



Polyphase Amazonian floods in the Olympica – Jovis Fossae channel system



EPSC abstract

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Introduction

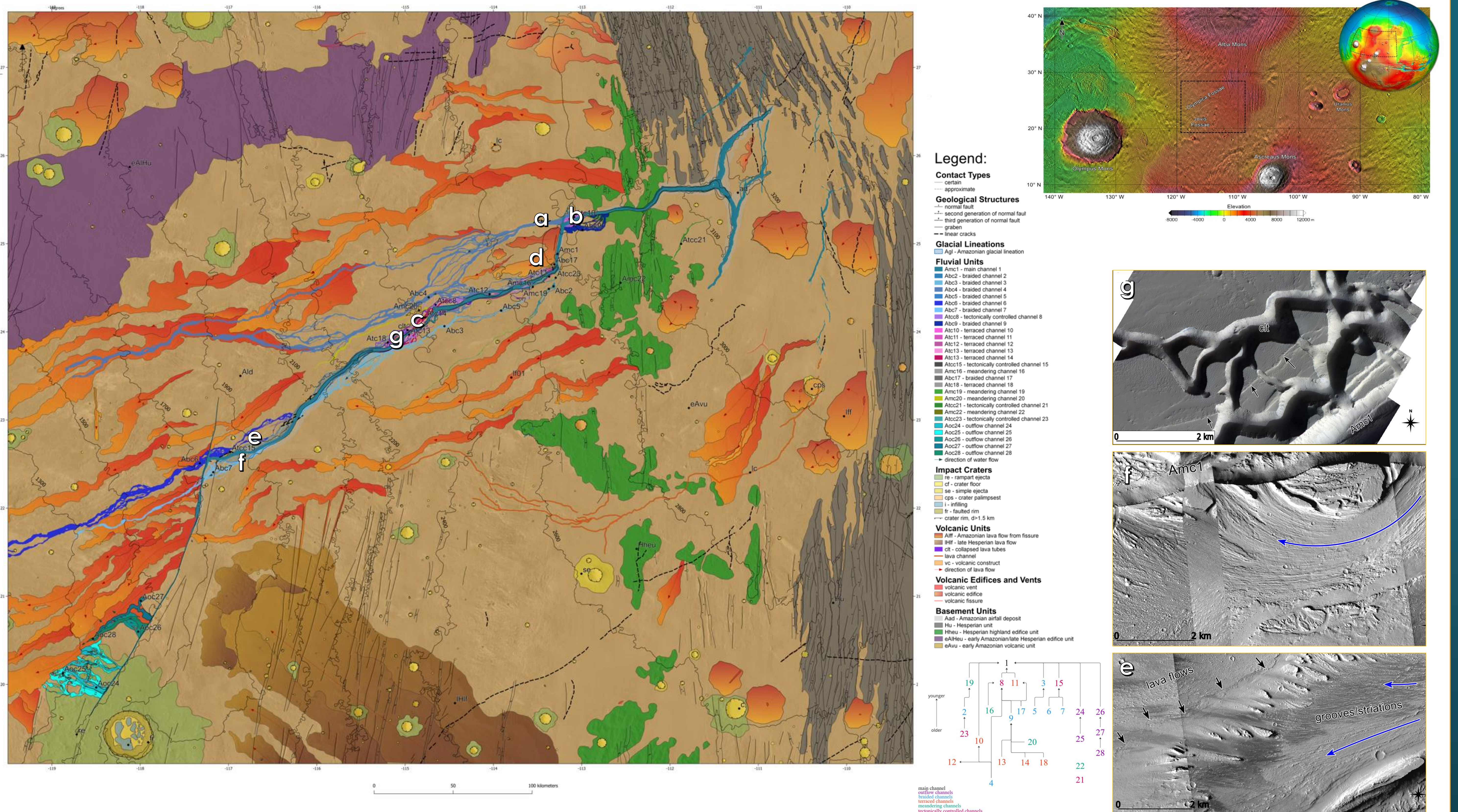
Outflow channels of the Tharsis on Mars testify to the drainage of groundwater reservoirs through magmatic and tectonic processes. We present a new geologic map of Olympica-Jovis Fossae where we mapped landforms and geological units, dated them relatively, and identified the chronological succession of outflow events and deposition of lava flows, airfall, and sedimentary units.

Cross-cutting relationships indicate a complex succession of 28 channel incisions in the volcanic stratigraphy. The six identified morphological channel types indicate different flow dynamics at different times and places. The map highlights the role of intensive volcanic activity in releasing water from overpressurized aquifers and possible subglacial flows, and will assist in a better understanding of the general conditions for catastrophic groundwater flooding in a volcanic context.

Methodology

- To map morphological and geological features we used visible imagery from the MRO CTX with a resolution of 6 meters per pixel and MRO HiRISE camera with a resolution of 25 cm per pixel.
- To constrain the intersecting relationships between channels we used the newest DEM topography data (2021 – 2023) from the Mars Express HRSC with a resolution of ~20-40 meters per pixel, and from the ExoMars Trace Gas Orbiter, Colour and Stereo Surface Imaging System with a resolution of 4.8 meters per pixel.
- To identify and map the outlines of lava flows we used the daytime infrared global mosaic from the Mars Odyssey THEMIS with a resolution of ~100 meters per pixel, in combination with the CTX images.

Results



Discussion

The map presents major groups of geological and morphological units: geological structures defined as linear cracks, grabens, and three generations of faults; volcanic features which include volcanic vents, lava flows, lava channels, and the direction of lava flows; basement units, from oldest to youngest: Noachian/Hesperian highland edifice unit, Hesperian unit, Hesperian/Early Amazonian unit, the most extensive Early Amazonian volcanic unit; **twenty-eight fluvial units** grouped into six morphological types of channels: braided systems, tectonically controlled channels, terraced channels, meandering channels, outflow channels, and the main channel; **glacial lineations**; **impact craters**; and **airfall deposit unit**.

A general investigation of the mapped units reveals the correlation between volcanic activity on Tharsis and Amazonian flooding events. The geological map displays many lava flows and volcanic edifices, partially covering the Tharsis volcanic plains, that formed after the Amazonian flooding events.

Conclusions

- The Olympica – Jovis Fossae system extends over 500 km, but the significance of its deep and narrow channels for the Tharsis dome is poorly understood.
- The observed morphological features, such as **channels with terraces and teardrop-shaped islands**, suggest that water flowed in the Olympica and Jovis Fossae channels. The existing literature on the morphology of channels and valleys on Mars indicates **long-lasting fluvial activity**. It is the most likely dominant mechanism for eroding channels and leaving terrain with typical fluvial marks.
- As the documented channels are incised in, but also partially covered by lava flow sequences, water must have flown after the emplacement of a large-scale Tharsis lava field, and before renewed lava flow emplacement.
- Our map documents not one flooding event by water, but a **prolonged history of repeated outflow** from various fissures, possibly related to an intensive period of distributed volcanism.

Acknowledgement & References

A manuscript of this work is under revision after review at the Journal of Maps as Zambrowska et al. „Geologic map of the polyphase Olympica – Jovis Fossae channel system.” This work is part of the doctoral studies of A. Zambrowska supported by the GeoPlanet Doctoral School and the Space Research Centre of the Polish Academy of Sciences. S. Poppe acknowledges support from the internal research fund of the Space Research Centre of the Polish Academy of Sciences.

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