

# Eruption and Interplanetary Evolution of a Stealthy Streamer-Blowout CME Observed by PSP at $\sim 0.5$ AU

Sanchita Pal<sup>1</sup>, Emilia Kilpua<sup>1</sup>, Simon Good<sup>1</sup>, Benjamin Lynch<sup>2</sup>, Erika Palmerio<sup>3</sup>,  
Eleanna Asvestari<sup>1</sup>, Jens Pomoell<sup>1</sup>, and Michael Stevens<sup>4</sup>

<sup>1</sup>University of Helsinki, Finland

<sup>2</sup>University of California, Berkeley, USA

<sup>3</sup>Predictive Science Inc., San Diego, CA 92121, USA

<sup>4</sup>Smithsonian Astrophysical Observatory, Cambridge, Massachusetts, USA

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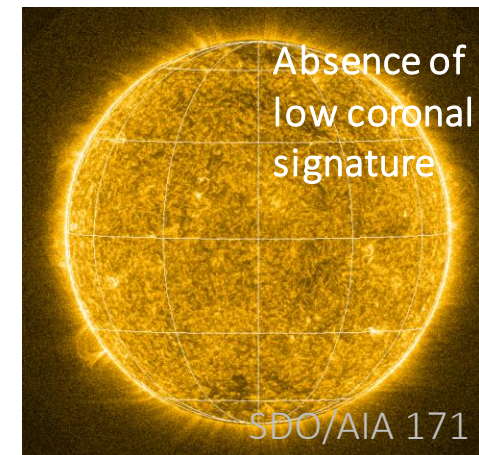
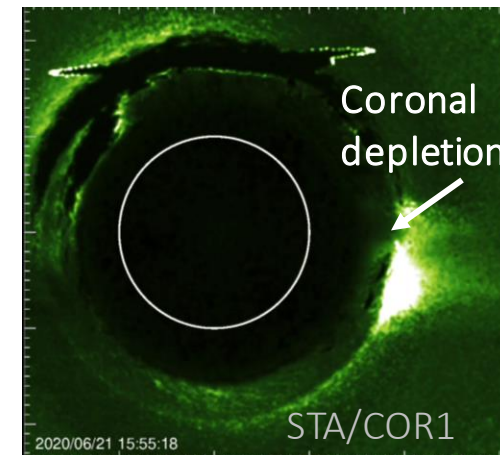
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Contact: [sanchita.pal@helsinki.fi](mailto:sanchita.pal@helsinki.fi)

# Solar observation of a **stealthy** Streamer blowout CME

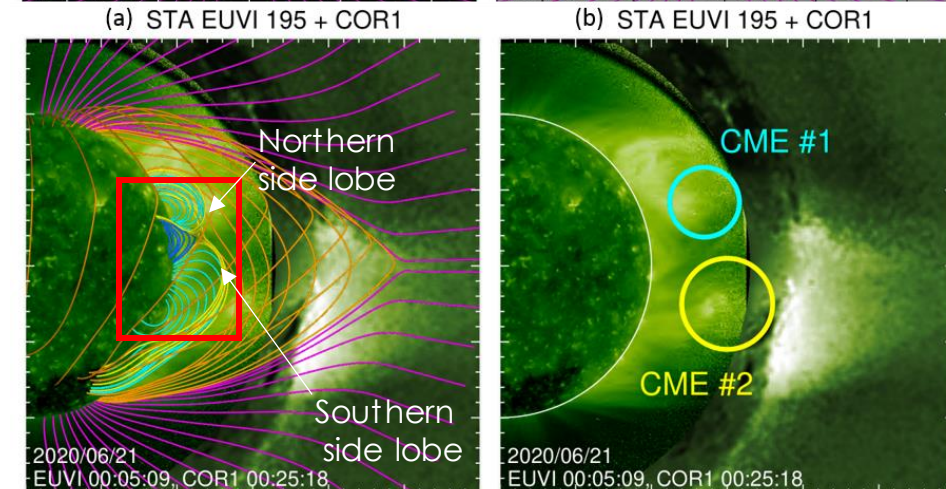
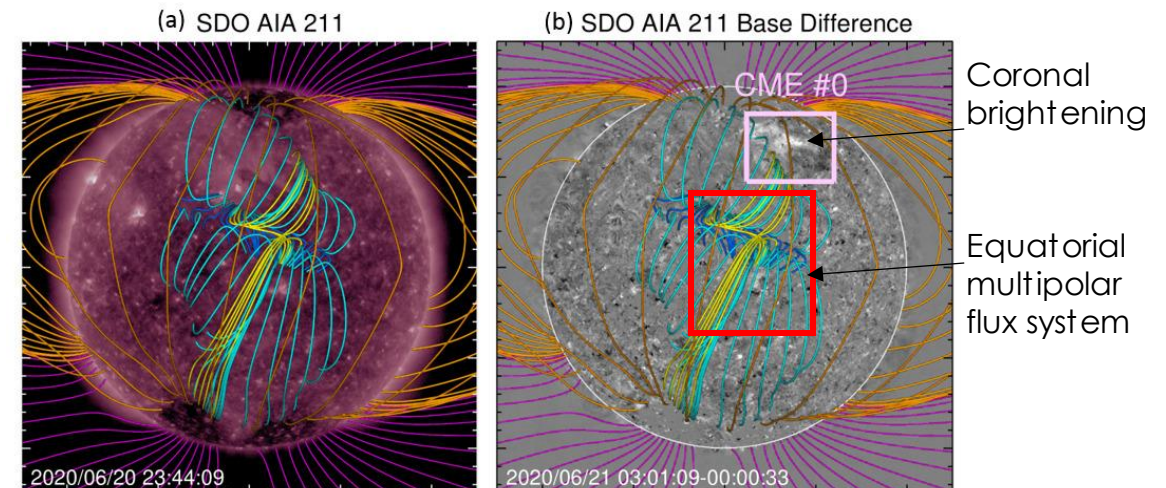
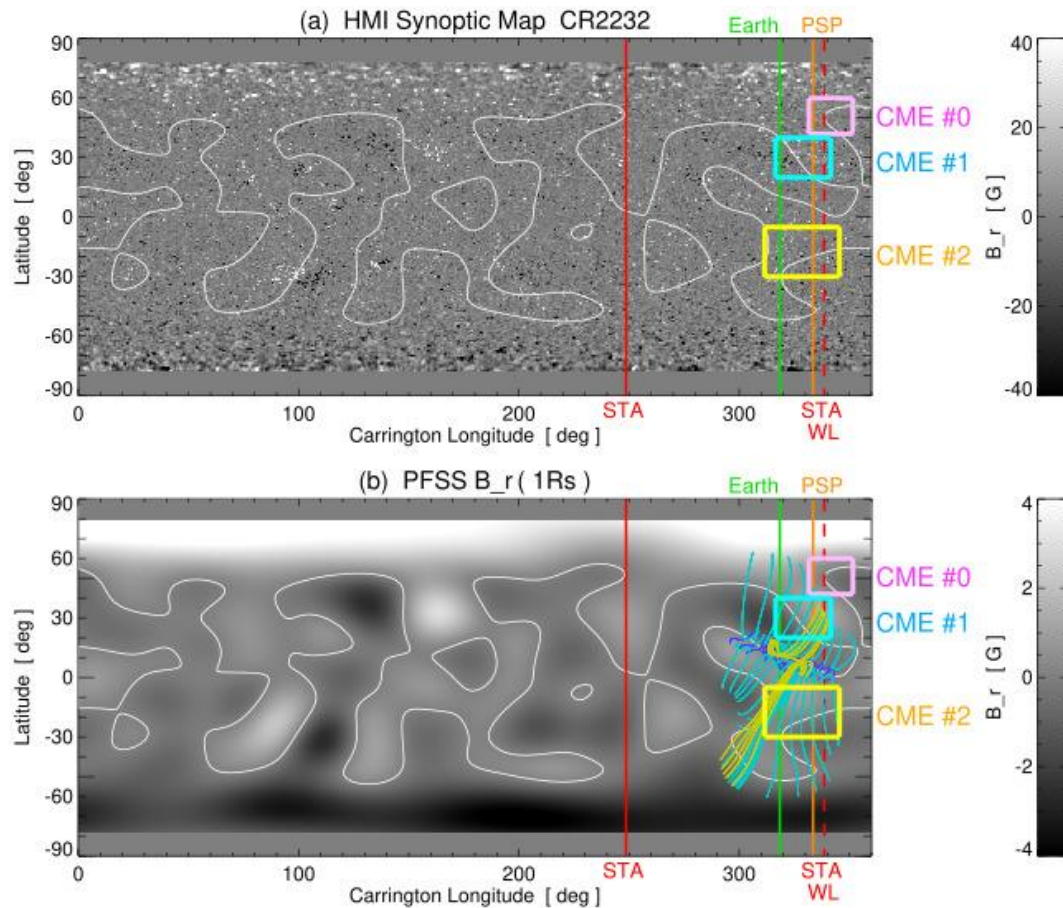
- In the solar corona prior to the event, the overlying streamer gradually swelled (STA/COR1 and COR2 observation shows)
- Afterwards a depletion region was detected in the solar corona (STA/COR1 observation shows)
- CME location mostly followed the tilt of the **heliospheric current sheet** (GCS and ForeCAT coronal modelling shows (Palmerio et al. 2021))
- The CME lacked classic low-coronal signatures (SDO/AIA observation shows)--stealth CME
- The CME appeared as a limb event with classical 3-part structure (STA/COR2 observation shows) on June 22, 2020





# Initiation of the SBO-CME preceded by sequential eruptions

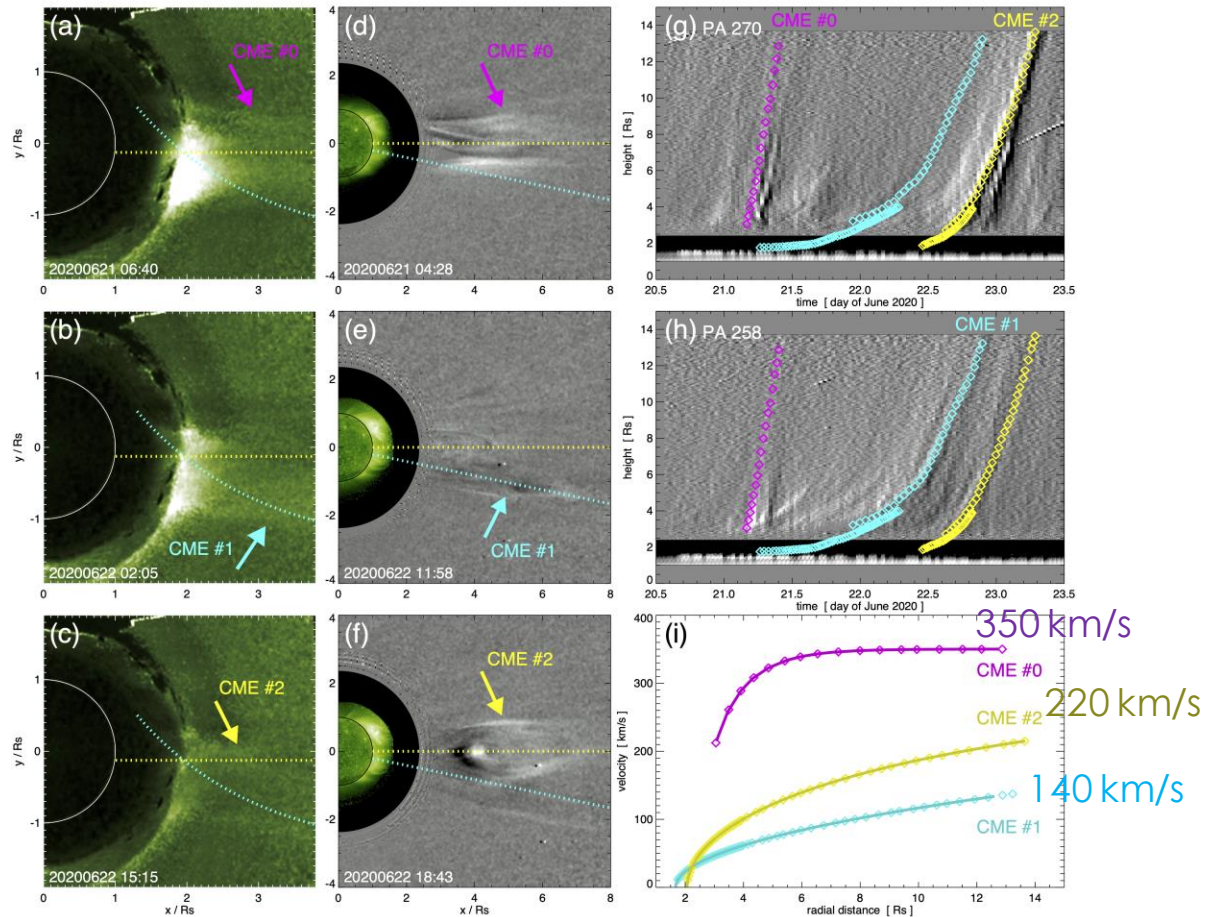
- Off-limb STA/EUVI, COR1, and COR2 imagery and the photospheric magnetic field extrapolation reveal that The SBO CME (CME #2) was preceded by two sequential eruptions -- CME#0 and CME#1.
- Source of CME #0 located outside and to the northwest of the equatorial multipolar flux system originating CME #1 and #2.



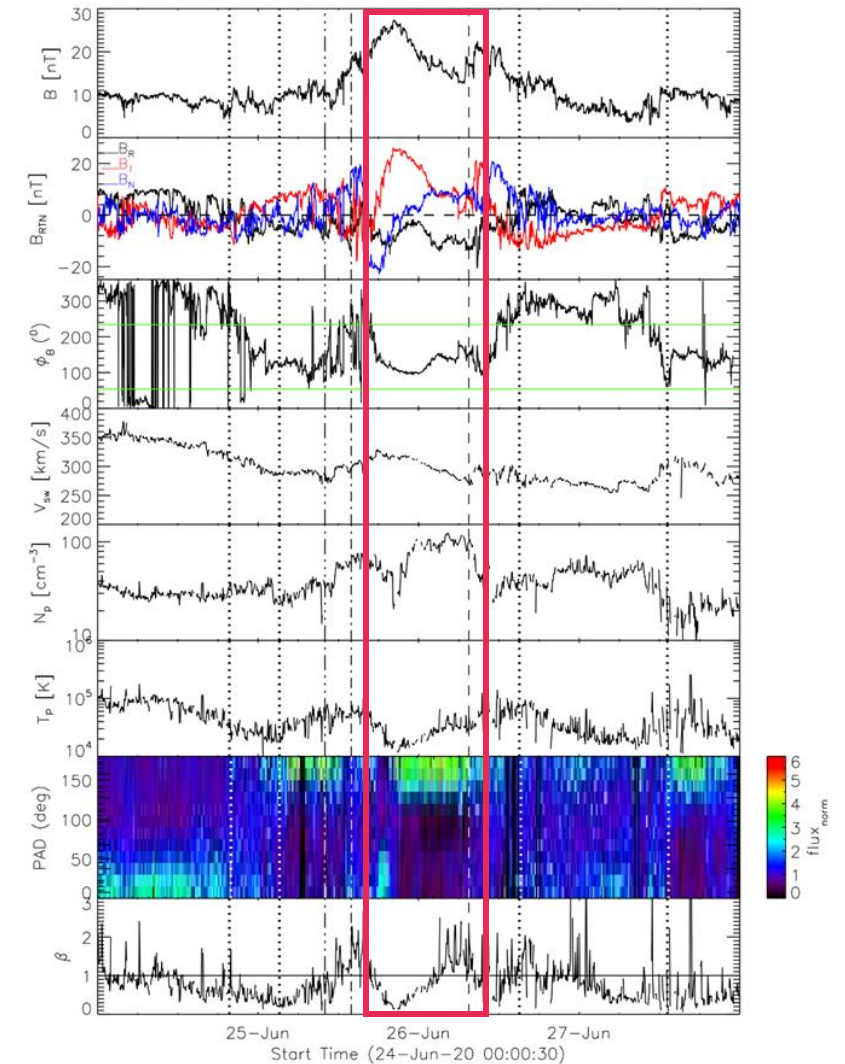


# Heliospheric observations of the sequential CMEs

- STA/HI observation shows CME #0 speed > CME #2 speed > CME #1 speed at 14 Rs.
- Only the SBO CME was identified with flux rope structure by the [Parker Solar Probe](#) at ~0.5 AU



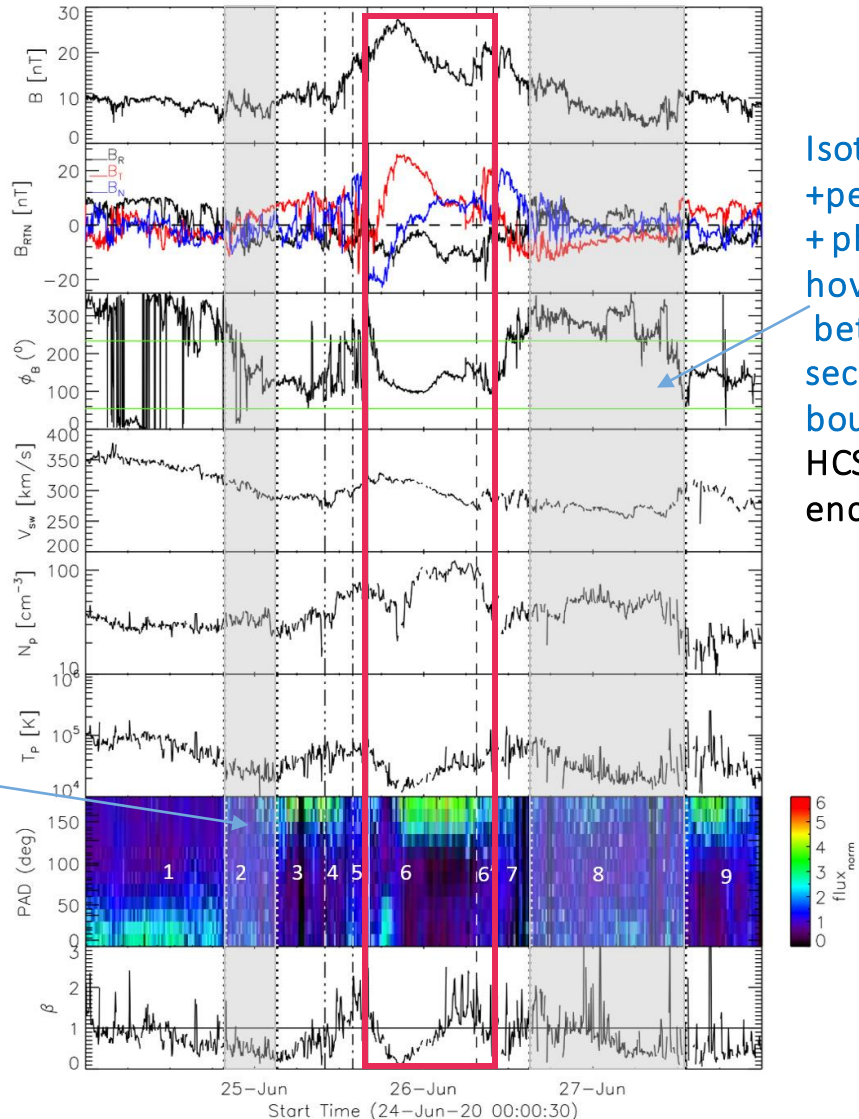
Coronal dynamics for each of the three sequential CMEs



PSP's In situ observation of the SBO-CME at ~0.5 AU

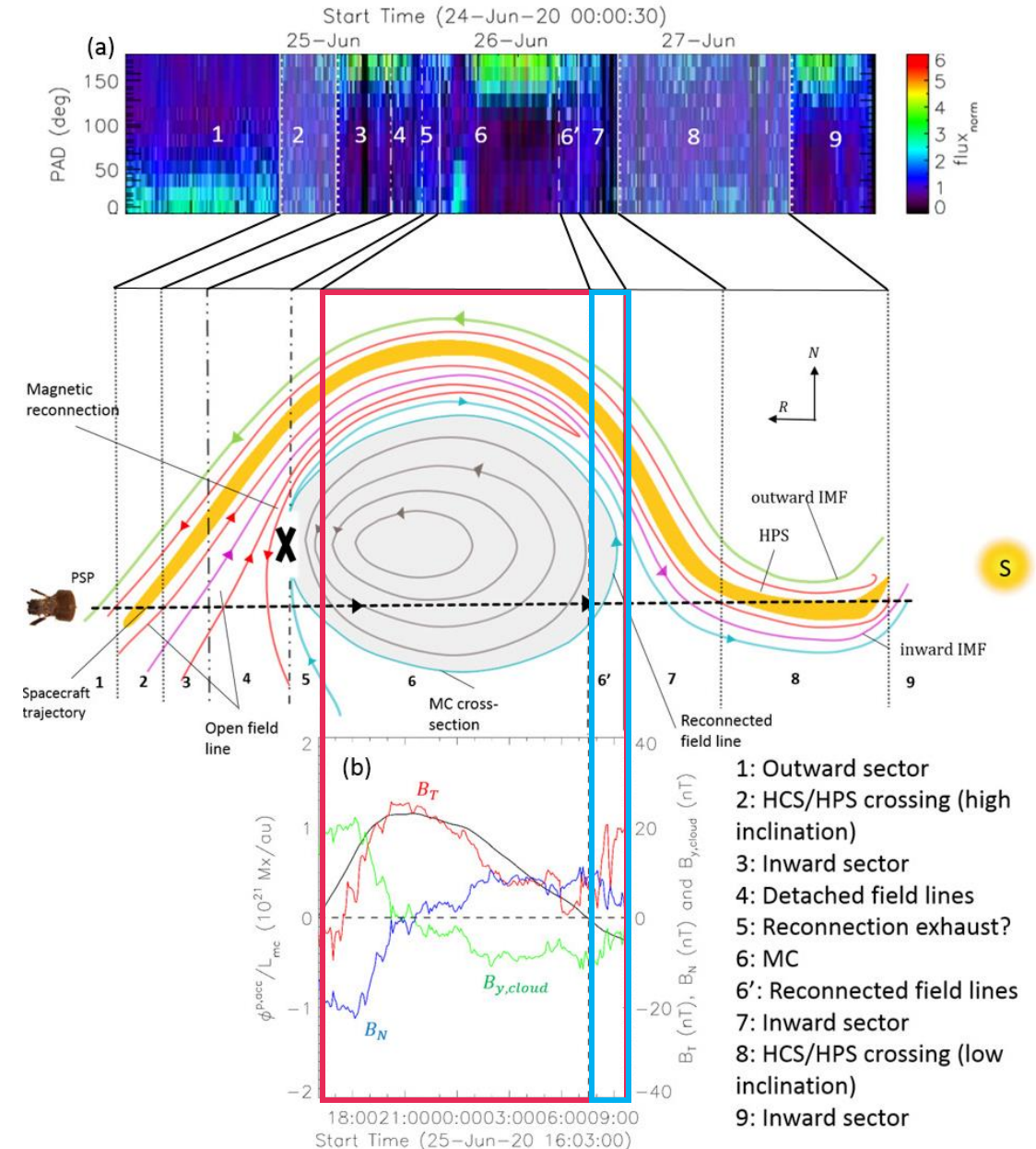
# SBO-CME adjacent interplanetary structures at ~0.5 AU

- In situ probing shows heliospheric large-scale structure – **heliospheric current sheet, plasma sheet** -- draped about the CME flux rope resulting in CME erosion.



Isotropic PAD  
+ peak in beta  
+ phi\_B  
hovering  
between  
sector  
boundary -->  
HCS/HPS  
encounter

True sector  
boundary +  
isotropic PAD  
+ elevated  
beta with  
duration  
of mins -->  
HCS/HPS  
crossing



Schematic of IMF draping and flux asymmetry analysis at ~0.5 AU



# Summary and Conclusions

- The SBO CME eruption was part of a multi-stage, **sequential (and most likely sympathetic)** eruption scenario (**Lynch and Edmondson 2013**).
- PSP witnessed the draping of heliospheric field lines and HCS/HPS about the SBO CME flux rope at 0.5 AU.
- Inclination ( $\sim 29^\circ$ ) of HCS behind the CME was smaller than the inclination ( $\sim 41^\circ$ ) of plasma sheet in front of the CME – the CME had an asymmetric, expanding, and non-circular FR structure.
- Draped heliospheric field lines had magnetic reconnection with the CME flux rope resulted in **erosion of  $\sim 18 \pm 11\%$  of the CME's azimuthal flux**.
- Analyzing the MC's back region populated with reconnected field lines, we estimated that the **reconnection initiated after a heliocentric distance of  $\sim 0.35$  AU**.

For more details on this work, please scan



Contact: [sanchita.pal@helsinki.fi](mailto:sanchita.pal@helsinki.fi)

This event has importance from several different perspectives for more details visit

- [Palmerio et al. 2021, ApJ, 109](#)
- [Möstl et al. 2022, ApJL, 924:L6 \(ST1.10 EGU22-1964\)](#)