

Tools for sharing and evaluating the CHAOS geomagnetic field model and the shc-file format for time-dependent spherical harmonic models

Clemens Kloss

Together with: Chris Finlay and Nils Olsen

DTU Space, National Space Institute, Technical University of Denmark

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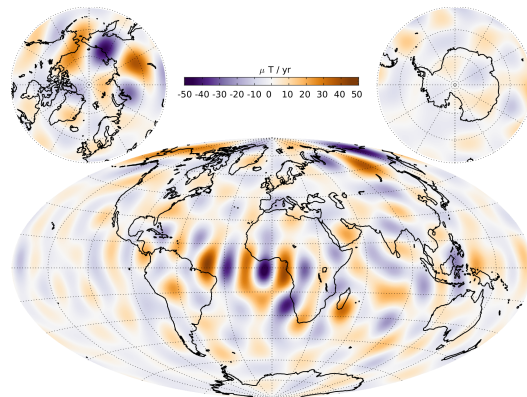


Introduction

- Geomagnetic field models are essential in the study of physical processes that contribute to Earth's magnetic field.
- Several groups build models: differences in magnetic data and mathematical methods used during model estimation (e.g. need not be spherical harmonic models).
- Choice between models depends on geophysical signals of interest.
- Currently no single platform for collecting models in standardised form.
- Here: details about the CHAOS geomagnetic field model and its distribution to users.

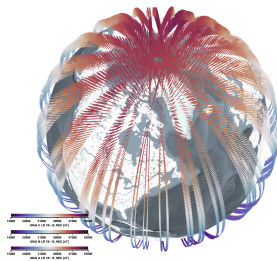
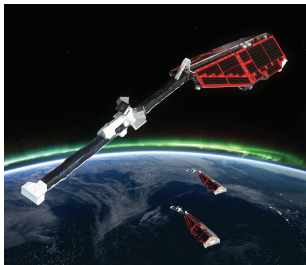
CHAOS - A Geomagnetic field Model

- Derived approx. every 4 months by the Division of Geomagnetism and Geospace at DTU Space using the latest ground and satellite data.
- Provides estimates of the recent internal field (generated in the Earth's core and crust) and the quiet-time external field (generated in the radiation belts, magnetotail and at the magnetopause).
- Current version covers the period between 1997 and 2022 (CHAOS-7.10).



Rate of change of the radial magnetic field in 2021 for $n \leq 16$ at the core-mantle boundary as given by CHAOS-7.9.

Magnetic Observations Used for Deriving CHAOS



Left: Artistic view of the three *Swarm* satellites in orbit. Right: Field intensity (color scale) of magnetic field measurements collected by the *Swarm* satellites over 18th October 2018 (images: [ESA](#) and [VirES](#)).

- Satellite magnetic survey missions are crucial by providing global observations of high temporal and spatial resolution.
- ESA's *Swarm* constellation consists of three identical satellites presently collecting high quality magnetic data.
- Also satellite data from previous survey missions (Ørsted, CHAMP, SAC-C) and platform magnetometer data (Cryosat-2).
- And monthly-means from a network of approximately 200 ground stations

Overview of Tools for Sharing and Evaluating CHAOS

Main repository of files and software is the [CHAOS website](#).

Model Coefficient Files

- Model coefficients stored in MATLAB's mat-format.
- Static internal field as shc-file (see following slides).
- Time-dependent internal field model in spline-coefficients format and as shc-file.

Computer Software for Evaluating CHAOS

- MATLAB code package (entire model).
- Python code package: ChaosMagPy (entire model, see next slide).
- Fortran-77 code for producing time series of Gauss coefficients and derivatives (int. field).

ChaosMagPy

- Open-source Python package for evaluating CHAOS and **many other** time-dep. spherical harmonic models
- Code can be viewed and downloaded on github.com.
- Extensive documentation available on readthedocs.org.



ChaosMagPy homepage.

```

1  chaos = load_CHAOS_matfile('CHAOS-7.10.mat')  # load latest CHAOS model from mat-file
2
3  dgnm = chaos.synth_coeffs_tdep(time, nmax=16, deriv=1)  # SV coefficients for n <= 16
4  B = chaos.synth_values_tdep(time, radius, theta, phi)  # time-dep. internal field
5
6  cals7k = load_CALS7K_txtfile('CALS7K.2')  # load CALS7K model file, usage like CHAOS

```

Python example code.

Introducing the SHC-file Format

- Developed for storing time-dependent spherical harmonic models.
- Adopted by the European Space Agency (ESA) for *Swarm* Level 2 and auxiliary product files (i.e. spherical harmonic field models that may be time-dependent).

Advantages

- ASCII file readable without installation of special software.
- Model snapshots in terms of Gauss coefficients can be directly read and used (no B-spline reconstruction).
- Provides enough information to properly reconstruct the continuous time-dependence of spline models.

Disadvantages

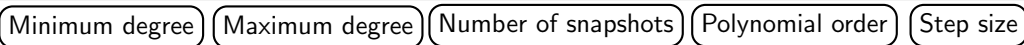
- Fixed time sampling unless full reconstruction is performed.
- Sensitive to the precision of the used floating-point format (since not specified in the shc-file definition).
- Lack of official reader/writer software.

SHC-file Layout Description

- Snapshot models at time instants t_i (decimal years) stored column-wise.
- Gauss coefficients stored row-wise in natural order ($g_1^0, g_1^1, h_1^1, g_2^0, \dots$).
- Step size S indicates extra snapshots needed for reconstructing the polynomial pieces and, hence, the continuous time-dependence.

# Comment line part: an arbitrary number of # lines starting with the hash key.					
n_{\min}	n_{\max}	N	K	S	
		t_1	t_N
n_{\min}	0	$g_{n_{\min}}^0(t_1)$	$g_{n_{\min}}^0(t_N)$
n_{\min}	1	$g_{n_{\min}}^1(t_1)$	$g_{n_{\min}}^1(t_N)$
n_{\min}	-1	$h_{n_{\min}}^1(t_1)$	$h_{n_{\min}}^1(t_N)$
...
n	m	$g_n^m(t_1)$	$g_n^m(t_N)$
...
n_{\max}	$-n_{\max}$	$h_{n_{\max}}^{-n_{\max}}(t_1)$	$h_{n_{\max}}^{-n_{\max}}(t_N)$

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...
n	m	$g_n^m(t_1)$	$g_n^m(t_N)$
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Summary and Conclusions

- SHC-file format developed for exchanging time-dependent field models of various types in the context of the *Swarm* mission.
- ChaosMagPy package developed primarily for the CHAOS geomagnetic field model.
- Nonetheless, both are usable more broadly for time-dependent spherical harmonic field models.
- Recommendation for strengthening cross-disciplinary sharing of models:
 - ① Provide open-source software for reading/writing model files of any format and for evaluating models (field values, Gauss coefficients, etc.).
 - ② Provide simple online tools for producing time series of Gauss coefficients and derivatives (downloadable, e.g., as csv-file).

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Thank you for your attention

