

The Hubble OPAL Program: 10 years of time-variable phenomena on Jupiter and the other giant planets (invited)

Supplementary Materials

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Introduction

The Outer Planet Atmospheres Legacy (OPAL) program began in 2014 as part of the Hubble 2020 legacy initiative (Simon et al. 2015). These observations were meant to cement long-term legacy of the Hubble Space Telescope (HST) by ensuring a regular cadence of giant planet observations to fill temporal gaps between individual programs (Fig. 1). The giant planets have highly dynamic atmospheres, so long-term trends tied to seasonal or other evolutionary cycles require regular data collected using the same instruments and filters.

In addition to building up a long data base of consistent observations on an annual cadence, serendipitous discoveries have been made along the way. Filters extend from the near-UV (F225W at 225 nm) to the near-IR (FQ889N at 889 nm), and each planet is imaged to cover all longitudes over a period of two planetary rotations. All raw data are immediately available to the public, and the team also hosts high

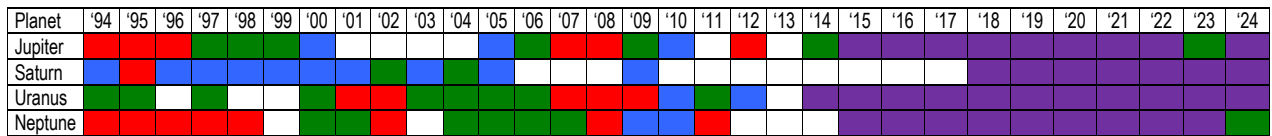
level science products in the form of global maps at the MAST Archive (Simon 2015).

In the References section, we list 64 publications based on OPAL data that were compiled using the official [MAST bibliographic search](#) tool, plus 6 recent additional publications not yet indexed by MAST. An automated search result list is maintained by the lead author on a [UCB-SSL page](#).

A review paper on this topic is currently in preparation, so the OPAL team encourages any authors not included in the references below to provide citations to their work using OPAL data.

OPAL at Jupiter

Figure 2 provides some examples of studies enabled by Hubble's exquisite spatial resolution and OPAL's global and temporal coverage. OPAL data have been central to detailed studies of Jupiter's



Color code: white: no data, blue: individual images, green: wind pairs, red: single maps, purple: two global/near-global maps

Figure 1. Temporal coverage in high-resolution UV/visible-wavelength coverage of the outer planets had significant gaps prior to the start of OPAL in late 2014 (with Saturn starting in 2018 after the end of the Cassini mission).

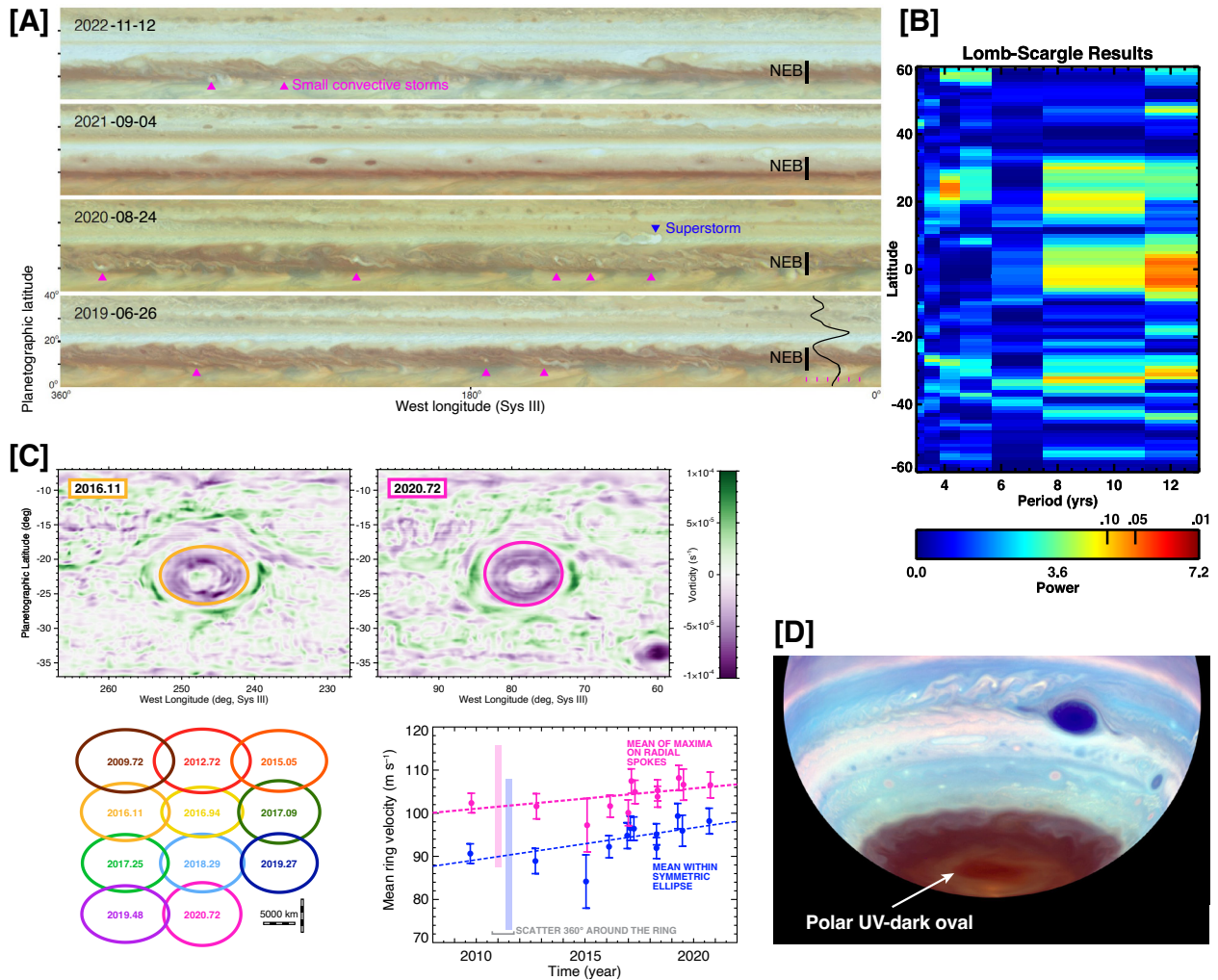


Figure 2. HST’s high spatial resolution and stable photometry are central to OPAL investigations of time-variable phenomena on Jupiter. **[A]** OPAL observed simultaneous changes in storm activity and the widths and colors of belts and zones, with anomalous conditions in 2021 (Simon and Wong 2024). **[B]** A strong 4.5-year periodicity in the 395-nm brightness near 23°N shows up in periodogram analysis, associated with superstorm eruptions like the 2020 event seen in panel A (Simon and Wong 2024). **[C]** The 2D wind field of the Great Red Spot (GRS) shows a dynamically “hollow” core (top two frames), irregular changes in shape and gradual decrease in size (lower left), and a slight increase in mean wind speeds (lower right; Wong et al. 2021). **[D]** UV-dark ovals in the polar hoods of stratospheric aerosols (like the northern dark oval seen in Cassini data, Porco et al. 2003) are more common in the south (Tsubota et al. 2025).

vortices including the Great Red Spot (Simon et al. 2015/2018/2024, Loeffler et al. 2016, Carlson et al. 2016, Bjoraker et al. 2018, Sánchez-Lavega et al. 2018/2021b, Iñurrigarro et al. 2020, Barrado-Izaguirre et al. 2021, Wong et al. 2021, Morales-Juberías et al. 2022, Anguiano-Arteaga et al. 2023, Sankar et al. 2024, Harkett et al. 2024), high speed narrow wind jets (Simon et al. 2015/2022, Hueso et al. 2017, Tollefson et al. 2017, Johnson et al. 2018, Marcus et al. 2019, Wong et al. 2020, Rogers et al. 2022), small atmospheric waves and turbulence (Simon et al. 2015/2018/2022,

Cosentino et al. 2017/2019, Tollefson et al. 2017, Fletcher et al. 2018, Giles et al. 2019), long-term color trends and aerosol layer variability (Medikoa et al. 2017, Tollefson et al. 2017, Braude et al. 2020, Pérez-Hoyos et al. 2020, Dahl et al. 2021, Fry and Sromovsky 2023, Simon and Wong 2024), and UV-dark ovals in the polar hoods (Tsubota et al. 2025).

Space missions

OPAL data have extended the science return of several space missions, with Jupiter observations commencing one year before Juno arrived at Jupiter. OPAL wind and cloud structure measurements have been used in diverse analyses of phenomena, from deep zonal atmospheric structure revealed by microwave emission (Oyafuso et al. 2020, Fletcher et al. 2021), to convective cycles in cyclonic vortices (Iñurriagarro et al. 2020/2022, Hueso et al. 2022). Wave, jet, and vortex features previously observed by Voyager, Galileo, and Cassini have also been studied in greater detail with the long-term OPAL program (Simon et al. 2015, Wong et al. 2021, Tsubota et al. 2025).

A selection of comparative studies using OPAL data and data from spacecraft and Earth-based observatories is shown in Fig. 3.

Earth-based observatories

High-resolution visible-wavelength observations from OPAL target the planets near solar opposition to maximize spatial resolution, as do many Earth-based programs. Multi-observatory studies include correlations between cloud color from OPAL and microwave brightness from the VLA (Costentino et al. 2017, de Pater et al. 2019), comparisons between Doppler velocimetry from the ground and time-series imaging from OPAL (Gonçalves et al. 2019, Schmider et al. 2024), calibration, validation, and context for spectral measurements (Mendikoa et al. 2017, Bjoraker et al. 2018, Braude et al. 2020, Dahl et al. 2021), and deep context for stratospheric winds and aerosol anomalies observed with JWST and Gemini (Hueso et al. 2023, Giles et al. 2019, Tsubota et al. 2025).

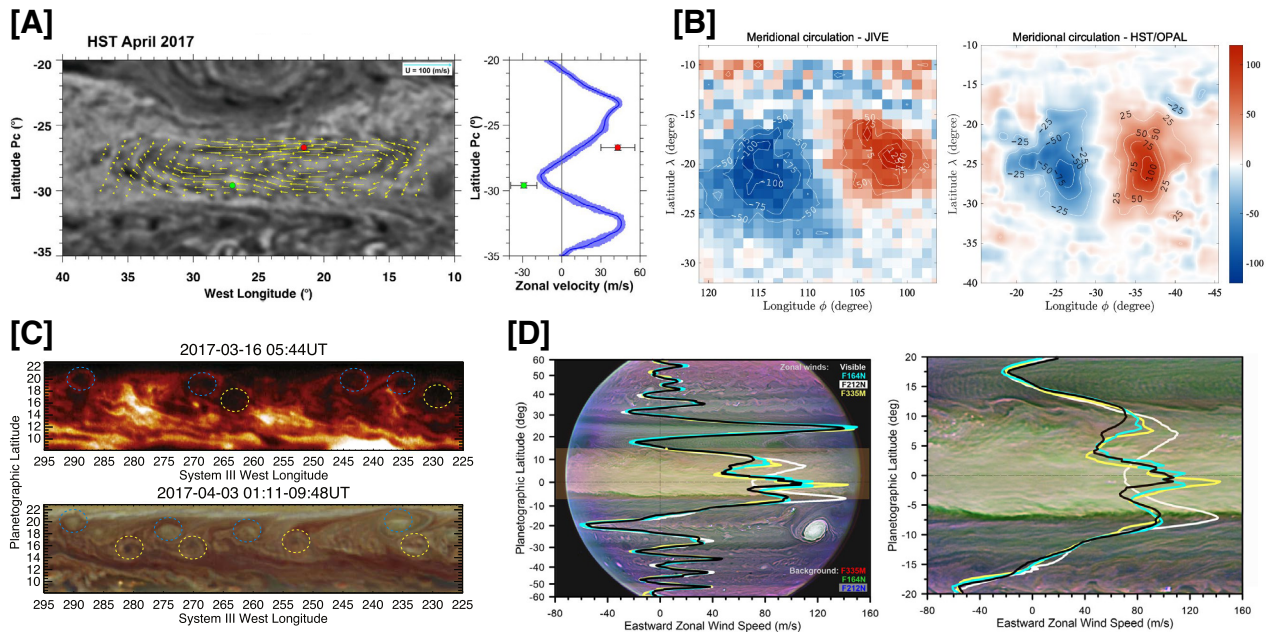


Figure 3. OPAL data provide comparisons, validation, and supplementation to other Earth-based and spacecraft datasets. **[A]** OPAL data enabled measurement of internal wind speeds in a cyclonic vortex, in images predating a convective outburst detected by JunoCam (Iñurriagarro et al. 2020). **[B]** Remarkable measurements of north-south velocities (red and blue respectively) within the GRS were performed using Doppler imaging, with OPAL data on the right used to validate the results (Schmider et al. 2024). **[C]** Transient mesoscale waves were detected in 2018 in OPAL data, as well as in thermal imaging by the VLT and Juno JIRAM. The maps here show an epoch when the mesoscale waves were detected in visible light, but were not visible in the infrared (Fletcher et al. 2018). Mesoscale wave comparisons were also made between OPAL data and images from amateur observers and the Voyager spacecraft (Simon et al. 2015/2018). **[D]** JWST imaging data revealed a narrow, high-speed upper tropospheric jet very close to the equator; OPAL data demonstrate that this high-altitude feature is completely undetected in visible-wavelength cloud-top wind tracking (Hueso et al. 2023).

Other planets, other observatories

The results cited here are a small subset of the Jupiter results achieved with the OPAL monitoring of the outer planets, with additional discoveries at Saturn, Uranus, and Neptune (some examples in Fig. 4). With more than 10 years of data in hand, and continuing for the life of Hubble, we expect the scientific return to increase geometrically. OPAL serves as a model for future long-term programs at other observatories (Wong et al. 2019, Jha et al. 2024) and guides the definition of mission science objectives (Palumbo et al. 2025).

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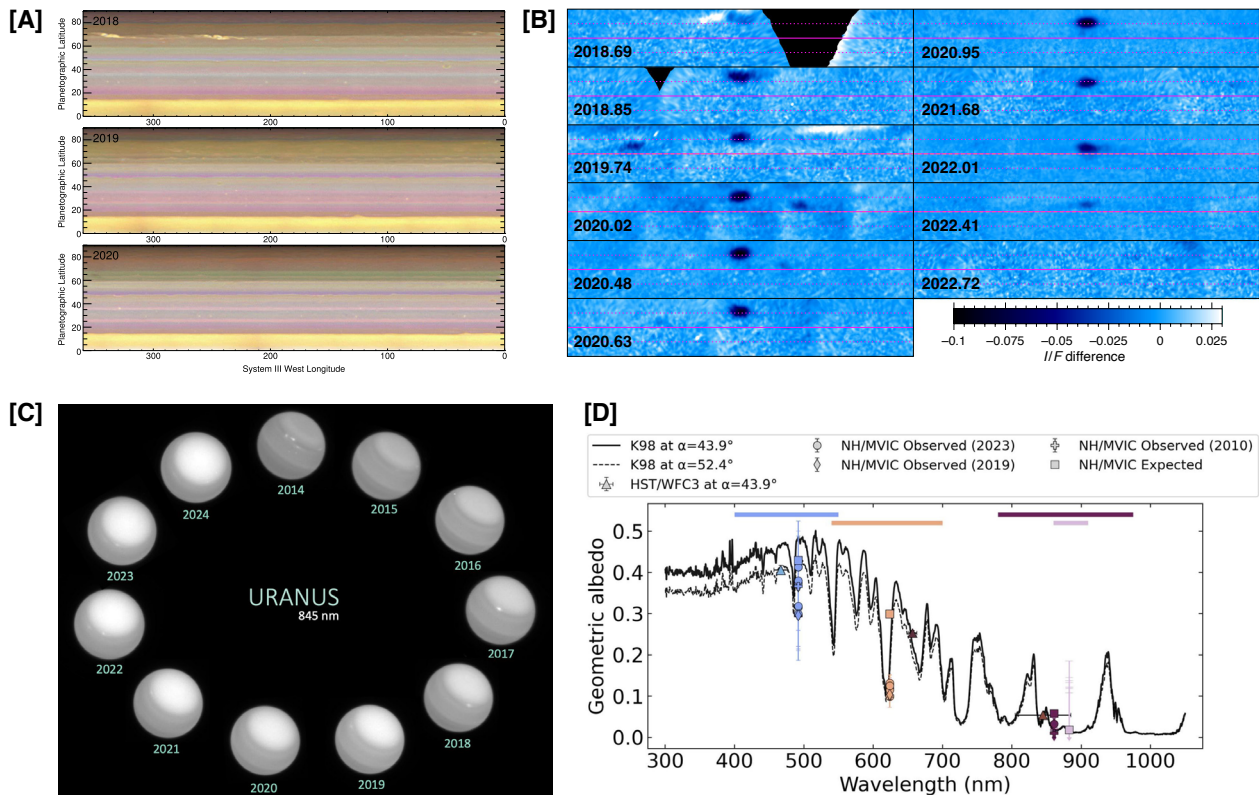


Figure 4. Many other OPAL results from Saturn, Uranus, and Neptune are beyond the scope of this Jupiter-focused report. **[A]** OPAL revealed numerous color changes in narrow banded structures at Saturn, with some potentially related to a convective outburst (Simon et al. 2021). OPAL also made new observations of Saturn ring spokes, tying to phenomena first observed by Voyager (Simon et al. 2023). **[B]** OPAL data enabled the full life cycle of a dark spot on Neptune to be observed for the first time (Simon et al. 2019, Hsu et al. 2019, Wong et al. 2022, Wong et al. 2024). **[C]** OPAL captured dramatic changes in the polar atmosphere of Uranus following solstice (Sromovsky et al. 2015, Toledo et al. 2019, James et al. 2023, Irwin et al. 2024). **[D]** Simultaneous OPAL and New Horizons observations of Uranus were used to quantify the dependence of disk-averaged brightness on solar phase angle, with the goal of providing local ground truth for interpreting directly-imaged exoplanet observations (Hasler et al. 2024).

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