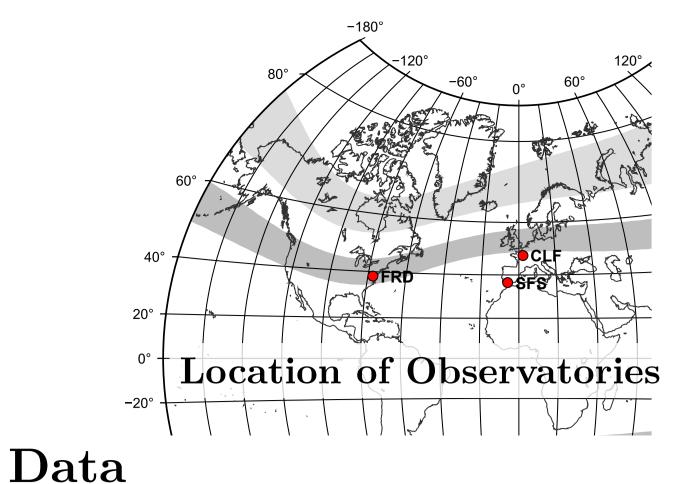




# Extraction of Solar Forcing Signatures in Ground Magnetometer Data from Sub-Auroral Regions

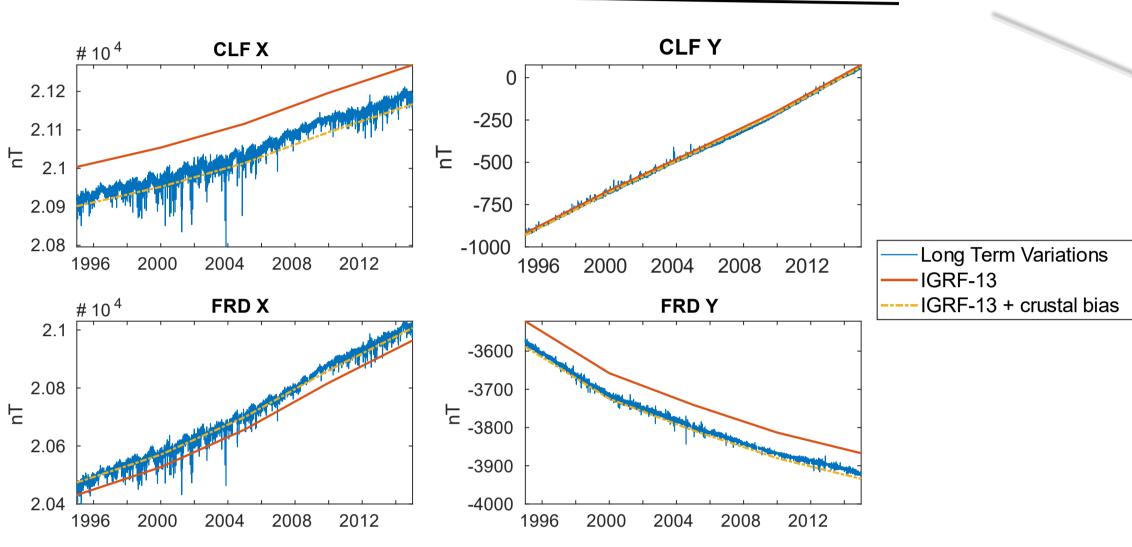
Veronika Haberle<sup>1,2</sup>, Aurélie Marchaudon<sup>1</sup>, Aude Chambodut<sup>2</sup>, and Pierre-Louis Blelly<sup>1</sup>





### Introduction

Strong solar events that reach Earth can significantly disturb the geomagnetic field. As the recorded magnetic signal is a superposition of sources that act on a broad amplitude-frequency spectrum, the isolation of storm signatures within geomagnetic field measurements becomes an important and nontrivial task. In order to determine the strength in real-time, the identification of storms needs to be automated.



agreement with the IGRF model (Alken et al. 2021), indicating that the filter catches the secular variation. Contributions of other internal and external sources are also present.

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—Storm Signature

Dec 19

Dec 19

Dec 19

Source Extraction

—Geomagnetic Field Measurement

**Long Term Variations** 

**Diurnal Variations** 

Residuals

-550

Dec 10

Dec 10

Dec 10



Geomagnetic field measurements are taken from observatories at midlatitudes that are part of INTERMAGNET and which provide high quality data in minute resolution in the local NED coordinate system:

X geographic North, Y geographic East, Z vertically down The extraction of sources is realised via temporal filters with periods of above 36h for the long-term and at periods of 24h, 12h, 8h,

### Y - geographic East and 6h for the diurnal variations, see Haberle et al. 2022. Diurnal Variations Dec 19 Dec 10 Dec 16

Dec 19

Dec 19

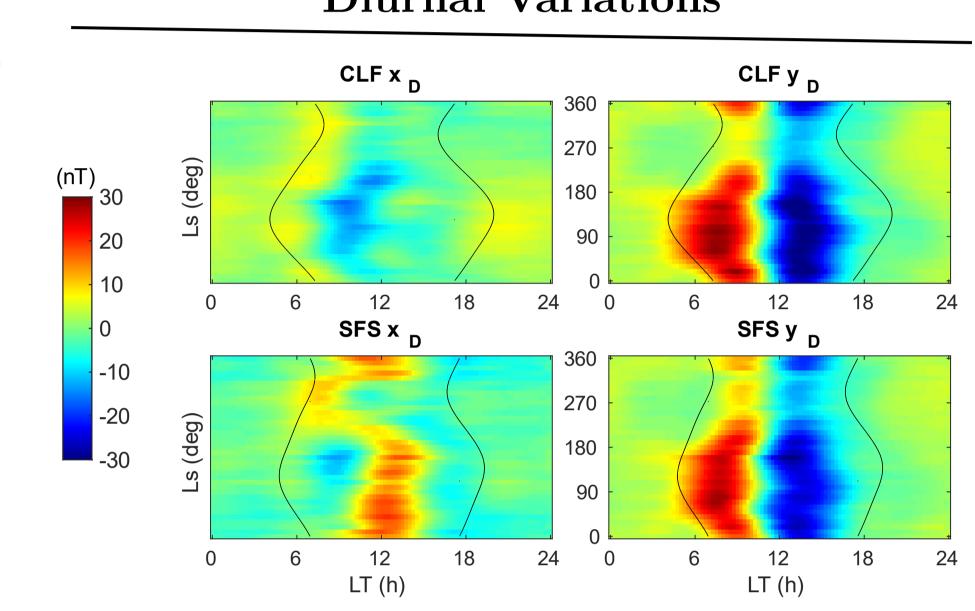
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2002

Dec 16

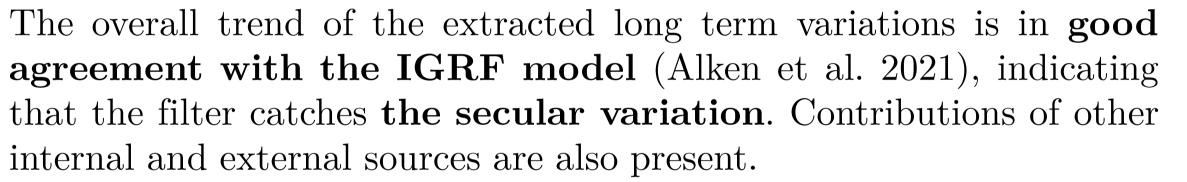
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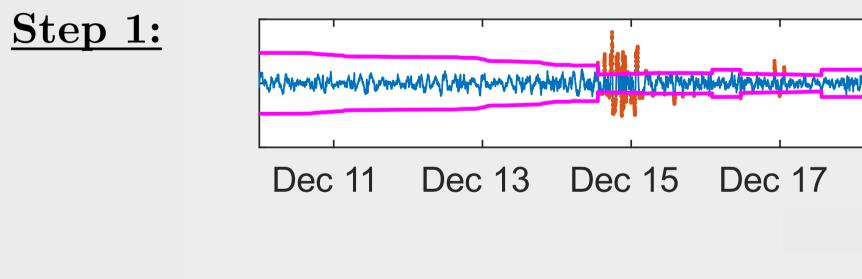
The superposed epoch analysis of diurnal variations in dependence of solar longitude and local time is generated over geomagnetically quiet periods. The characteristics are in good agreement with the solar quiet current system (Haberle et al. 2022).

# Long Term Variations

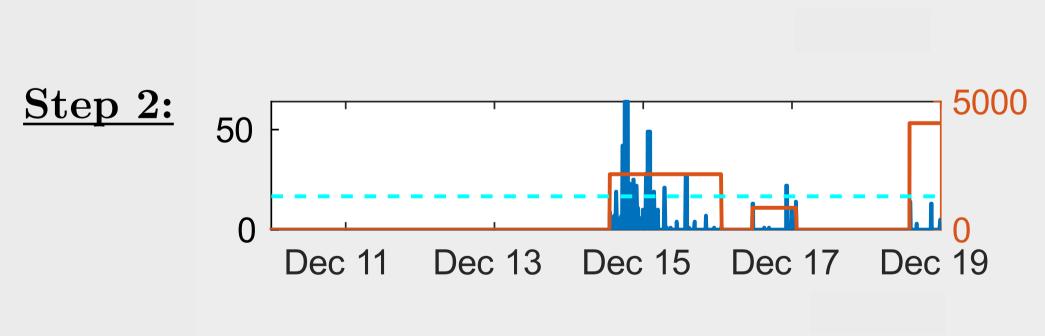


## Extraction of Storm Signatures

Removing long-term and diurnal variations from geomagnetic field measurements results in the residuals containing predominantly solar forcing contributions. These can then be used within the disturbances characterise geomagnetic field. The automatic identification of storm signatures is done in two steps. First, typical quiet levels are determined, followed by the counting of non-quiet residuals per time interval.



Dec 13



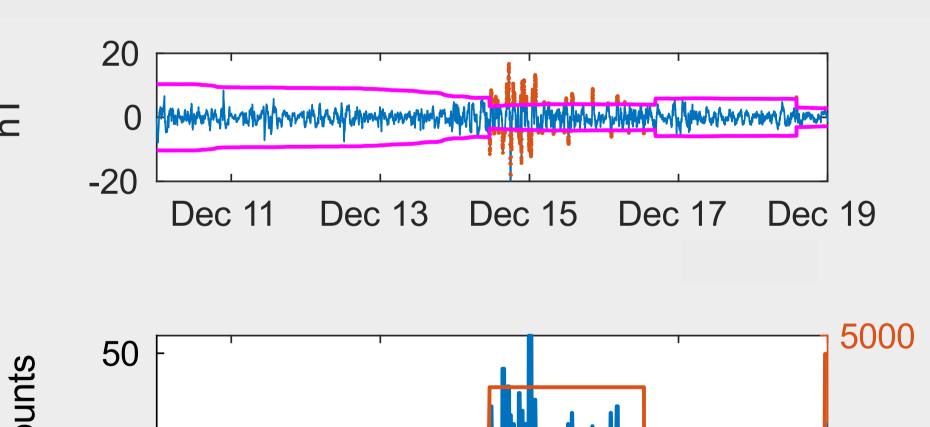
X - geographic North

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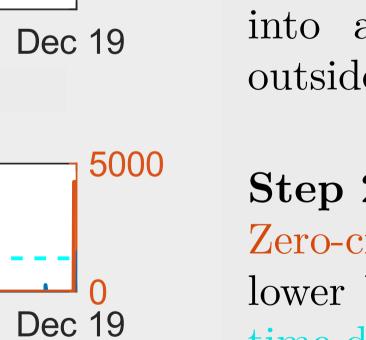
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Dec 15

Dec 13

Dec 13



### Step 1: Determining quiet levels

A quiet-time expectation boundary is defined by taking into account the previous three months. Residuals outside this boundary are considered storm candidates.

### Step 2: Defining storm intervals

Zero-crossings for storm candidates between upper and lower boundaries are taken care of. A minimum storm time duration threshold defines storm signatures.

### References

Alken, P. et al. Earth Planets Space 73, 49 (2021). https://doi.org/10.1186/s40623-020-01288-x Haberle, V. et al. (2022) Journal of Geophysical Research: Space Physics, 127. https://doi.org/10.1029/2022JA030407

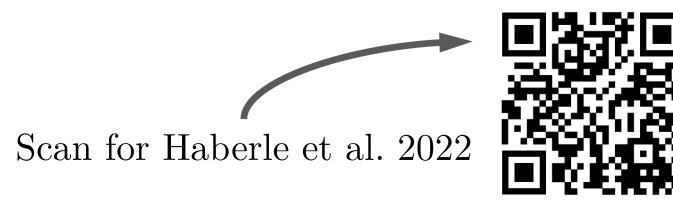
# Conclusion and Perspective

With these traditional methods stronger storms are identified correctly including recovery phases. Sometimes shorter, quiet periods are mistakenly marked as a storm but tuning of parameters is time-consuming. Can artificial intelligence help in finding optimal parameters and improve storm identification?

# Chambon-la-Foret BCMT data repository doi.org/10.18715/BCMT.MAG.DEF | doi.org/10.5880/INTERMAGNET.1991.2018

### Magnetic Observatory Data Access FRDSFSSan Fernando Fredericksburg INTERMAGNET data repository

École et observatoire des **sciences de la Terre** Université de Strasbourg



More questions? Reach out! veronika.haberle@irap.omp.eu





