

Resolving Discrepancies in Electron Ring Current Models: The Importance of a Pre-Midnight Loss Process

1 GFZ Helmholtz Centre for Geosciences, Potsdam, Germany; 2 Institute of Physics and Astronomy, University of Potsdam, Germany; 3 German Aerospace Center (DLR), Institute for Solar-Terrestrial Physics, Neustrelitz, Germany

Motivation

During geomagnetic storms, surface charging effects occurring on satellites can damage their solar panels and electronics. To predict and mitigate these effects it is essential to understand the particle population causing these phenomena: the electron ring current.

In this work, we try to improve our understanding of the storm dynamics of the ring current electrons with energies of around 10 keV, which are an important factor for determining the surface charging effects.



Simplified VERB-4D simulations

VERB-4D solves for the transport of ring current particles while accounting for the loss to the atmosphere due to interactions with chorus and hiss waves:

$$\frac{\partial \bar{f}}{\partial t} = -\langle v_{\varphi} \rangle \frac{\partial \bar{f}}{\partial \varphi} - \langle v_R \rangle \frac{\partial \bar{f}}{\partial R} - \frac{\bar{f}}{\tau_{\text{wave}}}$$

Empirical models are used to specify the input parameters:

Electric field: Magnetic field: Boundary condition: **Chorus lifetimes:** Hiss lifetimes:

Volland (1973), Stern (1975) Tsyganenko (1989) Denton et al. (2015) Wang et al. (2024) Orlova et al. (2014)

Calculation of precipitating flux

To calculate the precipitating flux from the simulations, we assume a steady-state solution of the local pitch-angle diffusion equation inside the loss cone:

$$j_{
m LC}(lpha_{
m eq}) = N z_0 rac{I_0\left(z_0rac{lpha_{
m eq}}{lpha_{
m LC}}
ight)}{I_1(z_0)} \qquad z_0 = rac{2lpha_{
m LC}}{\sqrt{ au_{
m Bounce}/ au_{LT}}} \qquad N = rac{j_{
m sim}^{
m prec}}{j_{
m LC}^{
m prec}\Big|_{N=1}}$$

The flux profile inside the loss cone is normalized by the total number of particles lost in the simulation for a given time step.

Fitting of empirical lifetimes

The empirical lifetimes were created by multiplying an energy-dependent factor to the lifetimes due to chorus waves.

$$\log 10(\tau_{\text{empirical}}) = \log 10(\tau_{\text{chorus}}) - \frac{a - E}{b}$$

The empirical lifetimes were found by manually adjusting the fitting parameters a and b while comparing the trapped population against Van Allen Probes measurements





Bernhard Haas^{1,2}, Yuri Shprits^{1,2}, Dedong Wang¹, Julia Himmelsbach^{1,2}, Alwin Roy¹, and Katja Stoll^{2,3}

Results **Ring current simulations for geomagnetic storms** Model difference of Model difference of precipitating population trapped population **Fig. 1** [R_E] DUSK 2 0 -2 -4 -4 -6 $X_{SM} [R_{E}]$ $X_{SM} [R_{E}]$ log₁₀(model) - log₁₀(observations **Empirical lifetimes to estimate the missing loss** Lifetimes due to chorus waves **Empirical lifetimes Fig. 2** -1.5 1.5 -0.5 0.5 -2 -1 log10 days Simulations including the empirical lifetimes Model difference of Model difference of precipitating population trapped population **Fig. 3** DAWN used.

Haas, B., Shprits, Y. Y., Allison, H. J., Wutzig, M., & Wang, D. (2023). A missing dusk-side loss process in the terrestrial electron ring current. Scientific Reports

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log₁₀(model) - log₁₀(observations

 $X_{SM} [R_{E}]$



 $X_{SM} [R_{E}]$

Model errors observed during strong geomagnetic storms. (a) Overestimation of the trapped population compared to a Van Allen Probes satellite. The black path shows the drift trajectory of electrons. (b) Underestimation of the precipitating flux compared to the POES satellites.

Electron lifetimes for 1-10 keV electrons. (a) Original lifetimes due to chorus wave scattering. (b) Fitted lifetimes with increased loss in the pre-midnight sector

Same format as Fig. 1, but now the model with increased loss in the pre-midnight sector is

Outlook

Possible physical explanations for the missing loss

Waves not accounted for in VERB-4D:

Electrostatic Electron Cyclotron Harmonic (ECH) waves Time-domain-structures (TDS) •

Potential inaccuracies in VERB-4D regarding electron loss:

- Average behaviour of chorus waves might be inaccurate during strong storms • due to lack of statistics
- Changes in the cold plasma density are not reflected in changes of the electron lifetimes



Conclusions

- current fluxes between 10 and 50 keV
- Simulations including the variability of cold plasma density might be necessary

Haas, B., Shprits, Y. Y., Himmelsbach, J., Wang, D., Drozdov, A. Y., Szabó-Roberts, M., & Hanzelka, M. (2024). Modeling Pitch Angle Dependent Electron Precipitation Using Electron Lifetimes. Journal of Geophysical Research: Space Physics







Fig. 4

Diffusion coefficients for average cold plasma conditions and depletions at L=4.5 and MLT = 21. Statistical density is calculated based on Sheeley et al. (2001) and low density is assumed to be a factor of 5 smaller than Sheeley.

Increased loss in the pre-midnight sector is required to reproduce trapped and precipitating electron ring

The theoretical upper limit of electron scattering (strong diffusion) must be reached over a broad region

