

Convection in Europa's icy shell: composite rheology and dynamic grain size evolution

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MOTIVATION

Europa is a prime target for exploring extraterrestrial habitability. Convection and tectonics in its ice shell control heat and material transport, thus surface-ocean exchange. The style of convection and its link to surface observables is poorly understood though. Here we are interested in:

- How does Europa's ice shell deform (in which deformation mechanism?)
- How sensitive are the ice shell dynamics to grain-size evolution (GSE)?

APPROACH

Code StagYY [3]; incompressible extended boussinesq; 160x20 km box (1024x128 cells), 64 markers/cell. $Ra = 2.57 \times 10^7$, $n = (1, 2.4, 1.8, 4)$ for (dif, bsl, gbs, dis), $m = (2, 0, 1.4, 0)$, $E = (60, 60, 50, 60)$ kJ/mol [1,2], $p = 4$, $D^* = 4$ mm, $f = 10^{-2}$, $k = 4.4 \times 10^{-11}$ m⁴/s.

- **Composite rheology** as experimentally inferred for ice [1]. 4 different deformation mechanisms i [2]: diffusion creep (dif), basal slip (bsl), grain boundary sliding (gbs), and dislocation creep (dis), with different dependencies on stress and grain size (e.g. exponents n_i, m_i).

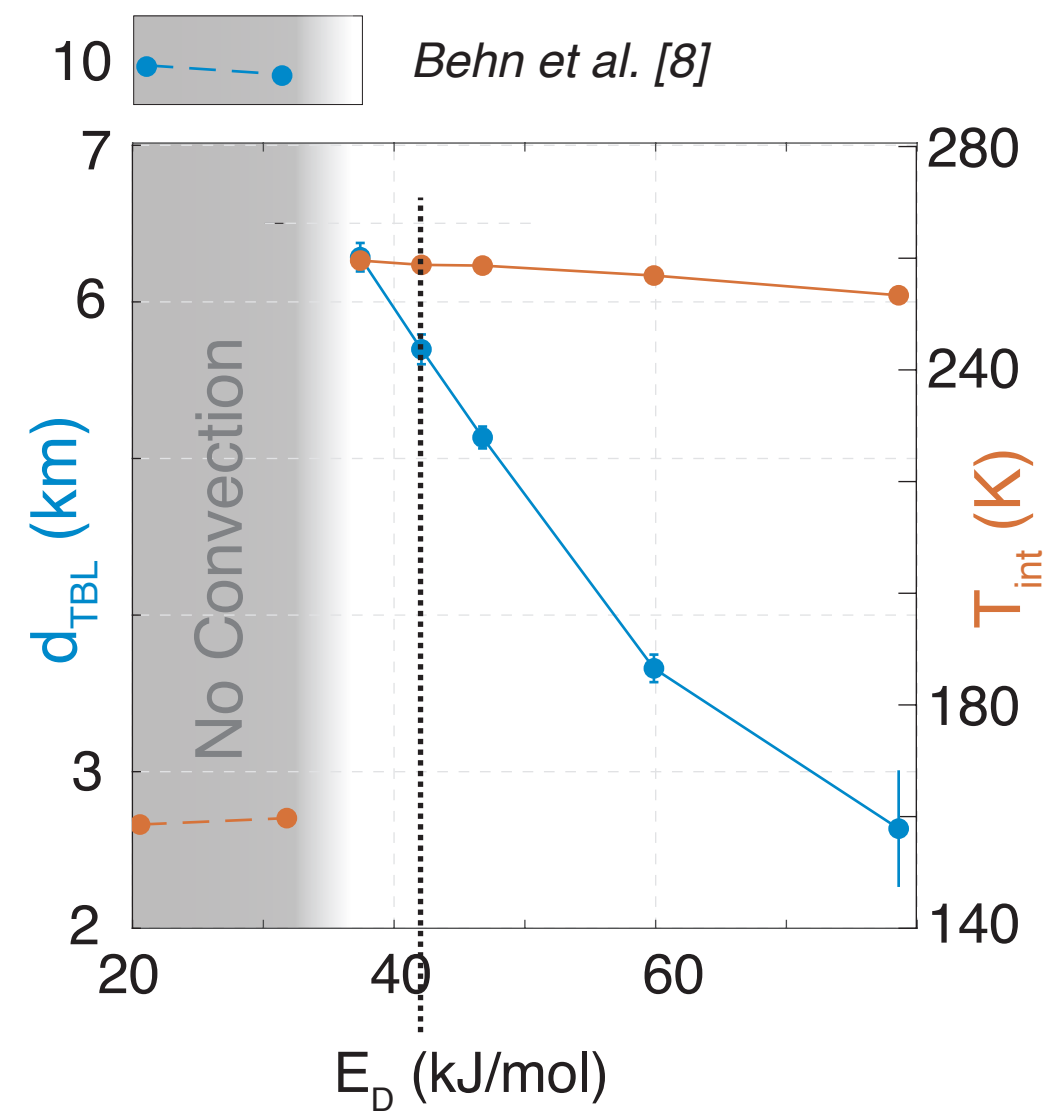
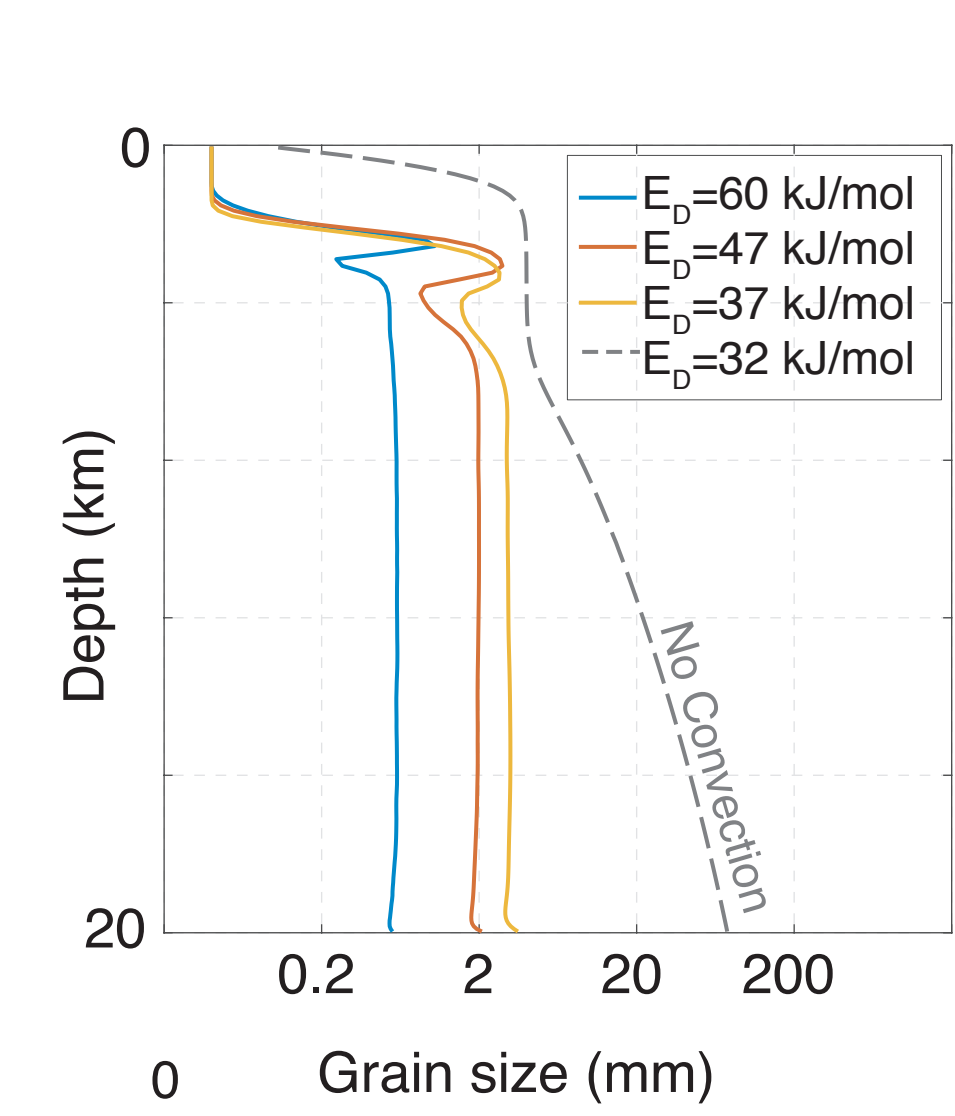
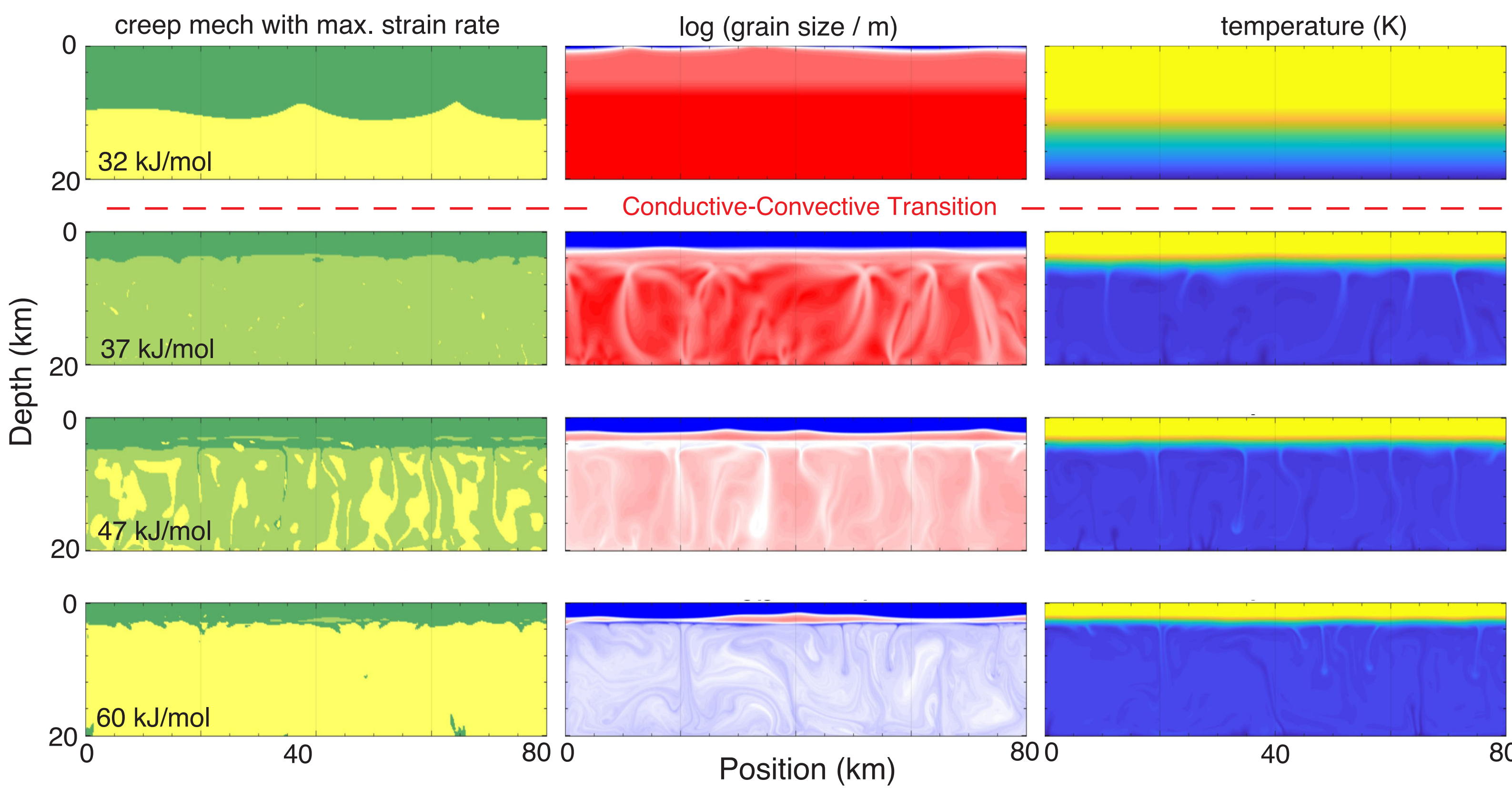
$$\eta_i = B_i \left(\frac{\tau}{\tau_i^*} \right)^{1-n_i} \left(\frac{D}{D_i^*} \right)^{m_i} \exp \left(\frac{E_i}{RT} \right) = \tilde{B}_i \tau^{1-n_i} D^{m_i} \exp \left(\frac{E_i}{RT} \right)$$

$$\eta = \left(\frac{1}{\eta_{dif}} + \frac{1}{\eta_{bsl} + \eta_{gbs}} + \frac{1}{\eta_{dis}} \right)^{-1}$$

- **Grain size evolution (GSE)** [4,7]: on tracers; balance of grain growth and reduction; Ψ = viscous dissipation (without the fraction done in DIF creep):

$$\dot{D} = \frac{k \exp(-E_D/RT)}{p D^{p-1}} - C f D^2 \Psi$$

GSE PARAMETERS (here only E_D)



- **Reducing E_{GS}** : Grain growth is less sensitive to temperature
- Faster grain growth in interior, therefore higher viscosity that can eventually cause **convection to cease (conductive ice shell)**.
- Other GSE parameters (k, f) indicate similar, but less pronounced effects. The system should also be very sensitive to the grain size exponent (p , to be tested).

d_{TBL} : Thickness of the thermal boundary layer (~the immobile layer)

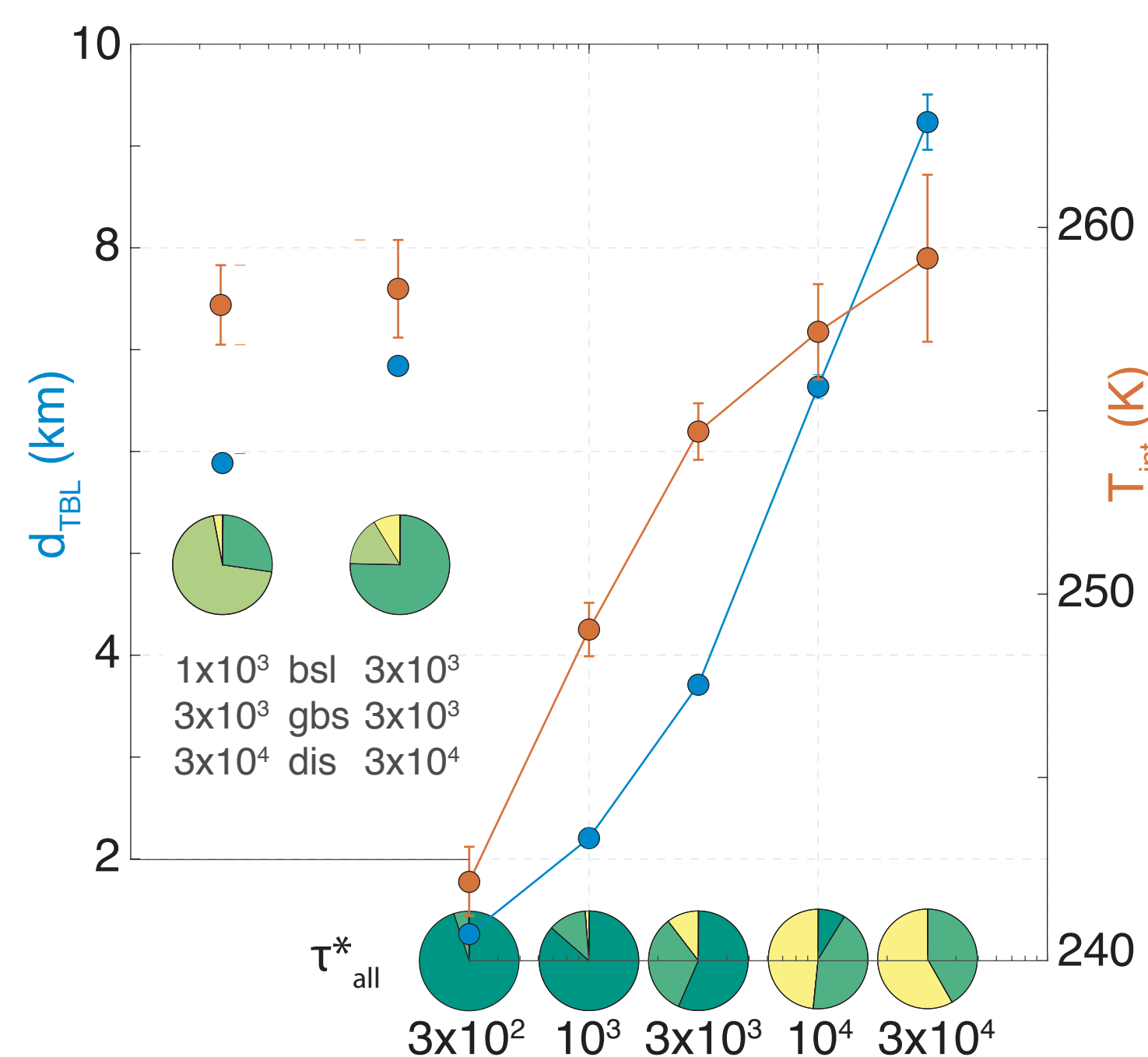
T_{int} : Characteristic temperature of the convecting interior. Value not meaningful if purely conductive!

TRANSITION STRESSES

- Reference stresses **crucial for dominant creep mechanism**. Equal values for all mechanisms promote the one with largest stress exponent.

- **Immobile layer thinner with more abundant nonlinear creep** (especially DIS) [see 4,5], because of lower viscosity and more vigorous convection

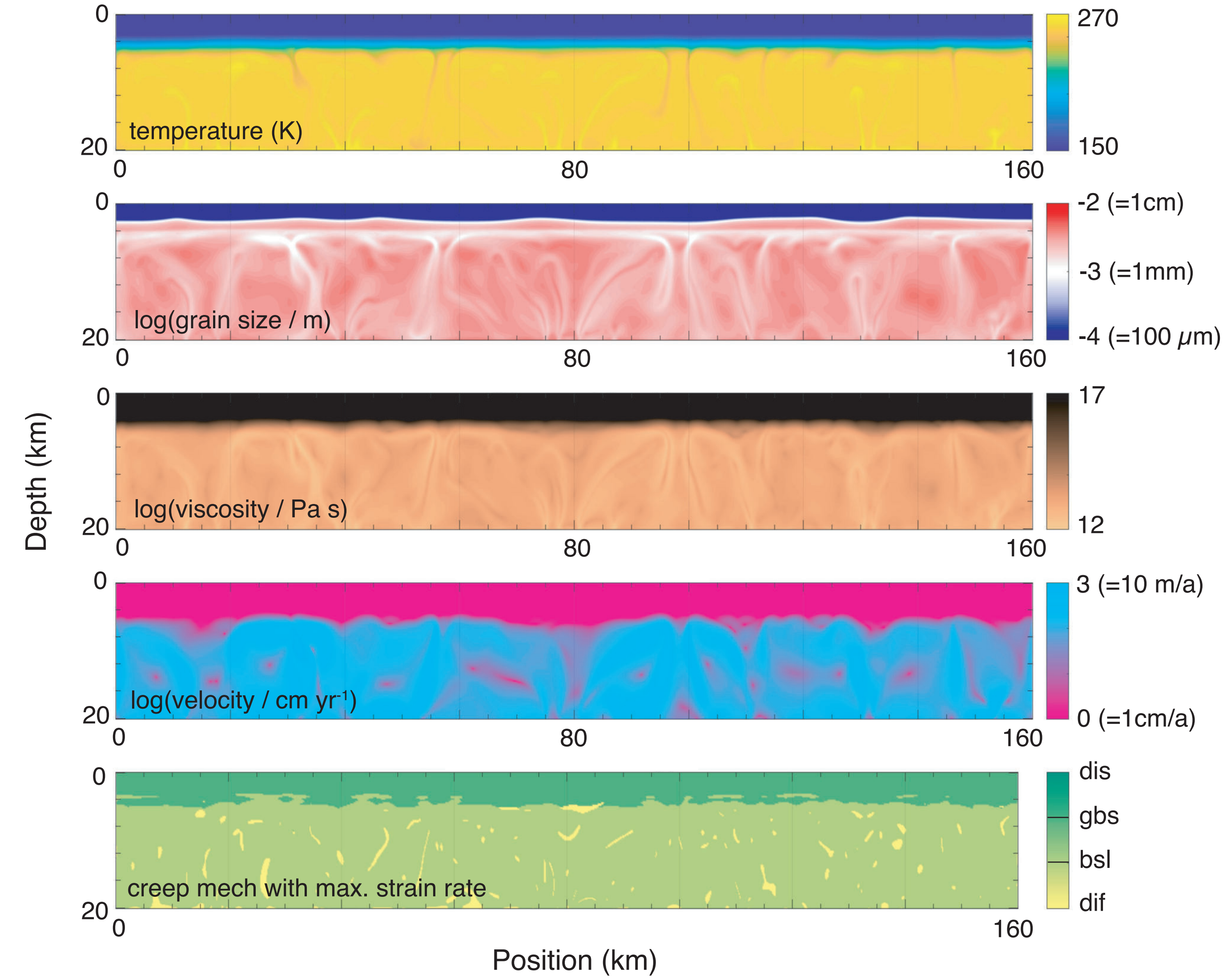
- Laboratory data [1] suggests that here $\tau_{bsl}^* \ll \tau_{gbs}^* \ll \tau_{dis}^*$ as already used. But results are very sensitive to chosen ratios and values! Values matching the laboratory constraints are to be tested!



EXAMPLE

- **Immobile surface layer**: ~6 km thick, small grains (<100μm), high viscosity (since strongly T-dependent viscosity), max. strain rates occur in GBS creep

- **Convecting interior**: Larger grains (several mm), mild variation in viscosity (~1.5 order of magnitude) and grain size (<1 order), highest strain rates in BSL creep, or in DIF (some pockets)

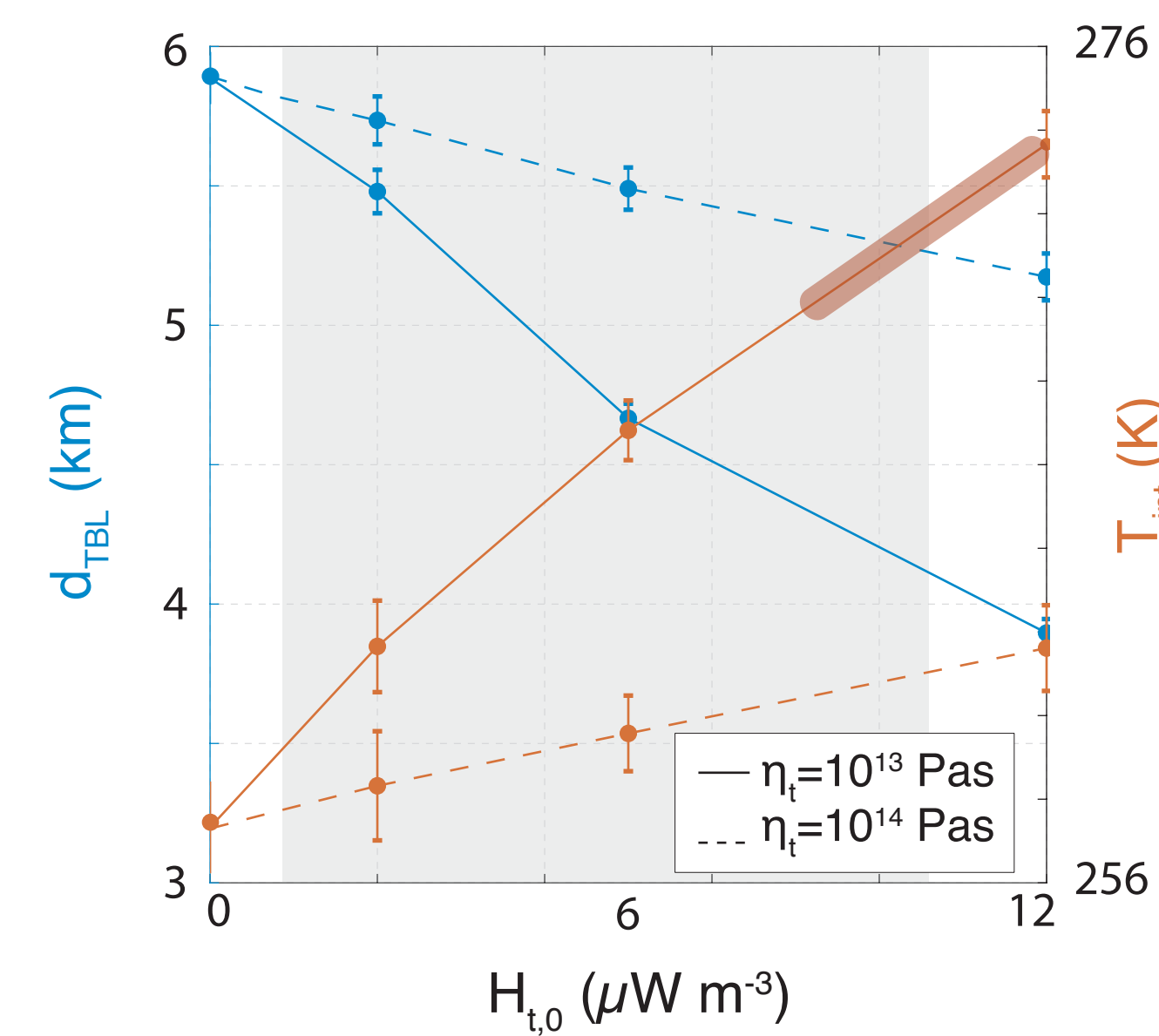


TIDAL HEATING

$$H_t = H_{t,0} \left(\frac{2\eta\eta_t}{\eta^2 + \eta_t^2} \right)$$

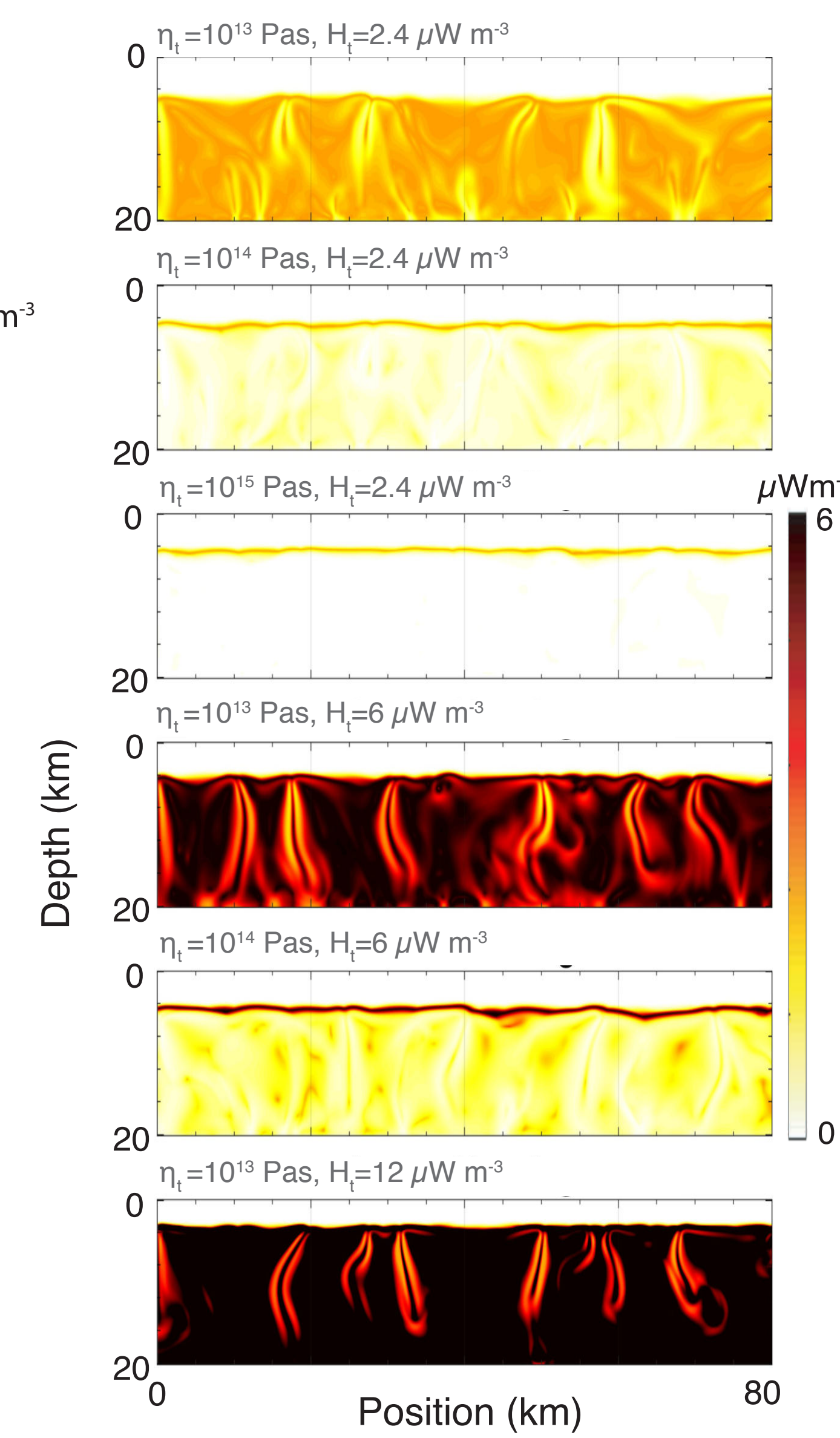
$H_{t,0}$: Optimal tidal heating rate at η_t , Europa: ~1-10 $\mu W m^{-3}$

η_t : Optimal viscosity for tidal heating, ~ratio of shear modulus & orbital period. Europa: ~10¹³ - 10¹⁴ Pas.



- No tidal heating in immobile layer, but higher $H_{t,0}$ in interior **increases internal temperature (T_{int})** and **thins the immobile layer (d_{TBL})**.

- Higher T_{int} tends to **promote diffusion creep** compared to the other creep mechanisms.
- For small $\eta_t = 10^{13}$ Pas and large H_t , large portions of the ice shell would just melt!



TAKE HOME

- Convection in Europa's ice shell features a conductive, immobile layer over a convecting interior. The dominant creep mechanism and the amount of tidal heating determine immobile layer thickness.
- GSE affects the dominant creep mechanism, the thickness of the immobile layer and also the possibility of convection. It should not be omitted when investigating the dynamics of planetary ice shells!
- While laboratory experiments neither confirm nor reject the occurrence of diffusion creep, our simulations predict DIF to be a feasible mechanism in the convecting deeper interior.
- Composite rheology coupled to GSE does not generate a mobile surface by itself; no presented cases display any notable surface mobility. Plastic yielding may still be required as an additional weakening mechanism.

TO-DO

- Transition stresses in line with laboratory experiments
- Plastic Yielding (for surface mobilisation)
- Ice impurities and their effect on GSE