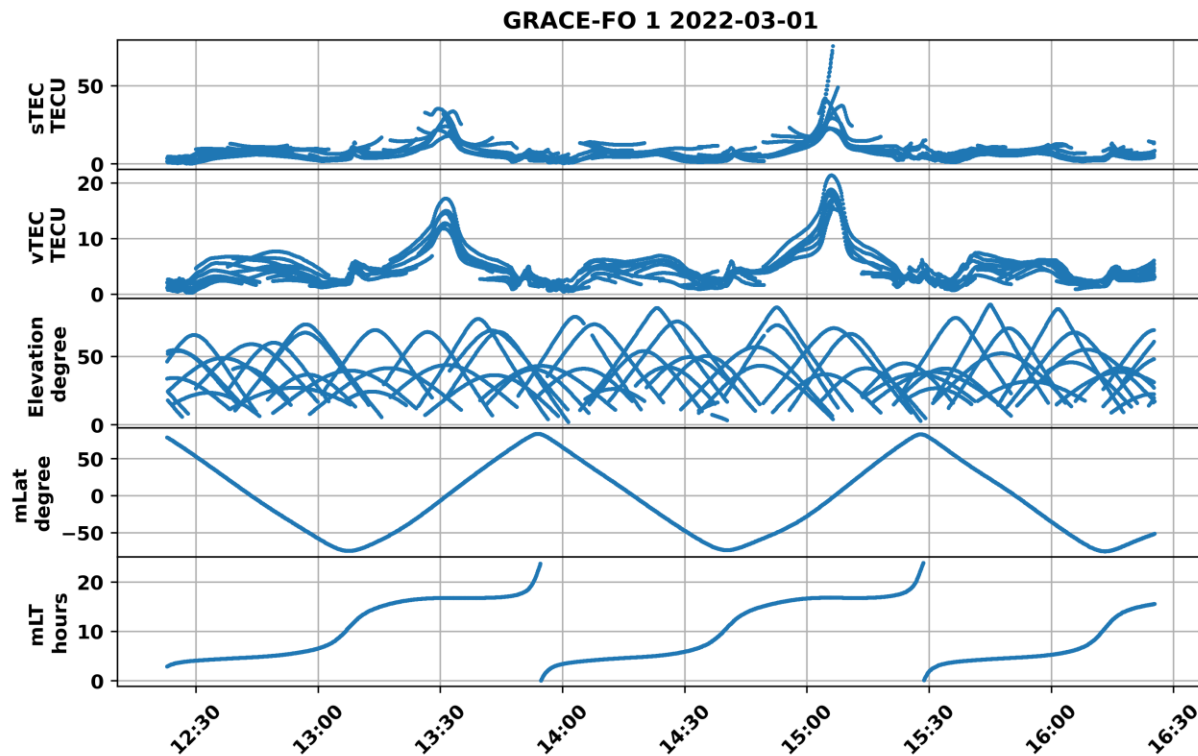
Vertical TEC is obtained using an elevation dependent mapping function  $M(\epsilon)$ .

The receiver DCB must be estimated. The following system is solved using least squares:

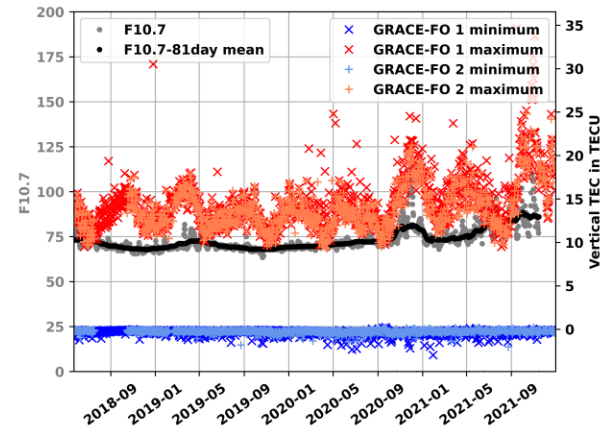
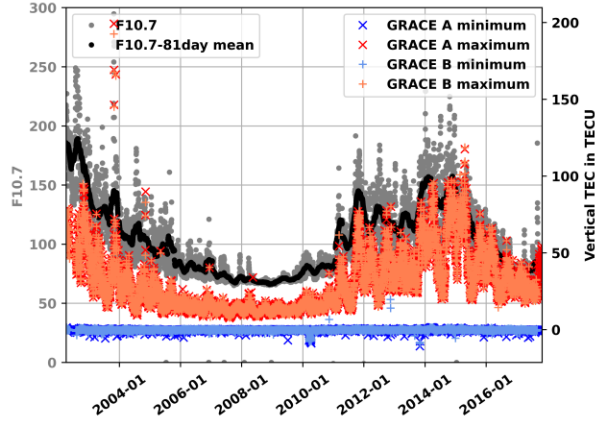
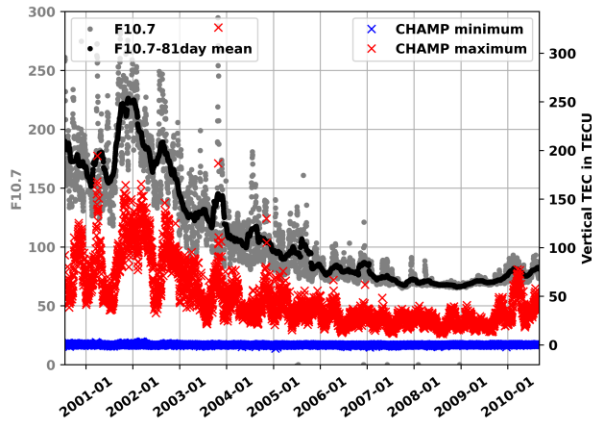
$$M(\epsilon_i) \cdot (sTEC_i + DCB_R) = M(\epsilon_j) \cdot (sTEC_j + DCB_R)$$

$L_i$ : phase measurement in m,  $P_i$ : code measurement in m,  $DCB_R$ : receiver  $P_1 - P_2$  DCB,  $DCB_T$ : transmitter  $P_1 - P_2$  DCB,  $MP$ : near-field multi-path,  $mn_{arc}$ : mean over a continuous arc of observations (no gaps, cycle-slips, or tracking issues)

# TEC time-series

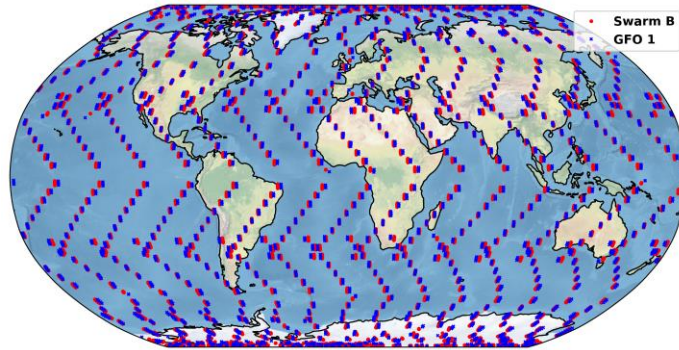


# Quality indication CHAMP



Correlation between F10.7 and max TEC (both as 81d mean):  
 0.93 (CHAMP), 0.81 / 0.82 (GRACE A / GRACE B), 0.74 / 0.71 (GRACE-FO 1 / GRACE-FO 2)  
 Local time dependent oscillations in TEC.

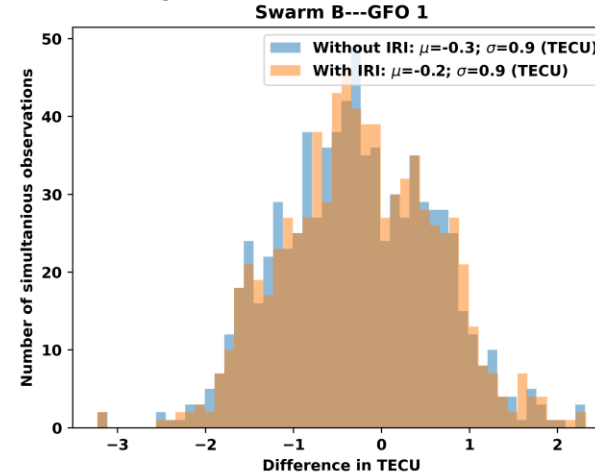
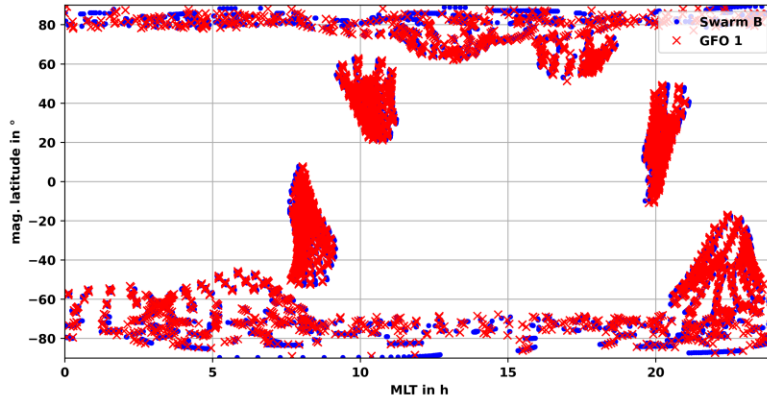
# TEC conjunctions



Swarm B and GRACE-FO 1 conjunctions:  
counterrotating December 2019. Less than 20 km  
difference in altitude.

Compensated for differences in TEC due to  
altitude differences using vertical TEC from IRI-  
2016.

Also serves for electron density validation  
(Swarm B Langmuir probes – GRACE-FO KBR)



# TEC conclusions

- Highly consistent time series
- Robust estimation of receiver DCB
- Excellent agreement during conjunctions with standard deviations below 2.4 TECU (below 1 TECU possible, if solar activity is low).

	CHAMP GRACE A	CHAMP GRACE B	GRACE A Swarm A	GRACE A Swarm B	GRACE-FO 1 Swarm A	GRACE-FO 2 Swarm B
Mean (TECU)	-0.3	-0.4	0.3	0.1	0.9	0.3
Std. (TECU)	2.4	2.1	1.1	1.0	1.2	0.7

# KBR electron density estimation

KBR provides dual-frequency phase measurements on K (24 GHz) and Ka (32 GHz) frequency. The ionospheric phase advance for Ka is stored in the  $Iono_{corr}$  variable

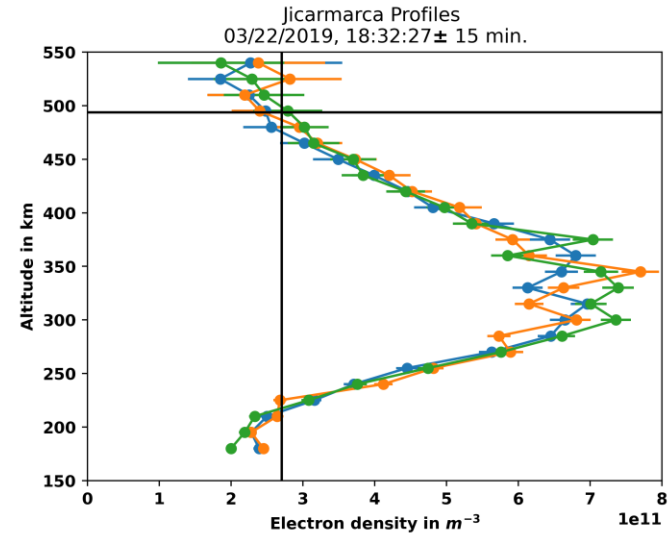
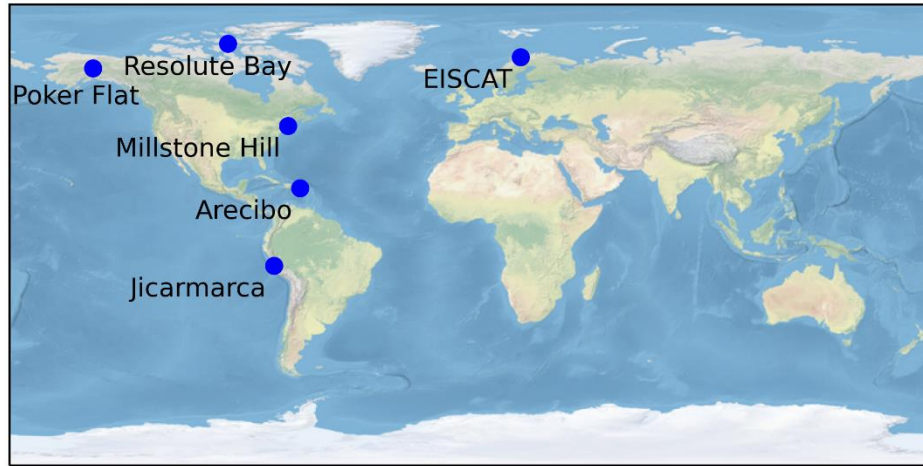
$$rTEC = Iono_{corr} \cdot \frac{f_{Ka}^2}{K}$$
$$K \approx 40.3 \text{ m}^3 \text{ s}^{-2}$$

Since the distance  $d$  between the satellites is only 200 km, this may be used to approximate the local electron density

$$rNe = rTEC / d$$

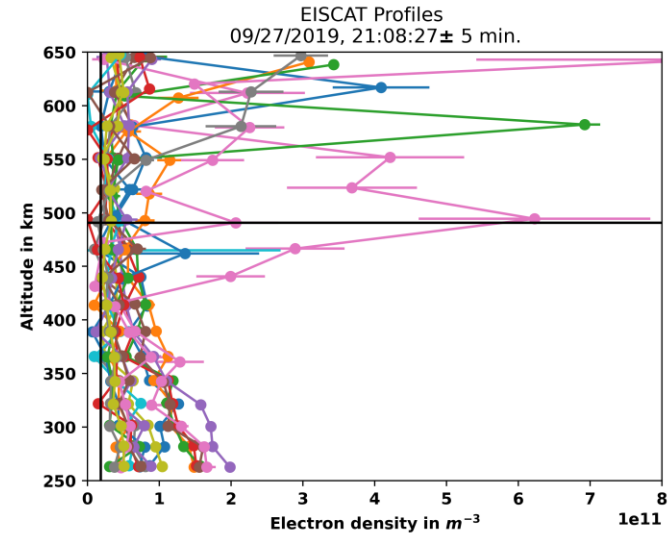
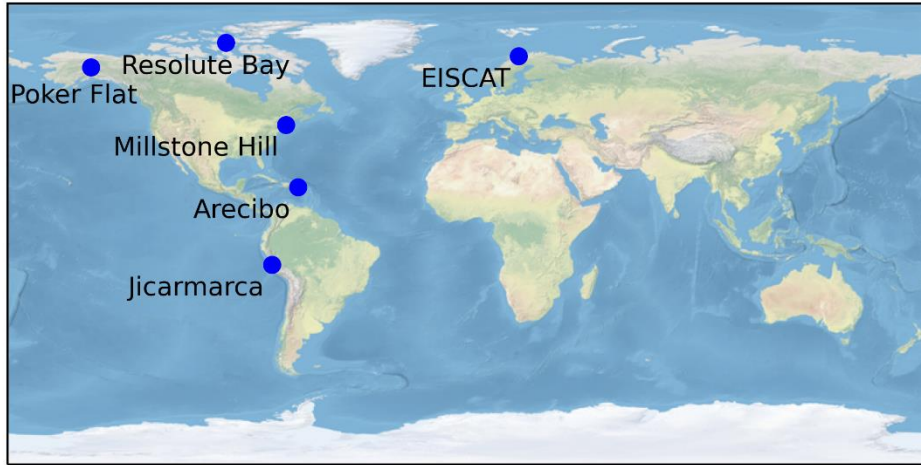
Absolut calibration needs to be provided externally (e.g. IRI-2016 model). Arbitrary offset in  $Iono_{corr}$  constant if KBR is not interrupted.

# KBR validation against radars



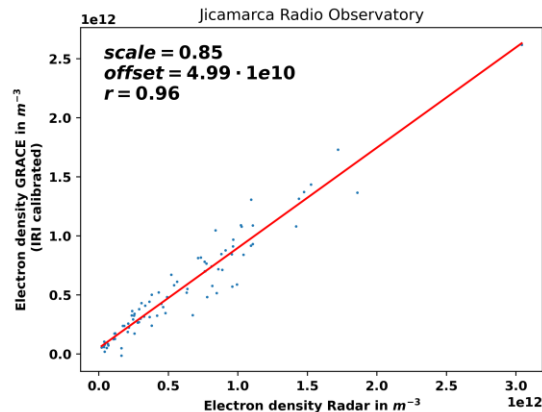
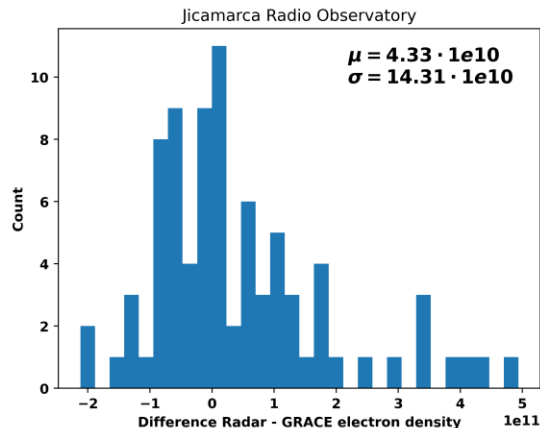
Criteria are  $\pm 5^\circ$  in lat/lon,  $\pm 15$  min, at least one radar observation within 20 km altitude of the satellite. Radar observations are interpolated between altitudinal sampling points using a quadratic function fitted on  $\log_{10}(Ne)$  on radar observations  $\pm 100$  km  $h_{sat}$ .

# KBR validation against radars



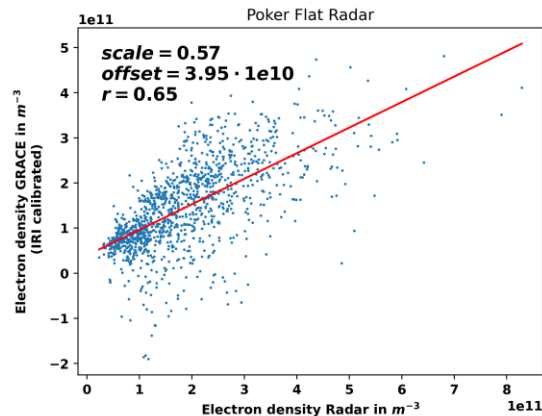
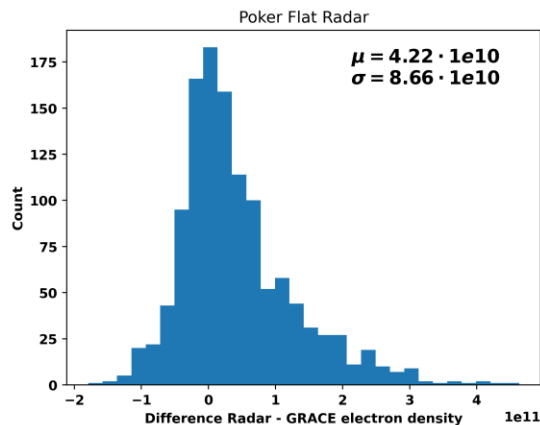
Criteria are  $\pm 5^\circ$  in lat/lon,  $\pm 15$  min, at least one radar observation within 20 km altitude of the satellite. Radar observations are interpolated between altitudinal sampling points using a quadratic function fitted on  $\log_{10}(Ne)$  on radar observations  $\pm 100$  km  $h_{sat}$ .

# KBR validation against radars



$\mu$ : mean calibration difference  
 $\sigma$ : standard deviation of calibration difference

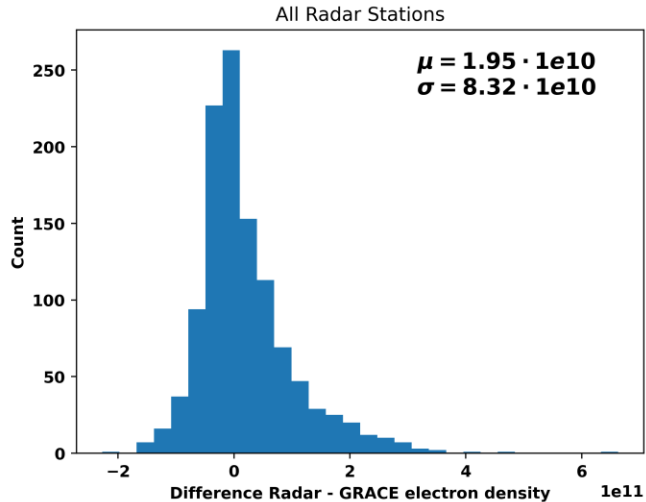
Best fit for low-latitude stations, e.g., Jicamarca.



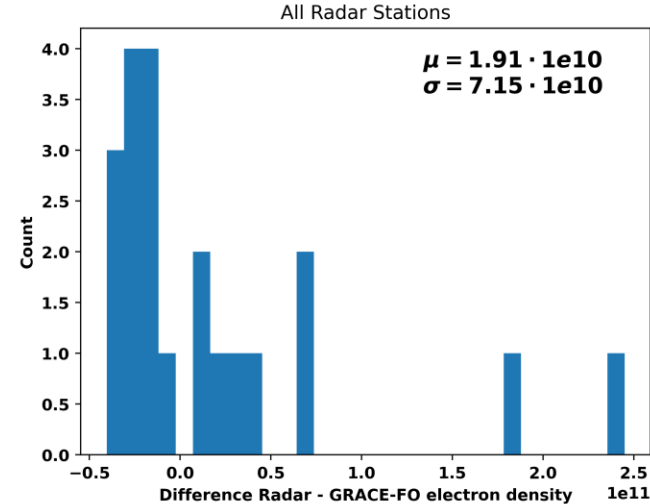
Challenging for near polar stations. Irregular profiles and large scatter in higher altitudes, e.g. Poker Flat.

# KBR validation against radars

GRACE



GRACE-FO



Small mean offset and standard deviation within the observational limits of the radar systems.

# Summary and conclusions

- TEC and electron density time series covering two solar cycles.
- Excellent agreement with Swarm data products.
- #days TEC: 3649, CHAMP  
5613, GRACE 1  
5552, GRACE 2  
1378, GRACE-FO 1 (until 31.3.2022)  
1207, GRACE-FO 2 (until 31.3.2022)
- #days Ne: 4879, GRACE  
1202, GRACE-FO (until 31.3.2022)
- Data available at <ftp://isdctp.gfz-potsdam.de>  
and soon via ESA at <http://swarm-diss.eo.esa.int>