



Mitigating the effects of remanance in magnetic data processing and inversion

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Natural remanance in magnetic exploration



RTP field calculated in inducing magnetization direction (left) and total magnetization direction (right).(Zhang et al., 2018)



Two magnetic inversion models without (top) and with (button) remanance. (Li et al., 2021)

Processing and inversion with remanance



- Total magnetization direction estimation by multiple correlation (Jian et al., 2022).
- Magnetization direction estimated by comparing the correlation between RTP field and total gradient of magnetic potential (Jian et al., 2023).
- High-precision magnetization vector inversion with sparse constraints (Li et al., 2022).



Outline

- Methodology
- Synthetic examples
- Field examples



Magnetization direction estimation use cross-correlation





 (a) Total gradient of magnetic potential (TGMP, solid red line) is the envelop of vertical derivatives of magnetic potential (DMP, dashed blue lines) of (b) two rectangular sources



Total gradient of magnetic potential (TGMP, red line) and RTP field (black line) tend to achieve max correlation in the true total magnetization direction.

Estimation using cross-relation





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Estimation using multiple correlation



Computation of multiple correlation

Step 1: select *p* direction insensitive fields.

Step 2 : establish a linear regression between RTP field and these direction insensitive fields. The regression coefficients can be computed through the least squares method.

Step 3 : The multiple correlation coefficient can be written as right.

 ω is the cross-correlation matrix of these *p*+1 quantities and ω_{11} is the algebraic cofactor of the first row and column element of ω .



Estimation use multiple correlation

Method abbreviations	Direction insensitive fields		
M2	TA, Q		
M3	TA, Q, NSS		
M4	TA, Q, NSS, DNSS		
M5	TA, Q, NSS, DNSS, R		
M6	TA, Q, NSS, DNSS, R, L		
M7	TA, Q, NSS, DNSS, R, L, Gravity		

- Magnitude magnetic transforms: TA, Q, R, L (Stavrev and Gerovska, 2000)
- Normalized source strength (Clark, 2009, 2012; Beiki et al., 2012)
- Derivative of Normalized source strength (Zhang et al., 2018)





Magnetic susceptibility inversion with remanance



Sharing

Magnetization vector inversion





The forward of total magnetic intensity in magnetization vector inversion is (Liu et al., 2017):

$$\begin{bmatrix} \mathbf{G}_{M_x} & \mathbf{G}_{M_y} & \mathbf{G}_{M_z} \end{bmatrix} \begin{bmatrix} \Delta \mathbf{M}_x \\ \Delta \mathbf{M}_y \\ \Delta \mathbf{M}_z \end{bmatrix} = \Delta \mathbf{B}$$

- Suitable for non-uniform magnetized situation.
- Inverse the magnitude and directions of magnetization.
- More computation consume and more serious nonuniqueness.

$$\mathbf{m} = \arg\min_{\mathbf{m}} \{P^{a}(\mathbf{m})\}$$
$$= \arg\min_{\mathbf{m}} \left\{ \|\mathbf{W}_{d}(\mathbf{Gm} - \mathbf{d}_{obs})\|_{2}^{2} + \alpha^{2} \|\mathbf{W}(\mathbf{m} - \mathbf{m}_{apr})\|_{2}^{2} \right\}$$

$$\mathbf{W} = \mathbf{W}_{depth} \mathbf{W}_{hard} \mathbf{W}_{L}$$

$$\mathbf{W}_{lp} = \begin{bmatrix} \mathbf{D}_{Lp} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{D}_{Lp} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{D}_{Lp} \end{bmatrix}$$

$$\mathbf{D}_{Lp} = \frac{1}{\left((\mathbf{J} - \mathbf{J}_{apr})^{2} + \epsilon^{2} \right)^{\frac{2-p}{4}}}$$

The model with sharp boundary is obtained by vector inversion based on sparse constraints(Li et al., 2022)





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RTP-L

VG-TG

M3

New method

RTP-NSS

RTP-TMA

Sharing is



RTP-NSS

RTP-TMA

RTP-L

VG-TG

M3

New method

Sharing is encouraged



RTP-NSS

RTP-TMA

RTP-L

VG-TG

M3

New method

Sharing is

encouraged

robustness of the different methods.



times for each source magnetization direction.



Magnetization vector inversion



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Field example: Yeshan, East China



 (left) Total field anomaly, (right) RTP anomaly calculated using inducing magnetization direction.
 The white line shows the boundary of the basalt rocks outcropping.



Contour map of the multiple correlation coefficient calculated with the multiple correlation method for the Xuezhuang area. The red star represents the position of the maximal coefficient.





Field example: Yeshan, East China



Comparison of the estimated total magnetization direction with previous study results from Liu et al. (2018b) and Nurindrawati and Sun (2020) of the Xuezhuang area, eastern China. RTP anomaly of Xuezhuang using the magnetization direction from (a) Liu et al. (2018b), (b) Nurindrawati and Sun (2020), and (c) multiple correlation method. The white line shows the boundary of the basalt rocks outcropping.





Field example: Yeshan, East China



Field example: Weilasito, North China



Area	Inclination	Declination	Area	Inclination	Declination
M1	–24°	-92°	M7	77 °	-10 °
M2	33°	–29°	M8	30°	135°
М3	25°	-19°	M9	83°	94°
M4	21°	–12°	M10	65°	91°
М5	40°	21°	M11	81°	–10°
M6	36°	–52°	M12	80°	64°

Total field anomaly (top) and estimated total magnetization directions in the Weilasito region.



RTP fields obtained using the estimated total directions in the Weilasito region.

Sharing is incourage

Field example: Haba River, North China



Total magnetic intensity of Haba River and the inverted magnetization model. The inverted remanent magnetization vector of Haba River area data with (left column) Q = 0.5and (right column) Q = 2.

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Example: magnetic data inversion of the Lunokhod-2 rover



Landing area

abundance (FeO wt%) map based on Clementine UVVIS, and (d) titanium abundance (TiO2 wt%) map



Le Monnier Bay with Digital Terrain Models (DTM) from Quickmap Hong et al., 2023







No dynamo Paleointensities Non-zero T Upper limit Time Before Present (Ga)

Paleointensity-age map based on Wieczorek et al., (2022), and Lunar sampe data from Lawrence et al., (2008), Cournède et al., (2012), Shea et al., (2012), Tikoo et al., (2012). Our intensity and age results are indicated by the green color.

The thermoremanent magnetization acquired by most rocks is approximately proportional to the magnetizing field for field strengths less than about 50-100 µT. Hong et al., 2023

$$M_{tr} = \chi_{TRM} H$$

 $\chi_{TRM:} 1.58^{+3.14}_{-1.05} \times 10^3$



Copernican

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Thanks for your attention!

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