# One-dimensional Microphysics Model of Venusian Clouds from 40 to 100 km

Impact of the Middle-atmosphere Eddy Transport and SOIR Temperature Profile on the Cloud Structure

<sup>1,2</sup>Hiroki Karyu, <sup>1</sup>Takeshi Kuroda, <sup>3</sup>Takeshi Imamura, <sup>1</sup>Naoki Terada, <sup>2</sup>Ann Carine Vandaele, <sup>2</sup>Arnaud Mahieux, <sup>2</sup>Sébastien Viscardy <sup>1</sup>Tohoku University, <sup>2</sup>Royal Belgian Institute for Space Aeronomy, <sup>3</sup>University of Tokyo Contact: hiroki.karyu.q4@dc.tohoku.ac.jp



# Takeaways

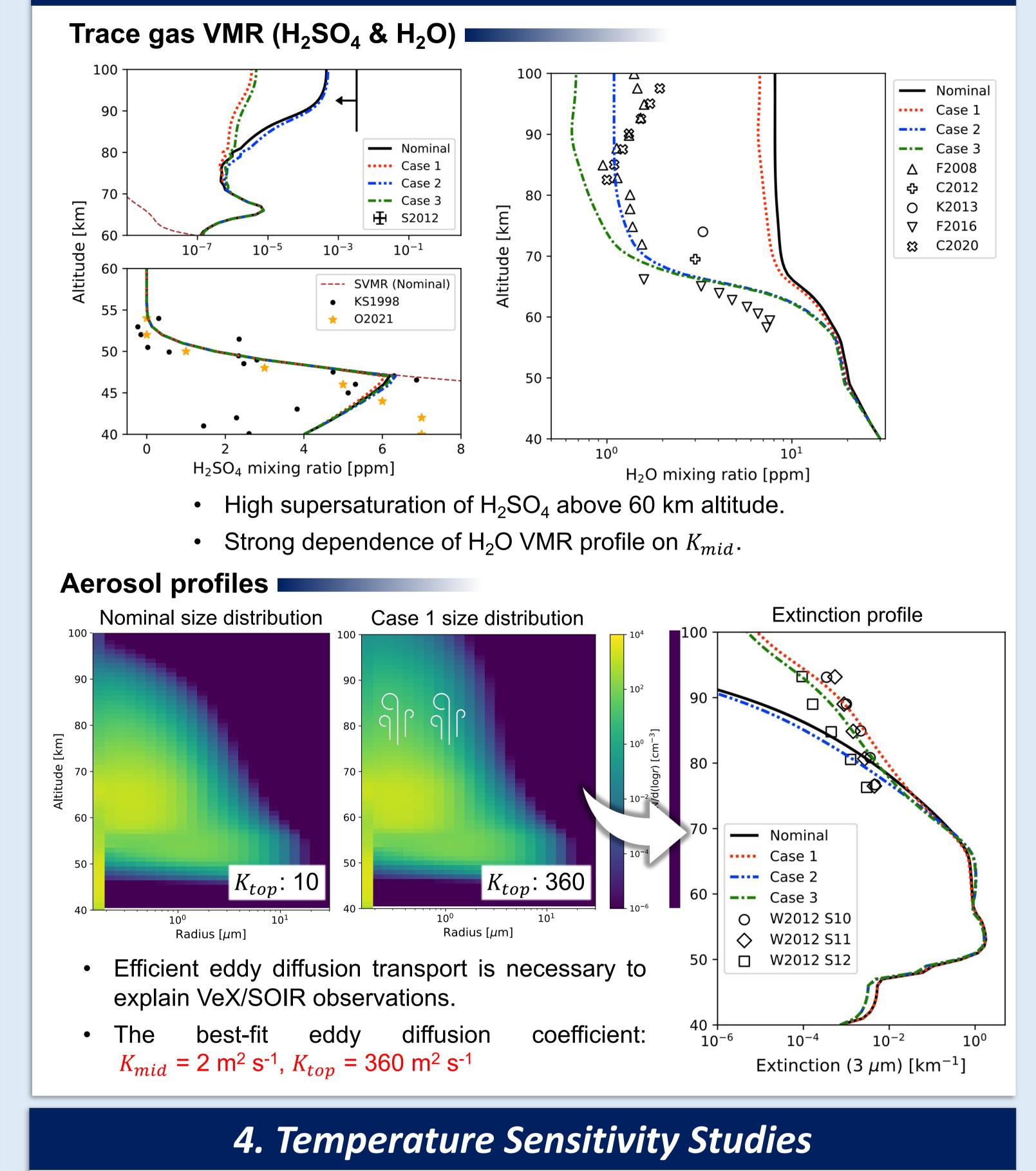
- 1. We constrained eddy diffusion profiles by reproducing vertical distributions of aerosol,  $H_2SO_4$  gas VMR, and  $H_2O$  gas VMR.
- 2.  $H_2SO_4$  vapor is highly supersaturated in the mesosphere, suggesting the possibility of sulfur source and homogeneous nucleation.
- mesospheric  $H_2O$  abundance is significantly influenced by 3. The condensation and evaporation of aerosols, suggesting that hydrogen escape is likely affected by the aerosol layers.



4. For more information, please see Karyu et al. (2024), PSJ.

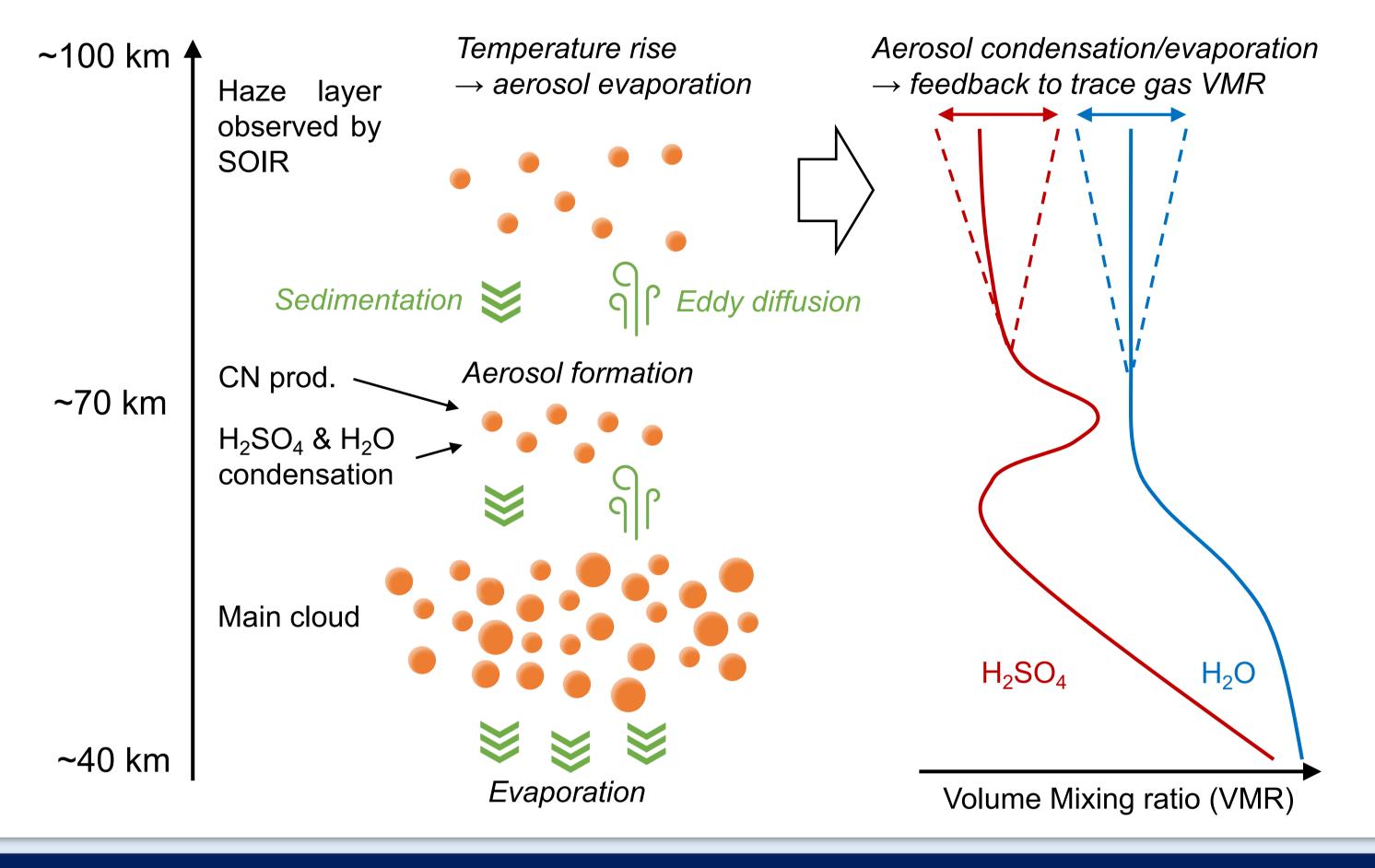
# 1. Motivation

## **3.** Eddy Diffusion Sensitivity Studies

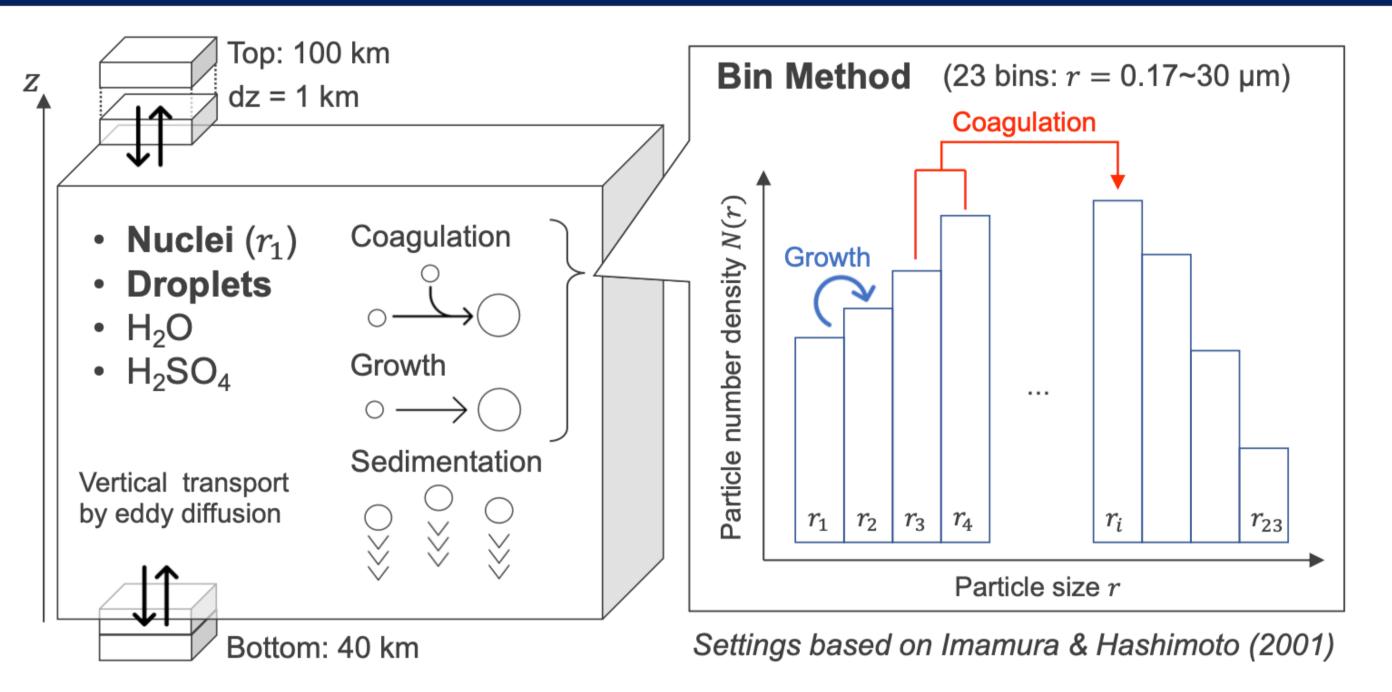


## Q. Which eddy diffusion profile reproduces aerosol, $H_2SO_4$ , and $H_2O$ distributions?

### Q. What role do the aerosols play in determining the abundance of condensational species?



## 2. Model & Settings



- The model used in this study is described in Karyu et al. (2024).
- $H_2SO_4$  and condensation nuclei (CN) production rate is prescribed by the photochemical model by Krasnopolsky (2012).
- The lower boundary conditions for CN,  $H_2O$ ,  $H_2SO_4$  are 4 × 10<sup>7</sup> m<sup>-3</sup>, 30 ppm, and 4 ppm, respectively. The upper boundary conditions are set as no VMR gradient for all species.

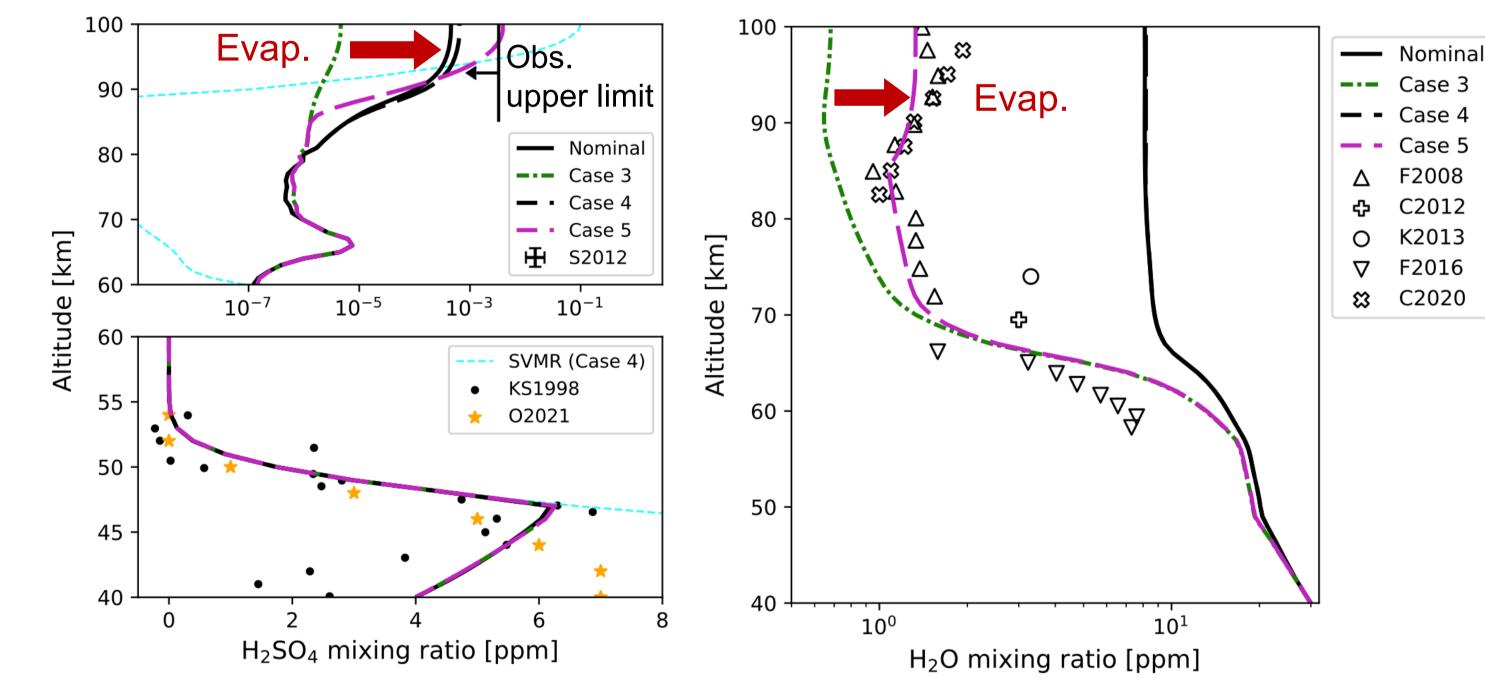
Eddy diffusion

Temperature

#### Sensitivity studies

• 4 types of eddy

#### Trace gas VMR (H<sub>2</sub>SO4 & H<sub>2</sub>O) I

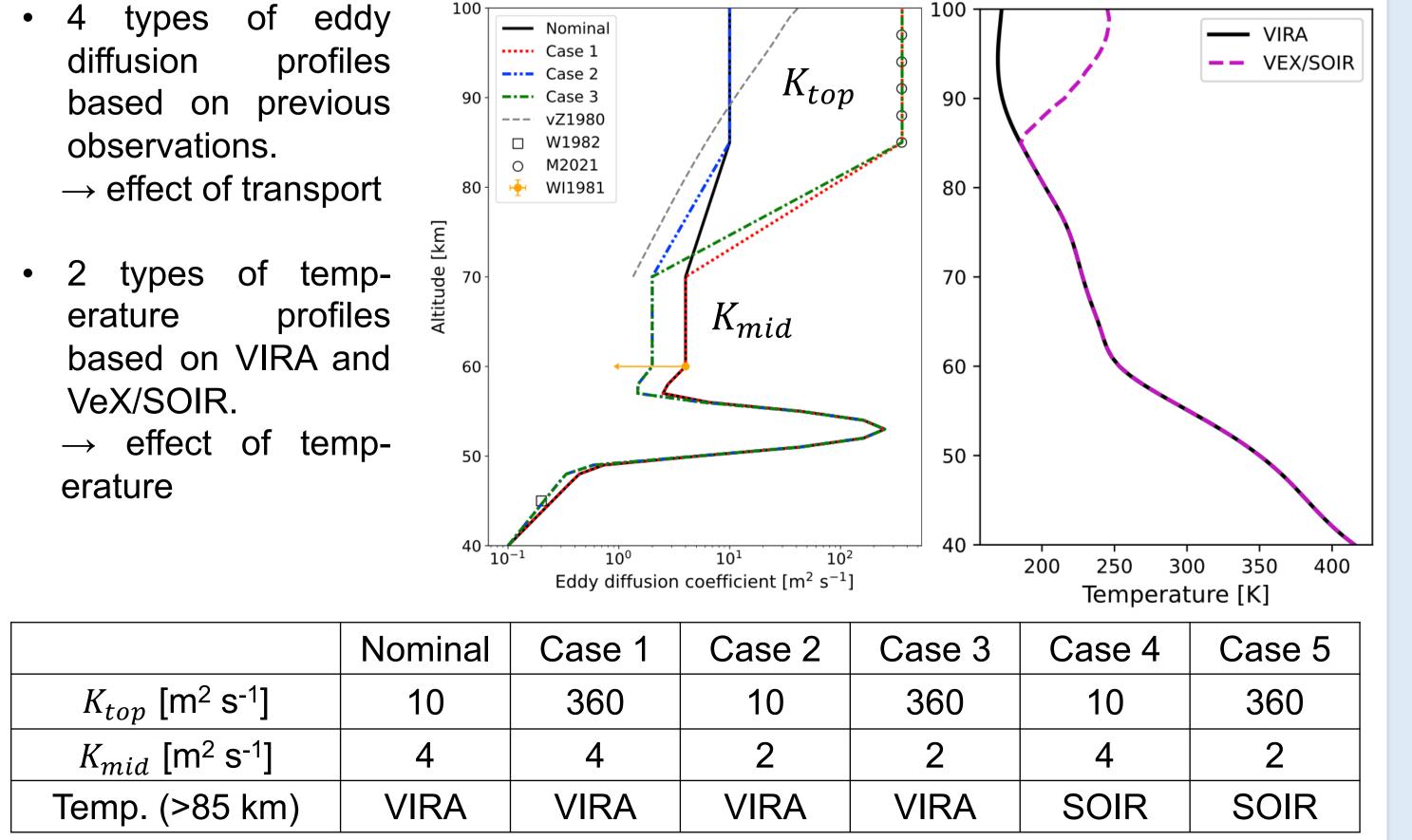


The VMR of condensational species increase due to aerosol evaporation.

## 5. Summary & Discussion

#### We found....

- Decreasing  $K_{mid}$  by half leads to tenfold decrease of the mesospheric H<sub>2</sub>O VMR.  $\rightarrow$  The H<sub>2</sub>O abundance is strongly influenced by the competition between condensation and transport at altitudes between 60 and 70 km.
- The supersaturation degree of H<sub>2</sub>SO<sub>4</sub> exceeds unity by several orders of magnitude.



• The aerosol evaporation increases the abundance of condensational gas species  $\rightarrow$  ~3 orders of magnitude for H<sub>2</sub>SO<sub>4</sub> and more than a factor of 2 for H<sub>2</sub>O.

#### Implications

Atmospheric evolution: Aerosol processes affect the H<sub>2</sub>O abundance below the homopause, and thus, the escape rate of H atoms. This effect should be considered in atmospheric evolution models.

**Cloud microphysics:** Homogeneous nucleation may be possible in the Venus mesosphere since  $H_2SO_4$  is highly supersaturated.

Sulfur chemistry: The evaporation of aerosols may serve as a sulfur source. However, the simulated value is <3 ppb, which is orders of magnitude lower than the value used in photochemical model (Zhang et al., 2010) to explain the observed SO<sub>2</sub> abundance by Balyaev et al. (2012).

References	Acknowledgments
Karyu+ (2024), PSJ; Imamura & Hashimoto (2001), JAS; Woo & Ishimaru (1981), Nature; von Zahn+ (1980), JGR; Mahieux+ (2021), Icarus; Seiff+ (1985), PSS; Knollenberg & Hunten (1980), JGR; Wilquet+ (2012), Icarus; Sandor+ (2012), Icarus; Oschlisniok+ (2021), Icarus; Kolodner & Steffs, (1998); Fedorova+ (2008), Icarus; Cottini+ (2012), PSS; Krasnopolsky (2013), Icarus; Fedorova+ (2016), Icarus; Cottini+ (2020), Icarus; Dai+ (2020), JGR; Stolzenbach+ (2023), Icarus; Zhang+ (2010), NatGeo; Balyaev+, (2012), Icarus	Graduate Program in Earth and Environmental Sciences, Tohoky University (GP-FES), and the