

write our model for future $s \ge t_0$ through Sklar's theorem as:

 $\mathcal{M}\left(oldsymbol{\mathcal{Y}}_{s} | oldsymbol{\mathcal{P}}_{t_{0}}
ight)$ Joint Density

Now the question is: "How to model marginal densities f_i and how to model the copula c to capture the spatio-temporal nature of the data?". We propose novel solutions for both parts.

1. Joint Generalised Neural Models for N

Have to take particularities of rainfall data into account when modelling $f_i(y_{i,s}|\mathcal{P}_{t_0})$.

- Rainfall data is **zero-inflated** (over-representation of 0-valued observations).
- Rainfall can be > 0 with a **continuous** probability density function on $\mathbb{R}_{>0}$.
- Rainfall can also never occur, resulting in a **discrete** observation with mass at 0.

We use a mixture model, giving separate probability to 0 and positive rain as:

$$f_i(y|\boldsymbol{\mathcal{P}}) = [1-p] \cdot \delta_0(y) + p \cdot \mathcal{G}(y;\mu,\phi) = f_i(y|\mu,\phi,p).$$

Expand on existing GLM and NN literature to create JGNMs- a model to learn parametric densities with the combined strengths of both approaches:



Green box: Initial Input data, a subset of \mathcal{P}_{t_0} relevant to location *i*. **Blue box**: NN to refine the data into new predictors, captures non-linear trends. **Red box**: Simultaneous GLMs for parameters of model, using the refined predictors.

• Data is on a 100*140 grid covering the UK with 7 predictors and target rain for 40 years with a 20 years/20 years train/test split.



Joint Generalized Neural Models and Censored Spatial Copulas for Probabilistic Rainfall Forecasting

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Introduction

Want a model for future joint distributions of rainfall across space. Use the copula trick: (1) modelling marginal distributions at each locations separately and (2) uniting them with copulas to infuse a spatial dependence. For predictors $\boldsymbol{\mathcal{X}}_{i,t} = (\boldsymbol{\mathcal{X}}_{i,t}^{(1)}, \dots, \boldsymbol{\mathcal{X}}_{i,t}^{(6)})$ and target rainfall $\boldsymbol{\mathcal{Y}}_t = (y_{1,t}, \dots, y_{n,t})$ at times $t \in \mathbb{Z}$ and locations i, combined into present information $\boldsymbol{\mathcal{P}}_{t_0} = \{\boldsymbol{\mathcal{X}}_{i,t}, \boldsymbol{\mathcal{Y}}_t : t < t_0\}$, we

> $= f_1(y_{1,s}|\mathcal{P}_{t_0}) \cdot \ldots \cdot f_n(y_{n,s}|\mathcal{P}_{t_0}) \cdot \mathbf{c} (F_1(y_{1,s}|\mathcal{P}_{t_0}), \ldots, F_n((y_{n,s}|\mathcal{P}_{t_0})))$ Censored Copula JGNM

Marginals			

2. Censored Spatial Copula

Assume \mathbf{c} is a Gaussian copula, inducing a known form for its density and allowing us to learn the spatial dependence. But we have to take into account the zero-inflation of data, so we modify the classical Gaussian copula through **censoring** :

- **1.** Transform all observations to a common Gaussian scale using probability integral transforms with their JGNM marginals giving $x_{i,t} = \Phi^{-1}(u_{i,t}) = \Phi^{-1}(F(y_{i,t}|\mu_{i,t},\phi_{i,t},p_{i,t})).$
- **2.** All 0-valued observations $y_{i,t}$ get mapped to the censoring levels $d_{i,t} = \Phi^{-1}(1 p_{i,t})$. **3.** Modify the copula density **c** to reflect said censoring.

The resulting density has to integrate over the censoring levels of dry locations, yielding

$$\mathbf{c}(x_{1,t},...,x_{n,t}) = \frac{\int_{-\infty}^{d_{i_1,t}} \dots \int_{-\infty}^{d_{i_k,t}} \phi_n(z_{1,t},...,z_{n,t};\mathbf{0},\boldsymbol{\Sigma}) \, dx_{i_1}}{\prod_{\{i:y_{i,t}>0\}} \phi(x_{i,t}) \cdot \prod_{\{j:y_{i,t}=0\}} \Phi(x_{i,t})} \Phi(x_{i,t}) \cdot \prod_{\{j:y_{i,t}=0\}} \Phi(x_{i,t}) + \sum_{i_1,i_2 \in \mathbb{N}} \Phi(x_{i,t}) \cdot \prod_{\{j:y_{i,t}=0\}} \Phi(x_{i,t}) + \sum_{i_1,i_2 \in \mathbb{N}} \Phi(x_{i,t}) + \sum_{i_1,i_2 \in \mathbb{N}} \Phi(x_{i,t}) + \sum_{i_2,i_2 \in \mathbb{N}} \Phi(x_{i,t}) + \sum_{i_1,i_2 \in \mathbb{N}} \Phi(x_{i,t}) + \sum_{i_2,i_2 \in \mathbb{N}} \Phi(x_i) + \sum_{$$

where Σ is a covariance matrix that we have to estimate. Construct Σ using GPKs, relying on parameter $\boldsymbol{\theta}$ and distance \mathcal{D} : $\boldsymbol{\Sigma}_{(i,j)} = k(\mathcal{D}_{(i,j)}|\boldsymbol{\theta}) = \frac{2^{1-\nu}}{\Gamma(\nu)} \cdot \left(\sqrt{2\nu} \frac{\mathcal{D}_{(i,j)}}{\boldsymbol{\theta}}\right)^{\nu} \cdot K_{\nu}\left(\sqrt{2\nu} \frac{\mathcal{D}_{(i,j)}}{\boldsymbol{\theta}}\right)$ We specify \mathcal{D} with relevant spatial information and only need to estimate $\boldsymbol{\theta}$.

3. Copula Estimation via Scoring Rules

Estimation of Σ via MLE becomes expensive due to integration in **c**. Instead rely on **likelihood**free inference via Scoring Rules. Optimise Energy Score as divergence between simulations and observations:

 $SR = 2 \cdot \mathbb{E}_{\mathbf{X}' \sim \mathbf{c}_{\theta}} \| \mathbf{X}' - \mathbf{x} \|_{2} - \mathbb{E}_{\mathbf{X}'_{1}, \mathbf{X}'_{2} \sim \mathbf{c}_{\theta}} \| \mathbf{X}'_{1} - \mathbf{X}'_{2} \|_{2}$

For selected $\boldsymbol{\theta}$ construct Σ to simulate data from $\phi_n(.|\mathbf{0}, \Sigma)$ and censor with from $\mathbf{c}_{\boldsymbol{\theta}}$. Choose $\boldsymbol{\theta}^* = \arg \min SR(\boldsymbol{\theta} | \{x_{i,t} : i \in \{1, \dots, n\}, t < t_0\}).$

4. Application to UK Rainfall

(a) Observed rain



(b) Simulated rain with distance and topology

4. Application to UK Rainfall

- Compare our model **Cens-JGNM** against benchmark **ConvCNP** and **VAE-GAN**.
- Assess calibration (a), spatial-coherence (b) and precision (c) of forecasts.



 $a_{1.t} \dots dx_{i_k,t}$

$$\mathbf{X}_{2}^{\prime} \|_{2}$$

in $d_{i,t}$; these are simulations

(a) Observed



(c) ConvCNP





Contribution 1: JGNM which unite two powerful frameworks into a reliable learner of parametric distributions with complex data patterns such as rainfall. **Contribution 2**: Censored Gaussian Copulas as a rigorous treatment of zero-inflated joint density modelling, or any data with censoring. **Contribution 3**: Main contribution of this work, a methodology for the estimation of copulas without relying on the

likelihood, much cheaper in most cases especially in high dimensions.





(b) Censored JGNM



Cross-correlation maps, (b)



Observed and simulated frequencies of rainfall. (a).

Conclusion