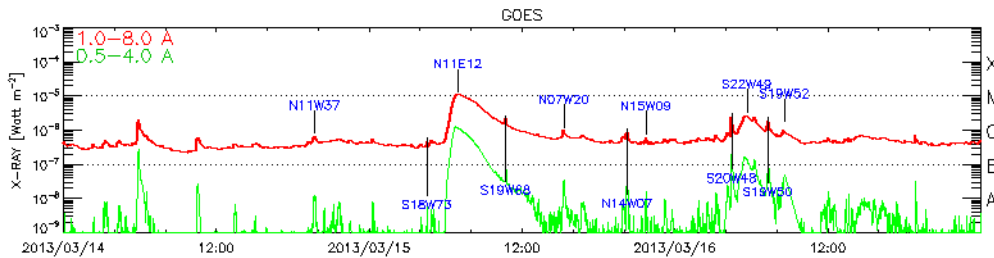


Analysis of the CME and associated gradual SEP event of March 2013

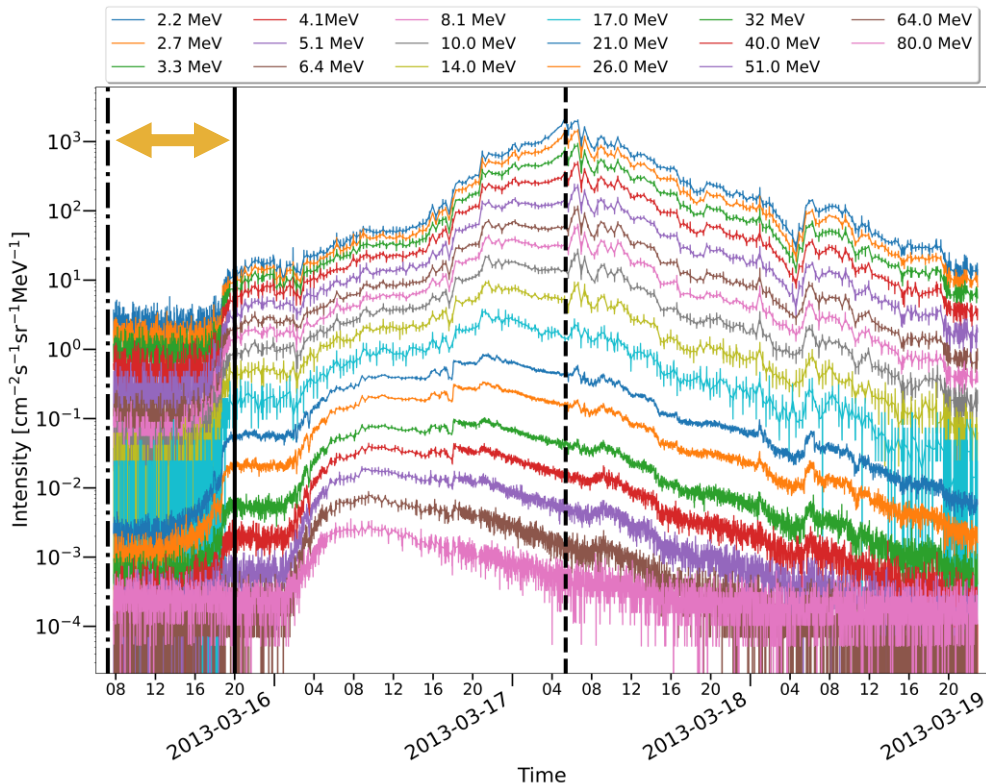
Antonio Niemela, Nicolas Wijzen, Angels Aran, Luciano Rodriguez, Jasmina Magdalenic, Stefaan Poedts



Event Analysis



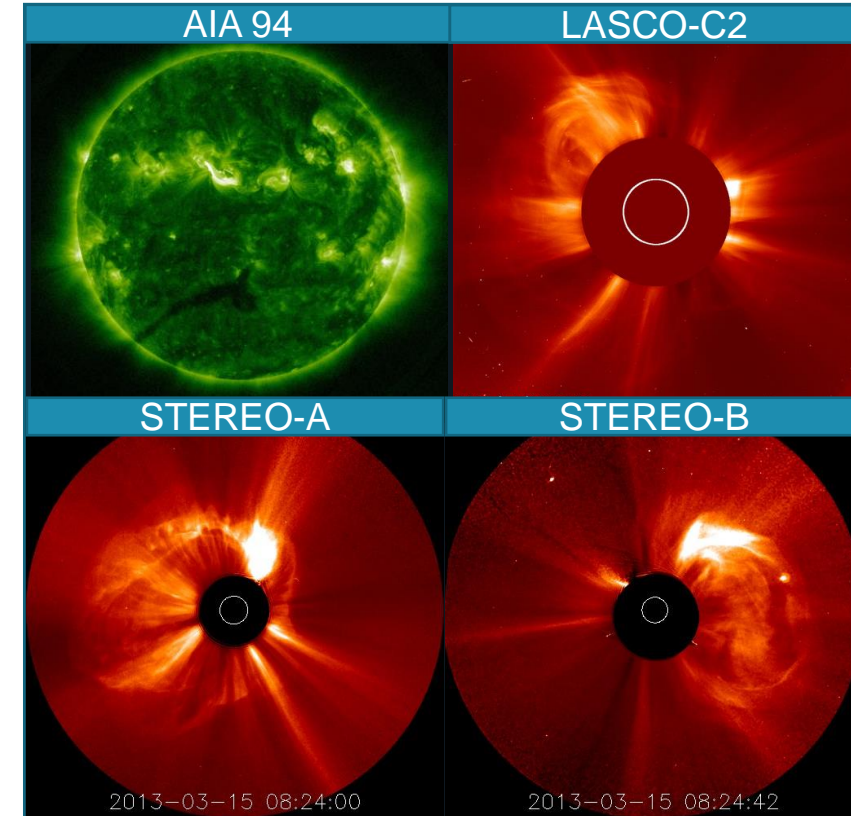
Event starts with a long duration M1.1 X-ray flare seen on the 15 March 2013 at 5:46UT.



Enhanced proton flux starting at 20:00UT.

Particles arrive at Earth with a 12-hour delay.

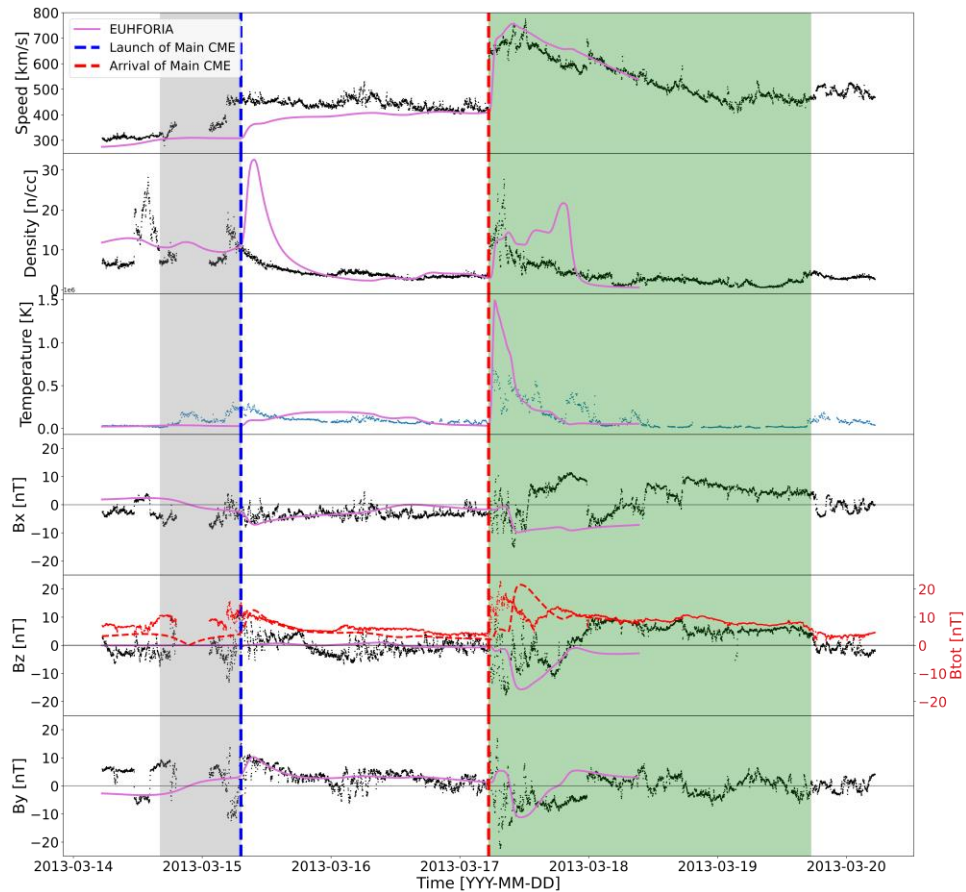
Intensities have different characteristics.



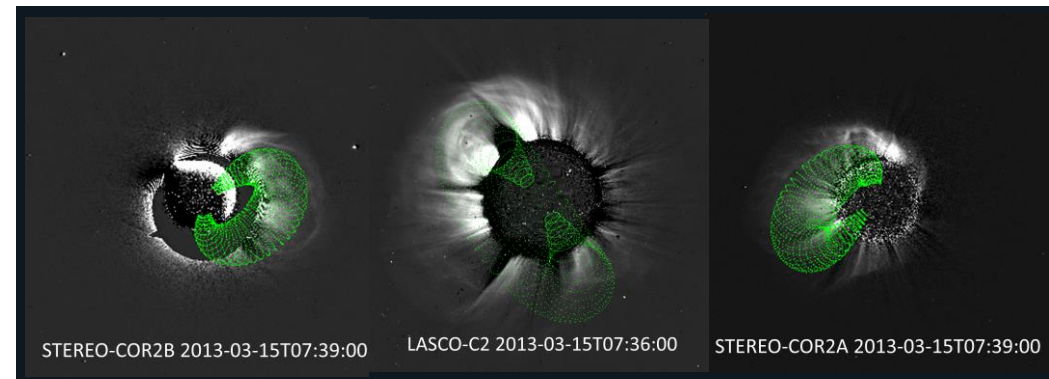
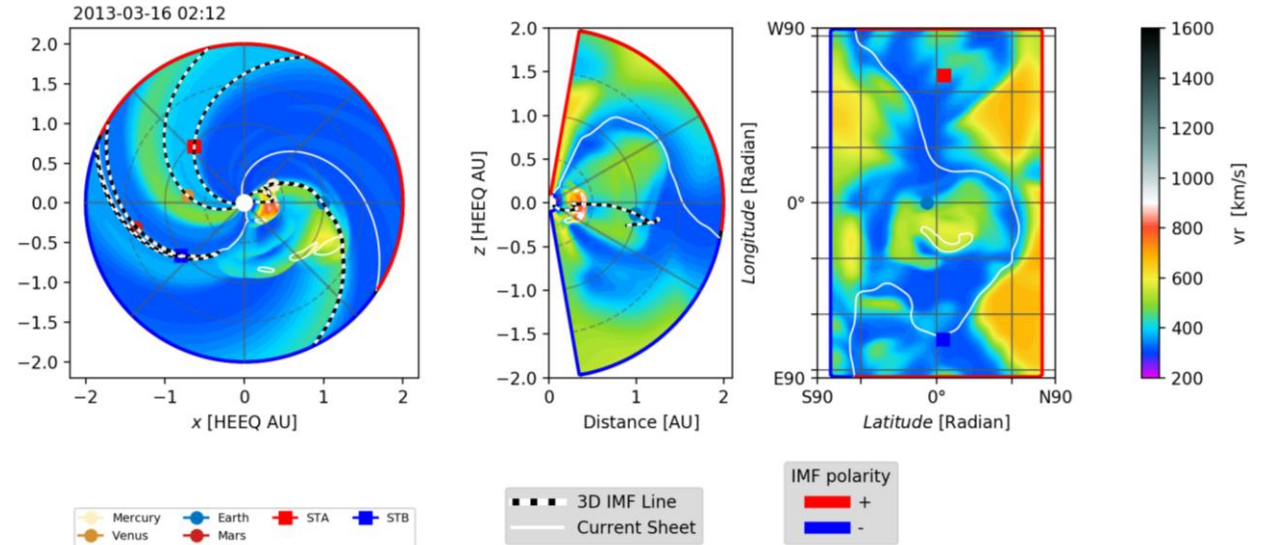
The Halo CME is captured by LASCO-C2 at 7:12UT with a small northward component seen with STEREO-A and B.

Simulations

EUHFORIA



Very active period, with 2 preceding CMEs

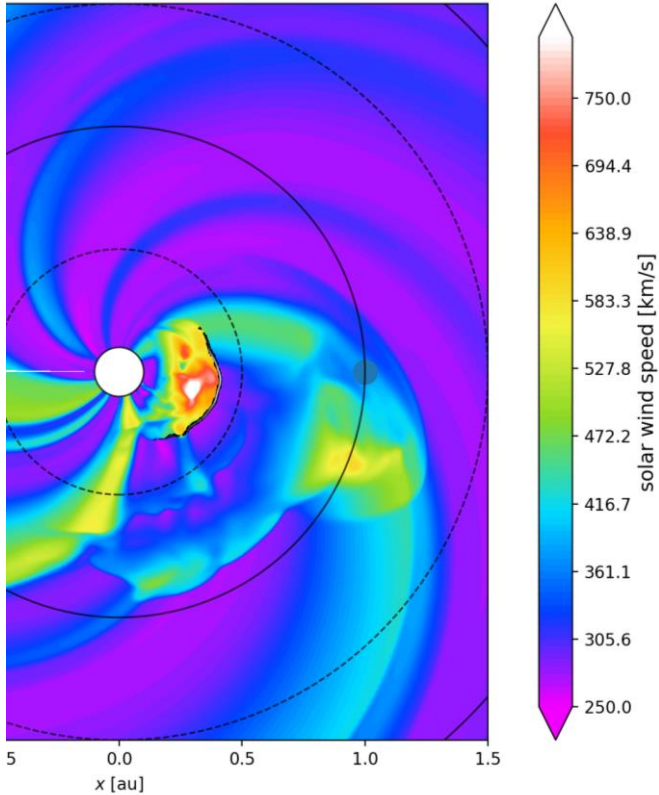


GCS reconstruction

Simulations

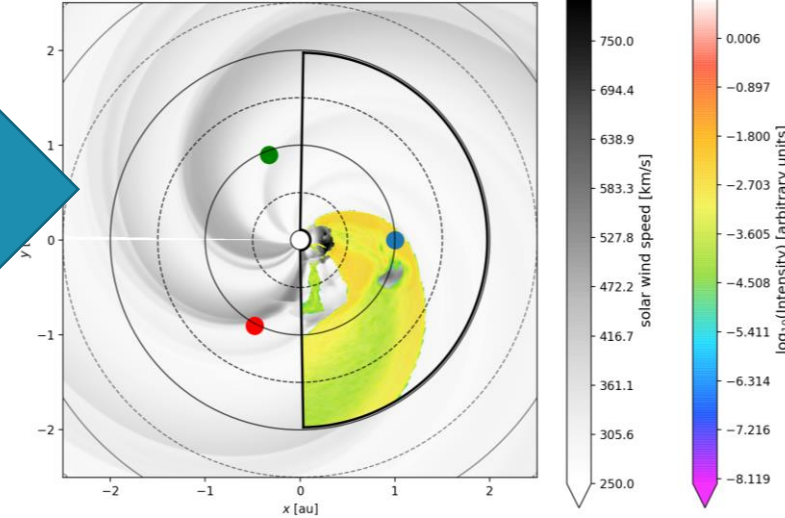
PARADISE

2013-03-16 02:14:30



Particle Injection

2013-03-16T00:59:30 Energy: 2.00 - 2.40 MeV

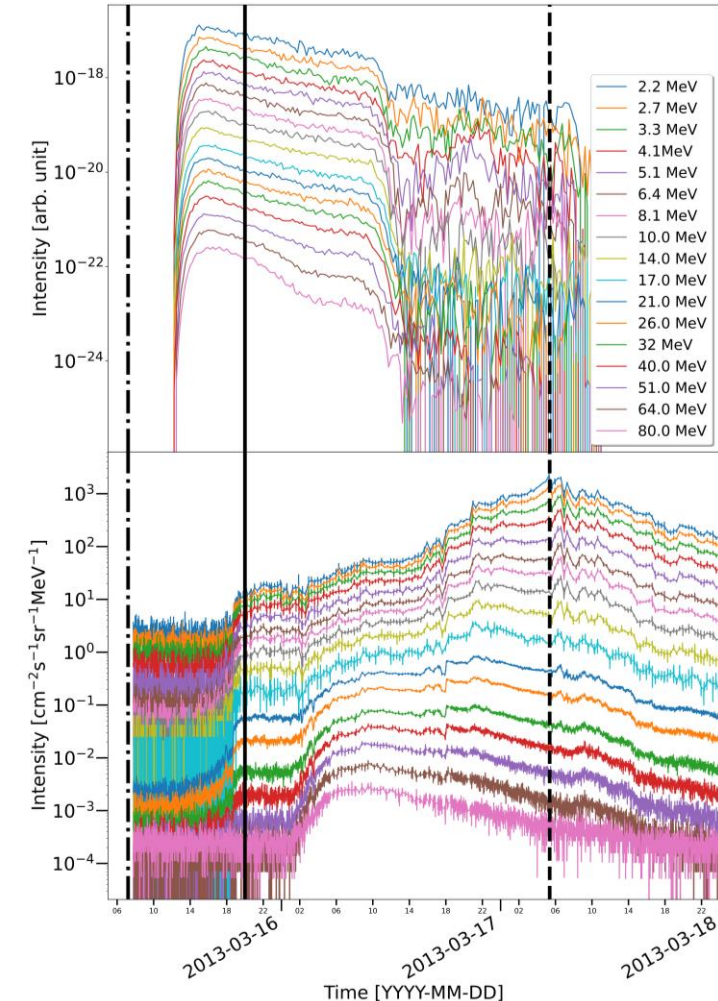


PARADISE captures a delay of 6 hs in the particle arrival

Intensities drop too early to “background” levels.

Blockade CME misses Earth

Shock tracing method does a good job at delimiting shocks.



Thank you for your attention

Summary

- EUHFORIA capture the background solar wind very accurately.
 - 3 spheromak CMEs injected into the domain.
- Shock tracing methods delimit the shock generated by the Main CME → Particles are injected on top of it to simulate the real shock acceleration.
- PARADISE captures a delay in the arrival of particles.
 - The preceding CME acts as a blockade for particles, but not close enough to Earth to generate an extra delay.

Future Work

- Better description of the particle population for more accurate predictions.
- Get the intermediate CME to affect Earth and quantify its effect on delaying particles
- Publication already in preparation.



euhforia

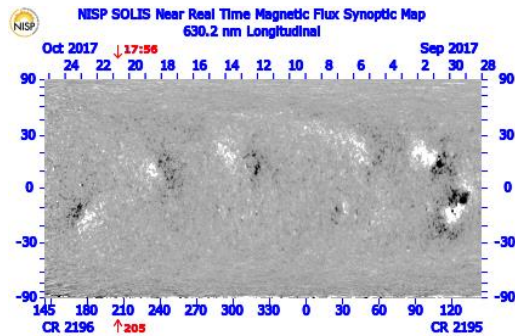
KU LEUVEN

EXTRA INFORMATION

References

- GCS model: Graduated Cylindrical Shell Model:
 - Thernisien, A., Vourlidas, A. & Howard, R.A. Forward Modeling of Coronal Mass Ejections Using STEREO/SECCHI Data. *Sol Phys* 256, 111–130 (2009). <https://doi.org/10.1007/s11207-009-9346-5>
- CCSOM constraining CMEs and Shocks by Observation and Modelling:
 - <https://www.bis.sidc.be/ccsom/>
- EUHFORIA: EUropean Heliospheric Forecasting Information Asset:
 - <https://www.euhforia.com>
 - Pomoell J, Poedts S. 2018. EUHFORIA: European heliospheric forecasting information asset. *J. Space Weather Space Clim.* 8: A35
- PARADISE: Particle Radiation Asset Directed at Interplanetary Space Exploration:
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- Spheromak Model for CMEs:
 - Gibson, S. E. and Low, B. C., “A Time-Dependent Three-Dimensional Magnetohydrodynamic Model of the Coronal Mass Ejection”, *The Astrophysical Journal*, vol. 493, no. 1, pp. 460–473, 1998. doi:10.1086/305107.

EUHFORIA basics



J. Pomoell and S. Poedts: J. Space Weather Space Clim. 2018, 8, A35

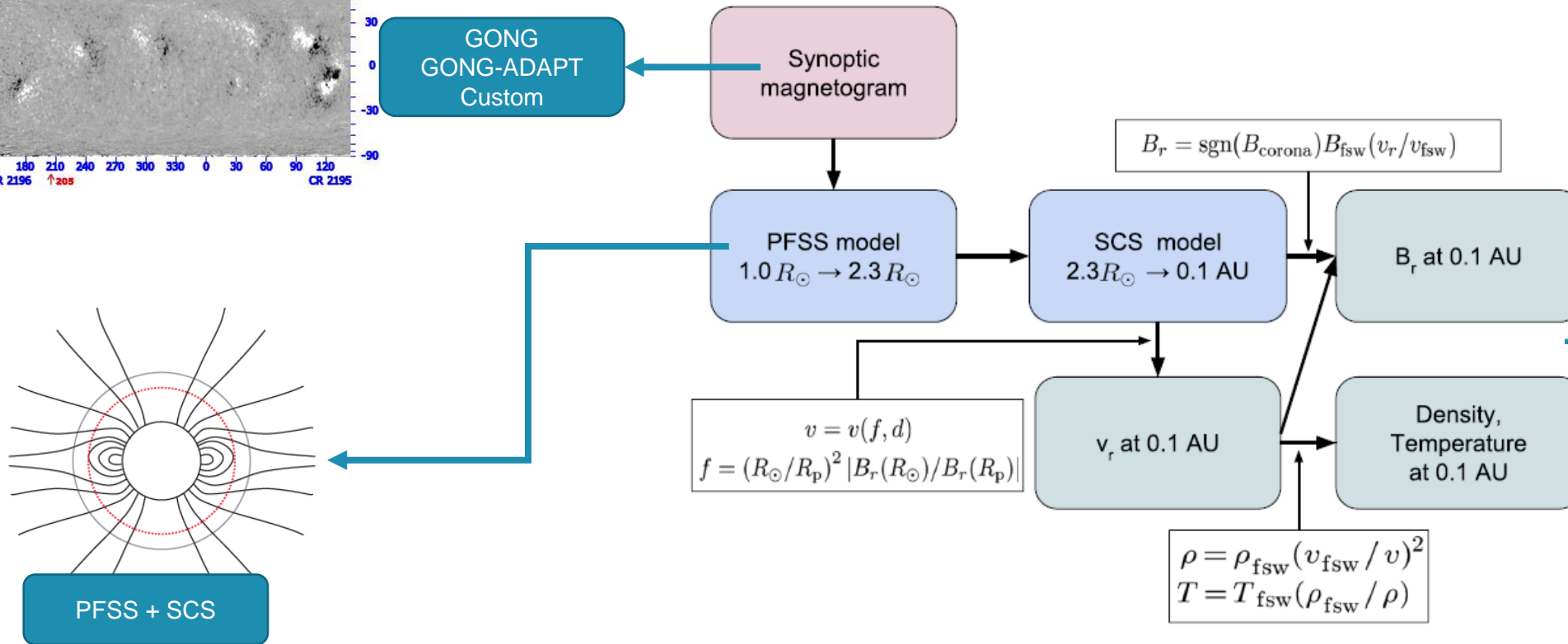


Fig. 1. Schematic of the steps involved in constructing the semi-empirical coronal model that provides the boundary conditions for the heliospheric MHD model.

PARADISE basics

‘Particle Radiation Asset Directed at Interplanetary Space Exploration’

- **PARADISE** solves the FTE by integrating the equivalent SDEs forward in time, i.e.,

$$\left. \begin{aligned} d\mathbf{x} &= \left(\frac{d\mathbf{x}}{dt} + \nabla \cdot \mathbf{D}_{\perp} \right) dt + \sqrt{2\mathbf{D}_{\perp}} d\mathbf{w}_{\mathbf{x}}, \\ d\mu &= \left(\frac{d\mu}{dt} + \frac{\partial D_{\mu\mu}}{\partial \mu} \right) dt + \sqrt{2D_{\mu\mu}} dw_{\mu}, \\ dp &= \frac{dp}{dt} dt, \end{aligned} \right\} \begin{array}{l} \text{Integrate and} \\ \text{sample } \sim 10^8 \\ \text{pseudo-particles} \end{array}$$

- The average SW velocity and magnetic field are obtained from the 3D ideal MHD model **EUHFORIA**
- The diffusion coefficient are derived from a composite slab/2D turbulence model with the assumptions of QLT or a non-linear theory (modular)