Motivation and Analysis

We analyzed multimission data from the **RBSP and MMS spacecraft** as well as relevant space weather parameters, and employed **Correlation Coefficient Analysis**, aiming to identify the drivers behind the sub-relativistic electron flux variations.

- We used electron flux data from 1 keV to 500 keV from: (1) RBSP-ECT (2012 2019) for L^{*} = [2 − 6], (2) MMS-FEEPS (2015 2024) for L^{*} = [6 − 8].
- We also used data from the OMNIWeb and SuperMAG databases 21 parameters: IMF, BT, Bz, Bs, Ey, Vsw, Psw, Np, HWR, Epsilon parameter, Newell's function, DST, SMR, SME, AL, SML, AE, AU, SMU, Kp, and Ap.
- We calculated the Pearson correlation coefficient between flux and solar wind/geomagnetic parameters in three ways: (a) lin(parameter) lin(flux), (b) lin(parameter) log(flux), and (c) log(parameter) log(flux).
- We also investigated possible time-lag up to 48 hours (2 days) in the past.

Time lag set to zero





- CCs increase with increasing energy, and as energy rises, strong CCs are observed over progressively lower *L** values.
- The CCs of seed electrons with all drivers **decrease** and are confined in a smaller L^* range, except for V_{sw} .



Solar Wind Speed



- *V_{sw}* is one of the most influential driving parameters for fluxes >50keV
- Energies >200keV correlate stronger with V_{sw} after a few hour lag and until 48h.



d) Flux at 102keV (RBSP/MagEIS) & 104.78keV (MMS/FEEPS) vs Vsw

c) Flux at \$4keV (RBSP/MagEIS) & 48.31keV (MMS/FEEPS) vs Vsw

AE Index



• Seed electrons correlate stronger with the index after a few hours. This delay becomes longer with increasing energy.



Important finding

- Max CCs of AE exhibit a consistently increased time-lag by 12 hours, compared to the max CCs of Vsw.
- At $L^* = [3.5 4.5]$ the response of seed electrons to Vsw variations is almost simultaneous, while the response exhibits an increasing time delay inward and outward from the heart of the outer belt.
- The increase of the time delay below $L^* = 3.5$ is much faster.
- CCs below $L^* = 3.5$ are below 0.5.

