

Proton precipitation in a hybrid-Vlasov simulation with southward interplanetary magnetic field driving: First 3D results

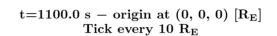
Maxime Grandin

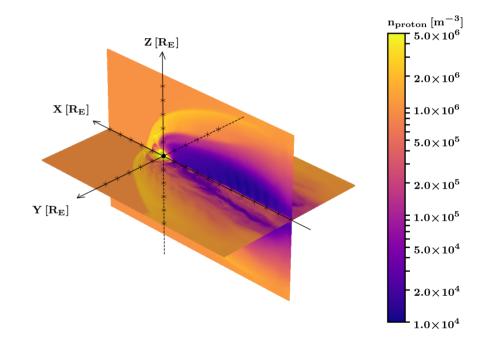
and the Vlasiator team (PI: Prof. Minna Palmroth)
University of Helsinki, Finland

EGU General Assembly – 23–27 May 2022

Vlasiator 6D run

- Vlasiator is a hybrid-Vlasov model of near-Earth space (ions as velocity distributions, electrons as fluid)
 - Vlasov equation for ions
 - Maxwell's equations for electromagnetic fields
 - Closure with Ohm's law incl. Hall term and $abla P_e$
- Same 6D run as described in the earlier talk by Minna Palmroth (EGU22-3573)
 - Southward IMF driving, 5 nT
 - $V_{\rm SW}$ = 750 km/s, $n_{\rm SW}$ = 1 cm⁻³
 - Inner boundary at 4.7 $R_{\rm E}$
 - Run duration: 1506 s
- Objective: Examine auroral ($E \lesssim$ 30 keV) proton precipitation in this first 6D simulation

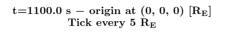


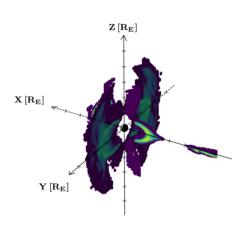


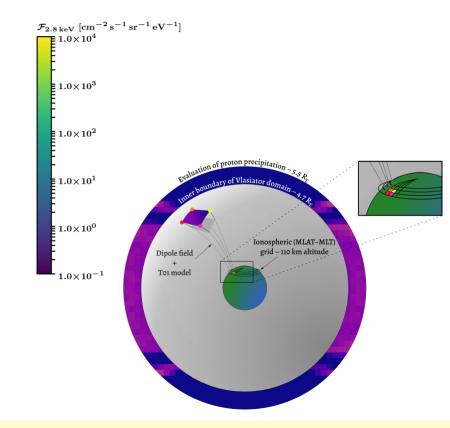


Precipitation data in Vlasiator

- From proton VDF, the differential number flux of precipitating protons is calculated (9 energy bins, 0.5–50 keV)
- The flux near the inner boundary is mapped to an ionosphere grid (MLAT-MLT) in both hemispheres
- From this differential flux, one can derive the integral energy flux and mean precipitating energy

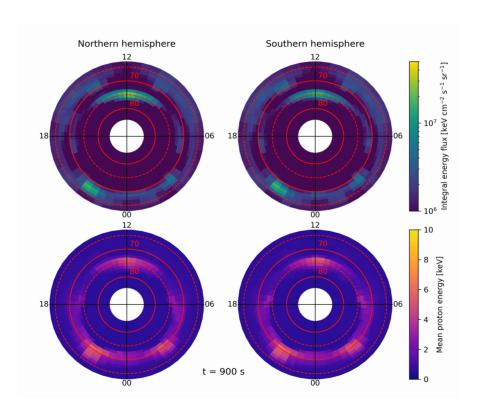








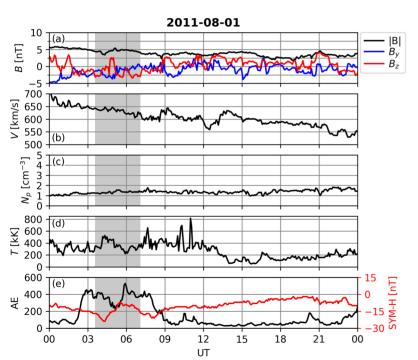
Global proton precipitation patterns

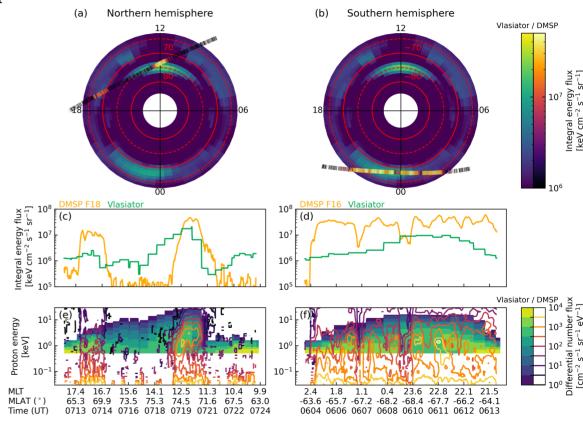




Comparison with satellite observations

 DMSP/SSJ observations during an event with comparable driving conditions







Key points and teaser

- First 3D simulation of auroral proton precipitation with Vlasiator (hybrid-Vlasov)
- Global precipitation patterns show interhemispheric symmetry on the nightside, not on the dayside
 - consistent with cusp precipitation being associated with flux transfer events transiting in it
- Comparison with DMSP observations indicate that the obtained fluxes are realistic

- Many analysis possibilities,
 e.g. keograms across MLT at selected MLATs
- Manuscript to be submitted soon!

