

Variations of upper thermospheric scale height based on neutral mass density measurements from coplanar low-Earth-orbit satellites Heikki Vanhamäki¹, Lei Cai¹, Anita Aikio¹, Marcus Pedersen¹, and Milla Myllymaa¹ 1. Space Physics and Astronomy, University of Oulu, Finland Contact: lei.cai@oulu.fi

Background

The mass density scale height H_{ρ} in the upper atmosphere is a measure of the vertical distance over which the neutral mass density ρ decreases by a factor of e (the base of natural logarithms). The change in H_{ρ} indicates changes in several key parameters of the upper atmosphere, including neutral temperature and composition.

The mass density scale height can be determined by the density measurements from a pair of co-planar satellites, where two satellites with similar inclinations are in the same local time sector but at different altitudes. Since 2014, multiple low-Earthorbit (LEO) satellite missions have provided neutral mass density measurements, including Swarm, GRACE, and GRACE-FO. Coplanar events have been identified as shown in the figure below.



Neutral mass density data

- Neutral density data products from ESA's Swarm Dissemination Server:
- DNS-ACC: every 10 s - DNS-POD: every 30 s

Satellite	DNS-ACC	DNS-POD
Swarm-A	N/A	Х
Swarm-B	N/A	Х
Swarm-C	Х	Х
GRACE-A/B	Х	N/A
GRACE-FO1	Х	N/A

Theory

Assuming hydrostatic equilibrium for an ideal gas, the neutral mass density scale height is defined as

$$H_{\rho} = -\frac{\rho}{d\rho/dz} \tag{1}$$

where z is the altitude. Deriving from Eq. (1), we get

$$\ln\frac{\rho_A}{\rho_B} = \int_{z_A}^{z_B} \frac{1}{H_\rho} dz \qquad (2)$$

where z_A and z_B are the altitudes of satellites A and B, respectively, and ρ_A and ρ_B are the neutral mass densities measured from the two satellites.

Taking the average of the integrand, the height-averaged reciprocal of the density scale height can be expressed as

$$HM(H_{\rho}) = 1/\left\langle\frac{1}{H_{\rho}}\right\rangle = \frac{z_B - z_A}{\ln(\rho_A/\rho_B)} \quad (3)$$

Here HM stands for harmonic mean.

Results





- Comparison with different empirical and simulation models.
- NRLMSIS 2.1
- DTM2020
- JB2008
- OTHITACS/TIEGCM
- Example: Event 2 in Oct-Nov
 - Divided into three latitudinal sectors.
- Averaged with a 6-h time window.
- Dawn sector
 - Smaller values at low latitudes from the measurements.
 - Prestorm increases associated
 - with substorms.
 - JB2000 reproduces better the Increases and decreases during storms
 - OTHITACS/TIEGCM gives a different trend from measurements and models
- **Dusk sector**
- Higher values than in the dawn sector.
- JB2000 gives a better performance.
- OTHITACS/TIEGCM gives smaller values than others.

Discussion and conclusions

- coplanar LEO satellites.
- storm activity.
- Comparison with different models. - JB2008 shows a similar storm dependence as measurements
- The scale heights from MSIS and DTM-2020 also show a dependence on storms, but the variations are less significant than those seen in JB2008.
- results.

This work is part of the ESA-funded project "Joule heating effects on ionosphere-thermosphere coupling and neutral density (JOIN)". The Python package GeospaceLAB is used to collect data and visualize the results.





Neutral mass density scale height is estimated using the measurements from

The scale height depends on latitude, local time, season, and geomagnetic

- None of the tested models can extract small-scale structures comparable to the measured

The role of the neutral temperature and composition changes in the observed scale height variations needs to be further investigated.

Acknowledgements



