Amplitudes of Magnetopause Surface Waves: **Comparison of THEMIS Observations with MHD Theory**

A. Pöppelwerth (1), N. Grimmich (1), R. Nakamura (2), F. Plaschke (1) (1) Institute of Geophysics and Extraterrestrial Physics, TU Braunschweig, Germany (2) Space Research Institute, Austrian Academy of Sciences, Graz, Austria a.poeppelwerth@tu-braunschweig.de | Phone: +49 531 391-5211

1. Motivation

- Earth's magnetopause (MP) in constant dynamic motion: e.g. solar wind variations and structures or velocity shear between magnetosheath and magnetosphere \rightarrow Waves on MP (e.g. Plaschke et al., 2009; Archer et al., 2024)
- Surface waves can be described by superposition of two evanescent waves – one on either side of MP
- According to magnetohydrodynamic (MHD) theory, amplitude ζ should **decrease exponentially** with distance *x* normal to the MP (e.g. Plaschke & Glassmeier, 2011; Plaschke, 2016):

$$\zeta(x) = \zeta_0 e^{ik_N}$$

• Comparison of S/C observations with MHD theory

2. Methods & Data • Magnetic field B and ion velocity V_i from $\swarrow_{i=1}^{2}$ **THEMIS** mission Σ • LMN - coordinate system with model MP (Shue et al., 1998)

- Phase velocity V_{ph} , period τ , and wave number k of MP wave from MP
- crossing (MPC) of S/C:
- Minimum Variance Analysis (MVA) for local normal and wave propagation direction
- \rightarrow Consecutive MPCs for τ
- > $V_{\rm ph}$ approximated with $V_{\rm i}$

Data Processing:

- Shift data from different S/C according to cross correlation
- Apply 60s **smoothing**
- Calculate analytic signal to get **amplitude**
- Calculate distance between S/C in N-direction



Technische Universität Braunschweig

References:

Plaschke, F., et al. (2016), Steepening of waves at the duskside magnetopause, Geophys. Res. Lett., 43, Archer, M. O., et al. (2024). Magnetopause mhd surface wave theory: progress challenges, Front. Astron. Space Sci. 11:1407172, https://doi.org/10.3389/fspas.2024.1407172 7373–7380, https://doi.org/10.1002/2016GL070003 Plaschke, F. and Glassmeier, K.-H. (2011). Properties of standing Kruskal-Schwarzschild-modes at Shue, J.-H., et al. (1998), Magnetopause location under extreme solar wind conditions, J. Geophys. Res., the magnetopause, Ann. Geophys., 29, 1793–1807, https://doi.org/10.5194/angeo-29-1793-2011 103(A8), 17691-17700, https://doi.org/10.1029/98JA01103



Time: 14:44:16 ---Time: 14:48:22 magnetosphere modified from Plaschke et al. (2016) Distance $[R_F]$ 5. Summary & Outlook Comparison of observation and theory:

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3. Case I -5.6 • Event on 2007-08-23 • all S/C observe low frequency Ğ −7.2 oscillations in V_i and B THB 10⁴ [nT] В [km/s] 10016:00 17:00 15:00 -THA -THB -THC -THD -THE



reference magnetopause

- 1) Case: agreement
- 2) Case: disagreement

Plaschke, F. (2016). ULF Waves at the Magnetopause. In Low-Frequency Waves in Space Plasmas (eds A. Keiling, D.-H. Lee and V. Nakariakov). https://doi.org/10.1002/9781119055006.ch12

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Possible reasons:

- Properties of MP wave not good enough determined
- MHD Theory only limited applicable
- Other sources of waves

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What's next?

• Expand case study to statistical study with further multi-spacecraft observations String-of-Pearl Configuration: 1) THEMIS (Coast Phase ~ 70 Events) 2) MMS (current campaign)

