

An innovative and automated vortex identification method based on the estimation of the center of rotation

Application to solar simulations

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What is a vortex? How do we identify it?



“A vortex consists in a multitude of material particles rotating around a common center”

Lugt, 1979

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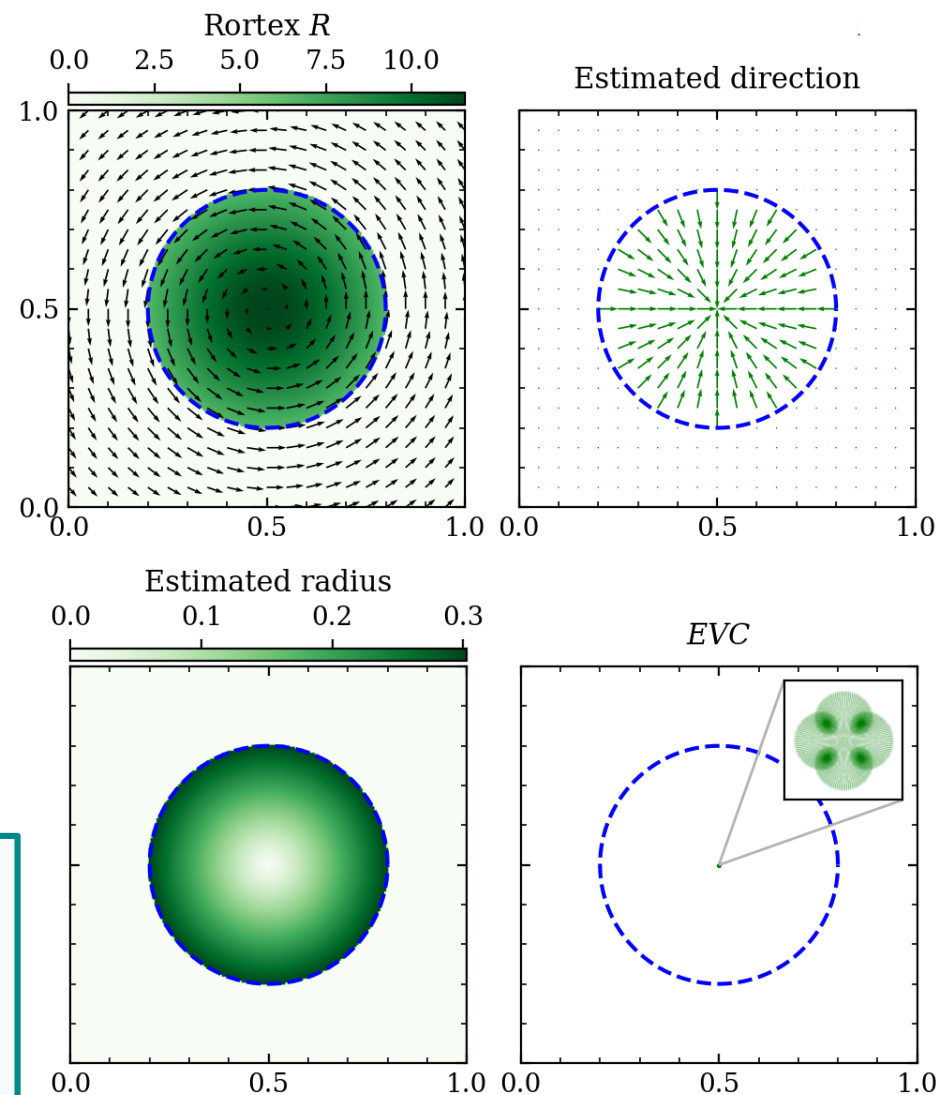


We merge the two strategies into one:

Estimated Vortex Center (EVC) Method

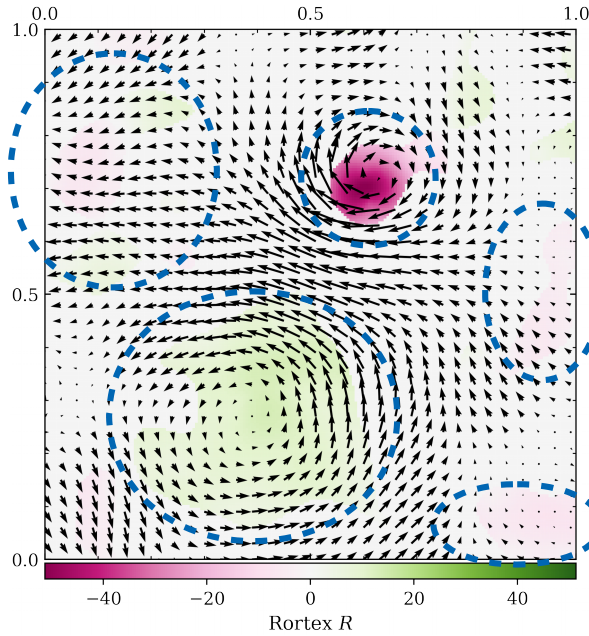
We estimate the **center of rotation** for every “rotating particle” (grid-cell) using the **Rortex criterion** (similar to vorticity) and the **flow characteristics**.

Vortices can then be identified by **clusters** of **Estimated Vortex Center (EVC)** points.



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Automated vortex detection with EVC method

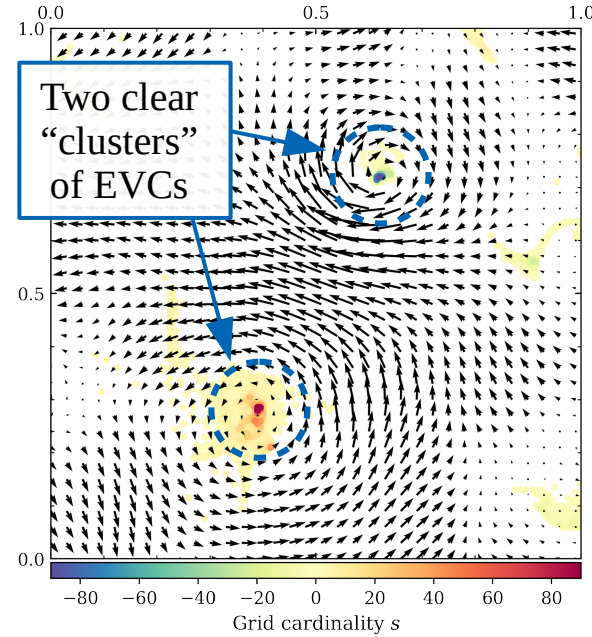
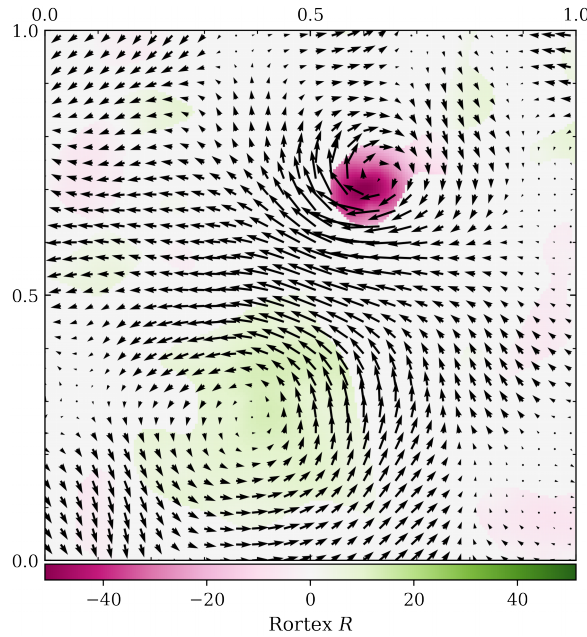


Many local
curvature
regions in the
velocity field

All vortices?
Noise?

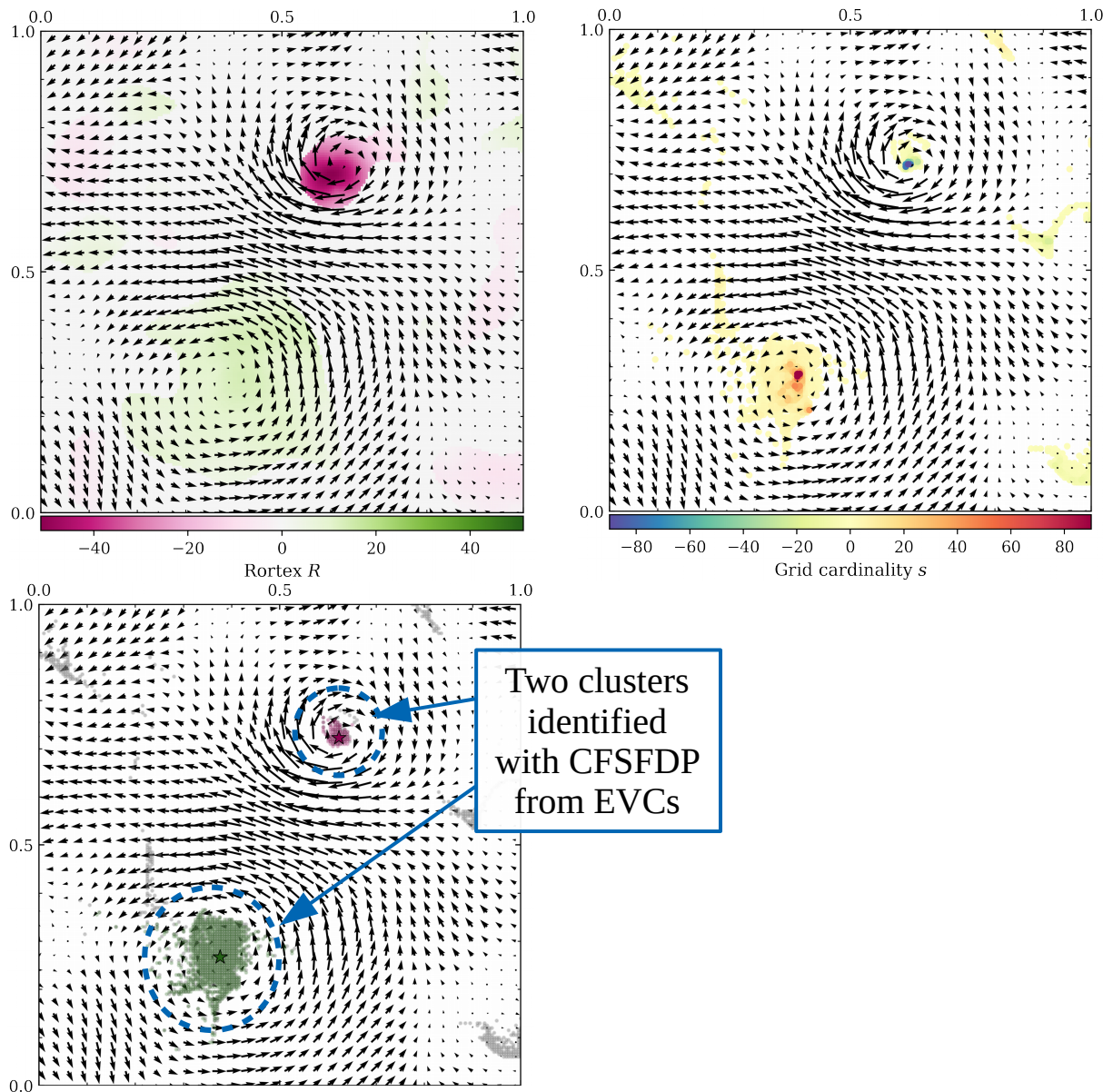
- ◆ Use the **Rortex criterion** (Tian et al., 2018) to identify **curvature** in local velocity field

Automated vortex detection with EVC method



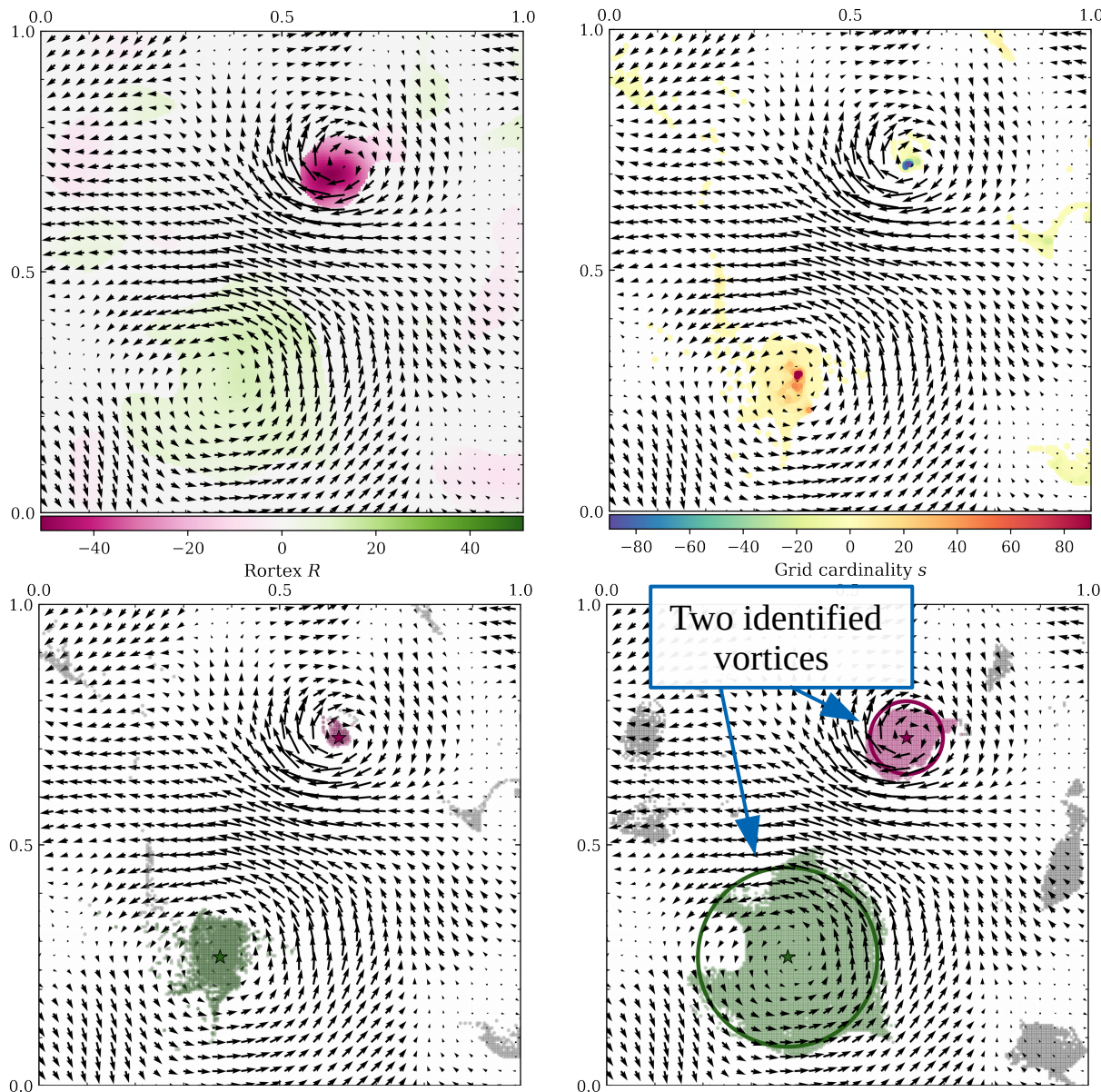
- ◆ Use the **Rortex criterion** (Tian et al., 2018) to identify **curvature** in local velocity field
- ◆ Use the **Rortex criterion** and the curvature of the **local velocity field** to estimate EVCs

Automated vortex detection with EVC method



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- ◆ **Cluster EVCs** with a grid-based adaptation of the Clustering by Fast Search and Find of Density Peaks (**CFSFDP**) algorithm (Rodriguez & Laio, 2014)

Automated vortex detection with EVC method



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- ◆ **Cluster** EVCs with a grid-based adaptation of the Clustering by Fast Search and Find of Density Peaks (**CFSFDP**) algorithm (Rodriguez & Laio, 2014)
- ◆ Each **EVC cluster** corresponds to a group of **grid cells** forming a **vortex**

➤ **Python implementation:**
<https://github.com/jcanivete/swirl>

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Application to numerical simulations



3D Radiative MHD numerical simulations of the solar atmosphere
(CO5BOLD code, [Freytag et al., 2012](#))

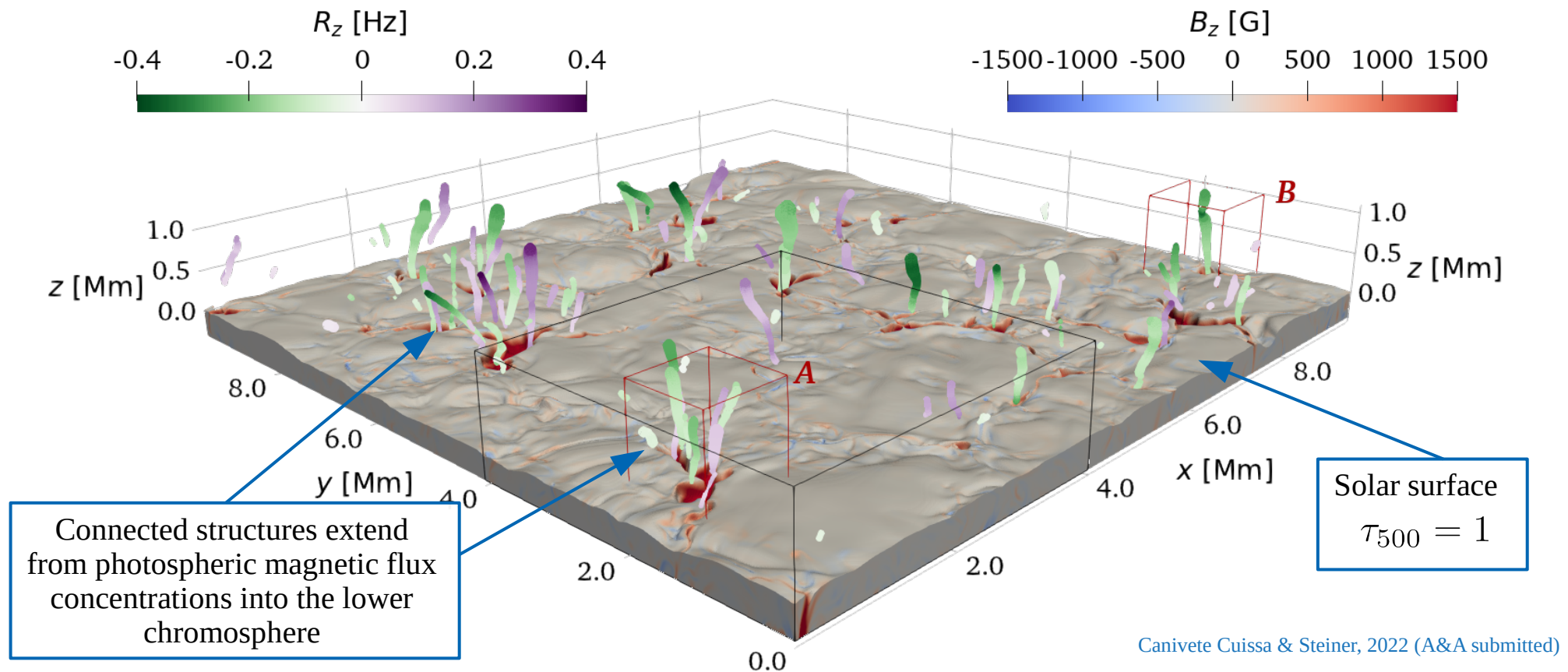
- ◆ $9.6 \times 9.6 \times 2.8 \text{ Mm}^3$ (960x960x280)
- ◆ Surface at $z \sim 1.2 \text{ Mm}$

Application to numerical simulations

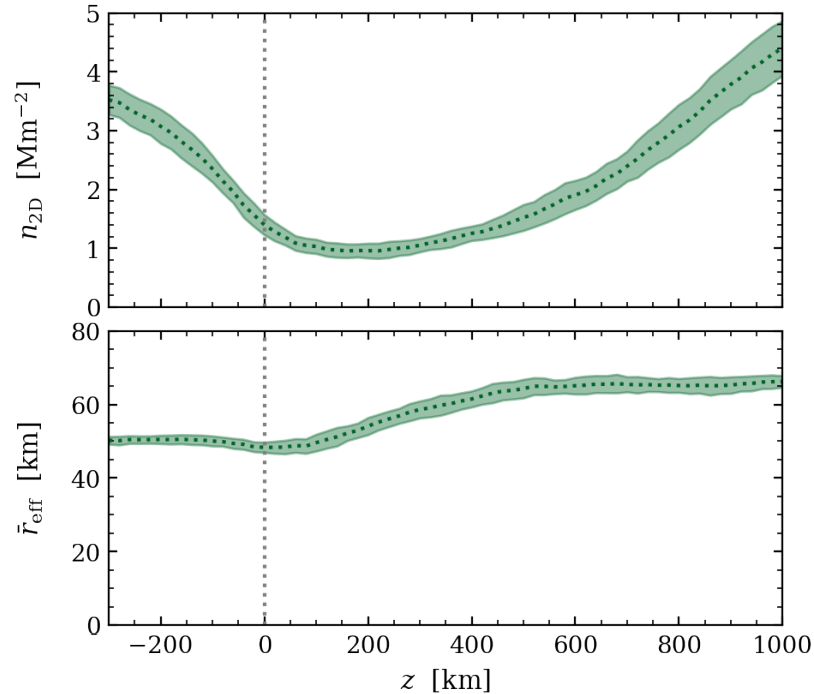


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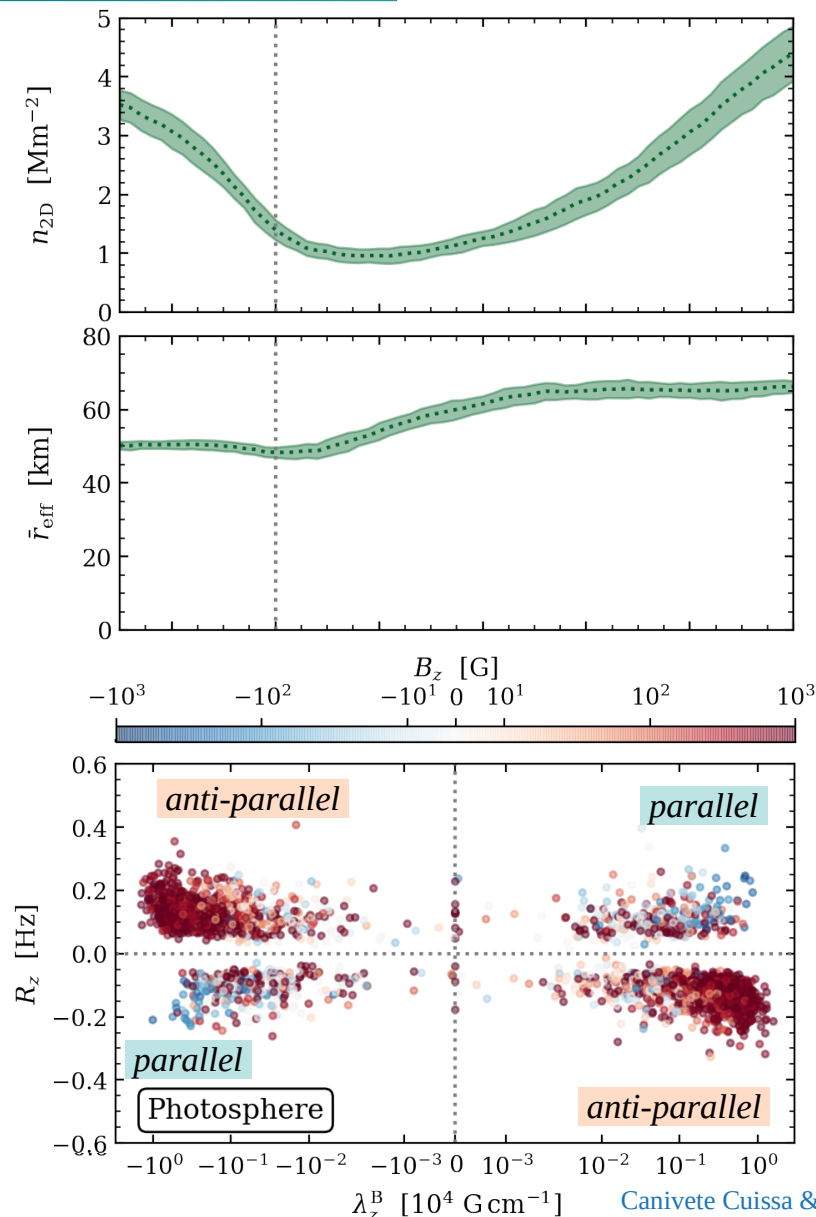
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 - **Smaller radii** ($\sim 50 \text{ km}$ in the photosphere)



- ◆ Statistics over 30 snapshots covering 2h of physical time show
 - **More swirls** (~ 1 Mm⁻² in the photosphere)
 - **Smaller radii** (~ 50 km in the photosphere)
 - **Photospheric swirls** correlate with **twists** in the **magnetic field**
 - **$\sim 80\%$** of photospheric swirls and magnetic field twists are compatible with **Alfvénic pulses** (see Battaglia et al., 2021)

$$\text{sign}(R_z \lambda_z^B) = -\text{sign}(B_z)$$

Plasma swirl

Magnetic field lines twist

Magnetic flux concentration

Conclusions



- ◆ The **EVC method** combines mathematical criteria and morphological methods to **accurately identify vortices** from 2D velocity fields
- ◆ A **publicly available** Python implementation is available
<https://github.com/jcanivete/swirl>
- ◆ Application to **CO5BOLD simulations** of the solar atmosphere shows:
 - **More swirls** with **smaller radii** than previous studies
 - Perturbations in the magnetic field compatible with **Alfvénic pulses**
- ◆ For more details:
Canivete Cuissa & Steiner (2022), A&A, submitted

Thank you for your attention