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# Time-variable gravity field determination from GRACE Follow-On data at the AIUB

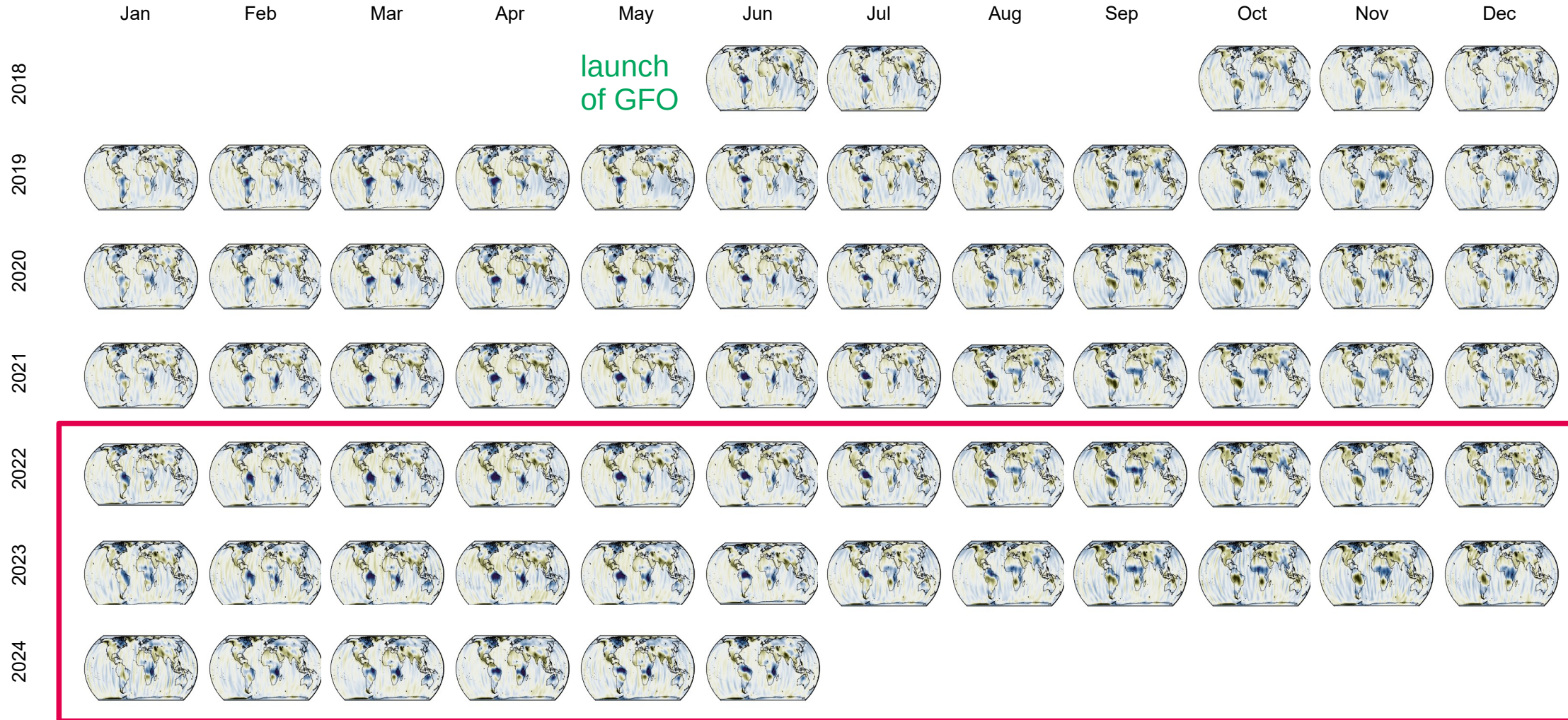
**Martin Lasser, Ulrich Meyer, Daniel Arnold and Adrian Jäggi**

GRACE Follow-On Science Team Meeting 2024, 8 – 10 October 2024, Potsdam, Germany

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# Operational GRACE Follow-On Solution

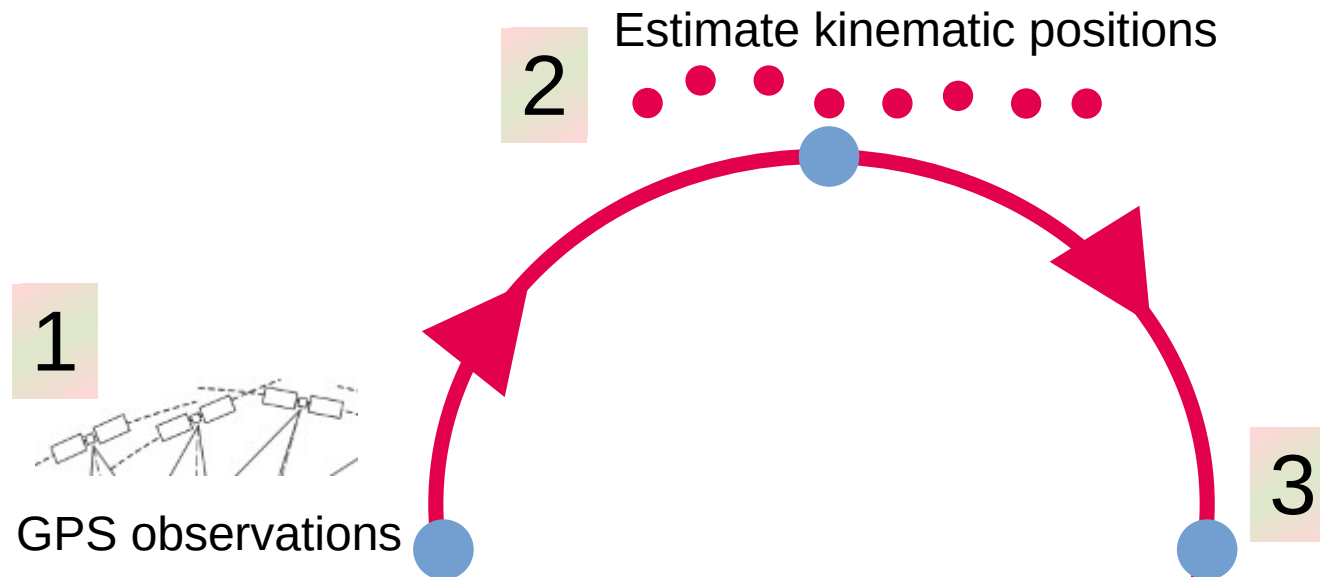
## Mosaic Jun 2018 – Jun 2024



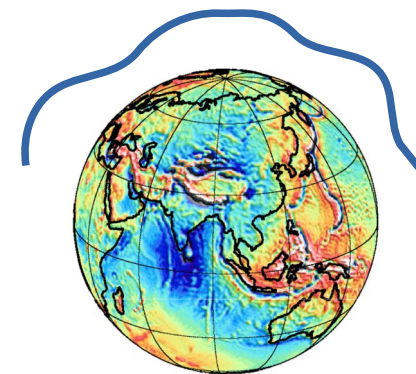
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# Gravity Field Recovery

## A little detour



Take kinematic positions and KBR/LRI observations and estimate (reduced) dynamic orbit + gravity field



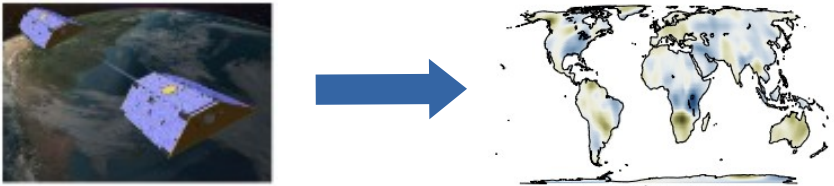
### Background information

- Celestial Mechanics Approach (CMA, Beutler et al., 2010) applied
- CODE GNSS products
- PCV maps used
- Ambiguities integer-fixed

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# Parametrisation

## For estimating gravity fields



### Basic parametrisation

- Initial conditions
- Bias in radial, along-track, cross-track
- ACC scaling factors
- Off-diagonal elements for ACC scaling  
→ since 2023 and for this study

### Additional parameters

- 15 min PCA per satellite in
  - radial
  - along-track
  - cross-track

in daily arcs

+ gravity field d/o=2..96

$\Sigma$  for 30 days:

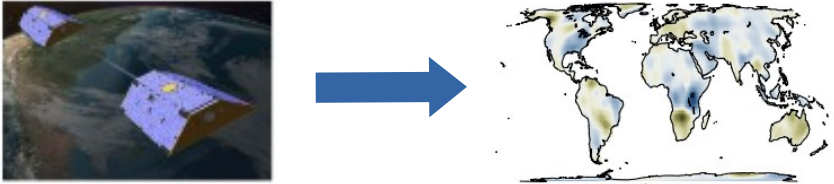
~ 18000 <orbit> parameters

+ 9405 gravity field d/o=2..96

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# Parametrisation

## For estimating gravity fields



### Basic parametrisation

- Initial conditions
- Bias in radial, along-track, cross-track
- ACC scaling factors
- Off-diagonal elements for ACC scaling

### Additional parameters

- 15 min PCA per satellite in
  - radial
  - along-track
  - cross-track

+ gravity field

### Force models

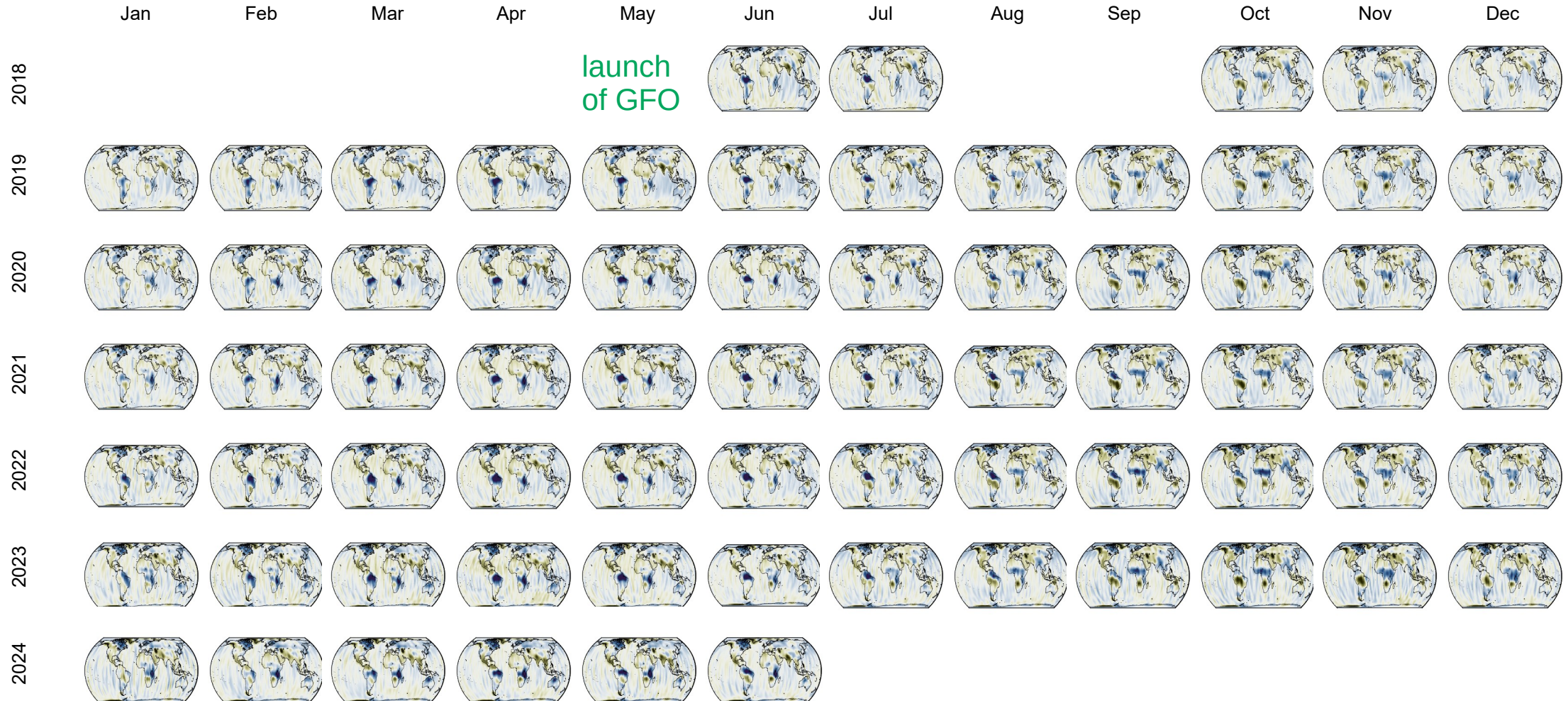
Gravity field	AIUB_APR
Astronomic bodies	JPL DE421 (all planets)
Mean pole	Linear
Solid Earth tides	IERS2010
Solid Earth pole tides	IERS2010
Ocean tides	FES2014b (+ admittances from TUG)
Ocean pole tides	Desai
Atmospheric tides	AOD RL06
Atmospheric & oceanic dealiasing	AOD RL06
Relativistic effects	IERS2010



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# Operational GRACE Follow-On Solution

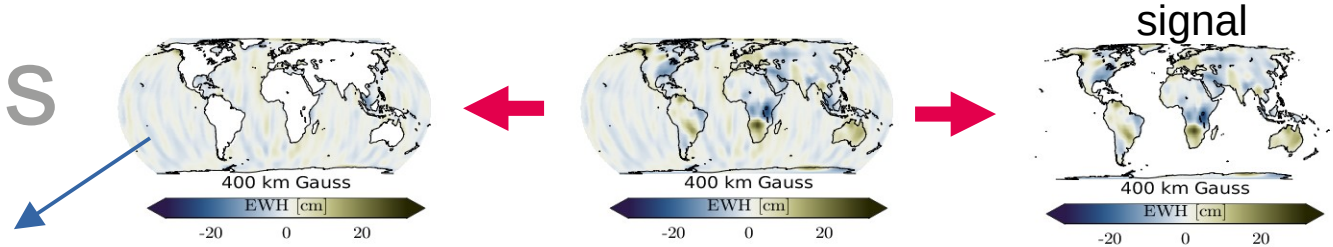
## Mosaic Jun 2018 – Jun 2024



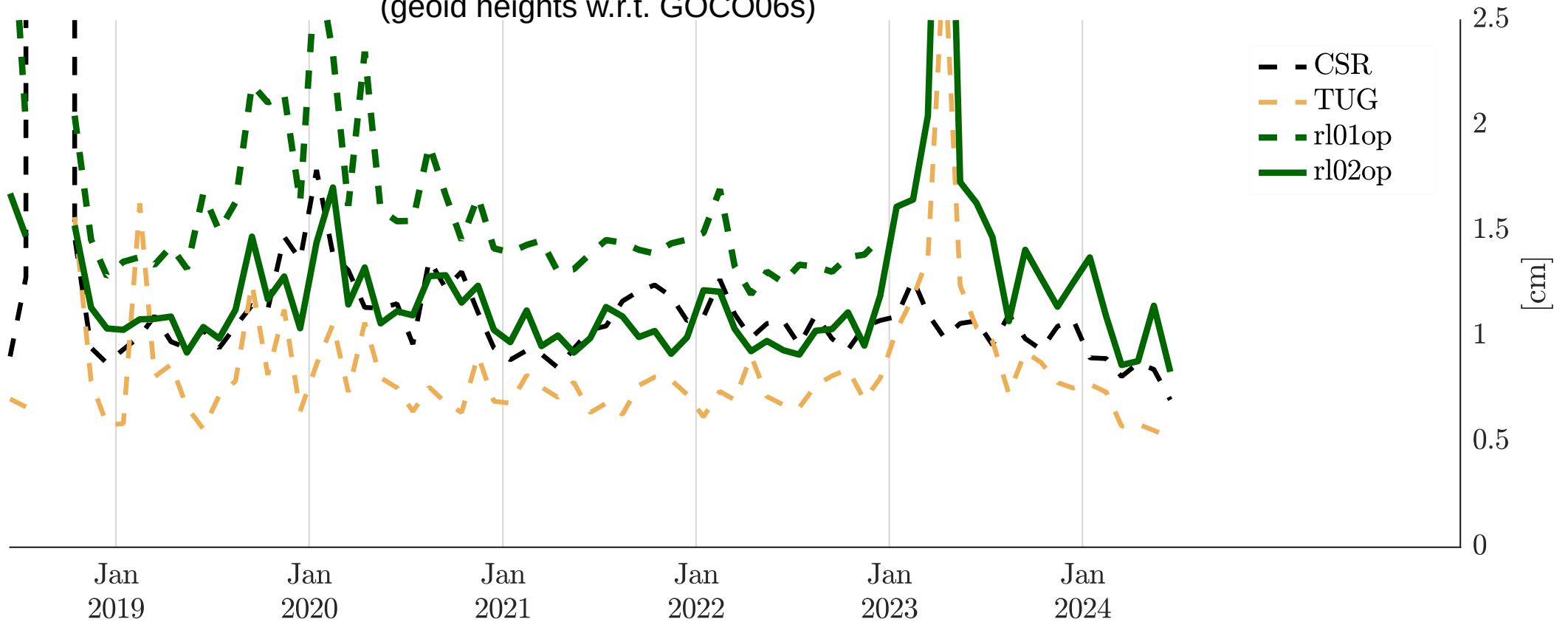
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# AIUB GRACE Follow-On Solutions

## RMS over oceans

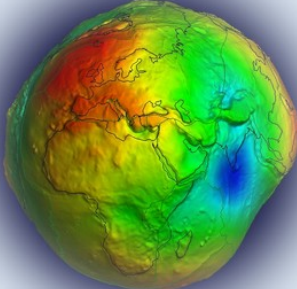





RMS over the oceans  
(geoid heights w.r.t. GOCO06s)





# $u^b$ GRACE Follow-On Observation concept

The processing standards to generate the GRACE Level-2 products of CSR, GFZ and JPL are also available in the Document Section of the GRACE archives at [GFZ ISDC](#) or [JPL PO.DAAC](#)

**ICGEM Home**

**Gravity Field Models**

- Static Models
- Temporal Models**
- Simulated Models
- Topographic Models

**Calculation Service**

- Regular grids
- User-defined points

COST-G (International Combination Service for Time-variable Gravity Field)				collapse all
FSM	DOI	quarterly	Fitted Signal Model	
Grace-FO RL01	DOI	monthly		
Grace-FO RL02	DOI	monthly		
Grace RL01	DOI	monthly		
Swarm	DOI	monthly		

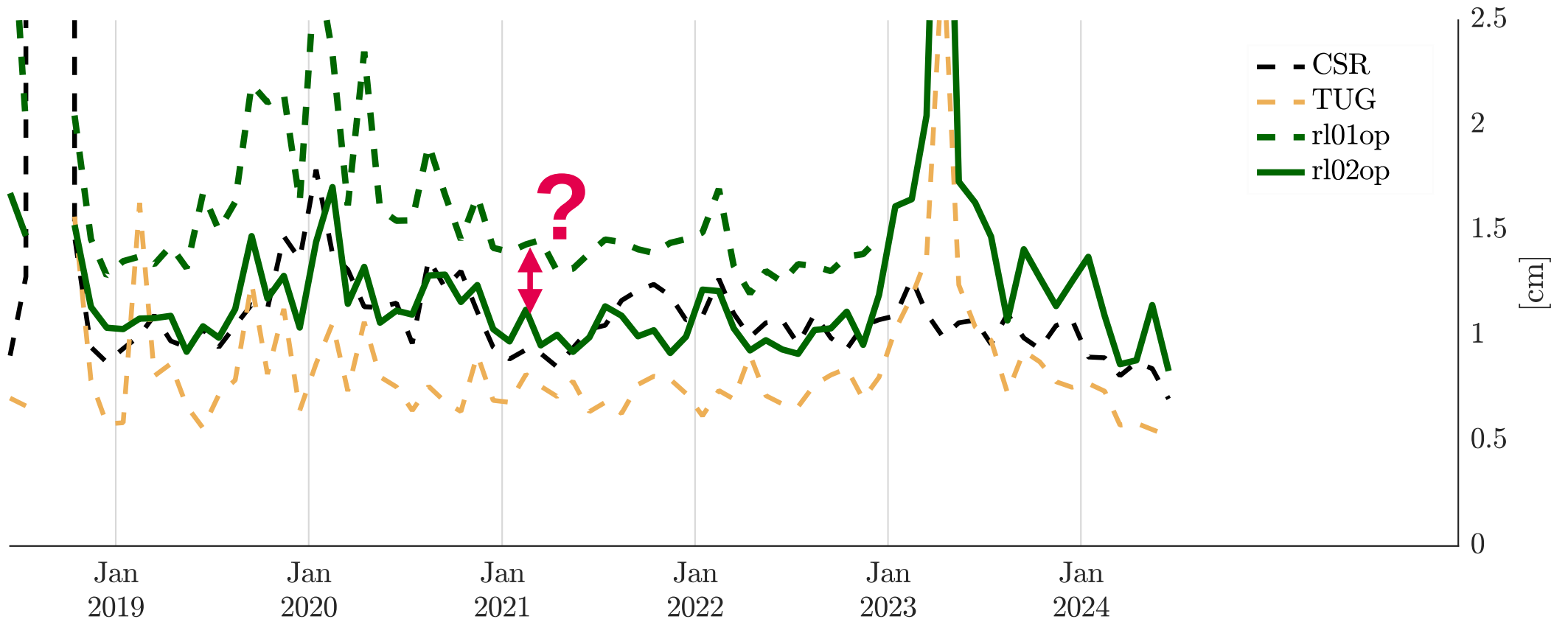
Other Models				expand all
+ AIUB				Astronomical Institute University Bern
AIUB C3P	DOI	monthly		
AIUB-GRACE-FO_op	DOI	monthly	Operational GRACE Follow-On monthly gravity field solutions from AIUB	
AIUB-GRACE-FO_rl02op	DOI	monthly	Operational GRACE Follow-On monthly gravity field solutions - RELEASE 02	
AIUB RL02		monthly	GRACE monthly solutions Release 2 from AIUB, more information can be found <a href="#">here</a>	
+ CNES				Centre national d'études spatiales



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# AIUB GRACE Follow-On Solutions

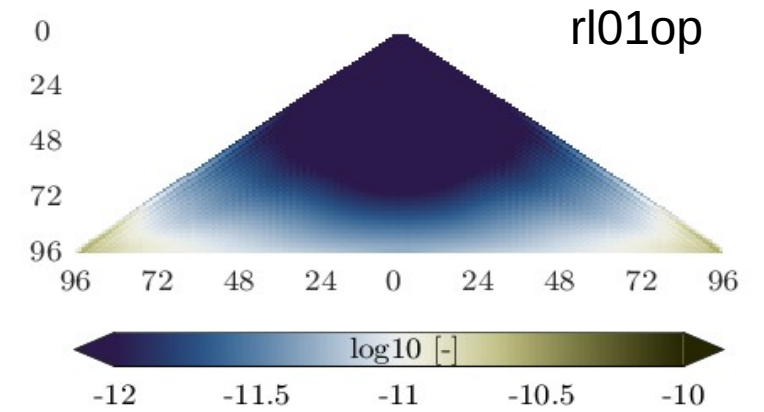
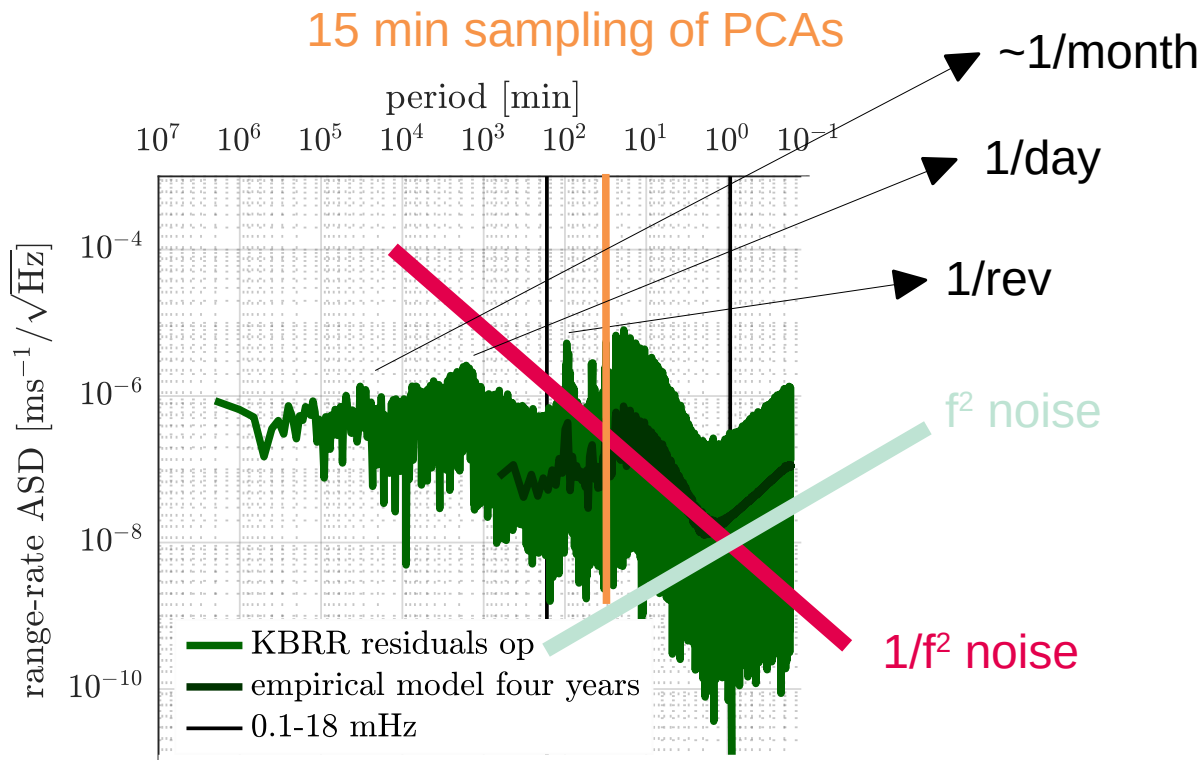
## RMS over oceans



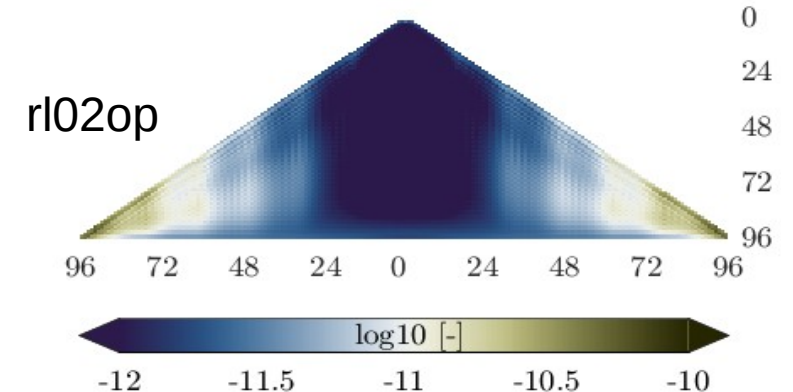
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# Noise modelling

## Based on post-fit residuals



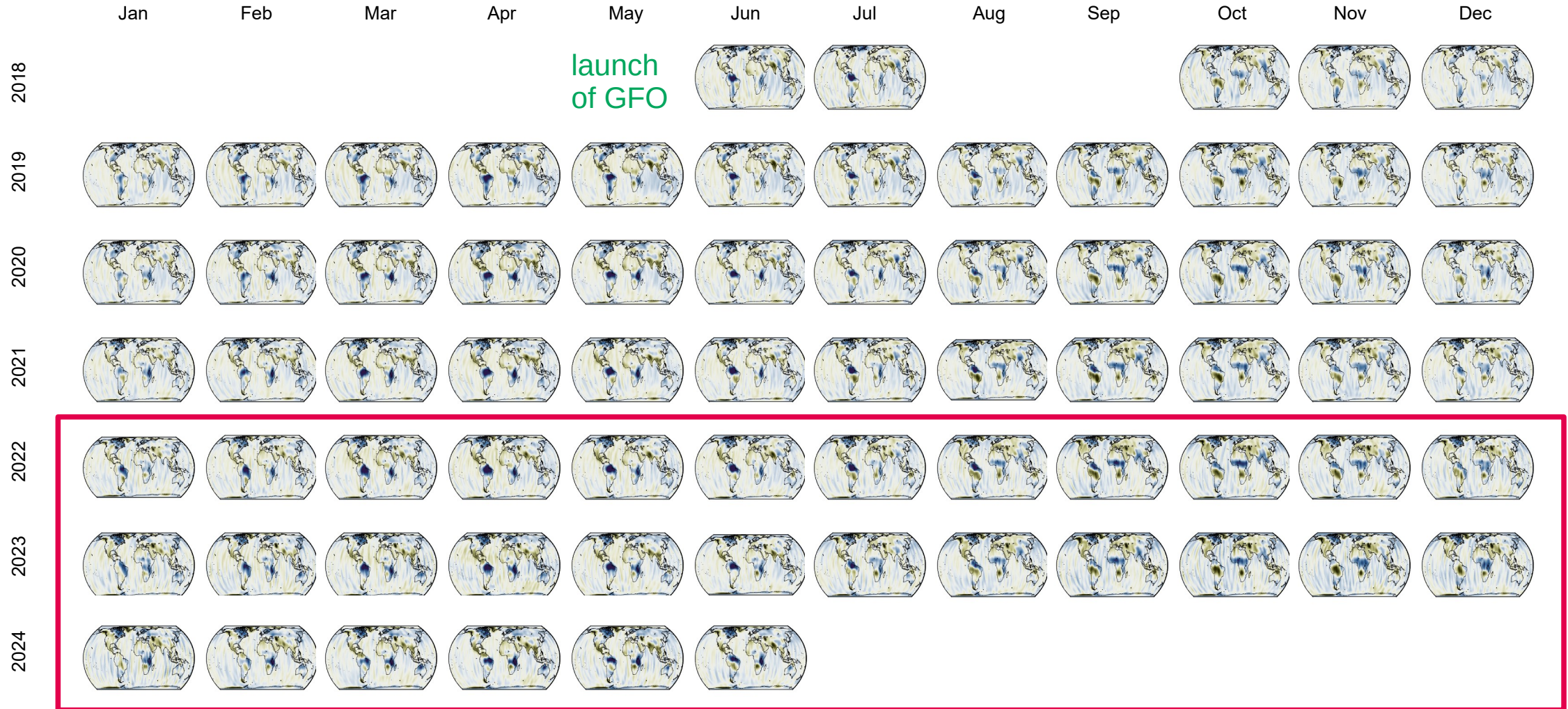
initial solution  $\rightarrow$  residuals  
auto covariance function  $\rightarrow$   
covariance matrix  $\rightarrow$   
weight matrix  $\rightarrow$  new estimation



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# GF2 ACC products

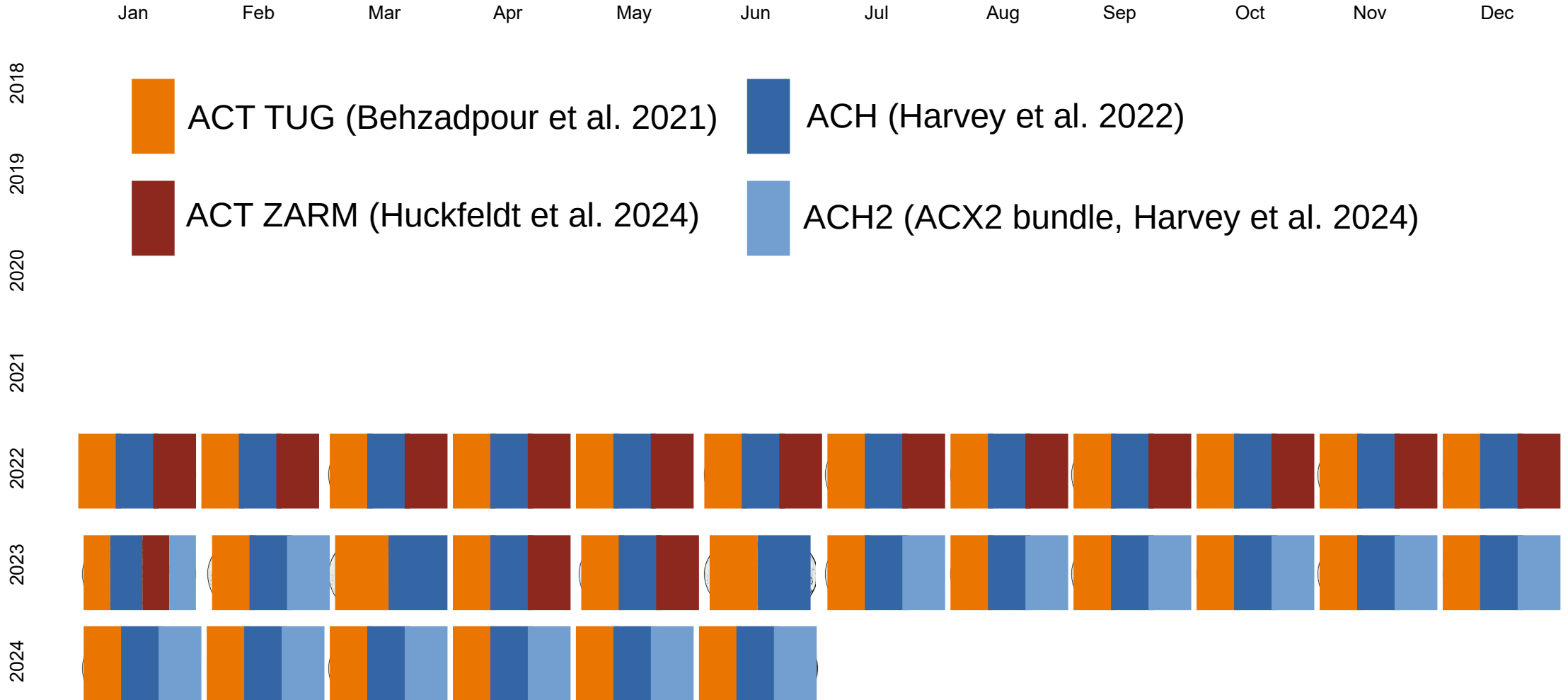
## Mosaic Jun 2018 – Jun 2024



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# GF2 ACC products

## Mosaic Jun 2018 – Jun 2024

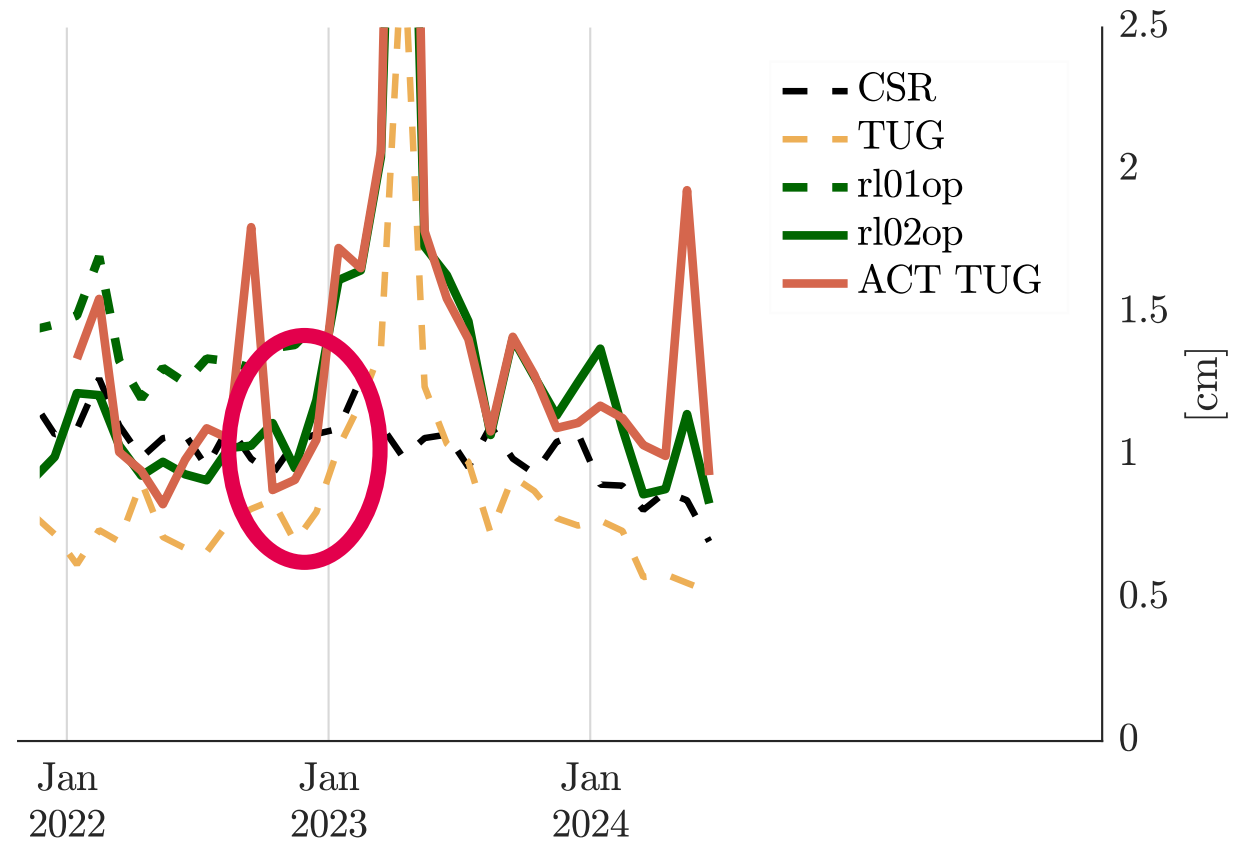




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# AIUB GRACE Follow-On Solutions

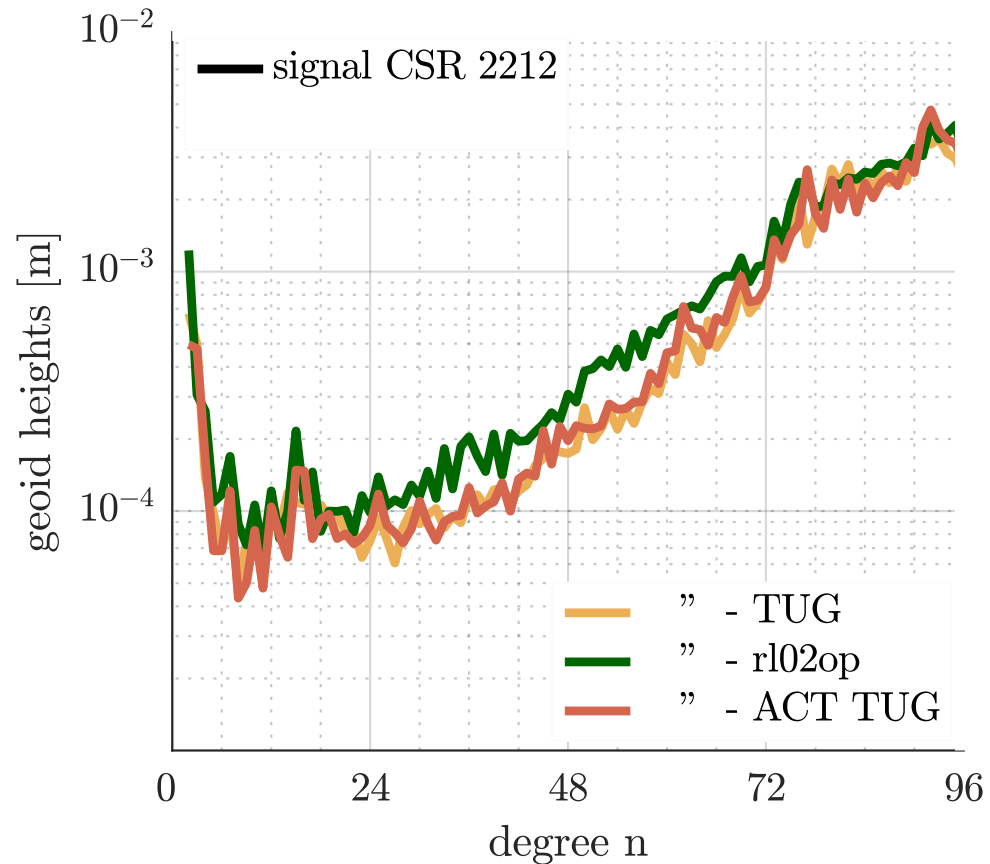
## RMS over oceans



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# December 2022

## Difference degree amplitudes



- Reason for the improvement: Application of variance component estimation to estimate the variance of the PCAs.

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# Piece-wise constant accelerations

## Constraining

$$\mathbf{N} = (\mathbf{A}^T \mathbf{P} \mathbf{A}) \quad \text{and} \quad \mathbf{b} = \mathbf{A}^T \mathbf{P} \mathbf{l} \quad \longrightarrow \quad \hat{\mathbf{x}} = \mathbf{N}^{-1} \mathbf{b}$$

$$\mathbf{N} = (\mathbf{A}^T \mathbf{P} \mathbf{A} + \mathbf{W})$$

$\mathbf{A}$  design matrix

$\mathbf{P}$  weight matrix

$\mathbf{l}$  observations

$$\mathbf{N} = \begin{bmatrix} \blacksquare & & \blacksquare & \blacksquare & \blacksquare & \blacksquare \\ & \blacksquare & & \blacksquare & \blacksquare & \\ \blacksquare & & \blacksquare & & \blacksquare & \blacksquare \\ & \blacksquare & & \blacksquare & & \\ \blacksquare & & \blacksquare & & \blacksquare & \\ & \blacksquare & & \blacksquare & & \\ \blacksquare & & \blacksquare & & \blacksquare & \blacksquare \\ & \blacksquare & & \blacksquare & & \\ \blacksquare & & \blacksquare & & \blacksquare & \blacksquare \\ & \blacksquare & & \blacksquare & & \\ \blacksquare & & \blacksquare & & \blacksquare & \blacksquare \end{bmatrix} + \begin{bmatrix} & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \end{bmatrix}$$

$\frac{\sigma_0^2}{\sigma_{PCA}^2},$

$\sigma_{PCA}^2 = \text{e.g., } 3 \times 10^{-10} \text{ ms}^{-2}$

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# Piece-wise constant accelerations

## Constraining

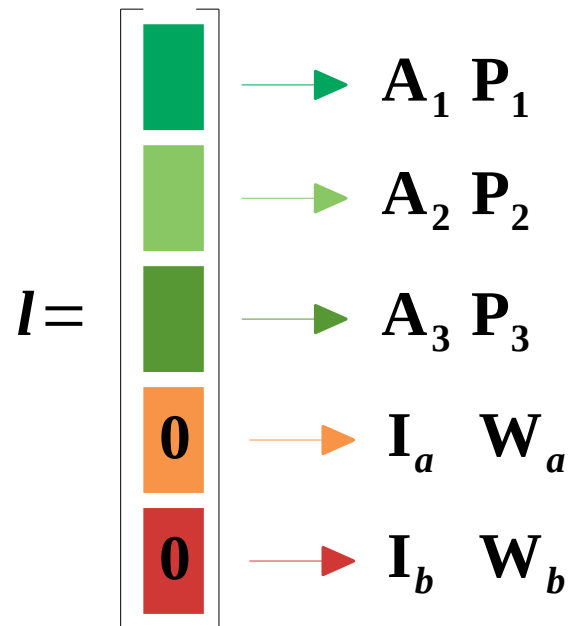
$$\mathbf{N} = (\mathbf{A}^T \mathbf{P} \mathbf{A}) \quad \text{and} \quad \mathbf{b} = \mathbf{A}^T \mathbf{P} \mathbf{l} \quad \longrightarrow \quad \hat{\mathbf{x}} = \mathbf{N}^{-1} \mathbf{b}$$

$\mathbf{A}$  design matrix

$$\mathbf{N} = (\mathbf{A}^T \mathbf{P} \mathbf{A} + \mathbf{W})$$

$\mathbf{P}$  weight matrix

$\mathbf{l}$  observations



- The observations of each arc are used to set up the normal equations (NEQs)
- Each arc is treated as being independent

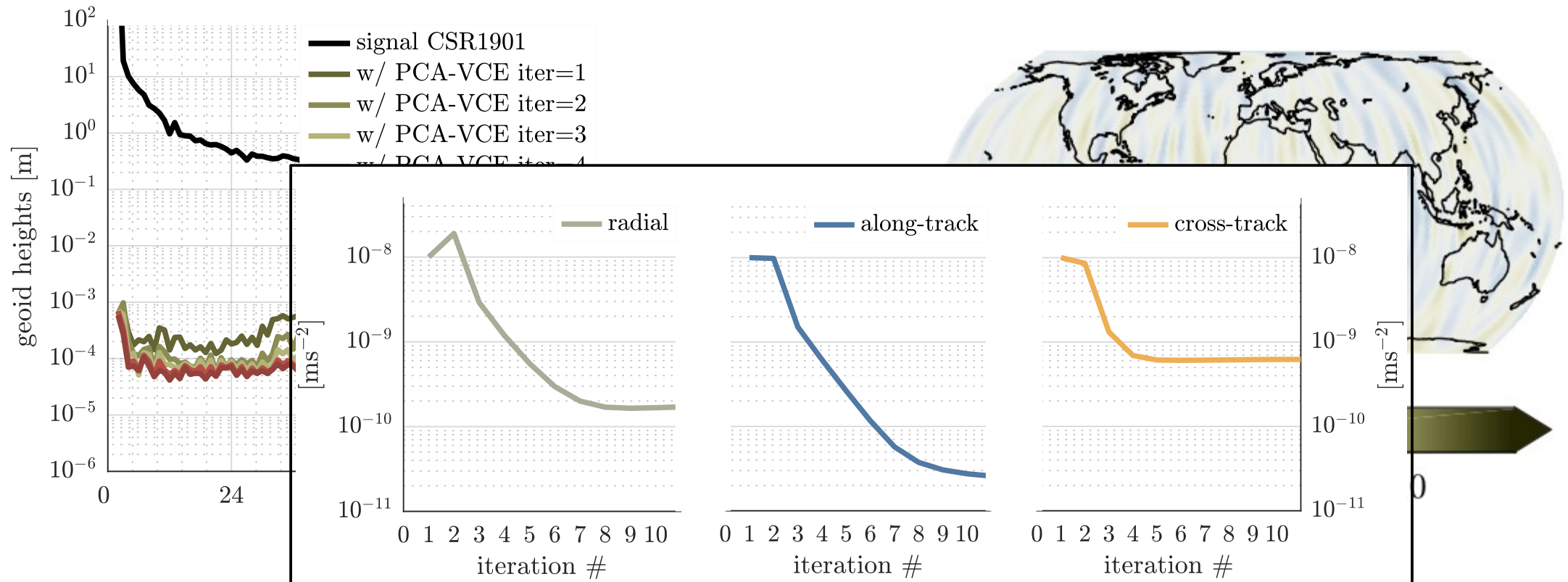
VCE: Each group of observations gets a weight based on its contribution to the final solution



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# Piece-wise constant accelerations

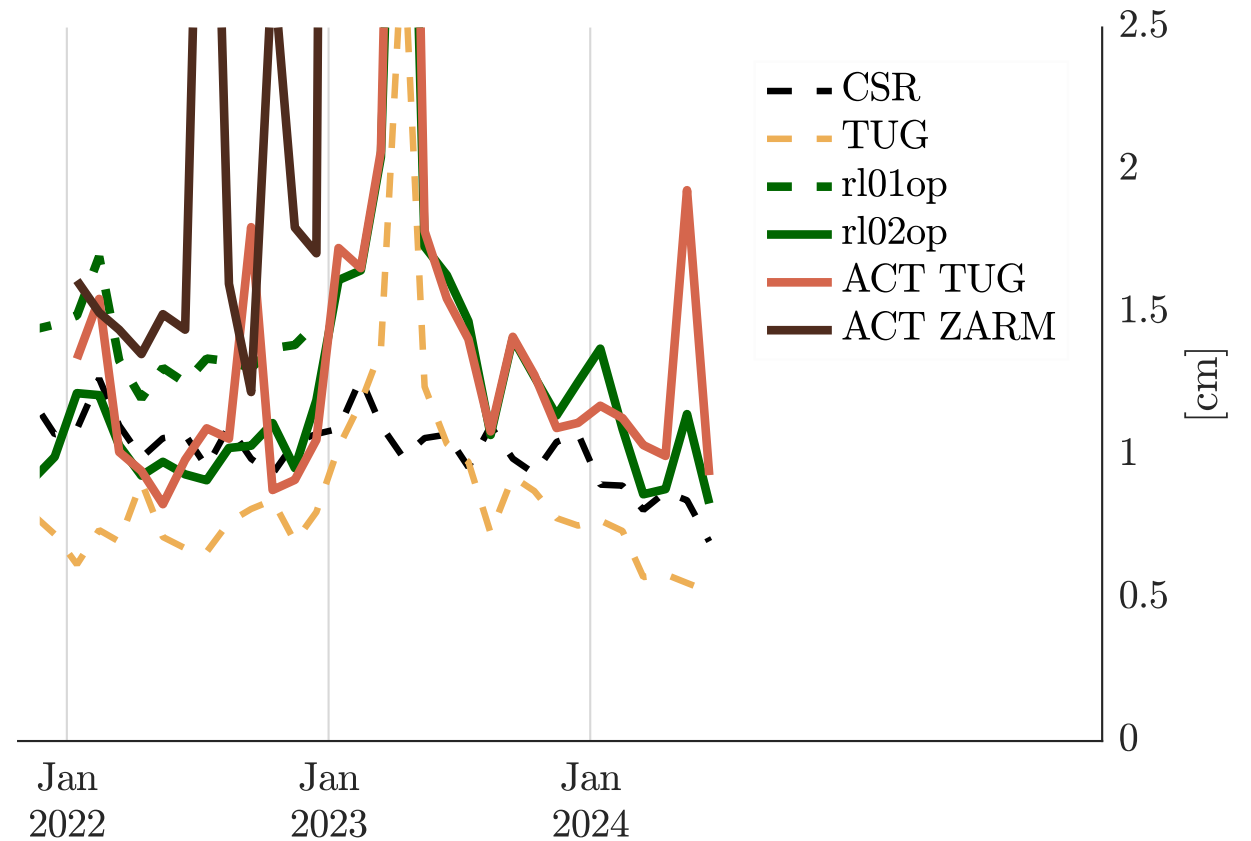
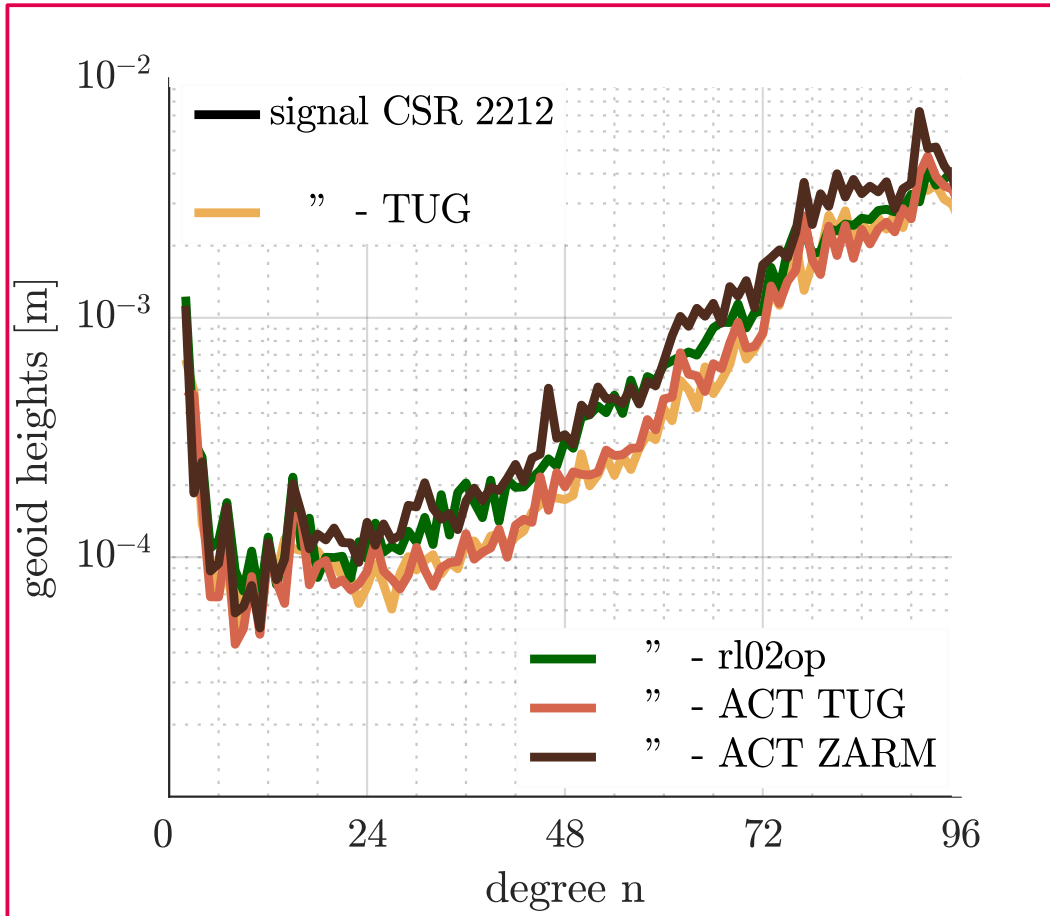
## Constraining



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# AIUB GRACE Follow-On Solutions

## RMS over oceans

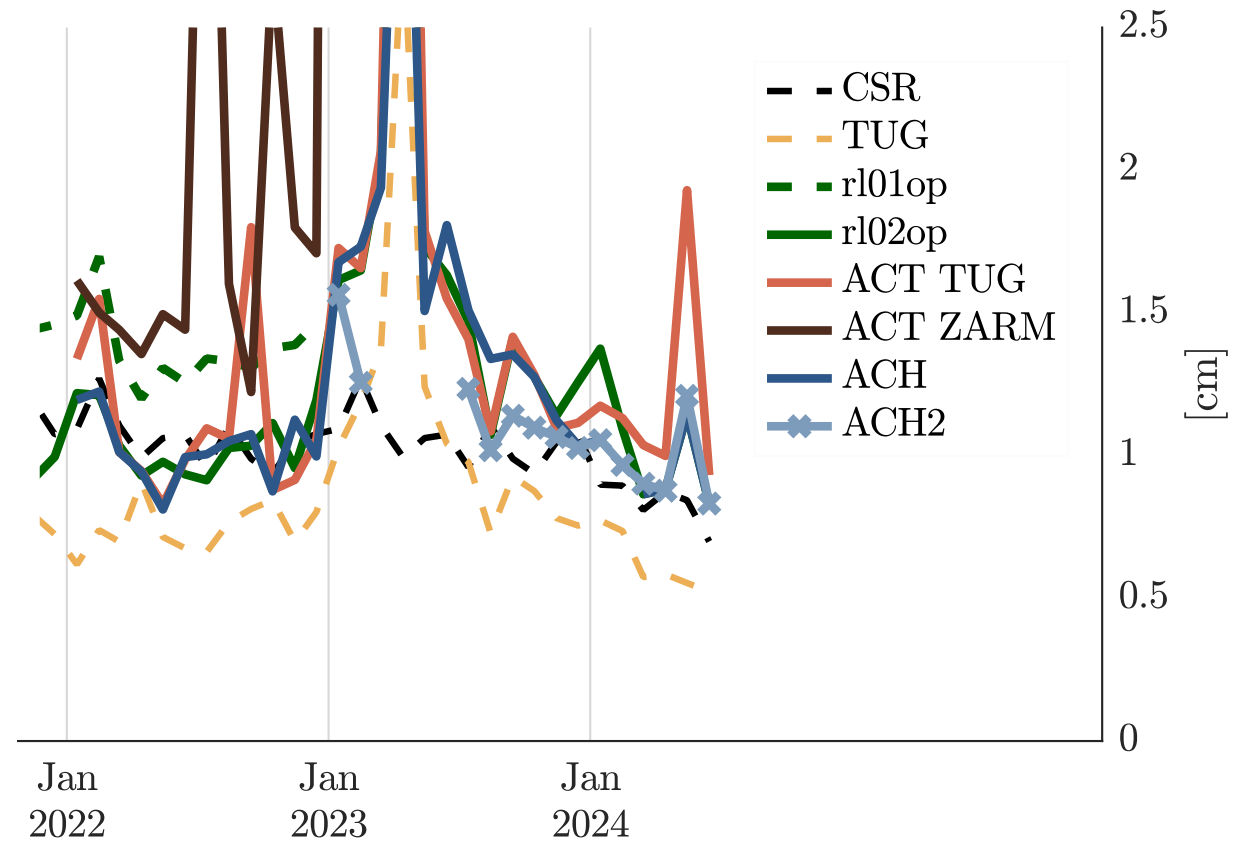


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# AIUB GRACE Follow-On Solutions

## RMS over oceans

- ACH → JPL's hybrid transplant
- ACH2 → ACH from ACX2 bundle



# $u^b$ GF2 – four ACT products

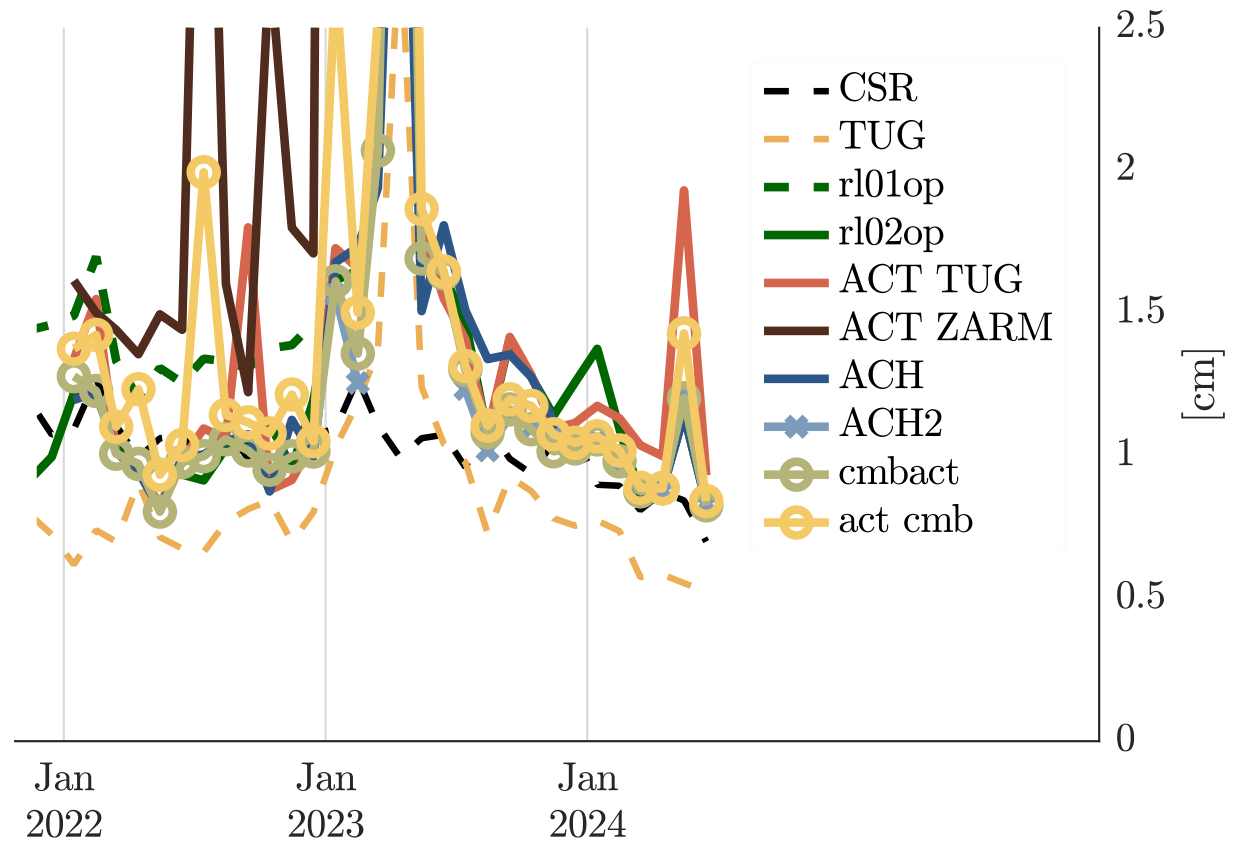
## What to do?

### Scenario #1 (cmbact)

Combine ACT products on NEQ level

### Scenario #2 (act cmb)

Combine ACT products by simple average





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# Summary & Conclusions

- Operational solution (rl02op) up to date (most recent for Jun 2024).
- Estimation of variance components for PCAs more important with more challenging data.
- ACH2 (ACX2) improves results (most until Nov 2023).
- Simple ACT product combinations seem to be a reasonable starting point.

Further steps:

- Screening procedure needs an update.
- Update of background models.

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# Thank you for your attention

## Contact

Martin Lasser

[martin.lasser@unibe.ch](mailto:martin.lasser@unibe.ch)

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## B

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