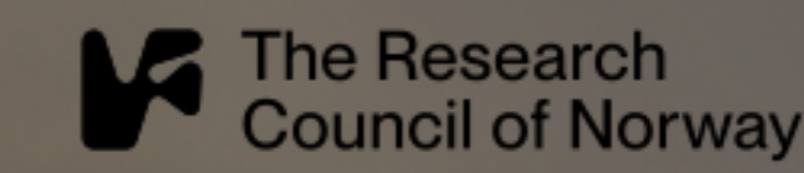




Exploring the feasibility of detecting seismically-generated infrasound waves on Venus using balloon platforms

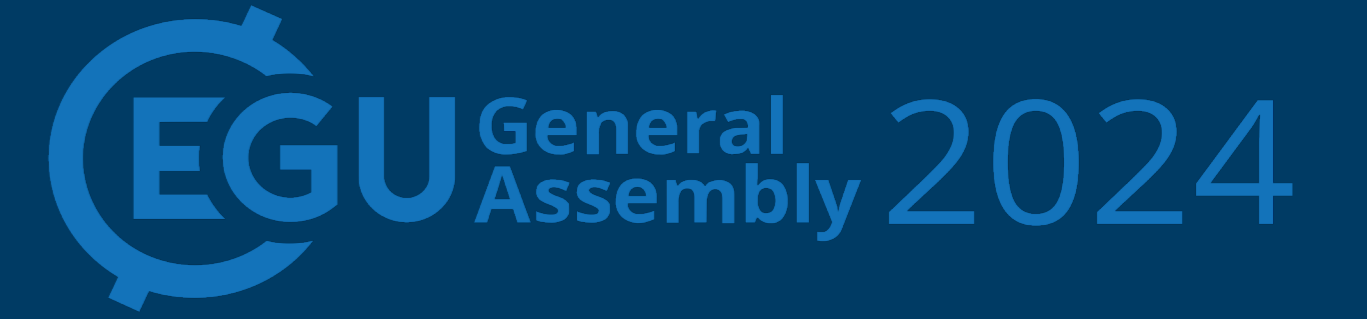
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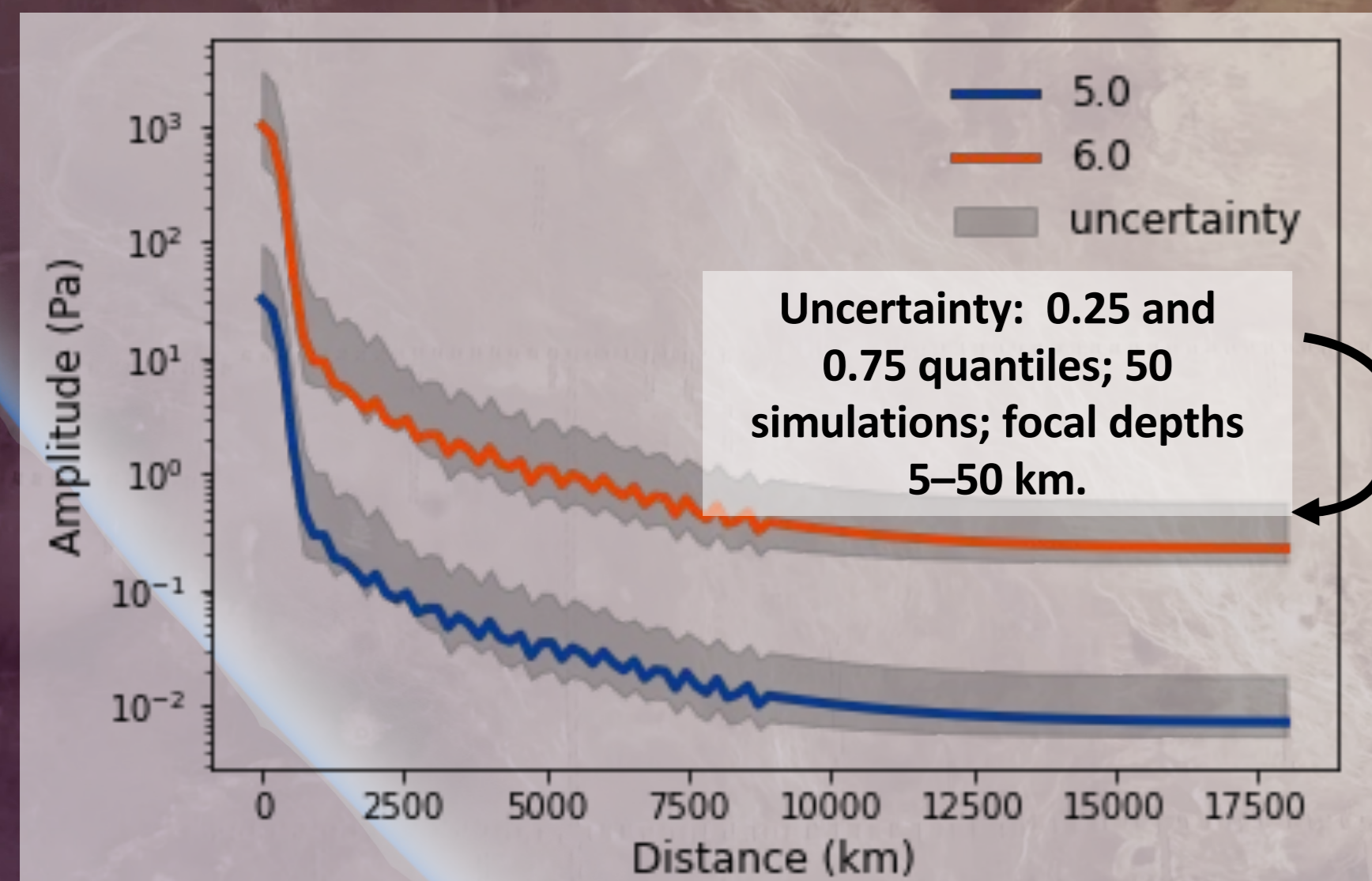
INTRODUCTION

- Due to the harsh surface environment with high pressure and temperature, balloon platforms might be one of the only realistic option to investigate Venus' seismicity [1].
- Seismoacoustic coupling is efficient on Venus due to its dense atmosphere: seismic waves couple to the atmosphere as infrasound which can be recorded by a balloon.
- Here we provide the first assessment of the global detectability of these seismic infrasounds at high altitude based on numerical modeling.

METHODS

#1 Estimate the spatial and temporal venusquake distribution λ_q in terms of magnitude, based on Earth scalings [2,3].

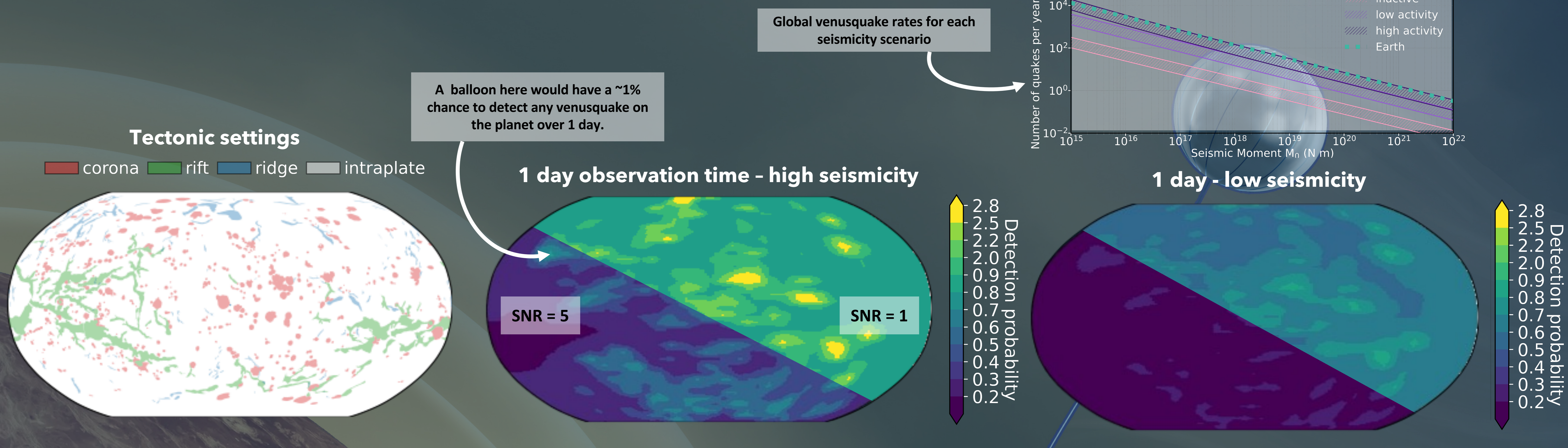
#2 Infrasound amplitude modeling using seismic Green's functions and ground-to-balloon scaling for a 2-layer Venus subsurface.



#3 Determine probability of observing at least one venusquake with $SNR > d$ over a time period t , i.e., the Poisson process:
 $\mathbb{P}(SNR > d, x_t^{obs}, t) = 1 - \exp[-\lambda(\lambda_q, SNR > d)t]$.

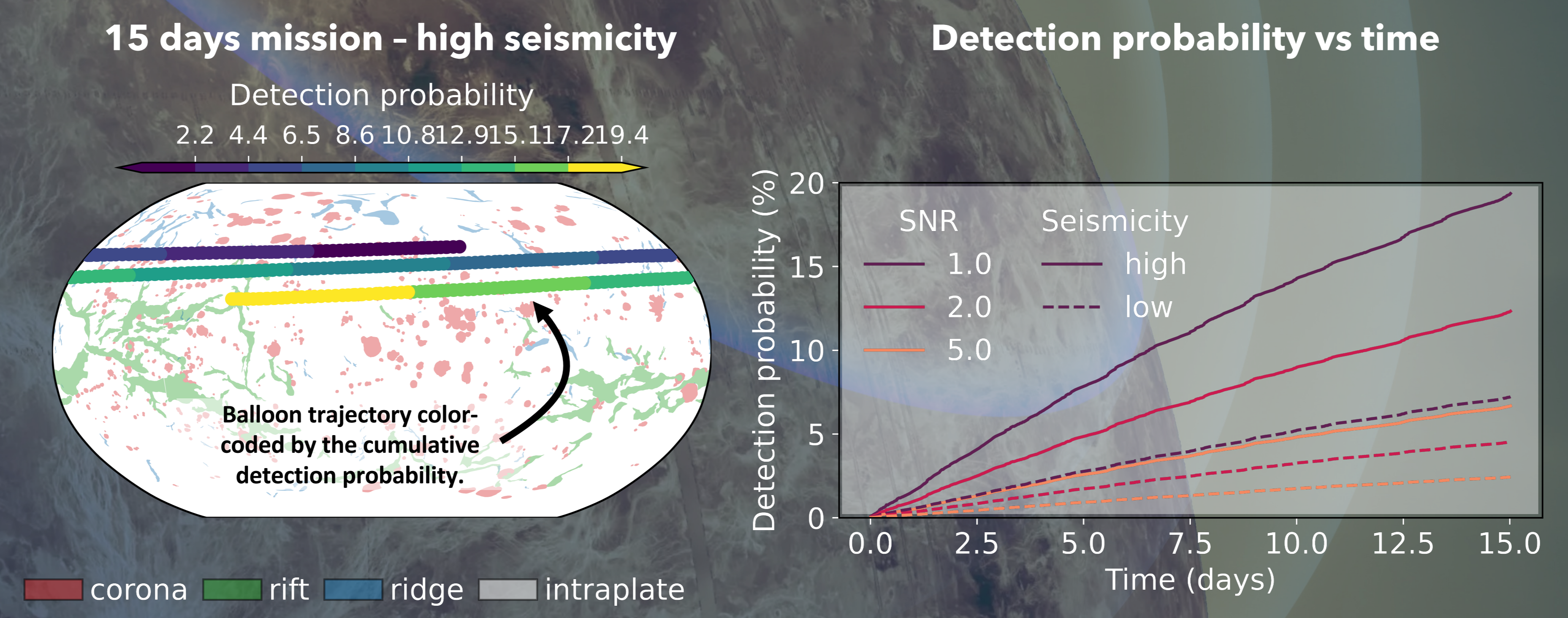
#4 Integrate probability #3 along a balloon trajectory freely drifting with horizontal winds
 $\mathbb{P}(SNR > d, x^{bal}) = 1 - \prod_{t \leq t_{max}} [1 - \mathbb{P}(SNR > d, x_t^{obs})]$

HOW DOES SEISMICITY INFLUENCE DETECTABILITY?



We estimate up to 12% chance to detect a venusquake over a 15-day balloon mission

CAN A BALLOON DETECT A VENUSQUAKE OVER THE ENTIRE DURATION OF A MISSION?



WHAT COULD THIS MEAN FOR FUTURE MISSIONS?

- Current seismicity models lead to **low detection probabilities (< 12.5 %)** for short duration missions.
- However, we have **large uncertainties** behind the predicted infrasound amplitudes **due to the choice of seismic velocities, attenuation, and atmospheric scaling.**
- **Several research questions should be addressed** to constrain the range of detectability:
 - a) How would amplitudes extracted from full-waveform simulations [4] affect detectability?
 - b) How sensitive are the predicted amplitudes on the choice of Venus subsurface models?
 - c) Can a balloon network vs a single balloon increase the detectability likelihood?
 - d) How accurately can we constrain the crust/mantle velocities from low-SNR infrasound?

[1] Krishnamoorthy, S. & Bowman, D. C. "A "Floatilla" of Airborne Seismometers for Venus." (2023) [10.1029/2022GL100978](https://doi.org/10.1029/2022GL100978)
 [2] van Zelst, Iris, et al. "Estimates on the possible annual seismicity of Venus." (2023) [10.31223/X5DQOC](https://doi.org/10.31223/X5DQOC)
 [3] Sabbeth, L., et al. "Estimated seismicity of Venusian wrinkle ridges based on fault scaling relationships." (2023) [10.1016/j.epsl.2023.118308](https://doi.org/10.1016/j.epsl.2023.118308)
 [4] Martire, Léo, et al. "SPECFEM2D-DG, an open-source software modelling mechanical waves in coupled solid-fluid systems: the linearized Navier-Stokes approach." (2022) [10.1093/gji/ggab308](https://doi.org/10.1093/gji/ggab308)

Coming, soon!
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