

# Vertical profiling of greenhouse gas mixing ratios above a coastal marsh using a laser heterodyne radiometer

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(Atmospheric Radiation Measurement)

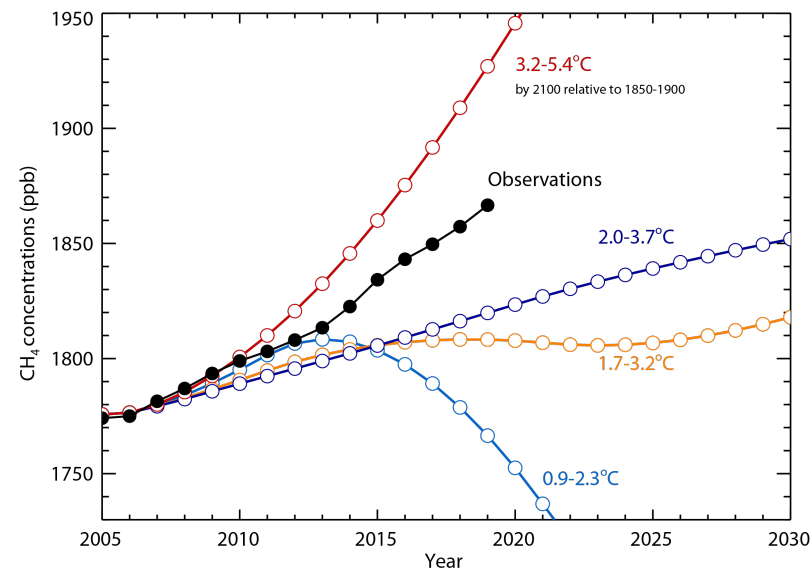
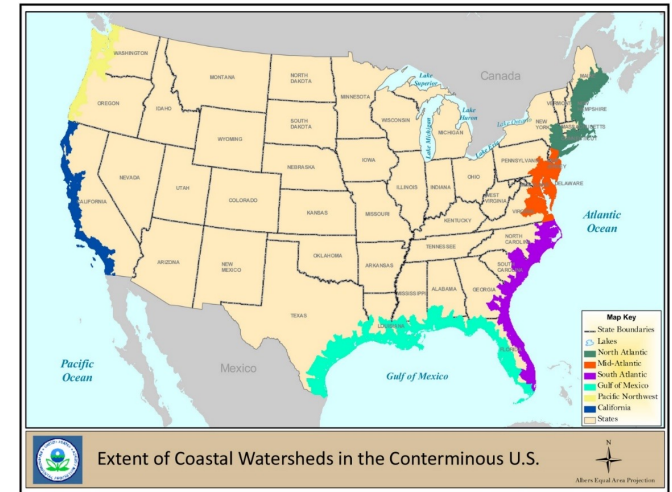


# Atmospheric Diagnostics

- Many options for “Point Sensors”
  - Ground: good temporal resolution, lousy spatial coverage
  - Aircraft: good temporal, slightly better spatial
- LIDAR, microwave sounding, TCCON
- Satellites
  - Typically report column averages
- Reliance on modeling to fill in the holes

# Wetlands and Carbon

- **Inland** wetlands: largest natural source of CH<sub>4</sub> emissions – 15-45%
  - Negligible contribution from *coastal* wetlands
- **Coastal** wetlands: net carbon sinks
  - ~300–600 Gt stored carbon
  - 5.1 °C Temp increase → >2x *coastal* wetland emissions
- Large uncertainty in land-air exchange
  - Complicated by carbon cycle
  - Data used in models are difficult to scale up

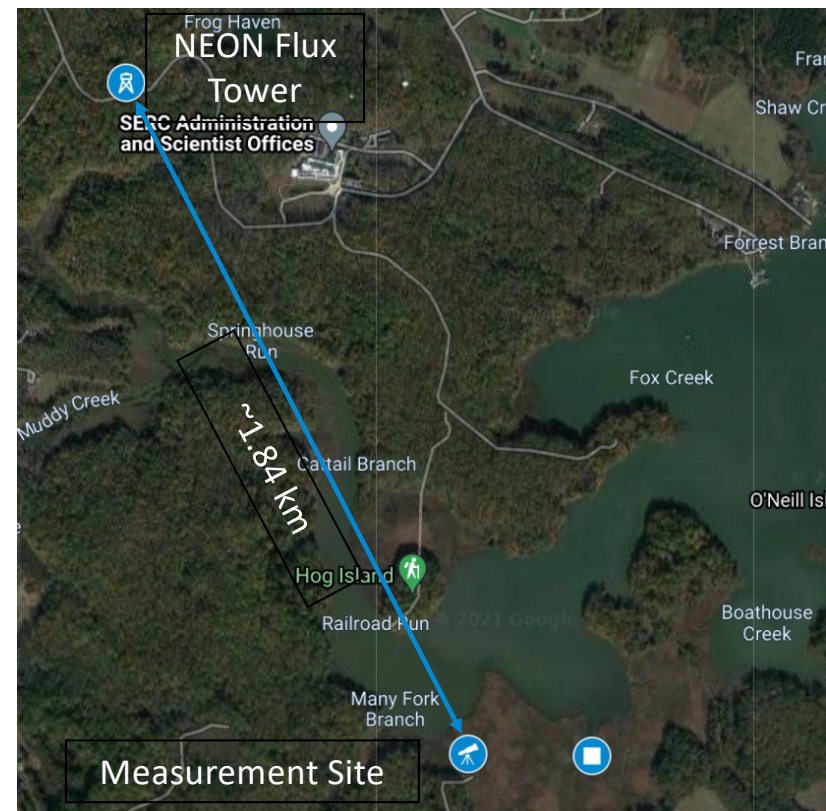
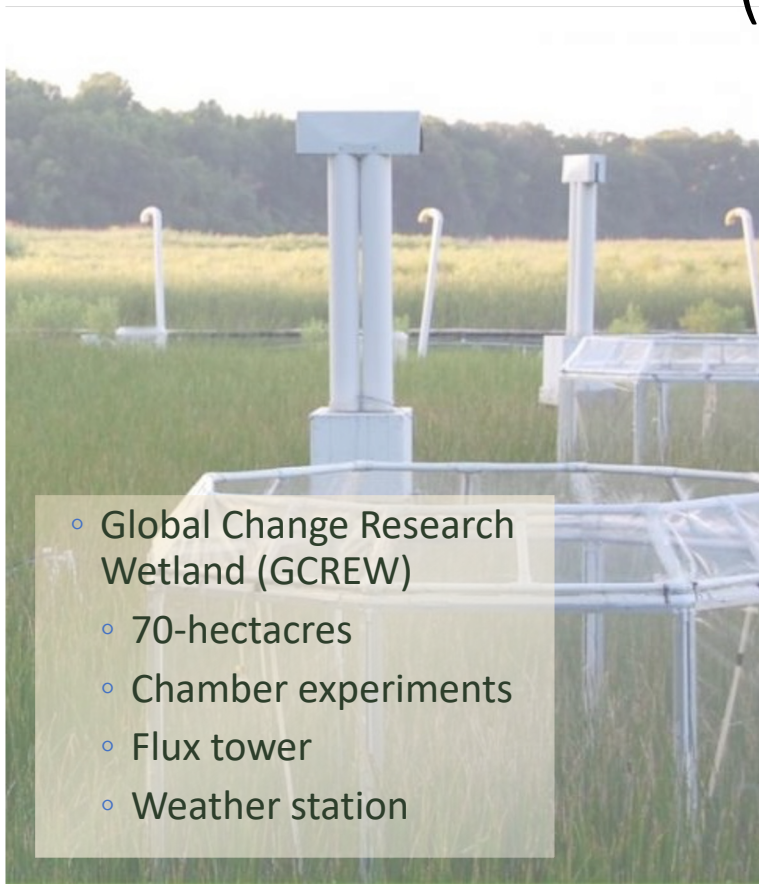


Global Carbon Project (2022).  
Global Methane Budget, 2022.  
[Updated from Saunio, et al.  
2016, ERL].

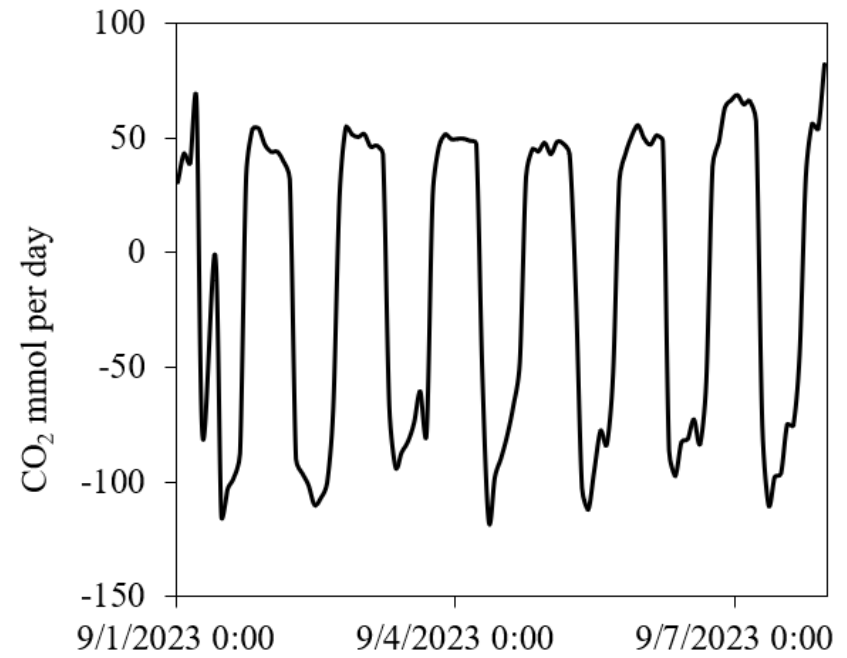
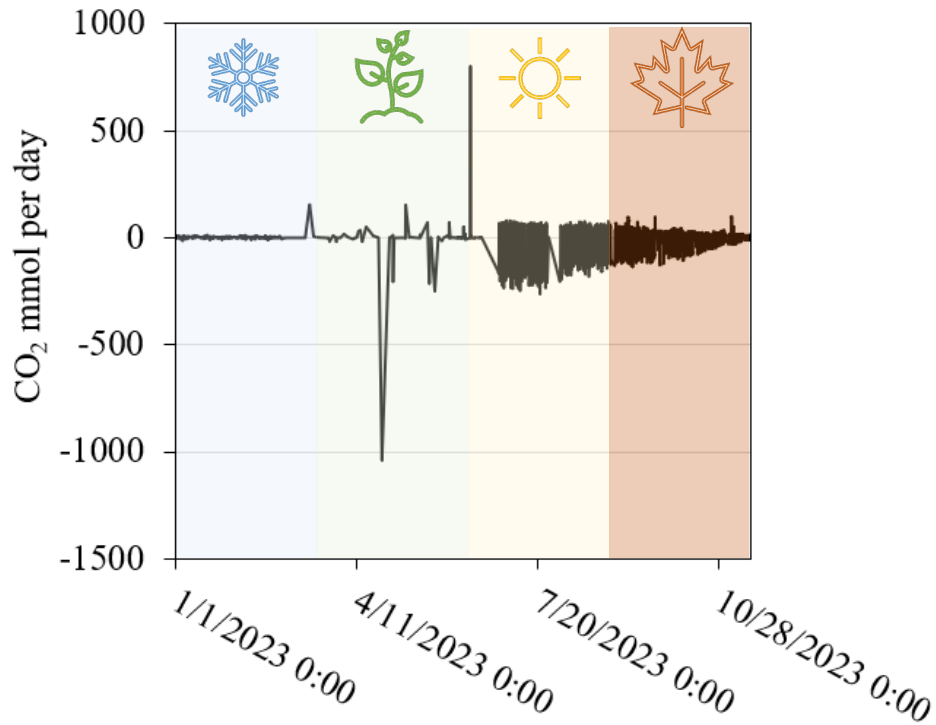
# Smithsonian *Environmental Research Center*



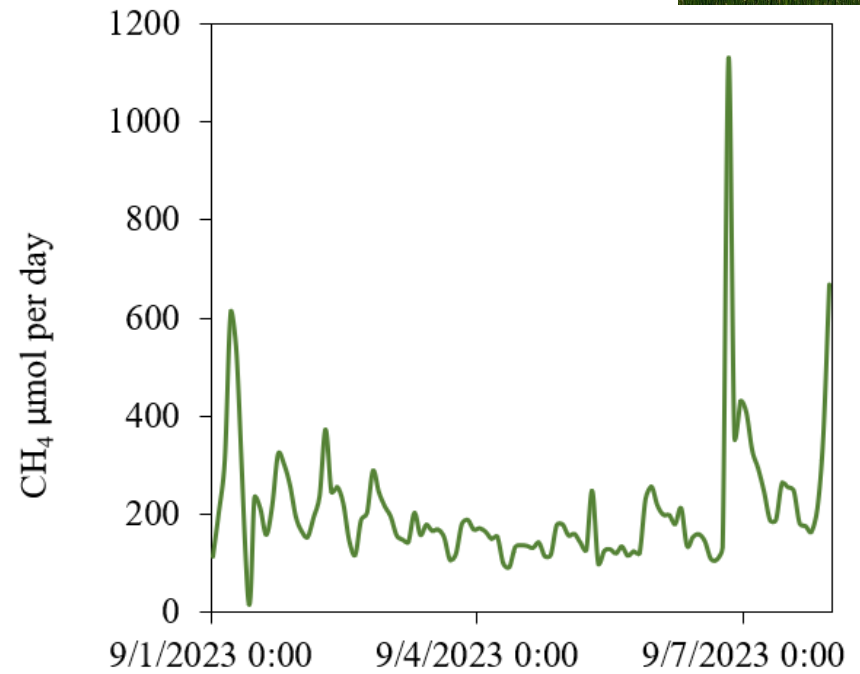
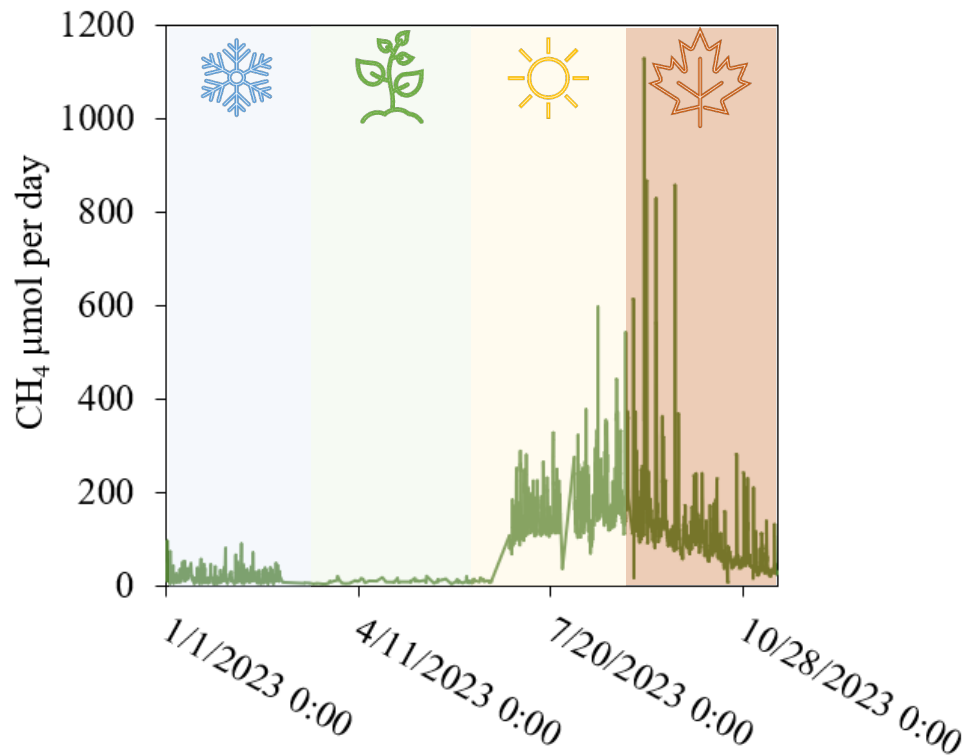
# Smithsonian Environmental Research Center (SERC)



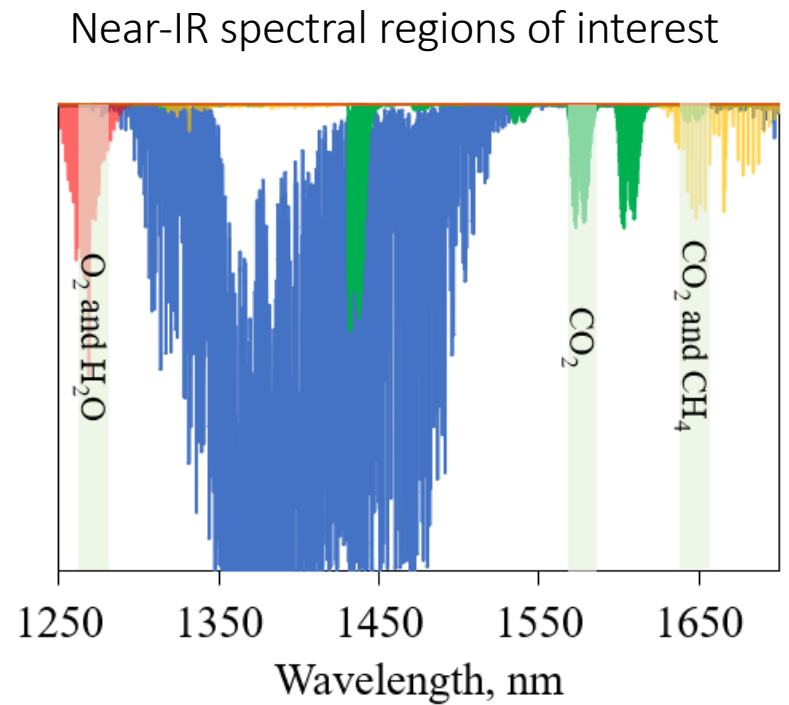
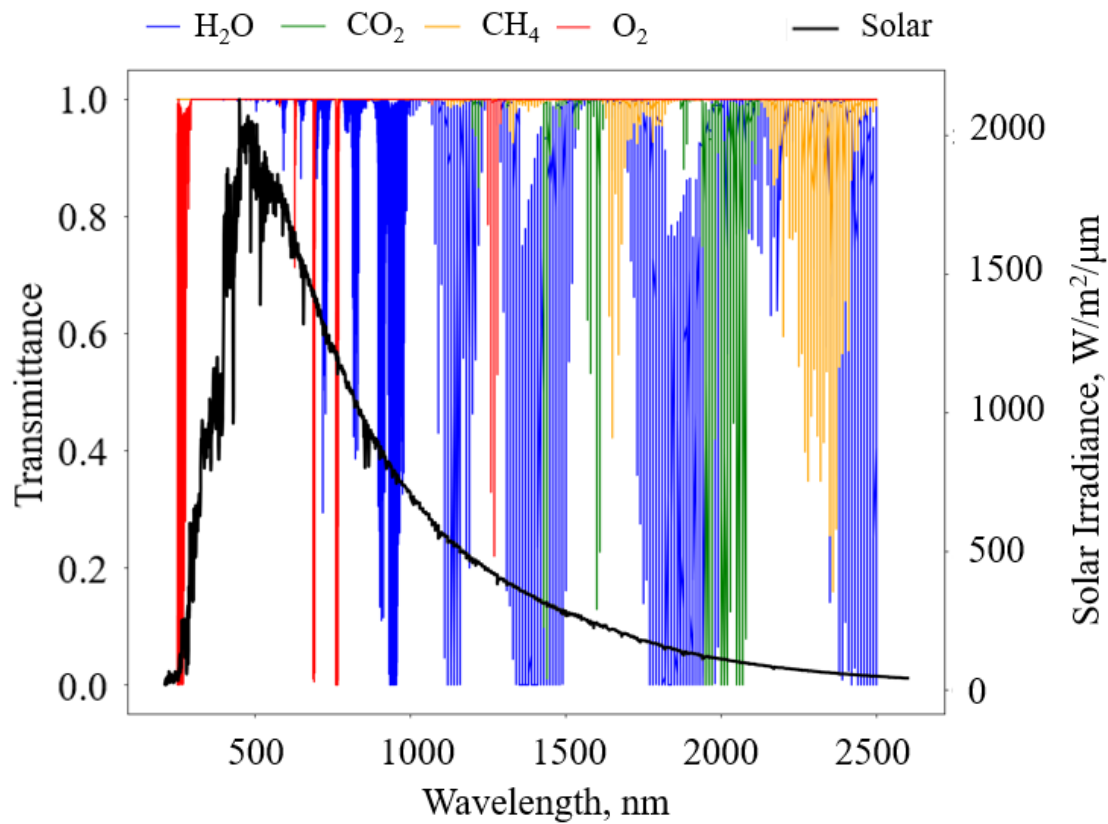
# GCREW CO<sub>2</sub> Chamber Flux Data



# GCREW CH<sub>4</sub> Chamber Flux Data



# Greenhouse Gas Absorption Spectroscopy



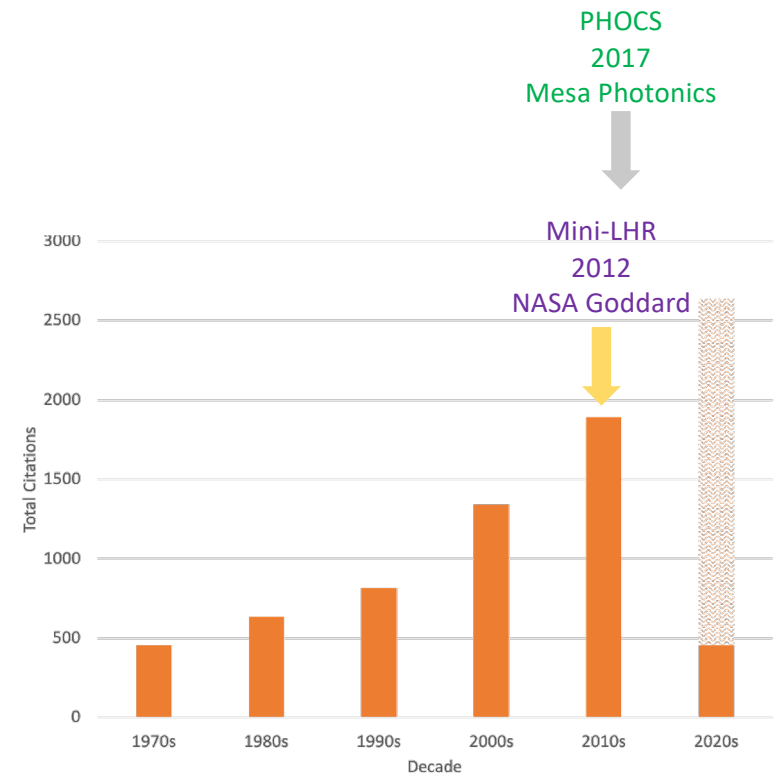


# Laser heterodyne radiometry (LHR) (a very brief history)

Robert Menzies  
(1970s)

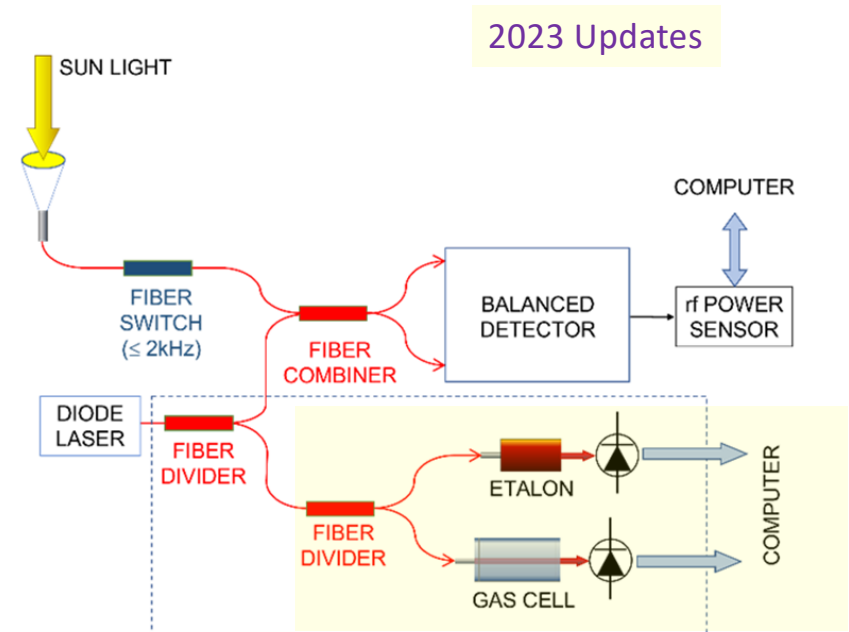


- “Remote sensing with infra-red heterodyne radiometers”, Opto-electronics volume 4, pages179–186 (1972).
- “Remote detection of SO<sub>2</sub> and CO<sub>2</sub> with a heterodyne radiometer“, Appl. Phys. Lett. 22, 592 (1973).
- “Laser heterodyne detection techniques.” in Laser Monitoring of the Atmosphere, pp 297–353 (1976).



# Laser heterodyne radiometry (LHR)

- Adapted from and analogous to radio receiver technology.
- Incoming light is combined with light from a narrow-band laser source (the local oscillator or LO) on a photodetector.
- The detector output contains AC electronic signals at the (optical) difference frequencies.
- PHOCS heterodyne signals are proportional to the solar spectrum at the LO wavelength.
- When the LO coincides with an optical absorbance, the heterodyne signal intensity will drop by an amount proportional to the absorbance.



# Laser heterodyne radiometry (LHR)



- It is an atmospheric absorption technique measuring the sun light intensity as a function of light frequency at the surface.
- There are lasers involved, but no light laser light leaves the instrument.
  - *“PASSIVE” laser technique.*
  - High spectral resolution of the laser and radiofrequency electronics enable vertical profiling.
- The characteristics of light absorption depend on pressure and temperature which change along the path through the atmosphere.
  - *We need a detailed model of atmospheric structure to interpret the spectra.*

# LHR installation at SERC/GCREW

## 0.6 m Rack

### Components:

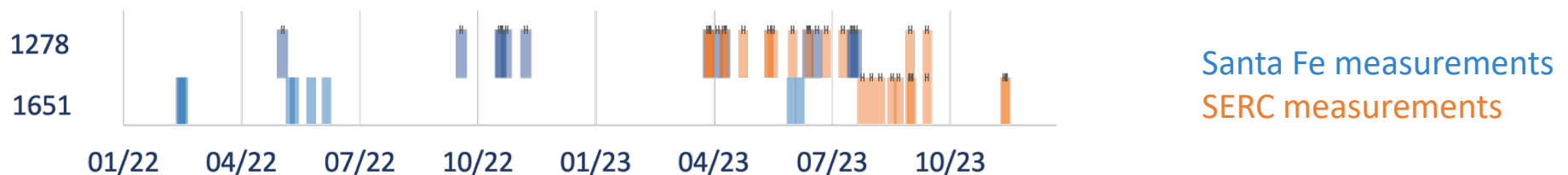
1. Power distribution & USB subsystems
2. Solar tracker electronics
3. CO<sub>2</sub>/CH<sub>4</sub>
4. O<sub>2</sub>/H<sub>2</sub>O
5. (Future?) Open Path Instrument



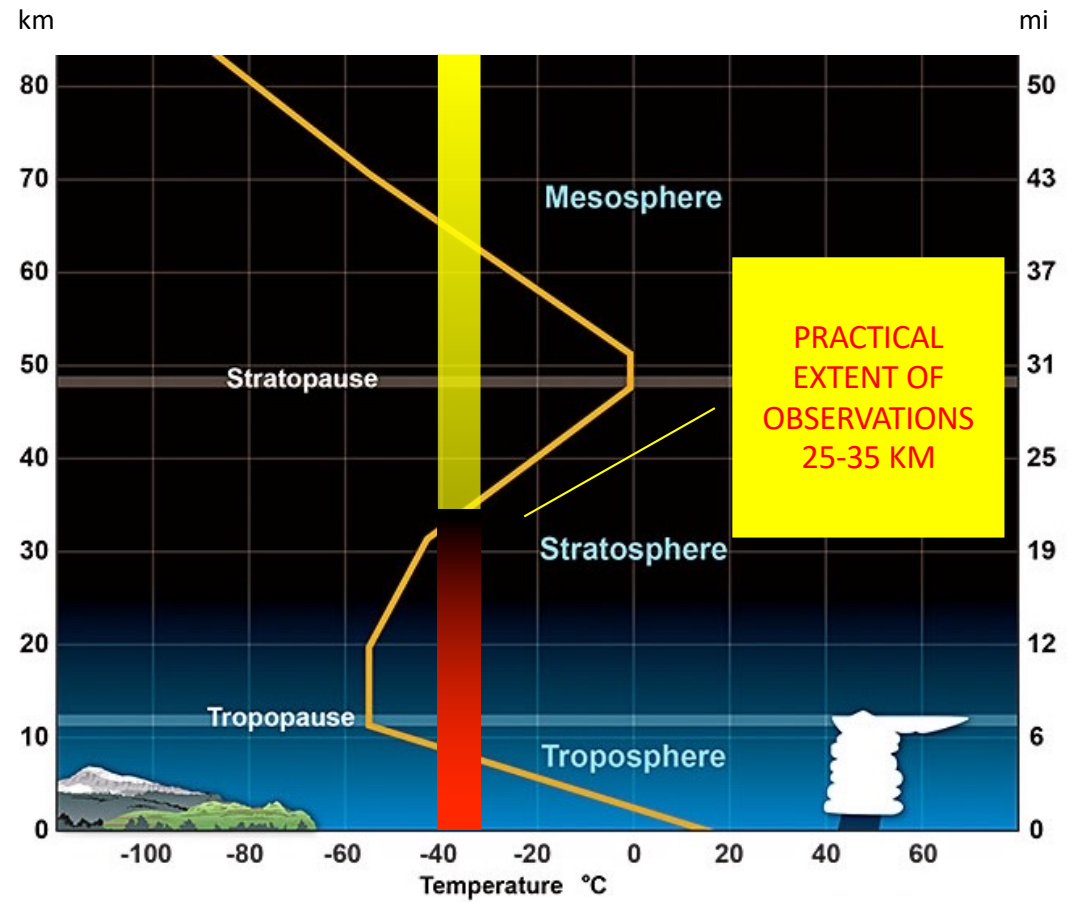
# LHR at SERC/GCREW

Not with out some bumps in the road...

- The Plague
  - Access limitations at GW or SERC.
  - Supply chain: 14-month delay in delivery of dome from the UK
  - Some assembly required.
- First deployed instrument was a “Franken-Fox” combination of parts from GW, NASA, and Mesa Photonics.
- First new heterodyne unit (at 1278 nm) delivered in Spring 2023
- Second (1651 nm) delivered in Summer 2023

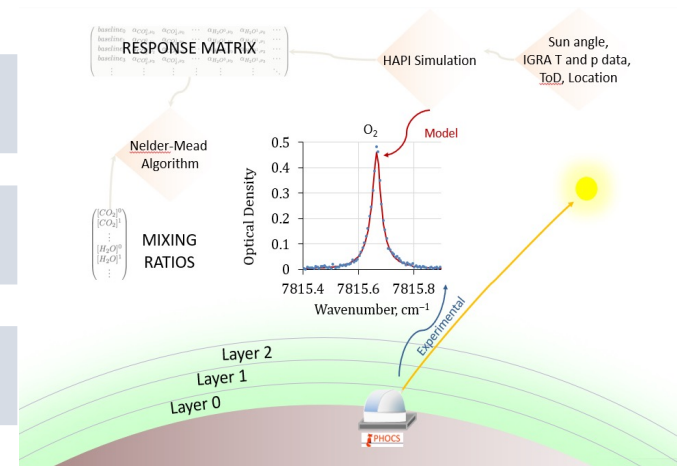
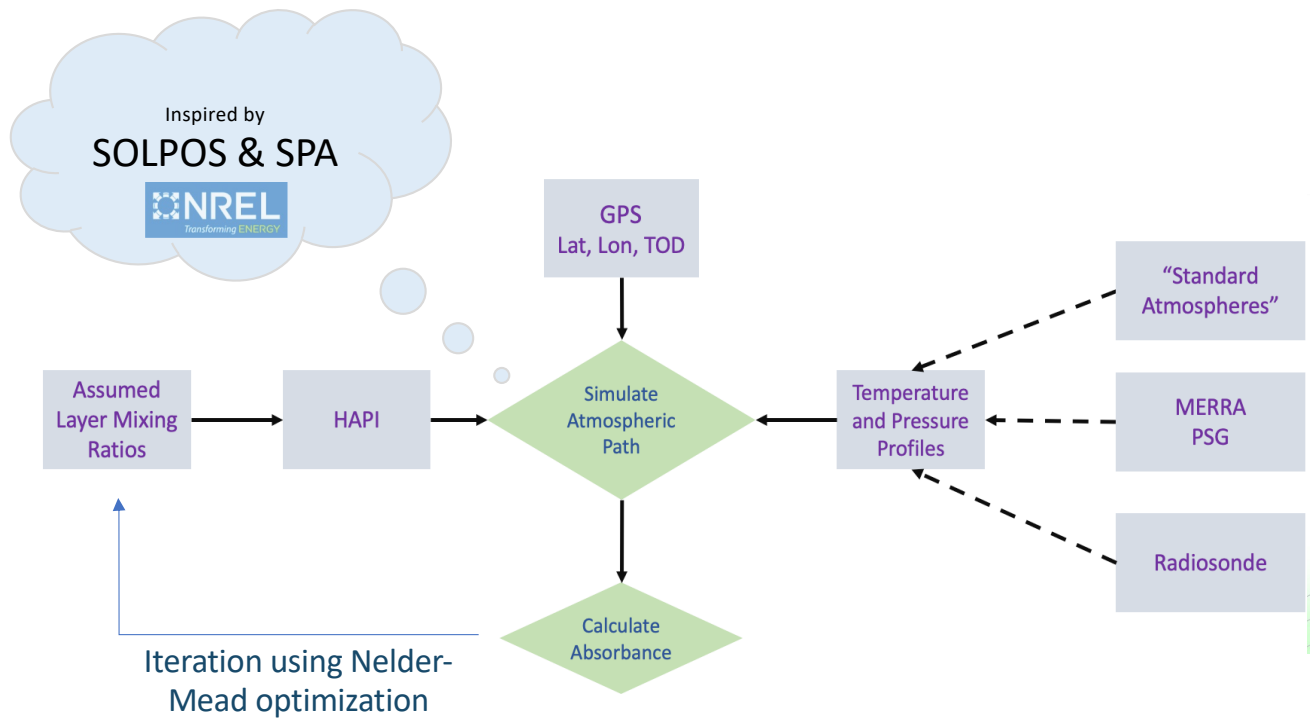


# The Atmosphere



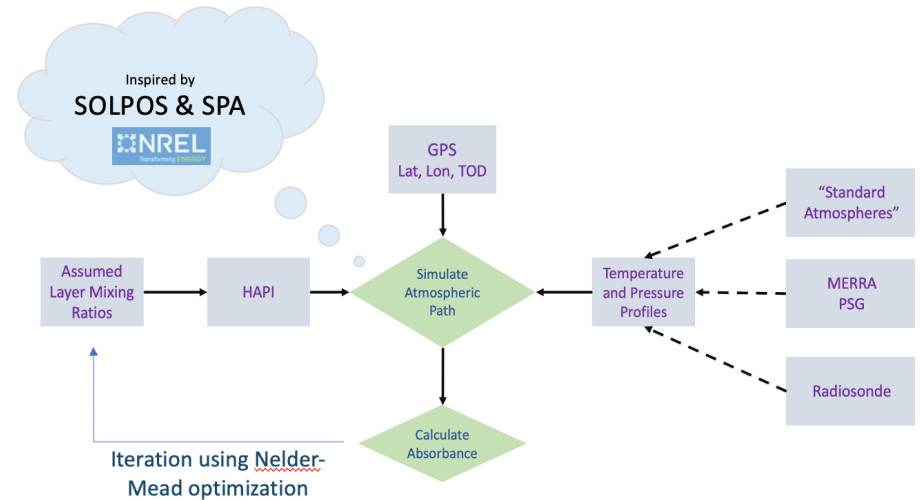
<https://www.weather.gov/jetstream/layers>

# Vertical Profiling Method



# 1651 nm Retrieval (redux)

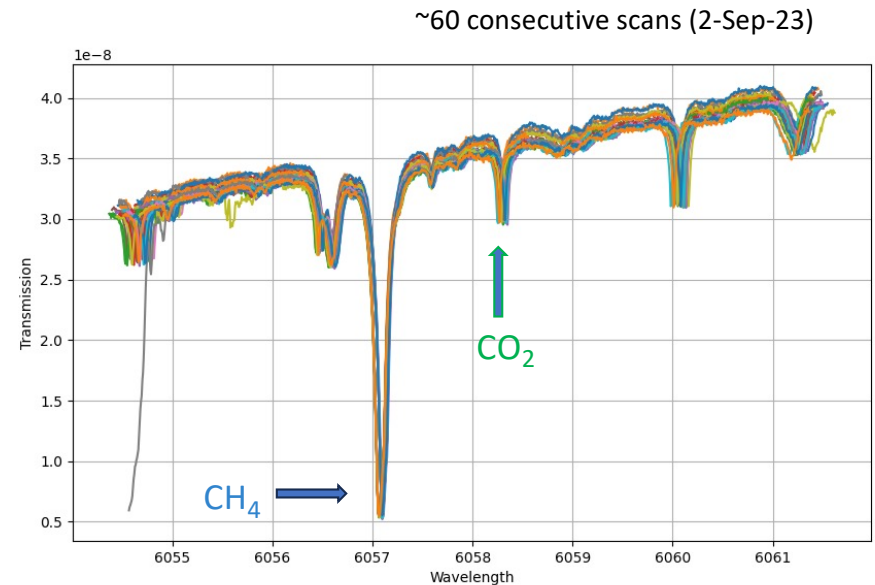
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2. Calibration
3. Absorbance Spectra
4. Isolate features from CO<sub>2</sub> and CH<sub>4</sub>
5. Refine Calibration Using *LahetraSim*
6. Calculate spectrum-specific absorption path coefficient matrix
7. Fit using Nelder-Mead or Truncated Newton Method





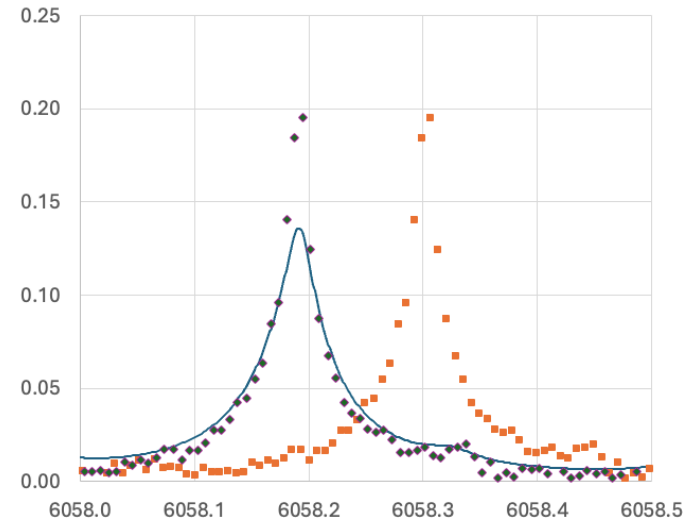
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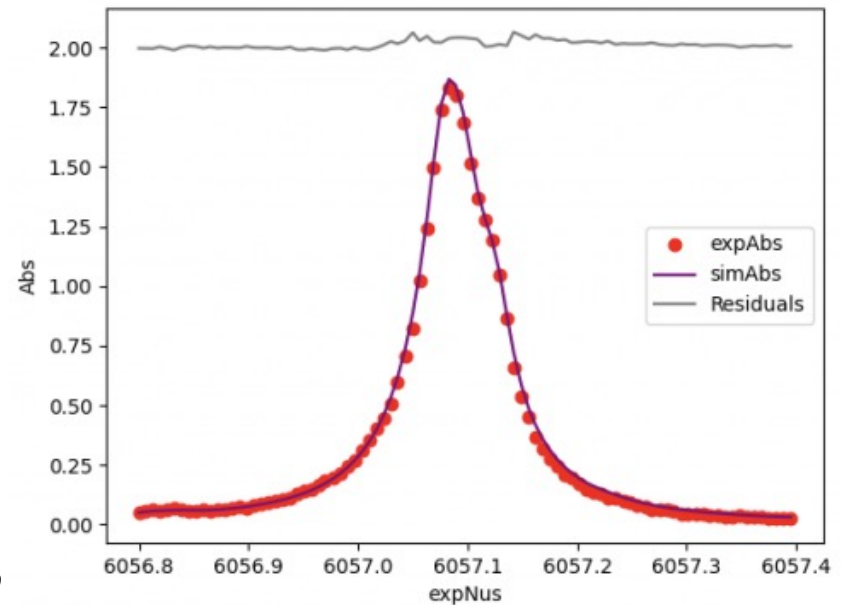
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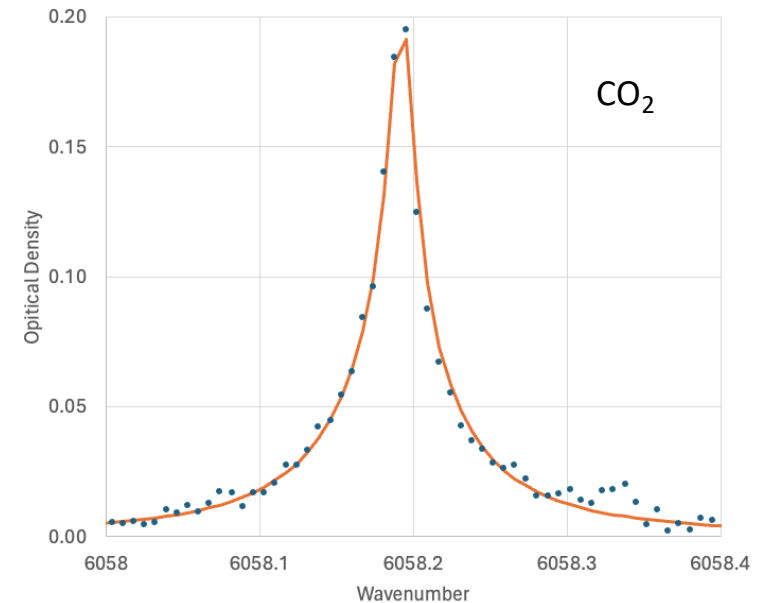
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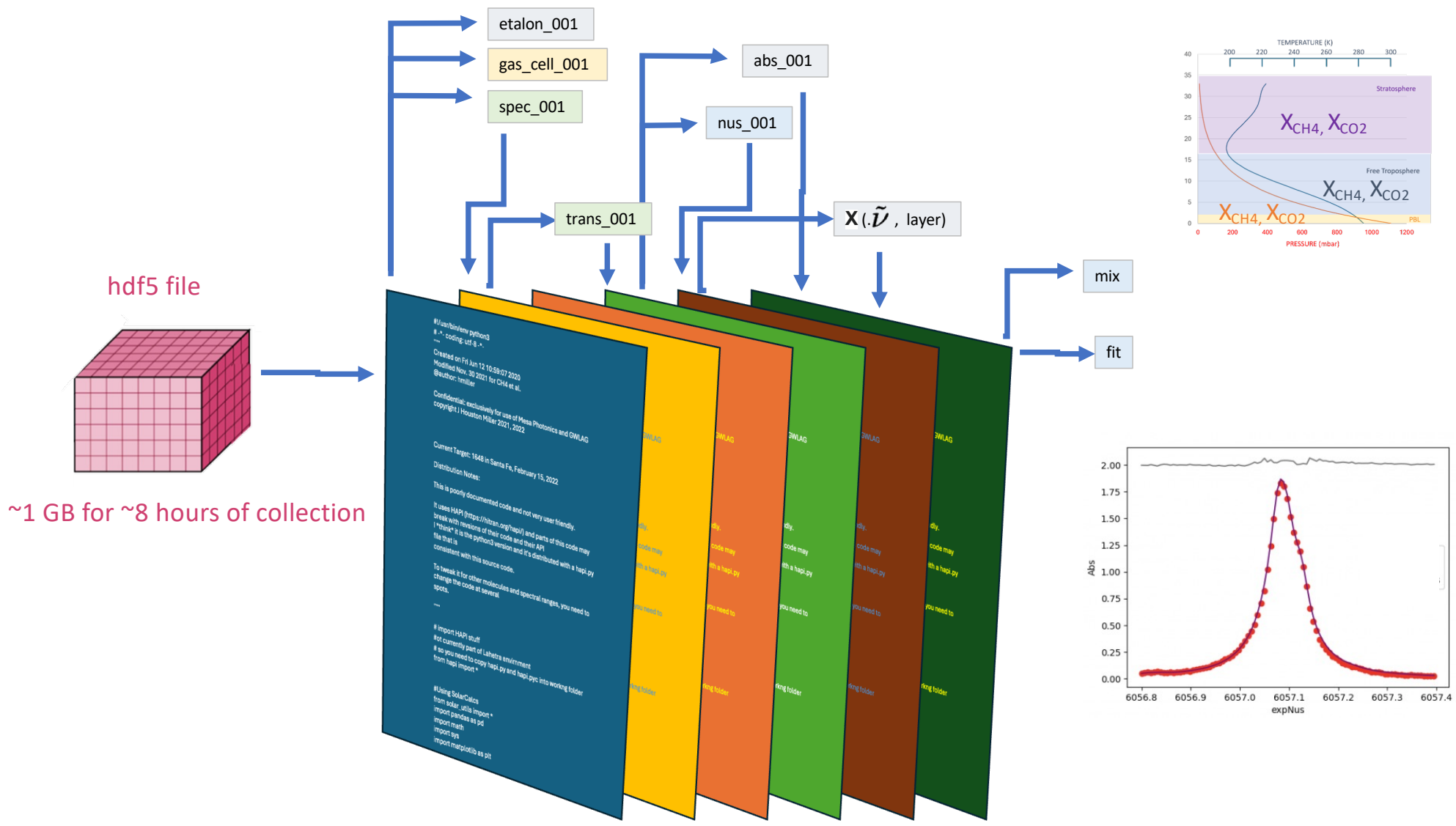
# Where are we (and where are we going)?

## The Good

- Arguably (one of) the best performing LHR instrument we are aware of.
  - Good performance (SNR > 100) at 100 MHz (0.0033 cm<sup>-1</sup>) spectral resolution

## The not so Good

- Current data record is temporally sparse and difficult to draw many meaningful scientific inferences
  - We need to do better in moving to autonomous operation
- Retrieval is still too cumbersome.



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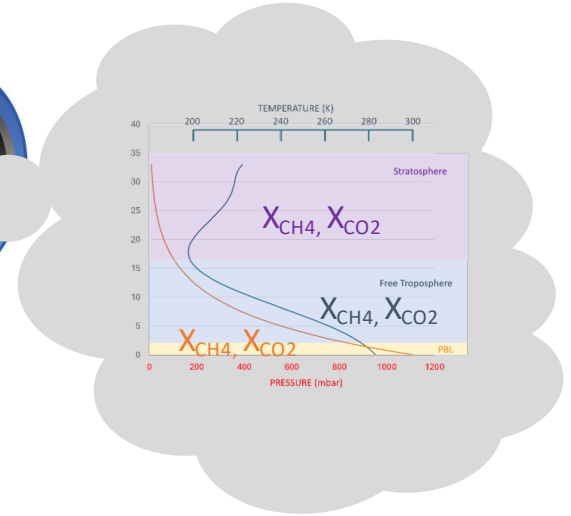
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  - The goal is *Drag and Drop* individual data records into a program queue

LHR Spectral Data,  $f(\text{time})$   
Temperature & Pressure,  $f(\text{time}, \text{location})$ ,  
GIS, local, surface weather





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- Retrieval is still too cumbersome.
  - The goal is *Drag and Drop* individual data records into a program queue
  - We don't yet fully understand the uncertainty in the retrievals: How sensitive are the fits to individual layer mixing ratios?
  - Absolutely, we need to constrain layer mixing ratios to deliver meaningful results and useful "priors" are difficult to come by.
  - Not as sensitive to (very near) surface emissions as we would like.
    - Averaging over the boundary layer
    - Co- location with one or more mixing ratio and/or flux *point* experiments would add valuable insight.

# What we didn't have time to talk about

- Santa Fe, New Mexico!
  - Lower humidity (1278 nm data) and higher elevation
- A novel way to refine temperature and pressure value using a trio of O<sub>2</sub> features (1278 nm data)
- Hydrogen Fluoride detection in the stratosphere (1278 nm data).

# Acknowledgements

- Smithsonian Environmental Research Center
  - Roy Rich
- Mesa Photonics and GWU were supported under a joint Small Business Technology Transfer grant from the Department of Energy, DE-SC0019543.

