

Dust aerosols' mineralogy in the chemical transport model COSMO-MUSCAT during JATAC and comparison with lidar data

Sofía Gómez Maqueo Anaya¹, Dietrich Althausen¹, Kerstin Schepanski², Matthias Faust¹, Bernd Heinold¹, Ina Tegen¹, Holger Baars¹, Julian Hofer¹, Ronny Engelmann¹, Anett Skupin¹, Martin Radenz¹, Albert Ansmann¹, Birgit Hesse¹, Ulla Wandinger¹, Eder Silva³, Elizandro Rodrigues³, Pericles Silva³, and Cordula Zenk^{3,4}

¹ Leibniz Institute for Tropospheric Research, Leipzig, Germany
² Free University of Berlin, Berlin, Germany

³ Ocean Science Centre Mindelo, Mindelo, Cape Verde
⁴ GEOMAR Helmholtz Centre for Ocean Research, Kiel, Germany



contact: maqueo@tropos.de

Motivation

Can the modelled minerals explain the observed optical properties?

Example case: 16 Sept 2021, 22:15 – 23:39 UTC at Mindelo

Dust particle size and mineralogy are crucial for understanding the **interaction between dust aerosols and radiation** (Balkanski et al., 2007; Kok et al., 2023). Despite the well-known mineral variations in dust aerosols, most atmospheric models consider them to have uniform composition.

Iron oxide content in minerals is correlated with a distinct interaction with the UV/VIS spectral region and varies depending on the **source region** (Formenti et al., 2011; Veselovskii et al., 2020).

Summary & Outlook

Polly^{XT} lidar ratio (LR) vertical profile:

- Differences between the LR at 355 nm and 532 nm are higher at 1 – 3 km than at 3.5 – 5 km.

COSMO-MUSCAT :

- Higher simulated iron oxide (hematite) mineral content at 1 – 3 km than at 3.5 – 5 km.

Outlook

Compare modelled minerals to elemental chemical analyses and / or mineralogical measurements.

Investigating further the relationship between the iron oxide minerals and radiation.

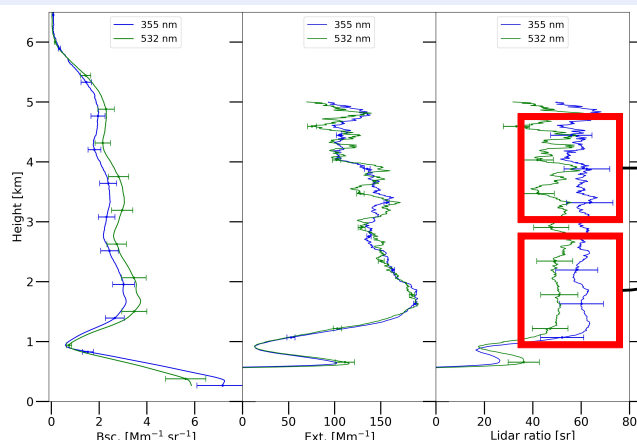


Fig. 1: Lidar ratio (extinction / backscattering) spectral difference is higher between 1 – 3 km than between 3.5 – 5 km.

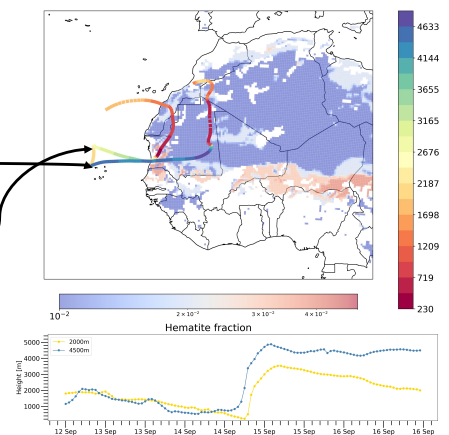


Fig. 2: Backtrajectory arriving at 2 km passes through a high hematite content mineral dust emission site.

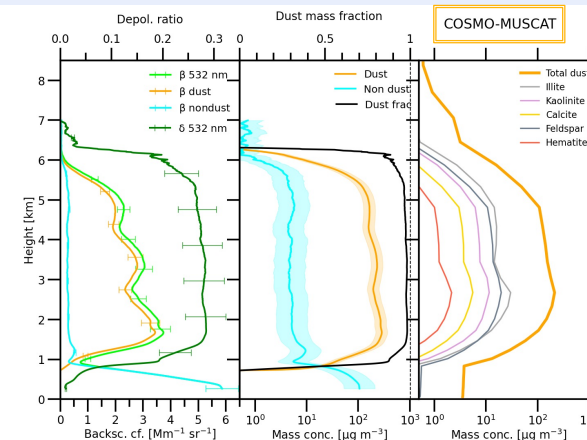


Fig. 3: Dust mass concentrations from applying the POLIPHON method compared to COSMO-MUSCAT's results. Higher simulated hematite content between 1 – 3 km.

Observations with the multi-wavelength polarization lidar, Polly^{XT} (3+3+3)

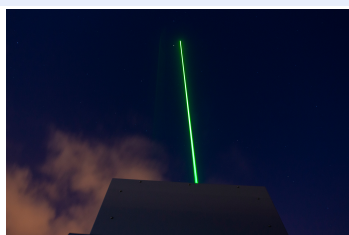


Fig.4 Polly^{XT} laser beam at OSCM, Mindelo. Edson Silva, Etfilmes/OSCM, July 2021

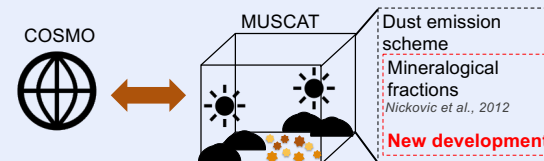
Retrieved particle parameters at three wavelengths (355 nm, 532 nm, and 1064 nm):

- Extinction coefficient
- Backscattering coefficient
- Lidar ratio
- Angstrom exponent
- Depolarization ratio
- Dust mass concentration (POLIPHON, Ansmann et al. (2019))



24/7 operation: <https://polly.tropos.de/>

Minerals estimations through COSMO-MUSCAT simulations



- Horizontal res: 0.25°x0.25°
- Vertical res: 40 sigma-p levels
- DWD meteorological fields
- Dust emission scheme follows Tegen et al., (2002)
- GMINER mineralogical database

COSMO-MUSCAT is a regional chemical and transport model used for:

- Emission, transport and deposition of dust in the Sahara Desert now considers mineral fractions (Gómez Maqueo Anaya, et al., *in prep*).
- Backtrajectories calculated using LAGRANTO help identifying source regions of the simulated minerals.

References

Ansmann, A., Mamouri, R.-E., Hofer, J., et al. (2019), doi:10.5194/amt-12-4849-2019; Balkanski, Y., Schulz, M., Clauquin, T., and Guilbert, S. (2007), doi:10.5194/acp-7-81-2007; Formenti, P., Schütz, L., Balkanski, Y., et al. (2011), doi:10.5194/acp-11-8231-2011; Kok, J.F., Storelvmo, T., Karydis, V.A. et al. (2023), doi:10.1038/s43017-022-00379-5; Nickovic, S., Vukovic, A., Vujadinovic, M., et al. (2012), doi:10.5194/acp-12-725845-2012; Tegen, I., Harrison, S. P., Kohfeld, K., et al. (2002), doi:10.1029/2001JD000963; Veselovskii, I., Hu, O., Goloub, P., et al. (2020), doi: 10.5194/acp-20-6563-2020